

# HST Focus in SMOV4: Strategy for OTA adjustment & SI Focus

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#### ABSTRACT

This report documents in summary the observatory-level effort to produce optimal focus during Servicing Mission Observatory Verification 4 (SMOV4). It describes the extrapolation of monitor data available prior to SMOV4, and the resulting decision to adjust the HST Secondary Mirror +3  $\mu$ m away from the Primary in concert with the active optical alignment of WFC3 and COS in order to maximize image quality and confocality among the instruments. It also provides the historical HST focus dataset used for such long term focus predictions.

# 1. Introduction

HST focus has been monitored and adjusted throughout the Observatory's life. The monitoring normally consists of phase retrieval determinations of the amount of defocus present in images taken with one or more Science Instruments (SI).

Since deployment, the HST Optical Telescope Assembly (OTA) has shrunk by ~150  $\mu$ m (3x10<sup>-5</sup> of its length). This has resulted in ~25 Secondary Mirror (SM) adjustments away from the Primary Mirror (PM) to maintain the focus to the SIs. On top of this long term trend, operational temperature fluctuations produce focus changes which add "noise" to our estimates of the overall focal length of the OTA, some of which can be modeled (Lallo et al. 2005, DiNino et al. 2008).

This long-term shrinkage is ongoing, and upon entering SMOV4, our trending analysis of ACS phase retrieval data indicated a defocus of  $-2.5 \,\mu\text{m}$  at the SM ( $1\mu\text{m}$  @ SM = 6.1nm rms wave front error @ SIs). A negative focus value indicates the SM is too close to the PM. Comparable amounts of defocus (positive or negative) are frequently experienced in normal science operations due to thermal variations, but it was desired to remove any systematic offset during SMOV, in order to pass through Cycle 17 and 18 with the new and existing SIs as close as possible to their best focus.

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To accomplish this, STScI produced a plan to offset the best-focus position for COS and WFC3 found during their SMOV4 optical alignment by an amount equivalent to the  $-2.5 \mu m$  estimated amount of defocus at ACS. This would leave the old and new SIs confocal within our errors, but ~  $-2.5 \mu m$  out of focus. This activity was followed by a ~  $3.0 \mu m$  OTA refocus on 20 July 2009, in order to leave the SIs at an estimated ~  $+0.5 \mu m$  focus, slightly overshooting best focus, with the intention of passing through zero and evolving slightly negative over the Cycle 17/18 period.

For a detailed discussion of subsequent camera focus and confocality estimates, and focus trending over this period see Niemi & Lallo (2010).

#### 2. The Data

The body of HST focus-monitor data spans back to a few months after HST deployment on 24 April 1990. It comprises data from various CCD imaging instruments over the mission life. With some exceptions prior to WFPC2 commissioning in 1994, the data is typically obtained approximately monthly, with multiple exposures taken over some significant fraction of an HST orbit. The targets are normally one or more isolated, bright, non-saturating point sources, observed through filters with central wavelengths between ~0.4 and ~0.8  $\mu$ m. Proposal numbers for the more recent incarnations of the routine focus monitoring program are CAL/OTA 11877, 11316, 11020, and 10752.

From observations such as these, we perform a parametric phase retrieval analysis to recover presumed optical aberrations. See appendix 1 and Niemi & Lallo (2010). The phase retrieval is usually configured to solve for focus, but is also used to fit coma and astigmatism when appropriate to the nature of a particular study. A phase retrieval analysis is performed on each exposure in the visit, often sampling much of an orbit. This has allowed an accurate characterization of the temperature-driven focus changes during HST orbits, and has facilitated the development of a temperature-based focus model (DiNino et al 2008). Using data obtained from the HST engineering telemetry, the model takes temperatures from a number of sensors throughout HST to predict focus. A purely temperature-based model gives some success in explaining short-term focus variations but does not follow long term trends, most or all of which are not thermally driven nor reversible. This requires the addition of a secular function that is obtained by fitting such trends. For the data and analysis discussed here, when we apply the model, we do not include the long term secular component since this is what we are seeking to determine. When used, the model is only applied to data from 2003 onward since this is the time period over which it was developed and fit to observations.

For the purpose of OTA focus maintenance, we desire independent (uncorrelated) measurements, so the data is processed to replace the multiple values for each visit with the single mean value for that visit. The resulting data file records mean focus values and corresponding time stamps. Over the duration of our monitoring, we have utilized various cameras so the file contains results from WF/PC-1, WFPC2/PC, ACS/HRC, and ACS/WFC, as well as a small number of STIS imaging observations. Since there are small focus offsets (non-confocality) between the SIs, the data in this file has been processed to express results with respect to best ACS/WFC focus. This required adding  $-1.0 \mu m$  to ACS/HRC data,  $-1.4 \mu m$  to WFPC2/PC, and  $+5.7 \mu m$  to STIS image data. WF/PC-1 results were used "as-is" since a reliable estimate does not exist for confocality between WF/PC1 and the other imagers. The large offset for STIS is due to STIS focus being optimized at the slit plane which is not confocal with the image plane, and it is this image focus that is measured by phase retrieval as reported here.

