



11678 - Resolved H alpha star formation in two lensed galaxies at z=0.9

Cycle: 17, Proposal Category: GO

(Availability Mode: SUPPORTED)

INVESTIGATORS

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VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) A2390ARCS	WFC3/IR	2	22-May-2009 21:05:31.0	yes

2 Total Orbits Used

ABSTRACT

We will obtain H alpha narrow-band images of two galaxies at z=0.912 that have been gravitationally lensed by the galaxy cluster Abell 2390. H alpha falls squarely into the F126N filter, and both galaxies fit in a single WFC3 field of view. Because these two galaxies are magnified by factors of 6.7 (+0.4) and 12.6 (+0.8), WFC3 IR pixels probe spatial scales of 150 and 80 pc. (Without lensing, the WFC3 pixels probe 1 kpc scales at these redshifts.) Thus, these two galaxies provide a rare chance to examine, in detail and at high S/N, the spatial distribution of star formation in average galaxies at z=1.

After lensing deprojection, we will study the spatial distribution of star formation, the star-forming disk properties and nuclear contribution, as well as the distribution of extinction (from the archival F55W to H-alpha ratio map). We will also compare integrated extinction--corrected H alpha to Spitzer-derived diagnostics of star formation rate.

OBSERVING DESCRIPTION

We use two filters: F126N to capture H-alpha at $z=0.912$, and F125W as continuum. Because the wide filter will continuum-subtract the narrow filter, identical dither positions are required. The two target galaxies are separated by just $19''$, and thus fit easily in a single WFC3 field of view.

Low background makes the F126N observations readnoise limited. To minimize readnoise, we use only 3 dither positions, and 3 long exposures in STEP200 mode. For pessimistic assumptions of H alpha distribution, the per-pixel F126N signal/noise ratios will be >4 and 6 for galaxies A and B. If H alpha is concentrated like the F814W light, then the S/N will be ~ 15 per pixel. We begin each orbit with an F126N exposure.

In F125W, we avoid saturation of the brighter ($J=17.31$) target by limiting exposures to <400 s. As such, we use STEP50 sampling, a good compromise since it avoids saturation and comes close to the sky/readnoise equality point.

For each filter, we use 3 dither positions, separated by $7''$, with integer pixel steps, to step over bad pixels and other detector defects. In such a small program, we are not able to dither sufficiently to overcome the undersampled PSF. The benefits of 4-6 dither positions would be overwhelmed by the increased readnoise in F126N.

Proposal 11678 - Visit 01 - Resolved H alpha star formation in two lensed galaxies at z=0.9

Sat May 23 01:05:35 GMT 2009

Visit	Proposal 11678, Visit 01, implementation Diagnostic Status: No Diagnostics Scientific Instruments: WFC3/IR Special Requirements: (none)									
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous			
	(1)	A2390ARCS	RA: 21 53 34.4600 (328.3935833d) Dec: +17 42 9.37 (17.70260d) Equinox: J2000	Redshift: 0.912	V=(?) J=17.31 and J=20.1	Reference Frame: ICRS				
Exposures	#	Label	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time/[Actual Dur.]	Orbit
	1	F126N 1	(1) A2390ARCS	WFC3/IR, MULTIACCUM, IR	F126N	NSAMP=14; SAMP-SEQ=STEP2 00	POS TARG -4.995,- 4.961		[==>]	[1]
	2	F126N 2	(1) A2390ARCS	WFC3/IR, MULTIACCUM, IR	F126N	NSAMP=14; SAMP-SEQ=STEP2 00	POS TARG 0.0,0.0		[==>]	[1]
	3	F126N 3	(1) A2390ARCS	WFC3/IR, MULTIACCUM, IR	F126N	SAMP-SEQ=STEP2 00; NSAMP=14	POS TARG 4.995,4. 961		[==>]	[2]
	4	F125W 1	(1) A2390ARCS	WFC3/IR, MULTIACCUM, IR	F125W	SAMP-SEQ=STEP5 0; NSAMP=12	SAME POS AS 3		[==>]	[2]
	5	F125W 2	(1) A2390ARCS	WFC3/IR, MULTIACCUM, IR	F125W	SAMP-SEQ=STEP5 0; NSAMP=12	SAME POS AS 2		[==>]	[2]
	6	F125W 3	(1) A2390ARCS	WFC3/IR, MULTIACCUM, IR	F125W	SAMP-SEQ=STEP5 0; NSAMP=12	SAME POS AS 2		[==>]	[2]
	7	F125W 4	(1) A2390ARCS	WFC3/IR, MULTIACCUM, IR	F125W	SAMP-SEQ=STEP5 0; NSAMP=12	SAME POS AS 1		[==>]	[2]

