## Cycle 21 Results and

## Cycle 22 Preparations

Andrew Fox, for the Science Policies Group STUC meeting, 18 October 2013

## Cycle 21 Summary Results

| Proposals | Requested Approved \% Accepted |  |  | ESA Accepted | $\frac{\text { ESA \% }}{\text { Total }}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| General |  |  |  |  |  |
| Observer | 822 | 192 | 23.4\% | 44 | 22.9\% |
| Snapshot Archival | 55 | 9 | 16.4\% | 2 | 22.2\% |
| Research | 142 | 35 | 24.6\% | 0 |  |
| AR Legacy | 13 | 2 | 15.4\% | 0 |  |
| Theory | 63 | 11 | 17.5\% | 1 | 9.1\% |
| Total | 1094 | $\underline{249}$ | 22.8\% | 47 | 22.9\% |
| Primary |  |  |  |  |  |
| Orbits | 19742 | 3308 | 16.8\% | 587 | 17.7\% |

## Over-subscription by Cycle



## Acceptance Fraction by Size



## Acceptance Fraction by Size over Cycles 17-21



## ESA Acceptance Fraction



ESA Approved Orbits


ESA Accepted Proposals


ESA Investigators


Proposal Institutional Acceptance Fraction


## Distribution of Science Categories

## Submitted Orbits by Science Category



## Instrument Usage

| Configurat. | Mode | $\begin{gathered} \text { Prime } \\ \% \end{gathered}$ | Coordinate <br> d <br> Parallel \% | Total | Instrument Prime Usage | $\begin{aligned} & \text { Instrument } \\ & \text { Prime + Parallel } \\ & \text { Usage } \end{aligned}$ | Pure <br> Parallel <br> Usage | Snap Usage |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| ACS/SBC | Imaging | 1.6\% | 0.0\% | 1.2\% |  |  | 0.0\% | 0.0\% |
| ACS/SBC | Spectroscopy | 0.029\% | 0.0\% | 0.0\% |  |  | 0.0\% | 0.0\% |
| ACS/WFC | Imaging | 10.4\% | 29.3\% | 14.7\% |  |  | 0.0\% | 37.5\% |
| ACS/WFC | Ramp Filter | 1.1\% | 0.0\% | 0.9\% | 3.1 | 16.8\% | 0.0\% | 0.0\% |
| ACS/WFC | Spectroscopy | 0.0\% | 0.0\% | 0.0\% |  |  | 0.0\% | 0.0\% |
| cos/Fuv | Spectroscopy | 23.7\% | 0.0\% | 18.3\% |  |  | 0.0\% | 4.6\% |
| cos/NuV | Imaging | 1.4\% | 0.0\% | 1.1\% | .5\% | 22.0\% | 0.0\% | 0.0\% |
| COS/NUV | Spectroscopy | 3.4\% | 0.0\% | 2.6\% |  |  | 0.0\% | 0.0\% |
| FGS | POS | 0.2\% | 0.0\% | 0.2\% | 0.2\% | 0.2\% | 0.0\% | 0.0\% |
| FGS | TRANS | 0.0\% | 0.0\% | 0.0\% |  |  | 0.0\% | 0.0\% |
| STIS/CCD | Imaging | 0.9\% | 1.6\% | 1.0\% |  |  | 0.0\% | 0.0\% |
| STIS/CCD | Spectroscopy | 2.8\% | 0.0\% | 2.2\% |  |  | 0.0\% | 0.0\% |
| STIS/FUV | Imaging | 0.5\% | 0.0\% | 0.4\% | 16.3\% | 13.3\% | 0.0\% | 0.0\% |
| STIS/FUV | Spectroscopy | 5.2\% | 1.6\% | 4.4\% |  |  | 0.0\% | 10.9\% |
| STIS/NUV | Imaging | 6.9\% | 0.0\% | 5.3\% |  |  | 0.0\% | 0.0\% |
| STIS/NUV | Spectroscopy | 0.0\% | 0.0\% | 0.0\% |  |  | 0.0\% | 10.6\% |
| WFC3/IR | Imaging | 10.6\% | 13.7\% | 11.3\% |  |  | 0.0\% | 6.8\% |
| WFC3/IR | Spectroscopy | 10.4\% | 9.5\% | 10.2\% | 41.8\% | 47.7\% | 60.0\% | 0.0\% |
| WFC3/UVIS | Imaging | 20.8\% | 44.2\% | 26.2\% |  |  | 40.0\% | 29.6\% |
| WFC3/UVIS | Spectroscopy | 0.0\% | 0.0\% | 0.0\% | $\square$ |  | 0.0\% | 0.0\% |
|  |  | 100\% | 100.0\% | 100.0\% | 100.0\% | 100.0\% | 100.0\% | 100.0\% |