The resulting file currently contains 363 independent measurements spanning from 1990 to March 2009. It is provided in appendix A3. ACS observations (WFC or HRC) are indicated by a "j" as the first character in the filename. WFPC2 data begins with a "u", WF/PC-1 is identified as "pre94", and in some instances (as well as for STIS) the points are named by the proposal and visit number or other descriptive label.

As already described, the HST OTA exhibits a long-term shrinkage, which is periodically corrected by commanding the SM away from the PM. Thus, in order to fit the long term behavior to make predictions used for effective control of the OTA focus, the measured focus values in this file must have accumulated offsets removed in order to reproduce the continuous focus trend rather than observed focus as maintained by periodic OTA refocusing. The history since 1994 of such mirror moves, their amounts and dates can be found at:

http://www.stsci.edu/hst/observatory/focus/mirrormoves.html

A peculiarity in the historical record should be pointed out here; mirror moves prior to 1994 are already captured in our input file. This can be seen in the first 44 rows identified as "pre94" (appendix 3). These values reflect the accumulated defocus up to 1994, rather than the measured values, so do not require piecewise processing to account for mirror moves prior to 1994. This processing of the file is accomplished by simply adding to each measured value the sum of all Secondary Mirror moves that occurred after the date of that observation. This recovers the smooth lifetime behavior whose predictive utility is discussed in the next section.

Finally, we maintain ancillary data files where aberrations such as coma and astigmatism are tracked, along with a number of other characteristics and correlations of interest, but for the purpose of OTA focus maintenance and prediction, the focus aberration is of interest, and the simple file reproduced in appendix 3 comprises the basic input.

#### 3. SMOV4 Focus Estimate

Plotting the mission-life focus data, after removing the OTA refocusings, produces a trend, with real focus variations creating "noise" about this otherwise monotonic behavior. In fig. 1 below we see this focus data evaluated prior to SM4. The best fitting double exponential,

Eqn. 1: 
$$\Delta SM = -5.04 + 56.26e^{-t/364.53} + 106.24e^{-t/2237.2}$$

(where  $\Delta$ SM is the change in SM position in microns and t is the number of days since 24 April 1990) produced a reasonable fit to the long-term data up to the cessation of HST science in May 2009 in anticipation of SMOV4 (last points plotted). The overall shrinkage of the HST OTA can be seen as ~150 µm. In practice, we have executed OTA refocusings listed in the provided URL to counteract this trend and maintain focus to typically within ~5 µm maximum at the SM.



**Figure 1:** Accumulated defocus of the OTA over the mission lifetime prior to SM4 (April 1990 - April 2009), and the best fitting double exponential.

Although the lifetime fit is illuminating for understanding long-term behavior, its rms residuals of 2.7  $\mu$ m (after focus model correction discussed in sect. 2) is too high to accurately base an OTA refocus determination. Systematic long-term displacements above and below the best fitting curve can be seen in figure 1. Because of this, fits to more recent data are often used when determining the overall OTA focus for the purposes of planning an OTA refocusing. This was the case for determining the SMOV4 focus. We examined only the portion of the focus data taken after the December 2002 OTA refocus. The data spans two OTA refocusings indicated in figure 2 by dotted vertical lines. These focus adjustments of +4.16  $\mu$ m (22 Dec 2004) and +5.34  $\mu$ m (31 July 2006) were added back into the observed focus values to produce a continuous dataset. This was then best fit with a single exponential,

Eqn. 2: 
$$\Delta SM = -6.16 + 201.64e^{-t/1754.4}$$

This fit had rms residuals of 1.7  $\mu$ m and predicted a focus value of -2.5  $\mu$ m for 1 August 2009 (rough date for the intended OTA refocus). A straight line fit had slightly higher rms residuals of 1.8  $\mu$ m and predicted -3.4  $\mu$ m on 1 August. For comparison, the double exponential fit to the lifetime data predicted -1.5  $\mu$ m with 2.7  $\mu$ m rms residuals, clearly underestimating the amount of actual defocus seen in monitoring data for the past 1.5 years. An exponential fit performed without model correcting the data increased the rms residuals to 2.0  $\mu$ m but did not result

in a significantly different fit or focus prediction, confirming that the temperature-based focus model was not introducing a systematic offset over the timescale being considered.

It is interesting to note that the constant term in both exponential fits given in equations 1 & 2 imply a total future shrinkage of the HST OTA of only another 5 to 6  $\mu$ m.

**Figure 2:** Accumulated defocus of the OTA from 2 December 2002 up to and including SMOV data points (identified with red arrows). Data have been model-corrected. Best fitting exponential predicted a focus of –2.5 microns on 1 August 2009. Best fitting straight line is dashed and included for reference.



