



11174 - A Spitzer/X-ray candidate cluster at $z > 2$: NICMOS imaging

Cycle: 16, Proposal Category: GO

(Availability Mode: SUPPORTED)

INVESTIGATORS

<i>Name</i>	<i>Institution</i>	<i>E-Mail</i>
Dr. Emanuele Daddi (PI) (ESA Member)	Commissariat a l'Energie Atomique (CEA)	edaddi@cea.fr
Prof. Alvio Renzini (CoI) (ESA Member)	Osservatorio Astronomico di Padova	arenzini@pd.astro.it
Prof. Nobuo Arimoto (CoI)	National Astronomical Observatory of Japan (NAOJ)	arimoto@optik.mtk.nao.ac.jp
Dr. Marcella Brusa (CoI) (ESA Member)	Max-Planck-Institut fur extraterrestrische Physik	marcella@mpe.mpg.de
Dr. Ranga-Ram Chary (CoI) (AdminUSPI)	California Institute of Technology	rchary@caltech.edu
Prof. Andrea Cimatti (CoI) (ESA Member)	Universita di Bologna	cimatti@arcetri.astro.it
Dr. Alexis Finoguenov (CoI)	University of Maryland	alexis@mpe.mpg.de
Dr. Xu Kong (CoI)	University of Science and Technology of China	xkong@ustc.edu.cn
Dr. Masato Onodera (CoI)	Yonsei University, Korea	monodera@galaxy.yonsei.ac.kr

VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) DADDIF-GAL-CLUS	NIC3	4	18-Jan-2008 02:01:27.0	yes
02	(1) DADDIF-GAL-CLUS	NIC3	3	18-Jan-2008 02:01:33.0	yes

7 Total Orbits Used

ABSTRACT

We propose deep H-band imaging with NICMOS of a remarkable $z > 2$ cluster of galaxy candidate. Over a 1000 arcmin² field imaged with Spitzer's IRAC and MIPS we have discovered a compact ($< 30''$ diameter) concentration of extremely red galaxies with a factor of > 40 overdensity over the adjacent field. Among these galaxies for which we can derive meaningful photometric redshifts, 17 are consistent with $z_{\text{phot}} = 2-2.5$, making very likely that the concentration is a real cluster at such high redshift. This is further supported by a 3.5 sigma detection of extended X-Ray emission on Newton-XMM data, by a likely color-magnitude sequence of red galaxies, and by the presence of a giant galaxy consistent with a BCG at the cluster redshift. While spectroscopic confirmation of the cluster might result prohibitive with current facilities, HST high resolution imaging will allow us to gain crucial information for the study and scientific exploitation of this hot gas hosting, record high- z cluster of galaxies. The HST high resolution observations will allow us to unveil the rest frame optical morphologies of the galaxies and confirm the presence of ellipticals in the structure, detect and characterize the color-magnitude relation, measure their effective radii and construct their Kormendy relation for the passively evolving subsample, improve the photometric redshift estimates to confirm the real cluster nature of the structure, estimate stellar masses and check for possible deviations from the local mass-size relation, search for mergers and AGNs, and establish a cluster benchmark for cluster-field comparisons at this highest redshift.

OBSERVING DESCRIPTION

Our goal is to obtain high resolution and high S/N imaging of the $z > 2$ cluster member candidates in the structure. Given the extremely red SEDs of the sources, the most efficient mean by far is observing in the near-IR with NICMOS in the F160W band. This also allows to obtain data at optical rest-frame wavelengths (between B and V bands rest frame for $z \sim 2-2.5$).

The angular size of the galaxy overdensity is about $30''$, which can be easily covered with a single pointing with the NICMOS3 camera. With a FWHM of about $0.2''$, corresponding to about 1.7 kpc (proper size) at $z = 2-2.5$, this will suffice to distinguish disks and irregular (i.e. dusty or merging) galaxies from spheroids and to constrain sizes to a limit of ~ 1 kpc. This size limit is comparable to the size of the smallest field $z > 1.4$

passive ellipticals (Daddi et al. 2005).

