



12219 - Multiple stellar generations in the Large Magellanic Cloud Star Cluster NGC 1846

Cycle: 18, Proposal Category: GO
(Availability Mode: SUPPORTED)

INVESTIGATORS

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VISITS

| <i>Visit</i> | <i>Targets used in Visit</i> | <i>Configurations used in Visit</i> | <i>Orbits Used</i> | <i>Last Orbit Planner Run</i> | <i>OP Current with Visit?</i> |
|--------------|------------------------------|-------------------------------------|--------------------|-------------------------------|-------------------------------|
| 01 | (2) NGC1846-OUT | ACS/WFC WFC3/IR WFC3/UVIS | 4 | 10-Jul-2010 00:51:04.0 | yes |

4 Total Orbits Used

ABSTRACT

The recent discovery of multiple stellar populations in massive Galactic globular clusters poses a serious challenge for models of star cluster formation and evolution. The finding of multiple main sequences in the massive clusters NGC 2808 and omega Centauri, and multiple sub-giant-branch in NGC 1851 and many other globulars have demonstrated that star clusters are not as simple as we have imagined for decades. Surprisingly the only way to explain the main sequence splitting appears to be Helium enrichment, up to an astonishingly high $Y \sim 0.40$. A unique angle on this problem can be provided by intermediate-age clusters in the Magellanic Clouds with peculiar main-sequence turn-off morphologies. Recent

discoveries, based on ACS data of unparalleled photometric accuracy, have demonstrated that the CMDs of a large fraction of these clusters (~70 %) are not consistent with the simple, single stellar population hypothesis.

Explanations for what conditions could give rise to multiple populations in Galactic Globular Clusters remain controversial; this is even more the case for LMC clusters

To properly constraint the multipopulation phenomenon in Magellanic Cloud star clusters, we propose deep UV/IR imaging of NGC 1846, a star cluster where multiple populations have already been identified. The proposed observation will allow us to accurately measure the age difference between the stellar populations providing fundamental clues on the formation mechanism. Our simulations of WFC3 performance suggest that we will be able to detect even the main sequence splitting caused by small He differences ($\Delta Y \sim 0.02$).

OBSERVING DESCRIPTION

UVIS and NIR WFC3, and ACS WFC observations of the Large Magellanic Cloud stellar cluster NGC1846

| | | | | | | |
|--------------|--------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|-------------|-------------------------------------------------------------------------------------|---------------------------|---------------------------------|-----------------------|
| Visit | Proposal 12219, Visit 01 Sat Jul 10 04:51:13 GMT 2010 Diagnostic Status: No Diagnostics Scientific Instruments: WFC3/IR, ACS/WFC, WFC3/UVIS Special Requirements: ORIENT 124D TO 130 D | | | | | |
| | Fixed Targets | # | Name | Target Coordinates | Targ. Coord. Corrections | Fluxes |
| (2) | | NGC1846-OUT | RA: 05 07 40.8000 (76.9200000d) Dec: -67 26 56.70 (-67.44908d) Equinox: J2000 | | V=11.31 | Reference Frame: ICRS |

Proposal 12219 (STScI Edit Number: 0, Created: Friday, July 9, 2010 11:51:13 PM EST) - Overview