# Approved TAC Proposals (7 Large/1 Pure Parallel/2 AR Legacy) 

| ID Last Name | First Name | Resources | Institution | Title |
| :---: | :---: | :---: | :---: | :---: |
| 1178. Ayres | Thomas | 230 | University of Colorado at Boulder | Advanced Spectral Library II: Hot Stars |
| 1827.\|Bean | Jacob | 150 | University of Chicago | Follow The Water. The Ultimate WFC3 Exoplanet Atmosphere Survey |
| 1241. Calzetti | Daniela | 154 | University of Massachusetts Amherst | LEGUS: Legacy ExtraGalactic UV Survey |
| 1777. Churchill | Christopher | 110 | New Mexico State University | A Breakaway from Incremental Science: Full Characterization of the $z<1$ CGM and Testing Galaxy Evolution Theory |
| 1421. illingworth | Garth | AR Legacy | University of California - Santa Cruz | High level science products from deep ACS and WFC3/IR imaging over the CDF-S/GOODS-S region |
| 1345. Malkan | Matthew | 375 Pure Parallel | University of Califomia - Los Angeles | WFC3 Infrared Spectroscopic Parallel Survey WISP: A Survey of Star Formation Across Cosmic Time |
| 1096. Peterson | Brad | 179 | The Ohio State University | Mapping the AGN Broad Line Region by Reverberation |
| 1484.\|Piotto | Giampaolo | 131 | Universita degli Studi di Padova | The HST Legacy Survey of Galactic Globular Clusters: Shedding UV Light on Their Populations and Formation |
| 1042. Treu | Tommaso | 140 | University of California - Santa Barbara | The Grism Lens-Amplified Survey from Space (GLASS) |
| 1226. 'van der Marel | Roeland | AR Legacy | Space Telescope Science Institute | Proper Motions of Distant Halo Stars: New Clues to Milky Way Structure, Evolution and Mass |

## Medium Proposals (35-74 orbits) Did they work?

- 400 orbits were available to TAC for Mediums in Cycle 21
- 109 Medium programs submitted, 13 approved (8:1 oversubscription) for 558 orbits.
- Compare 13 Mediums approved to 1 approved in Cycle 20
- Because the goal is to make proposal success rate ~independent of orbit request, for Cycle 22 we propose raising the Medium allotment from 400 to 600 orbit
- Feedback from TAC was positive: the Medium programs selected complemented the Large \& Treasury programs selected in terms of science areas.


