

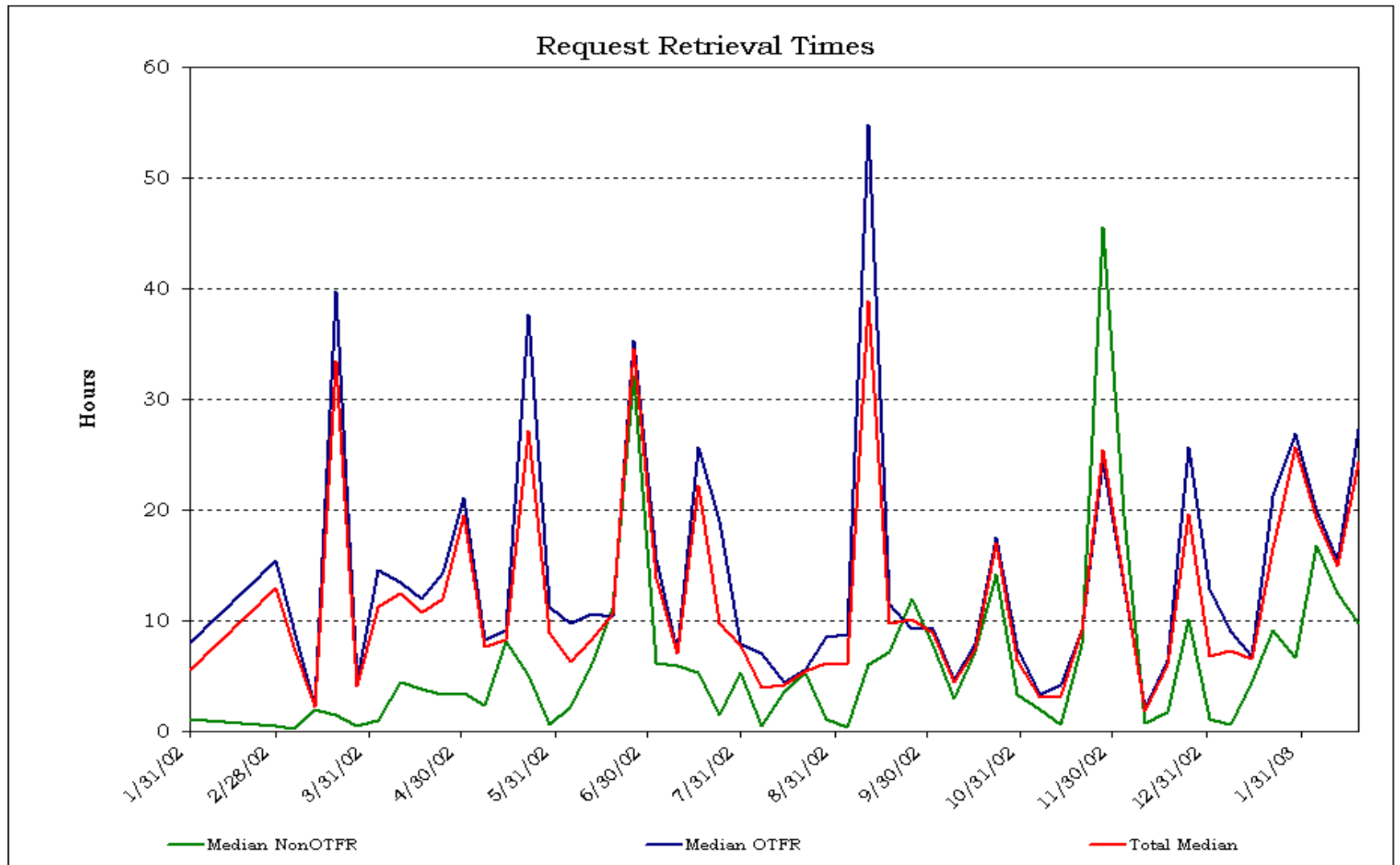
# Future Directions in HST Data Processing