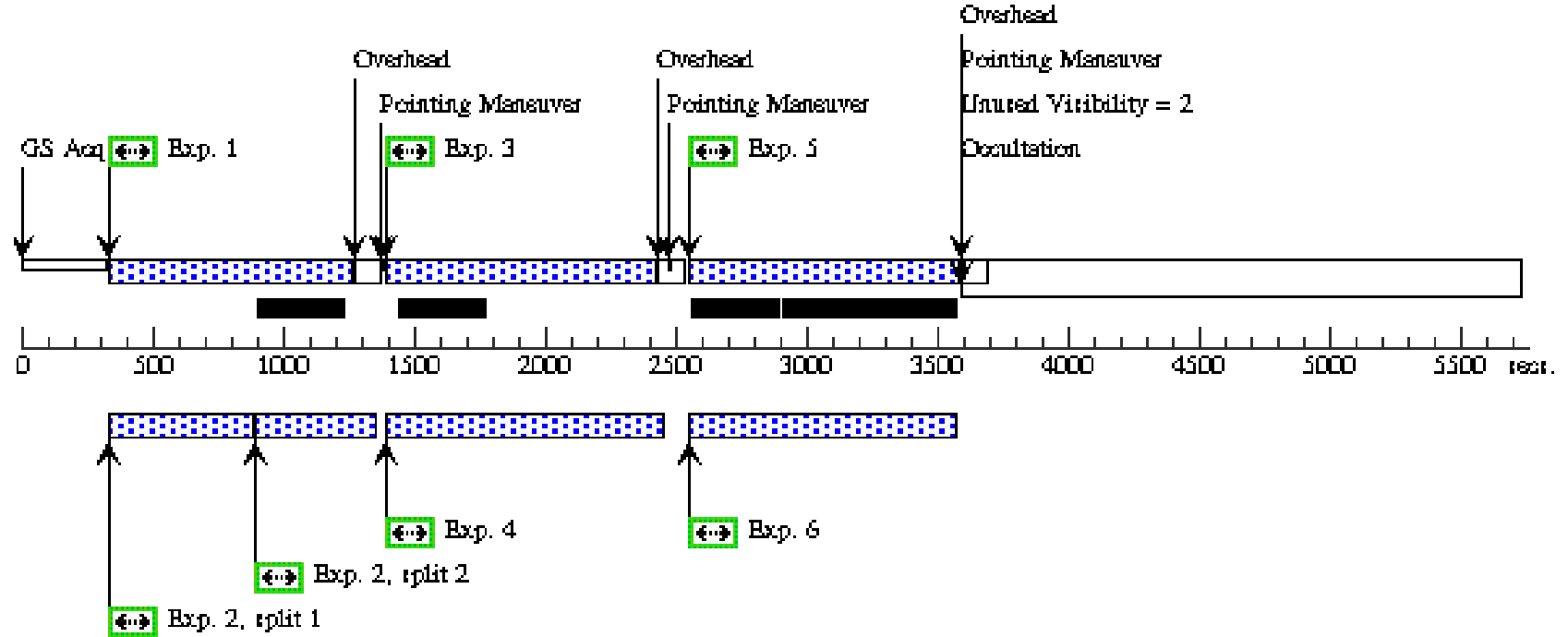
| # | Label | Target | Config,Mode,Aperture | Spectral Els. | Opt. Params. | Special Reqs. | Groups | Exp. Time/[Actual Dur.] | Orbit |
|-----------|-------|-----------------|-------------------------------|---------------|----------------------------------|--------------------|------------------------------|----------------------------------------------|-------|
| Exposures | 1 | (2) NGC1846-OUT | WFC3/UVIS, ACCUM, UVIS-CENTER | F336W | CR-SPLIT=NO | POS TARG -5.6,-4.2 | Prime + Parallel Group 1-2 | 900 Secs [==>] | [1] |
| | 2 | (2) NGC1846-OUT | ACS/WFC, ACCUM, WFCENTER | F814W | CR-SPLIT=2 | | Prime + Parallel Group 1-2 | 680 Secs [==>(Split 1)] [==>(Split 2)] | [1] |
| | 3 | (2) NGC1846-OUT | WFC3/UVIS, ACCUM, UVIS-CENTER | F336W | CR-SPLIT=NO | POS TARG -4.2,-2.8 | Prime + Parallel Group 3-4 | 1032 Secs [==>] | [1] |
| | 4 | (2) NGC1846-OUT | ACS/WFC, ACCUM, WFCENTER | F475W | CR-SPLIT=NO | | Prime + Parallel Group 3-4 | 900 Secs [==>] | [1] |
| | 5 | (2) NGC1846-OUT | WFC3/UVIS, ACCUM, UVIS-CENTER | F336W | CR-SPLIT=NO | POS TARG -2.8,-1.4 | Prime + Parallel Group 5-6 | 1032 Secs [==>] | [1] |
| | 6 | (2) NGC1846-OUT | ACS/WFC, ACCUM, WFCENTER | F475W | CR-SPLIT=NO | | Prime + Parallel Group 5-6 | 900 Secs [==>] | [1] |
| | 7 | (2) NGC1846-OUT | WFC3/UVIS, ACCUM, UVIS-CENTER | F336W | CR-SPLIT=NO | POS TARG -1.4,0.0 | Prime + Parallel Group 7-8 | 1032 Secs [==>] | [2] |
| | 8 | (2) NGC1846-OUT | ACS/WFC, ACCUM, WFCENTER | F475W | CR-SPLIT=NO | | Prime + Parallel Group 7-8 | 900 Secs [==>] | [2] |