Focus Trend Since Dec 2002 Mirror Move

The best estimate of  $-2.5 \,\mu$ m of defocus at ACS implied an equivalent amount of defocus at the STIS slit plane, since STIS and ACS were both originally focused to be confocal with WFPC2 and thus should remain confocal within the errors to each other. Early SMOV calibration results included a phase retrieval measurement of a STIS CCD OII image, which indicated that, like ACS, the resurrected SI had not seen a noticeable focus change since the last time it was operational. After correcting by the 5.7  $\mu$ m to express this measurement at the STIS slit focus (see sect. 2), this point was seen to be slightly more positive than expected, though it was not "out of family" with other outliers seen throughout our monitoring history. Additionally, STIS CCD phase retrieval results have carried higher errors than seen with the more frequently utilized imagers due to coarser calibrations of the algorithms consistent with their less frequent application to that SI, along with uncertainty associated with the focus offset between the imaging and

spectral modes. The STIS image-plane focus check made during SMOV was thus performed more as a gross indicator of any possibly new large focus offset within the instrument, rather than as a sensitive measurement of OTA focus state.

## 4. SMOV4 Focus Adjustments

The SMOV focusing plan that emerged from this estimate of  $-2.5 \ \mu m$  at the SM was as follows: offset WFC3 & COS foci during SI SMOV fine alignment by the equivalent of  $-2.5 \ \mu m$  to achieve confocality with ACS/WFC, then follow up with a +3.0  $\mu m$  OTA refocus (prior to WFC3 PSF specification tests and Early Release Science). This OTA adjustment would then leave the SIs at ~ +0.5  $\mu m$  focus, intentionally slightly overshooting best focus. Predictions using the best fitting exponential (eqn. 2) put best focus for the SIs near mid cycle 17 (summer 2010). Focus monitoring throughout cycle 17 would then establish the success to which confocality and best focus was achieved.

During SMOV4, between 7–9 July 2009, WFC3 UVIS & IR Fine Optical Alignment proposals 11434 & 11435, and COS NUV Optical alignment 11469 executed. These programs featured near-real-time "ops requests" to actively command the instruments' optics into their final positions for best overall image quality based upon analysis performed on images obtained earlier in the same programs. The final uploaded positions for the SI optics were biased at that time by the equivalent of  $-2.5 \mu m$  focus. The actual commanded values associated with the optical alignment of these instruments are described in Hartig (2009a, 2009b, 2010).

On 20 July 2009, a Secondary Mirror move of  $+3.0 \,\mu\text{m}$  away from the PM was executed by engineers at GSFC. The actual commanded amount of this move was  $+2.97 \,\mu\text{m}$  due to SM actuator quantization. This OTA refocus was smaller than most of the routine shrinkage-compensating adjustments that had been applied in the past. However, WFC3's imaging performance is believed to require tighter focus control than other HST imaging instruments, and the initial desire to stay within 2  $\mu$ m of defocus was adopted. After experience with WFC3 images in cycle 17 there is some evidence that a tighter focus control (e.g.  $+/-1 \,\mu$ m) may realize measurable improvement. This will be investigated further in cycle 18.

The timeline and decision tree specific to SMOV4 focus that was developed to prescribe contingency plans for focusing new or existing SIs is reproduced in appendix A2. In fact, other than the described biasing of the new-SIs followed by the OTA adjustment to optimize focus and confocality, no other active focusing activities were performed during (and since) SMOV4. The OTA and the repaired SIs (STIS and ACS) were found to not have experienced any gross changes across the Servicing Mission boundary, or while inactive prior to repair.

## 5. Post-SMOV4 Focus Monitoring

Since the SMOV focusing activities, proposal CAL/OTA 11877 has been used as the primary monitor of overall focus. The data are analyzed using phase retrieval analysis of images taken with WFC3/UVIS and ACS/WFC, either in parallel or within the same orbit. The phase retrieval

code requires SI-specific information in order to accurately characterize apparent optical aberrations like focus, and considerable investment was made during cycle 17 to adapt and calibrate the algorithms for WFC3. Niemi et al. (2010) give a thorough review of this.

Figure 3 plots focus monitoring data obtained since the SMOV4 period, illustrating the success of our SMOV4 focus initiative at restoring overall focus and producing practically confocal main imagers. A full treatment of findings & results since the SMOV4 focusing initiative can be found in Niemi & Lallo (2010).

Figure 3: Defocus of the OTA from 2 December 2002 up to June 2010, including data taken since the 2009 Secondary Mirror Move in late SMOV. Evident is the resulting overall state of WFC3/ UVIS and ACS/WFC near best focus through cycle 17 (mirror moves prior to the last are indicated but have been removed to show a continuous trend).