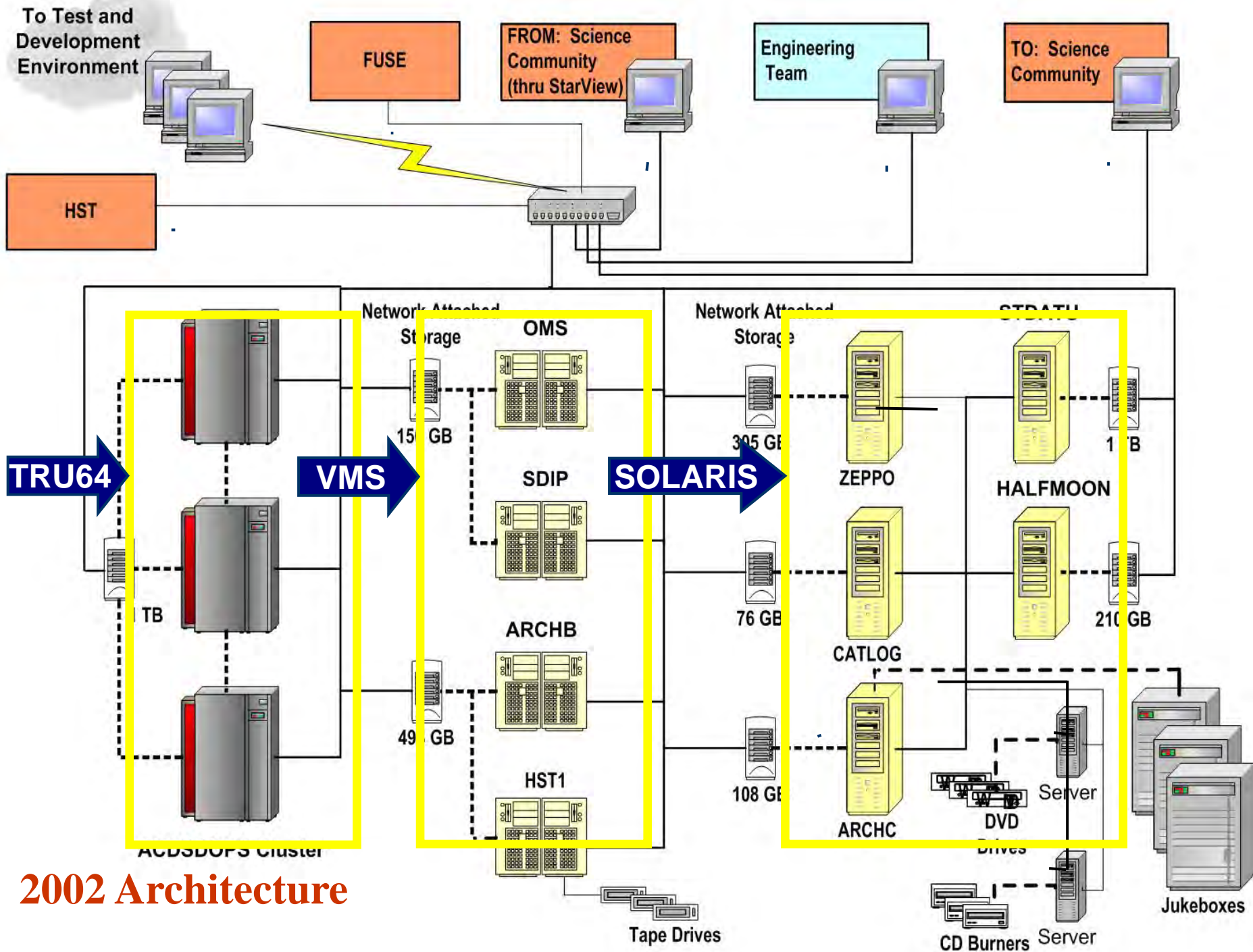
19 November 2004

# History (2002-2003)

- Several parts of the system exhibited unacceptable performance
  - ◆ Processing of data from HST to the Archive
  - ◆ Response time to user requests for data from Archive
- Several specific causes
  - ◆ NFS mount problems
  - ◆ Disk corruption in OPUS
  - ◆ Jukebox problems
  - ◆ Other specific hardware problems
- Symptomatic of more general problems with the data processing systems