| | 9 | (2) NGC1846-OUT | WFC3/UVIS, ACCUM, UVIS-CENTER | F336W | CR-SPLIT=NO | POS TARG 0.0,1.4 | Prime + Parallel Group 9-10 | 1032 Secs [==>] | [2] |
| | 10 | (2) NGC1846-OUT | ACS/WFC, ACCUM, WFCENTER | F475W | CR-SPLIT=NO | | Prime + Parallel Group 9-10 | 900 Secs [==>] | [2] |
| | 11 | (2) NGC1846-OUT | WFC3/UVIS, ACCUM, UVIS-CENTER | F336W | CR-SPLIT=NO | POS TARG 1.4,2.8 | Prime + Parallel Group 11-12 | 1032 Secs [==>] | [2] |
| | 12 | (2) NGC1846-OUT | ACS/WFC, ACCUM, WFCENTER | F475W | CR-SPLIT=NO | | Prime + Parallel Group 11-12 | 900 Secs [==>] | [2] |
| | 13 | (2) NGC1846-OUT | WFC3/UVIS, ACCUM, UVIS-CENTER | F336W | CR-SPLIT=NO | POS TARG 2.8,4.2 | Prime + Parallel Group 13-14 | 1032 Secs [==>] | [3] |
| | 14 | (2) NGC1846-OUT | ACS/WFC, ACCUM, WFCENTER | F475W | CR-SPLIT=NO | | Prime + Parallel Group 13-14 | 900 Secs [==>] | [3] |
| | 15 | (2) NGC1846-OUT | WFC3/UVIS, ACCUM, UVIS-CENTER | F336W | CR-SPLIT=NO | POS TARG 4.2,5.6 | Prime + Parallel Group 15-16 | 1032 Secs [==>] | [3] |
| | 16 | (2) NGC1846-OUT | ACS/WFC, ACCUM, WFCENTER | F475W | CR-SPLIT=NO | | Prime + Parallel Group 15-16 | 900 Secs [==>] | [3] |
| | 17 | (2) NGC1846-OUT | WFC3/UVIS, ACCUM, UVIS-CENTER | F336W | CR-SPLIT=NO | POS TARG 5.6,7 | Prime + Parallel Group 17-18 | 1032 Secs [==>] | [3] |
| | 18 | (2) NGC1846-OUT | ACS/WFC, ACCUM, WFCENTER | F814W | CR-SPLIT=2 | | Prime + Parallel Group 17-18 | 680 Secs [==>(Split 1)] [==>(Split 2)] | [3] |
| | 19 | (2) NGC1846-OUT | WFC3/IR, MULTIACCUM, IR-FIX | F160W | NSAMP=12; SAMP-SEQ=STEP5 0 | POS TARG -5.6,-4.2 | | [==>] | [4] |
| | 20 | (2) NGC1846-OUT | WFC3/IR, MULTIACCUM, IR-FIX | F160W | NSAMP=12; SAMP-SEQ=STEP5 0 | POS TARG -4.2,-2.8 | | [==>] | [4] |