# References

Burrows, C., Hubble Space Telescope Optical Telescope Assembly Handbook, V1.0, 5/1990

- Di Nino, D., Makidon, R. B., Lallo, M., Sahu, K., Sirianni, M., Casertano, S., *HST Focus Variations with Temperature*, ISR ACS 2008-03, 5/2008
- Hartig, G., Dressel, L., Delker, T., WFC3 SMOV Programs 11424, 11434: UVIS Channel Onorbit Alignment, ISR WFC3 2009-45, 12/2009
- Hartig, G., Dressel, L., Delker, T., WFC3 SMOV Programs 11425, 11435: IR Channel On-orbit Alignment, ISR WFC3 2009-46, 12/2009
- Hartig, G., Delker, T., Keyes, C., SMOV: COS NUV On-orbit Optical Alignment, ISR COS 2010-04, 2/2010
- Krist, J. E., Burrows, C. J., *Phase Retrieval analysis of pre- and post-repair Hubble Space Telescope images*, Appl. Opt. 34, 4951–4963 {1995}
- Lallo, M., Makidon, R. B., Casertano, S., Gilliland, R., Stys, J., *HST Temporal Optical Behavior* & Current Focus Status, ISR TEL 2005-03, 10/2005
- Niemi, S.-M. Lallo, M., Phase Retrieval to Monitor HST Focus: II. First Results Post-Servicing Mission 4, ISR TEL 2010-02, 10/2010
- Niemi, S.-M. Lallo, M., Hartig, G., Cox, C., *Phase Retrieval to Monitor HST Focus: I. WFC3* UVIS Software Implementation, ISR TEL 2010-01, 7/2010

## Appendix

#### A1. Phase Retrieval

Routine optical monitoring of HST is performed using IDL code to perform parametric (model-fitting) phase retrieval of a nearly in-focus PSF. The technique iteratively generates model PSFs and compares them with the observed data (Krist & Burrows 1995). The wavefront is characterized by the series of Mahajan Zernike polynomials modified for the 0.33 obscuration applicable to HST. They are listed below and discussed in greater detail in the OTA handbook (Burrows 1990). Current application and adaptation for WFC3 is described by Niemi et al. (2010).

The phase retrieval process produces an estimated wavefront described by the series:

(1) 
$$W(r, \theta) = \sum_{n} c_n \alpha_n Z_n$$

where  $Z_n$  are the Zernike polynomials,  $\alpha_n$  are the normalization factors sometimes seen included as part of the Zernike polynomial, and  $C_n$  are the solved-for coefficients representing wavefront error in microns. For n = 4 to 8,  $\alpha_n Z_n$  are given below:

$$\alpha_4 Z_4 = 3.89 (r^2 - 0.55445)$$
  

$$\alpha_5 Z_5 = 2.31 r^2 \cos 2\theta$$
  

$$\alpha_6 Z_6 = 2.31 r^2 \sin 2\theta$$
  

$$\alpha_7 Z_7 = 8.33 (r^3 - 0.673796r) \cos \theta$$
  

$$\alpha_8 Z_8 = 8.33 (r^3 - 0.673796r) \sin \theta$$

In the case of focus (n = 4) the  $c_4$  coefficient can be expressed as:

(2) 
$$c_4 = \frac{1+m^2}{\alpha_4 8F^2} \Delta_{\rm SM}$$

where F = 24 = HST focal ratio,  $\alpha_4 = 3.89$ ,  $m = \text{magnification} = F/f_{primary} = 24/2.3 = 10.43$ .

Thus  $c_4 = 0.0061 \cdot \Delta_{SM}$ 

Total rms wavefront error is  $(\Sigma c_n^2)^{1/2}$ 

## A2. SMOV4 Focus Decision Tree



# A3. HST Focus History Data File

Measured defocus, expressed with respect to best ACS/WFC focus (see section 2). No focus model correction.

ObsName,	Date, Mod	ified .	JD, Microns	of	Secondary	Mirror	despace
pre94	pre94	48110	-1.00				
pre94	pre94	48183	-23.00				
pre94	pre94	48219	-32.00				
pre94	pre94	48231	-28.00				
pre94	pre94	48292	-39.00				
pre94	pre94	48371	-51.00				
pre94	pre94	48402	-48.00				
pre94	pre94	48486	-57.00				
pre94	pre94	48511	-61.00				
pre94	pre94	48513	-55.00				
pre94	pre94	48536	-59.00				
pre94	pre94	48566	-58.00				
pre94	pre94	48584	-61.00				
pre94	pre94	48591	-57.00				
pre94	pre94	48621	-51.00				
pre94	pre94	48656	-66.00				
pre94	pre94	48671	-69.00				
pre94	pre94	48695	-73.00				
pre94	pre94	48698	-74.00				
pre94	pre94	48730	-66.00				
pre94	pre94	48761	-72.00				
pre94	pre94	48762	-72.00				
pre94	pre94	48/86	-72.00				
pre94	pre94	48803	-/6.00				
pre94	pre94	48840	-/8.00				
pre94	pre94	488/6	-//.00				
pre94	pre94	48901	-/8.00				
pre94	pre94	48909	-/3.00				
pre94	pre94	48936	-70.50				
pre94	pre94	48956	-78.00				
pre94	pre94	48986	- /8.00				
pre94	pre94	49016	-81.00				
pre94	pre94	49036	- / / . 00				
pre94	pre94	49000	-01.00				
prega	prega	49120	-84 00				
pre94	pre94	49161	-86.00				
pre94	pre94	49208	-86.00				
pre94	pre94	49236	-87 00				
pre94	pre94	49261	-90.00				
pre94	pre94	49279	-91.00				
pre94	pre94	49281	-89.00				
pre94	pre94	49301	-93.00				
pre94	pre94	49314	-88.00				
u2a70305t	8/03/94	49419	-2.40				
u2a70605t	20/03/94	49431	-0.50				
u2a70905t	25/03/94	49436	-0.70				
u2a70c05p	1/04/94	49443	-1.10				
u2a70i05t	21/04/94	49463	-2.30				
u2a70105t	1/05/94	49473	-12.40				
u2a70o05t	8/06/94	49511	-4.10				
u2a70r05t	14/06/94	49517	-3.50				
u2a70u05t	4/07/94	49537	1.10				
u2a70x05t	16/07/94	49549	-0.50				
u2a71005p	25/07/94	49558	-0.50				
1							