# History (2002-2003)





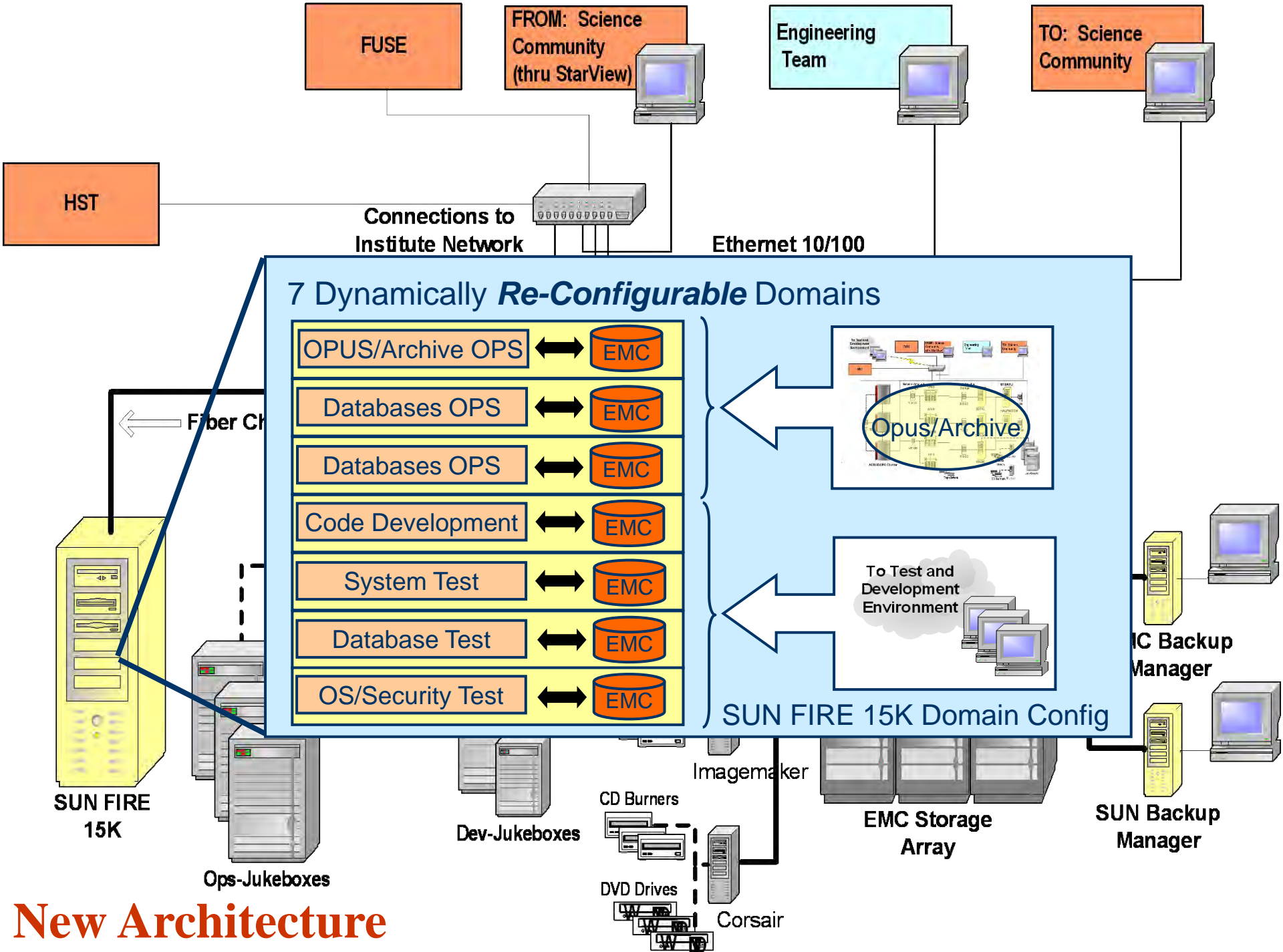
# 2002 Architecture

# Goals for re-implementation (2003-2004)

- Reduce total cost of data processing activities
- Simplify hardware and network architecture
  - ◆ Reduce Operating Systems from 3 to 1
  - ◆ Consolidate many boxes into two highly reliable boxes
  - ◆ Flexible allocation of computing resources
- Provide common development, test, and operational environments
  - ◆ Development and test systems could not replicate load of operational systems
- Improve ability to capture performance data, metrics, etc.