The IRAC brightest galaxy in the structure (candidate BCG) has $H=22.2$ AB, derived from the best fit SED (Fig.~\ref{fig:x}). The cluster member candidates are up to 3.5 magnitudes fainter in IRAC than the BCG (see Fig.~\ref{fig:1}). Conservatively assuming they are as red as the BCG in $(H-m_{\{3.6\mu m\}})$ color, we need to reach $H=25$ AB in order to secure high quality imaging of most cluster members (14/17 to this limit).

We have used the NICMOS exposure time calculator for estimating the required integration time to reach $S/N>10$ detection of a $H=25$ source. We consider a $1''$ diameter aperture and assume that at least 70\% of the flux falls within the aperture. We consider 2 sub exposures of 1300s per orbit (52 minutes visibility).

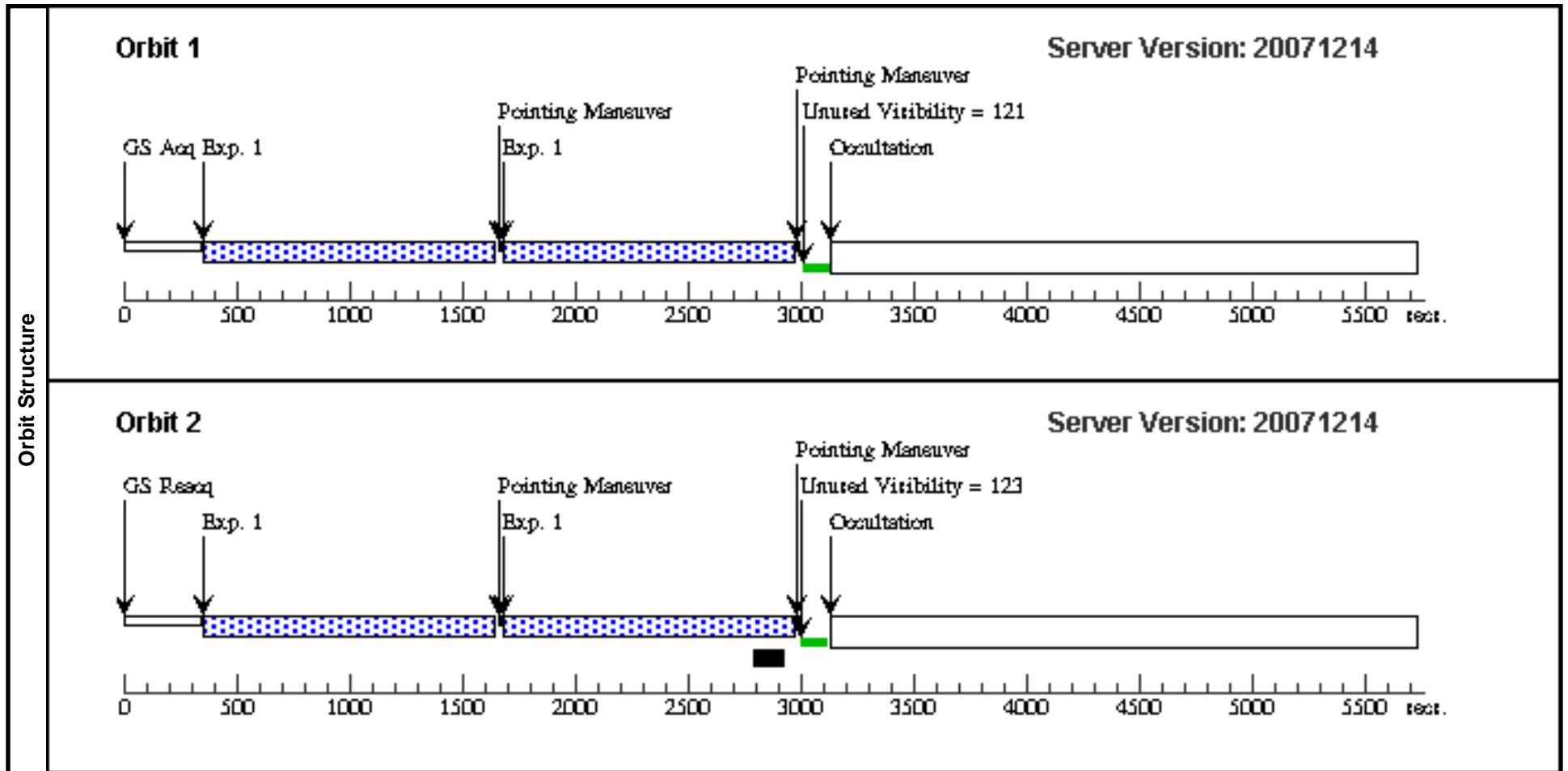
We find that in total 7 orbits are necessary to reach the required S/N ratio.