u2a71305t	4/08/94	49568	-0.70
u2a71605t	23/08/94	49587	-1.80
u2a71905t	2/09/94	49597	-0.30
112a71c05t	21/09/94	49616	-1 70
$112 = 71 \pm 0.05 \pm 0.0$	26/09/94	49621	-1 40
u2a711050	20/00/04	40021	2.00
uza/1105t	20/10/94	49645	-3.00
u2a/1105t	25/10/94	49650	-4.90
u2a71o05t	14/11/94	49670	-4.10
u2a71r05t	21/11/94	49677	-2.00
u2a71u05t	12/12/94	49698	-2.70
u2a71x05t	20/12/94	49706	-5.20
112a72605t	11/01/95	49728	-4 70
u2a72005t	21/01/05	10730	-0.90
u2a72905t	21/01/9J 11/02/05	49730	-0.00
uza/zcust	11/02/95	49759	-1.90
u2a/2105t	13/02/95	49/61	-3.00
u2n10203p	7/03/95	49783	1.20
u2n10403t	13/03/95	49789	-1.30
u2o00501t	7/05/95	49844	-1.50
u2s61101t	27/07/95	49925	-2.70
112s61201t	6/08/95	49935	-3.90
112861301+	21/08/95	49950	-6.30
u23013010	21/00/05	40000	2.20
u2S614010	31/00/93	49960	-2.20
u2s61501t	18/09/95	49978	1.00
u2s61601t	28/09/95	49988	-0.20
u2s61701t	16/10/95	50006	0.10
u2s61801t	29/10/95	50019	-1.20
u2s62101t	13/11/95	50034	-1.60
112s62201t	22/11/95	50043	-0.40
112s62301t	12/12/95	50063	-2 00
112662401+	18/12/95	50069	0 20
u2502401t	10/12/95	50009	1 00
u2S62501L	10/01/96	50092	-1.80
u2s62601t	15/01/96	50097	-5.20
u2s62701t	8/02/96	50121	-0.70
u2s62801t	15/02/96	50128	-2.90
u2s63101t	6/03/96	50148	-6.30
u2s63201t	17/03/96	50159	-0.60
u2s63301t	28/03/96	50170	-0.50
112s63401t	7/04/96	50180	2 90
u20001010	2/05/06	50205	0 70
u2505501t	2/05/90	50205	2 20
u2S63601L	7/05/96	50210	-2.20
u2s63/01t	27/05/96	50230	0.80
u2s63801t	6/06/96	50240	3.30
u2s63901t	17/06/96	50251	-0.20
u2s64001t	24/06/96	50258	2.20
u3dy0103t	27/07/96	50291	0.17
u3dv0203t	30/07/96	50294	-0.13
113dv0403t	8/09/96	50334	-2.57
113dv0503+	17/09/96	50343	-0 77
u3uy0303t	20/00/06	50345	0.77
u3dy0603L	29/09/96	50355	-0.87
u3dy0702t	15/10/96	50371	-1.85
u3dy0802t	19/10/96	50375	-1.45
u3dy0904t	11/11/96	50398	6.50
u3dy1004t	21/11/96	50408	1.23
u3dv1102t	9/12/96	50426	4.30
u3dv1403t	9/01/97	50457	4.27
u3dy1502t	5/02/97	50484	2 25
13dy1600+	0/02/9/ 0/09/07	50100	J.JJ / EE
u3uy1002t	9/UZ/9/	JU488	4.00
u3srU102r	23/02/9/	50502	3.90
uJsr0202r	25/02/97	50504	3.45
u3sa1107r	27/02/97	50506	3.92
u3sa1308r	28/02/97	50507	4.31
u3sr2602r	3/03/97	50510	4.13
u3t9810hr	4/03/97	50511	4.00