# Goals for re-implementation (2003-2004)

- Continue planned pipeline evolution
  - ◆ DADS Distribution redesign provides more flexibility to users and operators
  - ◆ Storing copy of Raw Data on EMC will dramatically reduce load and reliance on Jukeboxes
  - ◆ Ingest redesign provides opportunity to finally end the arbitrary boundary between OPUS and DADS



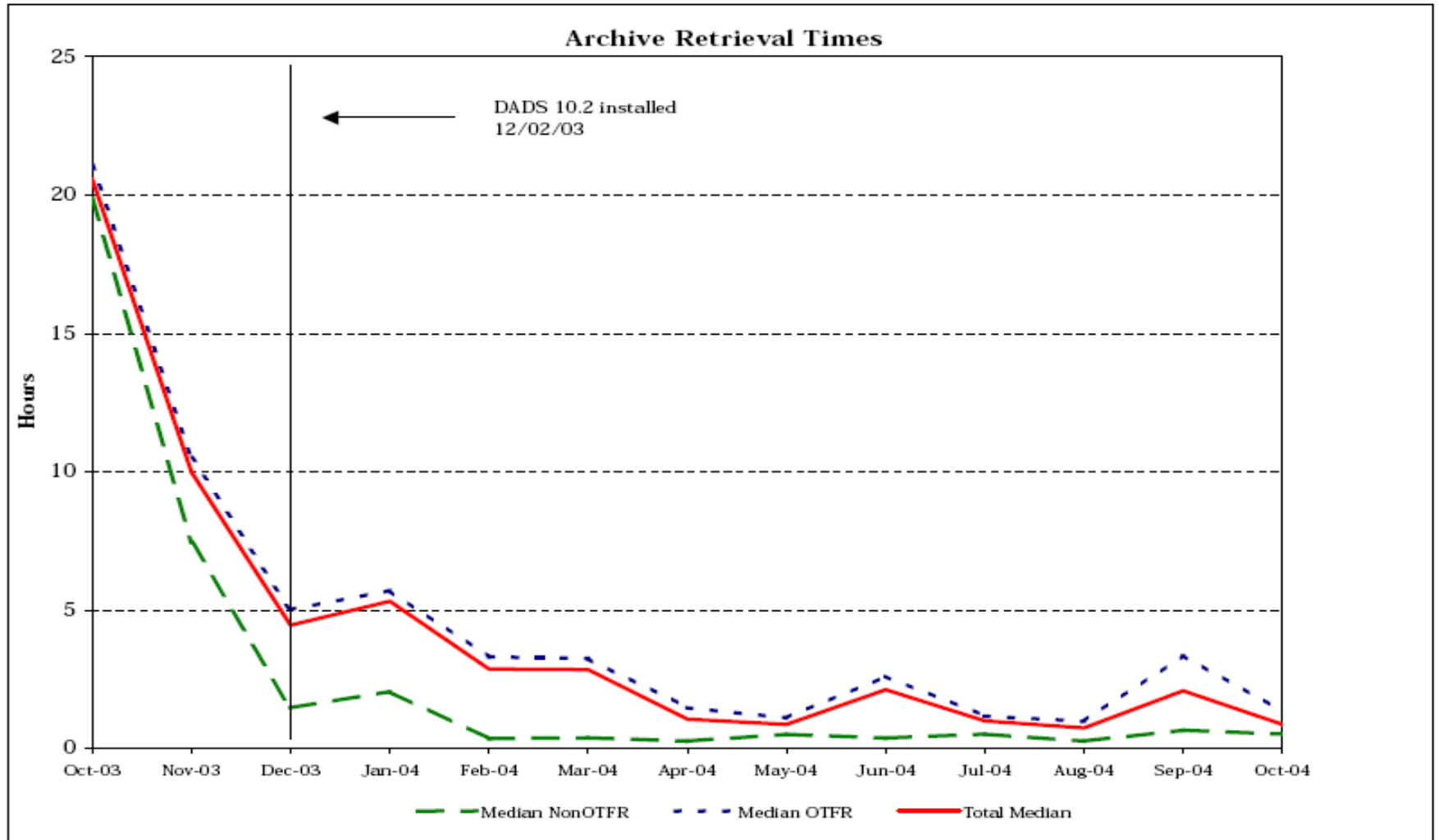
# New Architecture

# Re-implementation completed in 2004

- New Archive Ingest system installed June 2
  - ◆ Removes artificial boundary between OPUS pipeline and Archive
  - ◆ Removes VMS systems from input side of the Archive
- Upgraded to 64 bit compiler on June 21
- Tuned the EMC disk system over the summer
- Multidrizzle installed in ACS pipeline Sept. 22
  - ◆ Extra compute load is well within system capacity