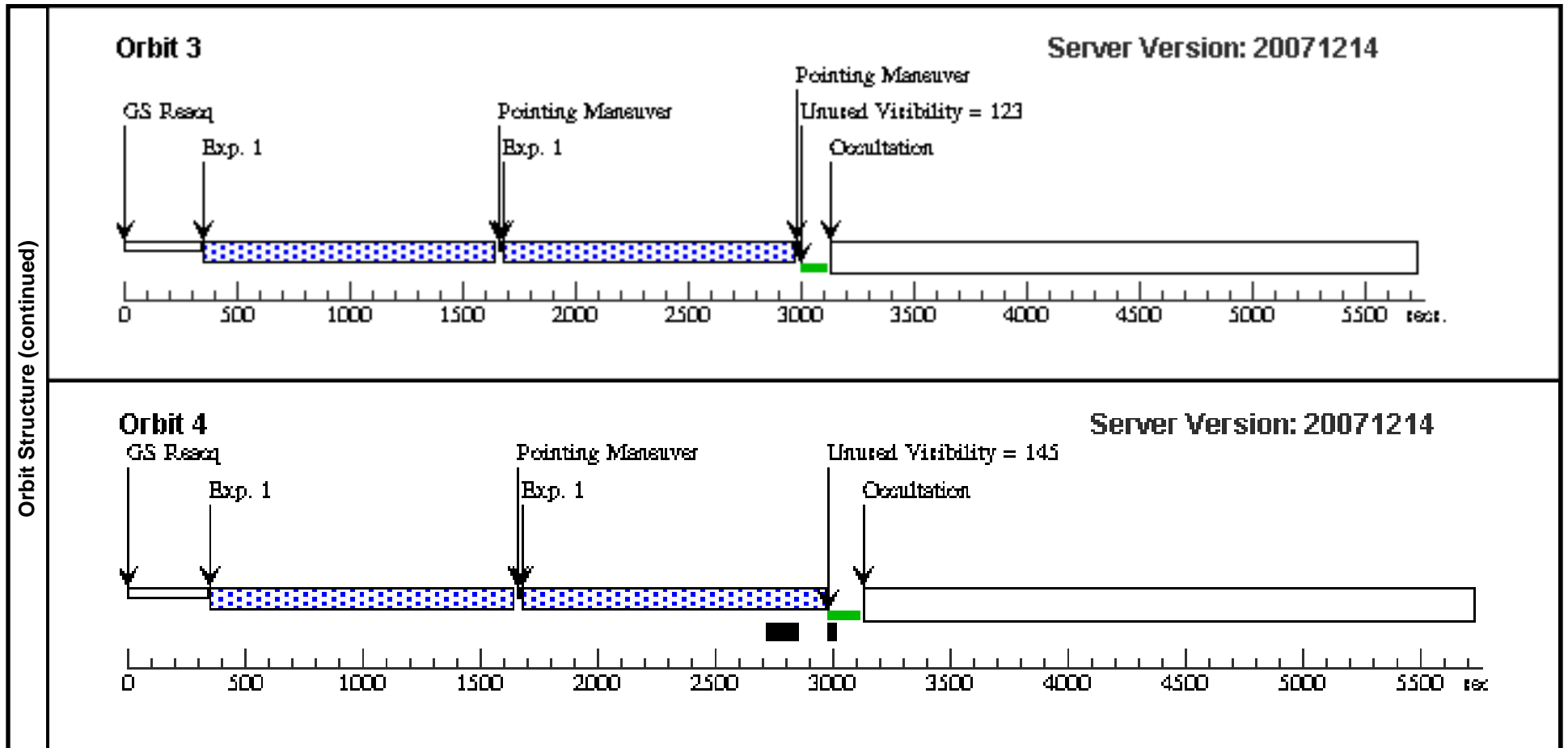
We will split the observations in a 4-orbit visit plus a 3-orbit visit. Each visit will perform a NIC-SPIRAL-DITH pattern with $2''$ (10 pixel) spacing (plus $1/2$ pixel or $1/3$ pixel extra shift for PSF sampling).