113sa0401r	5/03/97	50512	4 10
7021 15	6/03/97	50513	2 34
113er3002r	9/03/97	50516	_0 70
u33130021	12/02/07	50510	0.75
u35132021	15/05/9/	50520	-0.75
u3sr3402r	16/03/9/	50523	-0.05
u3sr3602r	20/03/97	50527	-0.45
u3sr3802r	21/03/97	50528	-0.30
u3sr3902r	24/03/97	50531	-1.50
u3sr4002r	27/03/97	50534	-2.60
u3sr4102r	30/03/97	50537	-5.80
u3dv1901r	31/03/97	50538	-2 90
u3ar4302r	1/01/07	50542	0 05
USS145021	4/04/9/	50542	0.95
u3dy2001r	5/04/9/	50543	1.40
u3dy2103m	23/04/9/	50561	4.10
u3dy2203m	2/05/97	50570	-1.40
u3dy2301m	14/05/97	50582	-1.40
u3dy2401m	22/05/97	50590	-0.50
u3dv250dm	6/06/97	50605	-0.82
u3dv260dm	8/06/97	50607	0.64
113dv2701m	24/06/97	50623	-1 00
113dy2801m	28/06/97	50623	_1 50
u3uy200111	20/00/97	50027	-1.50
u42w0103m	22/07/97	50651	0.80
u42w0203m	25/0//9/	50654	3.4/
u42w0402m	18/08/97	50678	-3.15
u42w0503r	17/09/97	50708	-0.10
u42w0603r	21/09/97	50712	-1.63
u42w0803r	18/10/97	50739	0.67
u42w0903r	10/11/97	50762	-4.13
1142w100dr	17/11/97	50769	-2.38
1142w1203r	14/12/97	50796	-1 87
u42w12031	E/01/00	50750	2 1 2
u42w13031	5/01/90	50010	-2.13
u42W1603r	6/02/98	50850	0.57
u42w1/03r	27/02/98	508/1	-1./3
u42w1803r	8/03/98	50880	-0.20
u42w2002r	5/04/98	50908	-0.65
u42w2102r	26/04/98	50929	-1.15
u42w2202r	3/05/98	50936	0.25
u42w2403r	31/05/98	50964	-3.23
1142w5604r	15/06/98	50979	16.10
1142w5801r	17/06/98	50981	12 30
11/2w580/r	18/06/98	50982	10 00
u42w50041	20/06/00	50002	14 60
u42w55051	20/00/90	50992	14.00
u4an030gr	1/0//98	50995	-2.70
u42w2803r	26/07/98	51020	-1.3/
u42w2903r	20/08/98	51045	0.57
u42w3002r	22/08/98	51047	-2.05
u42w3203r	16/09/98	51072	-1.97
u42w330dr	13/10/98	51099	-2.00
u42w3403r	17/10/98	51103	-1.23
u42w3602r	15/11/98	51132	0.90
1142w3703r	7/12/98	51154	-2 33
11/2w3803r	13/12/08	51160	_1 35
u42w30031	2/01/00	51100	1.55
u42w4003r	2/01/99	51100	1 00
u42w4103r	27/01/99	51205	-1.80
u42w4203r	29/01/99	51207	-2.70
u42w4401r	27/02/99	51236	2.70
u42w4503r	22/03/99	51259	-6.20
u42w4602r	26/03/99	51263	-1.95
u42w4803r	25/04/99	51293	-2.10
u5if1107r	18/05/99	51316	-3.72
u5if1807r	20/05/99	51318	-4.56
u5if2809r	20/06/99	51349	-2.60
115if3107r	12/07/00	51 271	_5 20
U) T T J T U / T	12/01/22	J T J I T	5.50