# Current Retrieval Performance



# Near-term Plans

- Clean up remnants of old architecture
  - ◆ Totally decouple Jukeboxes from retrieval/ingest
  - ◆ Determine future of safestore
    - ◆ Media
    - ◆ Jukeboxes (?)
- Science value-added improvements
  - ◆ Multidrizzle
  - ◆ Accuracy of coordinates
  - ◆ Calibration upgrades (e.g. STIS close-out)

# Multidrizzle

- Dither Working Group is leading development of Multidrizzle (A. Koekemoer et. al.)
- Recent steps:
  - ◆ Multidrizzle implemented in Pipeline for ACS associations
  - ◆ Same version released last week in STSDAS
- Next steps – more Instruments:
  - ◆ STSDAS version can handle STIS imaging, WFPC2 data
  - ◆ Extend association definitions for ACS, WFPC2 to cover “pos targ”, and other data groups
  - ◆ Implement WFPC2 and STIS imaging in pipeline (mostly requires additional scientific verification and validation)
  - ◆ Use for NICMOS and STIS spectroscopy will require some additional software development

# Multidrizzle

- Next steps – Tweakshift
  - ◆ **Currently:**
    - ◆ Pipeline Multidrizzle uses headers for offsets between images
    - ◆ STSDAS version can accept measured offsets
  - ◆ **Prototype Tweakshift (in STSDAS) to measure offsets from images**
    - ◆ Via object identification
    - ◆ Via cross-correlations
    - ◆ Via wavelets (TBD)
  - ◆ **After evaluation via STSDAS release, move to pipeline processing**
    - ◆ Eliminates small pointing errors between images and orbits
    - ◆ Provides opportunity to improve absolute position reference

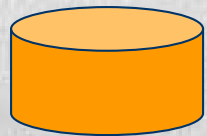
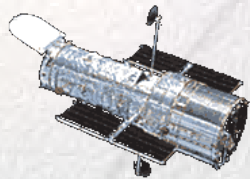
# Accuracy of Coordinates

- Accuracy of WCS coordinates is  $\sim 1\text{-}2$  arc-seconds
  - ◆ Original Guide Star Catalog limits accuracy
  - ◆ Slow Drift of FGSs relative to SIs has not been accurately calibrated
- Imaging data
  - ◆ Tying GSCII to Hipparchos frame provides opportunity to improve accuracy
  - ◆ Use objects found by Tweakshift to directly determine WCS coordinates – not affected by drift of FGSs relative to SIs
  - ◆ Evaluations have been successful
- Small FOV data sets (e.g. STIS)
  - ◆ Retroactively calibrate drift of apertures
    - ◆ Imaging data can provide FGS-to-FGS drifts
- Goal is accuracies of  $\sim 0.1\text{-}0.2$  arc-seconds

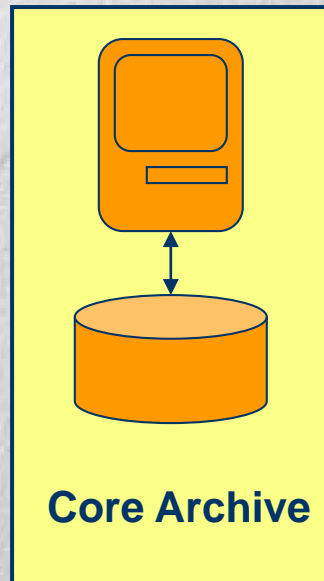
# Longer-Term Plans

- January, 2004 decision re SM4 led us to look for ways to augment the scientific value of the HST data processing and archive services
  - ◆ Stimulate an even broader re-use of the Hubble data
  - ◆ Support a more aggressive archival research program, in the event there is a major gap between HST and JWST
- Timed well with respect STScI planning for other missions we support (Kepler, JWST, NVO, etc.)
- We will be developing a proposal for enhancing the Hubble data services
  - ◆ Minimal disruption to the core architecture we have just finished
  - ◆ Add an outer “layer” that builds on the core
    - ◆ Higher level science products
    - ◆ Faster response time
    - ◆ NVO compatible