Proposal 11174 - Visit 01 - A Spitzer/X-ray candidate cluster at z>2: NICMOS imaging

Fri Jan 18 07:01:37 GMT 2008

Visit	Proposal 11174, Visit 01 Diagnostic Status: No Diagnostics Scientific Instruments: NIC3 Special Requirements: (none)										
	Patterns	#	Primary Pattern				Secondary Pattern			Exposures	
		(1)	Pattern Type=NIC-SPIRAL-DITH	Coordinate Frame=POS-TARG							(1)
		Purpose=DITHER	Pattern Orientation=0.0								
		Number Of Points=8	Angle Between Sides=								
		Point Spacing=2.06666	Center Pattern=false								
		Line Spacing=									
Fixed Targets	#	Name	Target Coordinates		Targ. Coord. Corrections	Fluxes		Miscellaneous			
	(1)	DADDIF-GAL-CLUS	RA: 14 49 14.4000 (222.310000d)	Dec: +08 56 18.40 (8.93844d)		V=28	J=23 AB brightest sources	Reference Frame: ICRS			
			Equinox: J2000								
	<i>Comments: The 30" diameter area of interest contain 10--20 targets with faint IR magnitudes, between AB 23 and 26</i>										
Exposures	#	Label	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time/[Actual Dur.]		Orbit
	1		(1) DADDIF-GAL-C LUS	NIC3, MULTIACCUM, NIC3	F160W	NSAMP=22; SAMP-SEQ=SPAR S64		Pattern 1-1 (1)	[=>(Pattern 1)]		[1]
									[=>(Pattern 2)]		
									[=>(Pattern 3)]		[2]
									[=>(Pattern 4)]		
									[=>(Pattern 5)]		[3]
									[=>(Pattern 6)]		
									[=>(Pattern 7)]		[4]
									[=>(Pattern 8)]		





Visit	Proposal 11174, Visit 02 Diagnostic Status: No Diagnostics Scientific Instruments: NIC3 Special Requirements: (none)									
	Patterns	#	Primary Pattern	Secondary Pattern	Exposures					
(2)		Pattern Type=NIC-SPIRAL-DITH Coordinate Frame=POS-TARG Purpose=DITHER Pattern Orientation=0.0 Number Of Points=6 Angle Between Sides= Point Spacing=2.1 Center Pattern=false Line Spacing=		(1)						
Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous				
	(1)	DADDIF-GAL-CLUS	RA: 14 49 14.4000 (222.3100000d) Dec: +08 56 18.40 (8.93844d) Equinox: J2000		V=28 J=23 AB brightest sources	Reference Frame: ICRS				
<i>Comments: The 30" diameter area of interest contain 10--20 targets with faint IR magnitudes, between AB 23 and 26</i>										
Exposures	#	Label	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time/[Actual Dur.]	Orbit
	1		(1) DADDIF-GAL-C LUS	NIC3, MULTIACCUM, NIC3	F160W	SAMP-SEQ=SPARS 64; NSAMP=22	POS TARG -6.0666 6,0,0	Pattern 1-1 (2)	[=>(Pattern 1)]	[1]
									[=>(Pattern 2)]	
									[=>(Pattern 3)]	[2]
									[=>(Pattern 4)]	
									[=>(Pattern 5)]	[3]
								[=>(Pattern 6)]		
Orbit Structure	<p>Orbit 1 Server Version: 20071214</p> <p>The diagram shows a horizontal timeline from 0 to 5500 seconds. Key events are marked with vertical arrows: 'GS Acq Exp. 1' at 0s, 'Exp. 1' at 50s, 'Pointing Maneuver' at 1600s, 'Unused Visibility = 121' at 1700s, 'Pointing Maneuver' at 3000s, and 'Occultation' at 3100s. A blue checkered bar represents the primary exposure from 50s to 1600s. A green bar represents a secondary exposure from 3000s to 3100s. A long white bar represents the final exposure from 3200s to 5500s.</p>									
	<p>0 500 1000 1500 2000 2500 3000 3500 4000 4500 5000 5500 sec.</p>									