115if3807r	15/07/99	51374	-4 30
LJ1130071	11/00/00	51371	1.50
u5114809r	11/08/99	51401	-4.60
u5if5808r	9/09/99	51430	-7.35
	17/00/00	E1 4 2 0	0 4 5
u5j/0502r	1//09/99	51438	0.45
u5if6107r	4/10/99	51455	-4.80
115; f(007m	0/10/00	F1460	7 20
UJII000/I	9/10/99	51460	-7.20
u5if7808r	4/11/99	51486	-4.10
11617050ar	28/12/00	51540	-3 54
uuiresugi	20/12/99	51540	5.54
u617270gr	30/12/99	51541	-2.93
11617290ar	30/12/99	51542	-3 27
uor7200gr	50/12/55	51542	5.27
u6013108r	3/01/00	51546	-5.29
1160f3208r	4/01/00	51547	-4.51
0.00202002	I/01/00	51540	2.00
u6013308r	5/01/00	51548	-3.89
u6175104r	16/01/00	51559	-2.20
1160n1209r	10/01/00	51561	_0 07
u001112001	10/01/00	JI J01	-0.07
u60n1408r	20/01/00	51563	-2.35
116176105r	30/01/00	51573	-4 10
u01701051	50/01/00	51575	4.10
u6176805r	31/01/00	51574	-0.83
1160f6108r	10/02/00	51584	-2.02
ucororoor	10/02/00	51501	2.02
u5154109r	22/02/00	21230	-3.60
u51s4809r	27/02/00	51601	-5.80
1151c5000m	27/02/00	51620	2 00
u)1800091	21/03/00	31030	-2.00
u5ls6106r	17/04/00	51651	-3.03
1151c6807r	19/01/00	51653	_1 88
uJIS00071	19/04/00	51055	4.00
u66k010fr	25/04/00	51659	-6.11
1166k020fr	28/04/00	51662	-3.55
C1 7100	1 5 / 0 5 / 0 0	51002	0.00
u51s/109r	15/05/00	516/9	-2.60
u51s7809r	21/05/00	51685	-3.50
115100000r	11/06/00	51700	_3 00
UJI500001	14/00/00	51709	-3.00
u68w010gr	30/06/00	51725	-1.26
1151s9109r	10/07/00	51735	0 90
	10/07/00	51735	0.00
u51s9809r	15/0//00	51/40	0.00
u68w020ar	30/07/00	51755	-1.34
	7/00/00	E17C2	0 25
U08XIIUCI	//08/00	21/03	-0.25
u68x180cr	15/08/00	51771	2.12
1168w030ar	26/08/00	51782	-2 25
uoowoooer	20/00/00	51702	2.25
u68x280br	10/09/00	51797	-0.30
1168w040er	19/09/00	51806	-0.56
	7/10/00	E1004	0.15
008X380CL	//10/00	51824	0.15
u68w050er	19/10/00	51836	-0.70
116b7080cr	10/11/00	51858	-3 65
	10/11/00	51050	5.05
u68w060er	16/11/00	51864	-5.33
11693580cr	29/11/00	51877	-1.15
	10/10/00	E1000	0 50
U08W0/0/1	12/12/00	21890	0.50
u68w070er	13/12/00	51891	-1.39
116936107r	25/12/00	51903	-4 80
00000071	23/12/00	51000	4.00
u693680cr	31/12/00	51909	0.60
1168w080ar	13/01/01	51922	-0 70
coo710	10/01/01	51022	0.05
u693/10Cr	22/01/01	21931	-2.95
u693780cr	28/01/01	51937	-1.20
116911000ar	10/02/01	51050	-0 49
uoowobogi	10/02/01	51950	-0.49
u6938107r	19/02/01	51959	-1.80
11693880br	24/02/01	51964	-0 10
CO-100	- 1, 02, 01	5107 F1072	C • ± 0
ubswi00gr	5/03/01	21813	-2.88
u69q110cr	19/03/01	51987	-1.45
1169019000	21/02/01	51002	_3 67
uusqroucr	24/UJ/UI	51992	5.0/
u68wl10gr	2/04/01	52001	-9.51
1169a280hr	22/04/01	52021	-0 62
		E 2 0 4 2	1 05
npadainci	14/05/01	J∠U43	-1.95
u69q380cr	17/05/01	52046	-4.05
11681130~~	28/05/01	52057	-2 60
uuowijugr	20/00/01	52057	-2.09
u69q480cr	17/06/01	52077	-1.50
1168w140ar	1/07/01	52091	-2 12
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u6hk010gr	21/07/01	52111	-0.49
u6hk020gm	18/08/01	52139	-2.35
u6hk030em	16/09/01	52168	-2.92
116hk040em	12/10/01	52195	0.39
116bk050em	11/11/01	52224	-3 57
uchkosocm	0/12/01	52221	_0 55
	9/12/01	52255	-0.55
u6hk0/0gr	6/01/02	52280	-1.66
u6hk080em	3/02/02	52308	-0.08
u6hk080gm	3/02/02	52309	-1.70
u8cz2108r	24/03/02	52357	-4.50
u8cz2508m	24/03/02	52358	-2.96
118c72708r	27/03/02	52360	-2 93
118 CV 01 07m	29/03/02	52362	3 16
u0CX0107m	29/03/02	52362	2.40
	29/03/02	52505	-2.70
u8CXU2UCm	3/04/02	52367	-1.55
u8cx030cm	29/04/02	52393	-3.60
u6hk120fm	28/05/02	52422	-7.31
u6hk130gm	30/06/02	52455	-1.95
u8hr0103m	12/07/02	52468	-4.83
u8hr020cm	16/08/02	52502	-5.59
i8hr03vva	18/09/02	52535	-5.66
118br030cm	18/09/02	52536	-4 55
uonii o Joem	10/00/02	52550	11 10
	10/10/02	52565 .	-11.49
J8nr05xuq	20/11/02	52598	-6.19
u8hr050cm	20/11/02	52599	-6.02
j8hr06n2q	14/12/02	52622	1.52
u8hr060cm	14/12/02	52623	0.83
j8hr07u2q	22/01/03	52661	1.01
u8hr070cm	22/01/03	52662	1.49
i8hr08hfa	24/02/03	52694	1.01
ughr080cm	24/02/03	52695	1 50
- Chroodaa	21/02/03	52720	-4 30
J 8111 0 9 0 9 9 9 9 9	31/03/03	52729	-4.50
ushr090cm	31/03/03	52730	-3./5
j8hr10r5q	29/04/03	52758	-1.87
u8hr100cm	29/04/03	52759	-1.11
j8hr12dfq	07/07/03	52827	-2.65
u8hr130cm	16/08/03	52867	-3.08
u8hr140cm	16/09/03	52898	-8.87
u8ta010cm	25/10/03	52937	-1.15
118±a020cm	15/11/03	52958	-1 74
118±0030cm	19/12/03	52992	-0 36
- 0+ - 0 4 0 -	10/01/04	52012	2 21
J8L404094	10/01/04	53013	-2.51
u8tq040cm	10/01/04	53014	-2.58
j8tq05j5q	08/02/04	53042	-1.49
u8tq050cm	8/02/04	53043	-1.16
u8tq060cm	7/03/04	53071	-3.97
u8tq070cm	4/04/04	53099	-1.22
j8tq08raq	02/05/04	53127	-4.65
i8tal0cla	27/06/04	53183	-3.74
i8ta11f6a	24/07/04	53210	-5 47
1000411104	21/00/04	53210	0 22
	21/00/04	53250	-0.55
ustq130cm	20/09/04	53268	-/.15
u94g010cm	18/10/04	53296	-5.35
u94g020cm	17/11/04	53326	-6.11
u94g030cm	13/12/04	53352	-4.29
u94g040cm	14/01/05	53384	-1.20
u94q050cm	13/02/05	53414	-1.13
1194a060cm	10/03/05	53439	-2 65
1194a070cm	6/04/05	53466	-0 24
- 01~00f	01/05/05	53400	_2 24
J949U8Ieq	01/05/05	53491 535491	-3.34
J9alu2skq	22/05/05	53512	-3.62
J94g10hqq	29/05/05	53519	-2.52
j94g10hqq	26/06/05	53547	-3.05