## Medium Programs Recommended by TAC and Panels

| ID Last Name | First Name | Institution | Orbits | Title |
| :---: | :---: | :---: | :---: | :---: |
| 1781. Barro | Guillermo | University of California - Santa Cruz | 56 | The progenitors of quiescent galaxies at $z \sim 2$ : precision ages and star-formation histories from WFC3/IR spectroscopy |
| 1691. Borthakur | Sanchayeeta | The Johns Hopkins University | 51 | Characterizing the Elusive Intragroup Medium and Its Role in Galaxy Evolution |
| 1319. Fesen | Robert | Dartmouth College | 39 | STIS Spectra of the Young SN Ia Remnant SN 1885 in M31 |
| 1995. Foley | Ryan | Smithsonian Institution Astrophysical Observatory | 36 | Understanding the Progenitor Systems, Explosion Mechanisms, and Cosmological Utility of Type la Supernovae |
| 1128. Fox | Andrew | Space Telescope Science Institute - ESA | 49 | The Closest Galactic Wind: UV Properties of the Milky Way's Nuclear Outflow |
| 1392. Geha | Marla | Yale University | 44 | A Non-Universal Initial Mass Function in the UltraFaint Galaxy Coma Berenices |
| 1764. Kallivayalil | Nitya | Yale University | $30+30 \text { in }$ <br> Cycle 23 | Proper Motion and Internal Kinematics of the SMC: are the Magellanic Clouds bound to one another? |
| 2015. Oestlin | Goeran | Stockholm University | 54 | eLARS - extending the Lyman Alpha Reference Sample |
| 1788. Rodney | Steven | The Johns Hopkins University | $20+20+20$ <br> for 3 cycles | Frontier Field Supernova Search |
| 1150. Shull | J. | University of Colorado at Boulder | 35 | Deep COS Spectra of the Two Brightest Quasars that Probe the He II Post-Reionization Era |
| 1430. Siana | Brian | University of California Riverside | 48 | The Ultraviolet Frontier: Completing the Census of Star Formation at Its Peak Epoch |
| 1706. Tripp | Todd | University of Massachusetts Amherst | 60 | Directly Probing $>10^{\wedge} 6 \mathrm{~K}$ Gas in Lyman Limit Absorbers at $\mathrm{z}>2$ |
| 1237. Wood | Brian | Naval Research Laboratory | 36 | Tracking the Winds of Red Giants from the Star to the ISM |

## The UV Initiative - did it work?

Reminder of motivation for UV Initiative:

- Hubble is the only current mission capable of undertaking detailed UV (900-3200 $\AA$ ) observations.
- These capabilities have a limited lifetime.
- New NASA UV missions will not be launched in the near future.
- A special UV initiative was introduced in Cycle 21 to emphasize this unique HST resource.


## UV Initiative (cont.)

- In Cycle 21, each panel was asked to aim to devote at least $\mathbf{4 0 \%}$ of its orbit allocation to UV-specific science
- The TAC was asked to aim to devote at least $\mathbf{5 0 \%}$ of its orbit allocation to UV-specific science
- The Initiative also extended to archival and theory proposals, aimed at producing UV-specific high level data products and tools for the Hubble archive, and models for interpreting them.
- These allocations were targets, not quotas. UV-specific proposals recommended for acceptance had to meet the usual requirement of high scientific quality.


## UV Initiative (cont.)

- Result 1: we received a large number of UV proposals: 25 Large, 50 Medium, 327 Small
- Result 2: 60\% of prime GO orbits in Cycle 21 (1986 of 3316 orbits) and $43 \%$ of approved proposals were devoted to UV observations. This included WFC3 U-Band (F336W filter).
- UV Initiative reflected in COS \& STIS instrument usage:
- COS prime usage in C21: $18.9 \%$; in C22: $28.5 \%$
- STIS prime usage in C21: 9.6\% ; in C22: $16.3 \%$
- Together COS and STIS accounted for $44.8 \%$ of C21 prime usage
- Overall spectroscopy percentage in C21:38\% ( $27 \%$ in C20)


## UV Initiative was applied across all science categories

UV Orbits by Panel


## Targets of Opportunity

Following discussion with STUC last year, definition of disruptive ToO was changed in Cycle 21 from two to three weeks. Did that make any difference?

Cy 20 ToO Disruptive: 13 submitted/6 approved
Cy 20 ToO Non-disrup: 17 submitted/5 approved
Cy 21 ToO Disruptive: 12 submitted/5 approved
Cy 21 ToO Non-disrup: 27 submitted/9 approved
Number ToO proposals submitted and approved has gone up More non-disruptive proposals submitted in C21

## Director's Discretionary (DD) Proposals

- DD proposals are received throughout the year, reviewed by internal committee and by external reviewers.
- In Cycle 20 (last complete cycle), 22 DDs were submitted, 15 were approved for 103 orbits ( $68 \%$ success rate)
- 6 Solar System (of which 5 were on Comet ISON)
- 2 Exoplanets
- 3 GRB
- 2 SN
- 1 AGN
- 1 ISM
- Successful DD proposals are listed immediately on public webpage (default is no proprietary time)