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| | | | | | | | | |
|----|-----------------|-----------------------------|-------|----------------------------------|--------------------|------------------------------|-------------------|-----|
| 21 | (2) NGC1846-OUT | WFC3/IR, MULTIACCUM, IR-FIX | F160W | NSAMP=12; SAMP-SEQ=STEP5 0 | POS TARG -2.8,-1.4 | | [==>] | [4] |
| 22 | (2) NGC1846-OUT | WFC3/IR, MULTIACCUM, IR-FIX | F160W | NSAMP=12; SAMP-SEQ=STEP5 0 | POS TARG -1.4,0 | | [==>] | [4] |
| 23 | (2) NGC1846-OUT | WFC3/IR, MULTIACCUM, IR-FIX | F160W | NSAMP=12; SAMP-SEQ=STEP5 0 | POS TARG 0,1.4 | | [==>] | [4] |
| 24 | (2) NGC1846-OUT | WFC3/IR, MULTIACCUM, IR-FIX | F160W | NSAMP=12; SAMP-SEQ=STEP5 0 | POS TARG 1.4,2.8 | | [==>] | [4] |
| 25 | (2) NGC1846-OUT | WFC3/IR, MULTIACCUM, IR-FIX | F160W | NSAMP=12; SAMP-SEQ=STEP5 0 | POS TARG 2.8,4.2 | | [==>] | [4] |
| 26 | (2) NGC1846-OUT | WFC3/IR, MULTIACCUM, IR-FIX | F160W | NSAMP=13; SAMP-SEQ=STEP5 0 | POS TARG 4.2,5.6 | Prime + Parallel Group 26-27 | [==>] | [4] |
| 27 | (2) NGC1846-OUT | ACS/WFC, ACCUM, WFC | F814W | | | Prime + Parallel Group 26-27 | 326 Secs [==>] | [4] |