j94g11k1q	27/07/05	53578	-0.62
u94g120cm	26/08/05	53608	-2.33
j94q07p7q	09/09/05	53622	-4.62
u9i4020cm	1/10/05	53644	0.53
u9i4040cm	11/12/05	53715	-1.00
i9i405mba	30/12/05	53734	-3 89
- 0 i 10 6 aug	22/01/06	53757	-3 14
j914005wq	22/01/00	53737	-3.14
J9140712q	23/02/06	53769	-0.70
J91408CIq	13/03/06	53807	-1.80
j9i409djq	19/04/06	53844	-8.40
j9i410a2q	25/05/06	53881	-4.04
j9i4llubq	09/06/06	53895	-3.99
j9i412jrq	05/07/06	53921	-5.05
j9i413bwq	07/08/06	53954	-1.54
u9i4140cm	9/09/06	53987	-1.11
u9vg010cm	6/11/06	54045	0.67
j9vg02d0q	27/11/06	54066	3.60
j9vg03usq	20/01/07	54120	1.02
j9vg03w0q	21/01/07	54121	-0.76
u9vg040mm	6/03/07	54165	-3.18
u9vq060hm	25/04/07	54215	-2.81
u9va070pm	28/05/07	54248	-4.02
ma080pveu	7/07/07	54288	-1.21
119va090km	7/08/07	54319	-1.12
119va100cm	18/09/07	54361	-1 58
119va110om	6/11/07	54410	-2 95
1197791200m	2/12/07	54436	2.90
119779120pm	15/12/07	54449	-0 48
11979130pm	17/01/08	54449	1 7/
u9vg140pm	20/01/00	54402	1./4
u9vg150pm	20/01/00	54495	1 1 2
u9vg100pm	2/03/00	54527	-1.13
u9vg170pm	10/03/00	54541	-0.10
u9vg180pm	20/04/08	54576	1.00
uastuluom	1/05/08	54587	-1.96
ua5t020om	13/06/08	54630	0.36
ua5t030nm	23/06/08	54640	-0.59
ua5t040om	31/07/08	54678	-0.93
ua5t050om	15/08/08	54693	-3.27
ua5t160om	24/09/08	54733	-3.07
ua5t180nm	19/12/08	54819	-2.87
ua5t270om	1/01/09	54832	-1.48
ua5t190nm	28/01/09	54859	-2.93
ua5t200nm	23/02/09	54885	-5.18
ua5t210mn	22/03/09	54912	-3.49
ua5t220om	3/04/09	54924	-4.07
ua5t280om	6/04/09	54927	-2.87
ua5t290om	10/04/09	54931	-1.53
ua5t300om	23/04/09	54944	-1.54
ua5t310om	25/04/09	54946	-1.39
ua5t320om	1/05/09	54952	-1.33
ua5t330om	7/05/09	54958	-1.40
STIS SMOV	4/06/09	54986	-2.30
ACS SMOV	14/06/09	54996	-0.62
ACS SMOV	9/07/09	55021	-2.20
ACS 11501	24/07/09	55036	0.40
	- 1/ 0 / / 0 0	00000	0.10