## Joint Observatory Programs

- Proposers can request observations on both HST and Chandra/ Spitzer/XMM-Newton/NOAO, to avoid double jeopardy of writing two proposals on same science
- Chandra allocates up to $\mathbf{1 0 0}$ HST orbits, in exchange for 400 ks of Chandra time
- Spitzer allocates up to $\mathbf{6 0}$ HST orbits, in exchange for 60 hours of Spitzer time
- XMM-Newton allocates up to 30 HST orbits, in exchange for 150 ks of XMM-Newton time.
- NOAO makes available up to $5 \%$ of its observing time to the HST TAC (one-way arrangement).
- NRAO is negotiating with us for a new joint HST/NRAO proposal category for Cycle 22 ( 30 orbits HST, up to 5\% NRAO)


## Joint Observatory Programs (cont.)

In general, the partner observatories give away (almost) all their available HST time, whereas we do not give away most of their time. E.g. the Cycle 21 HST TAC approved

- 20 ks of Chandra time (out of 400 ks available)
- 20.4 hours of Spitzer time (out of 60 hours available)
- 0 NOAO nights (out of $5 \%$ of their time available)
- 179 ks of XMM-Newton time ( 150 ks available), but 0 ks were approved in Cycle 20


## CYCLE 22 PREPARATIONS

(see next talk from Neill Reid on policy changes)

## Cycle 22 Proposal Review Schedule

- 1/8/14: Call for Proposals release
- 4/11/14: Phase I Proposal deadline ( $\sim 6$ weeks later than C21)
- Early May 2014: Proposals sent to reviewers
- Early June 2014: Preliminary grades due
- 8-13 June 2014: Panels and TAC meet (at STScI/JHU)
- 6/25/14: Notifications sent out
- 7/24/14: Phase II and Budget deadlines
- 10/1/14: Cycle 22 begins


## Cycle 22 Features

- All five instruments will be offered:
ACS, COS, FGS, STIS, WFC3
- Medium category (35-74 orbits) will be continued
- The UV Initiative will be continued
- Joint HST/Spitzer proposals will be continued contingent on the results of the NASA senior review
- Joint HST/NRAO proposals may be introduced


## Cycle 22 Tentative Orbit Allocation

- ~3400 orbits available for C22 GO Programs (up from 3200)
- Break-down: 1800 orbits for panels; 600 for Medium proposals; 1000 for the TAC (Large \& Treasury)
- Additional orbits:
- $\mathbf{1 9 0}$ for Chandra/Spitzer/XMM-Newton
- 30 for NRAO (TBD)
- 100 for routine DD
- $\mathbf{2 7 0}$ for Frontier Fields
- $\mathbf{2 5 0}$ for calibration
- carry-over, continuation programs, and repeats
- Also 1000 SNAPs ( 150 maximum for COS + STIS)


## Panel Structure

## 14 panels (unchanged)

- Planets 1/2: local and distant solar systems, exoplanets, debris disks
- Stars $\mathbf{1 / 2} / \mathbf{3}$ : cool and hot stars in any stellar evolutionary phase, star formation, IDM
- Stellar Populations $\mathbf{1 / 2}$ : resolved stellar populations in the Galaxy and the nearby universe
- Galaxies $\mathbf{1 / 2 / 3}$ : stellar content of galaxies, ISM in galaxies, dynamics, galaxy morphology, galaxy evolution
- AGN \& IGM 1/2: QSOs, AGN, IGM, QSO absorption lines
- Cosmology $1 / 2$ : galaxy clusters, lensing, GRBs, deep surveys


## Panel Structure (cont.)

- Expect $\sim \mathbf{6 0 - 8 0}$ proposals per panel (Small/Medium/SNAP/ Regular AR)
- Expect $\mathbf{\sim} \mathbf{6 0}$ in the TAC (Large/Treasury/Legacy AR)
- Chairs for all 14 panels are being contacted
- Panel Chairs and three At-Large members will form the TAC chaired by Pat McCarthy (Carnegie)
- Each panel will have 9 panelists plus the Chair
- Candidate panelists are currently being considered
- We pay particular attention to subject balance, diversity, ESA representation, and balance between senior and junior astronomers