Orbit 1

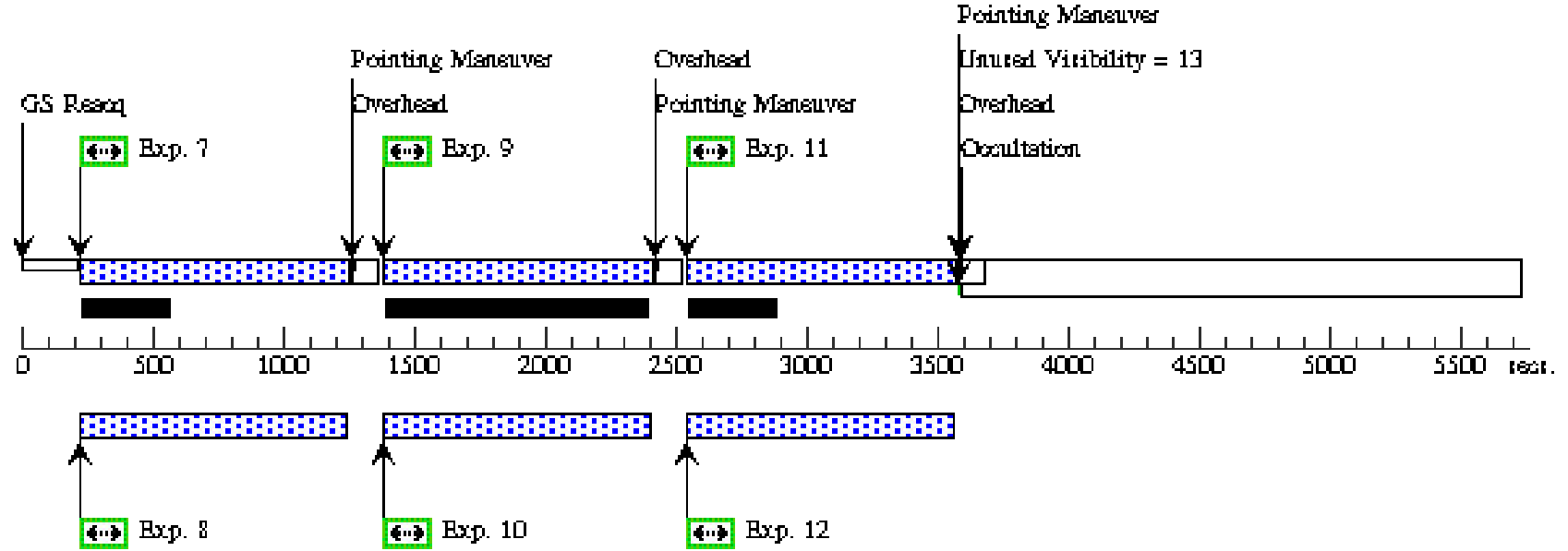
Server Version: 20100505



Orbit Structure

Orbit 2

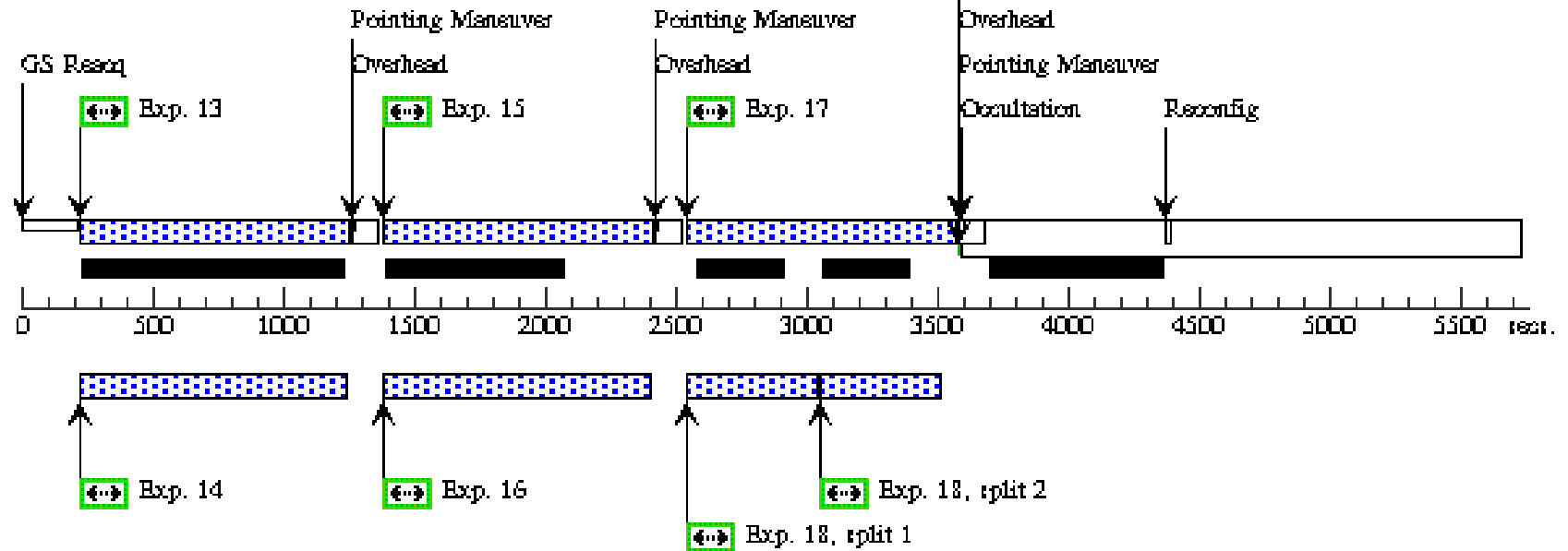
Server Version: 20100505



Orbit 3

Server Version: 20100505

Unated Visibility = 13



Orbit 4

Server Version: 20100505