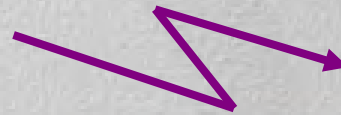
# Hubble Data Archive Current Situation



**Safe Store**



**Core Archive**

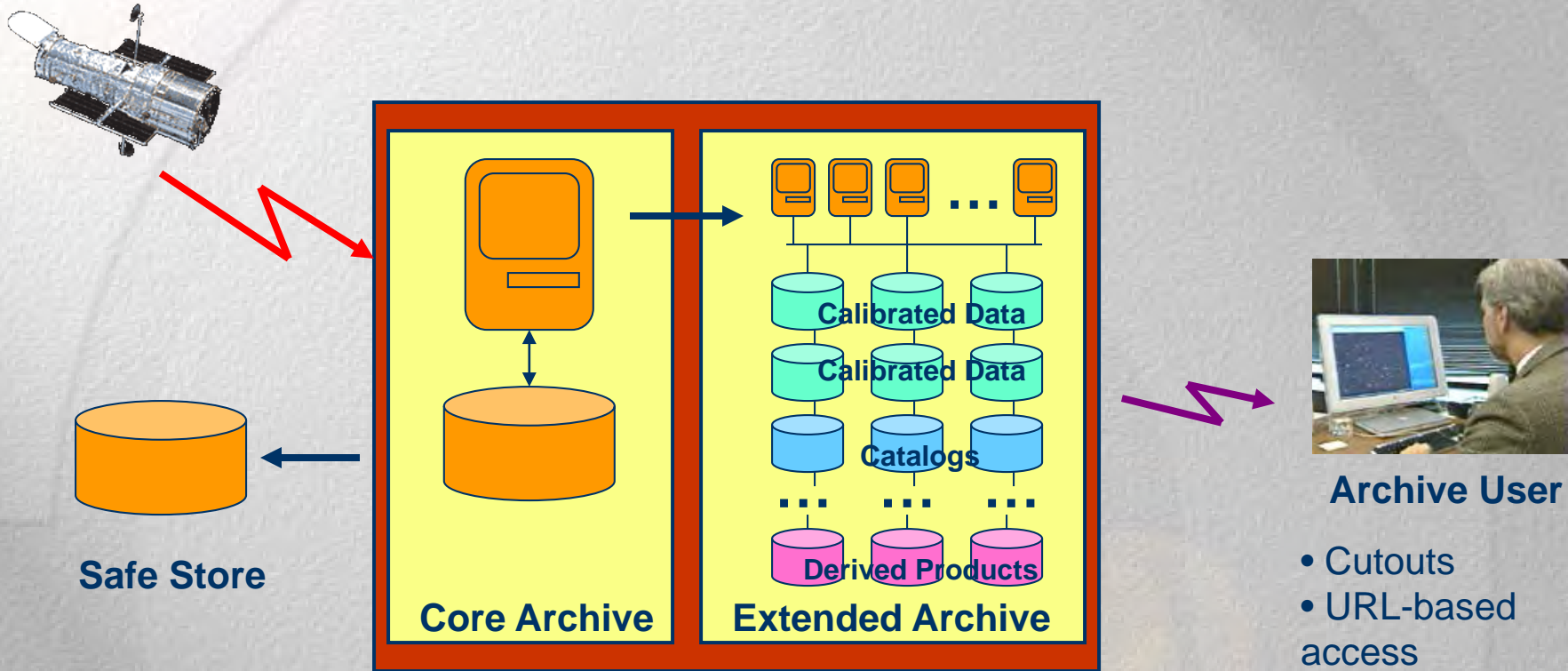


**Archive User**

- Process data from HST
- Re-process requested data for user
- All data is available

- Retrieval Times:  
Hours

# Hubble Data Archive Potential Enhancements



**Safe Store**

**Core Archive**

**Extended Archive**

**Archive User**

- Process data from HST
- Re-process data for extended archive
- All data is available

- Most calibrated data online
- Automatic updates from core
- Improved WCS coordinates
- Object catalogs
- Space/time “footprint” service
- Other high-level products...

- Cutouts
- URL-based access
- Retrieval Times: Seconds



# Next steps

- Formulate specific plans and schedules for near-term (~ next year) upgrades
  - ◆ Multidrizzle
  - ◆ Coordinate accuracies
- Document our concept for enhanced HST data services for HSTP and HQ
- We welcome STUC input to both processes