12758 - Alpha Cen: Climbing out of a Coronal Recession?

Cycle: 19, Proposal Category: GO

(Availability Mode: SUPPORTED)

INVESTIGATORS

Name	*Institution*	*E-Mail*
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VISITS

Visit	*Targets used in Visit*	*Configurations used in Visit*	*Orbits Used*	*Last Orbit Planner Run*	*OP Current with Visit?*
01	(1) HD128620 (2) HD128621-1	STIS/CCD STIS/FUV-MAMA	2	07-Sep-2011 21:29:16.0	yes
02	(1) HD128620 (3) HD128621-2	STIS/CCD STIS/FUV-MAMA	2	07-Sep-2011 21:29:24.0	yes

4 Total Orbits Used

ABSTRACT

Nearby Alpha Centauri contains the two best characterized G and K dwarfs, next to the Sun itself, thanks to the accurate orbit, resolved angular diameters, and well understood co-evolutionary state. Alpha Cen A & B also have the best studied stellar X-ray activity cycles, extending back to the 1970's. Present proposal is to continue tracking the evolving multi-decadal high-energy narrative of Alpha Cen with semiannual HRC-I pointings in Cycles 13-15, as solar twin A is expected to be rising to cycle maximum from an extended coronal recession. STIS E140M spectra will support and leverage the broad-band X-ray measurements by probing subcoronal dynamics and providing a low-T boundary condition for DEM modeling, with connection to HRC through the FUV Fe XII coronal forbidden line.

OBSERVING DESCRIPTION

The STIS observations are straightforward. Alpha Cen falls in the HST Continuous Viewing Zone numerous times during the year, allowing the two stars to be captured in a single visit of just two orbits. Two such visits (four orbits total), about six months apart, are contemplated for Cycle 19, compatible with the semiannual pointings by Chandra. There is no need to strictly coordinate the FUV and X-ray visits, because the FUV Fe XII coronal forbidden line can tie the STIS observation into the X-ray framework.

In each HST/STIS visit, the binary companions are observed sequentially, beginning with Alp Cen A, the brighter of the two stars. The target is acquired with the CCD and F25ND5, followed by a 4.3 ks exposure with the E140M-1425 medium-res echelle through the photometric slot (0.2 ×0.2), which delivers R40,000 and good sensitivity (peak S/N 60 per resol at the tops of the important Si IV and C IV resonance lines). The observation is split into three subexposures to mitigate instrumental drifts. The exposure depth is sucient to capture the key Fe XII 1242 coronal forbidden line, which as mentioned earlier is used to tie the STIS FUV measurements into the X-ray timeline. Following the E140M exposure, a ~5" offset manuever is performed to Alp Cen B, and a peak-up is performed in dispersed light (G430L) with the 0.3x0.05ND filtered slit to precisely locate the target. This particular slit was chosen because the default peak-up sweeps out a sufficient area (0.3"x0.3") to capture the B component, given that the time-dependent separation of the binary companions is accurately known (to the 0.1" level; based on five years of measurements by Chandra HRC-I). After the peak-up, a second 4.3 ks E140M exposure through the photometric aperture is obtained on B. Given that the AB orbital separation is changing relatively rapidly in the present epoch, Alp Cen B was specified as two separate targets, with different offsets, to reflect the orbital evolution over the six months between the two STIS visits.

Proposal 12758 - Visit 01 - Alpha Cen: Climbing out of a Coronal Recession?

Visit

Proposal 12758, Visit 01, implementation Thu Sep 08 01:29:30 GMT 2011

Diagnostic Status: No Diagnostics

Scientific Instruments: STIS/CCD, STIS/FUV-MAMA

Special Requirements: CVZ; BETWEEN 29-FEB-2012:20:00:00 AND 01-MAR-2012:20:00:00

Comments: This is first CVZ opportunity after 1 Mar 2012, and to allow visit 2 to be scheduled six months later, but still within Cycle 19.

Fixed Targets

#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
(1)	HD128620	RA: 14 39 31.2880 (219.8803667d)	Proper Motion RA: -3.80 arcsec/yr	V=-0.1+/-0.1	Reference Frame: ICRS
	Alt Name1: ALP-CEN-A	Dec: -60 49 59.06 (-60.83307d)	Proper Motion Dec: +0.35 arcsec/yr	FUV peak fluxes are 5E-12 per Angstrom (excluding Lyman-alpha, which is ~ 8E-11)	
		Equinox: J2000	Parallax: 0.747"		
			Epoch of Position: 2010.0		
			Radial Velocity: 24 km/sec		

Comments: The target coords for epoch 2009.95 were measured directly from a high-precision Chandra HRC-I image, and include the parallactic shift at that time. The effective proper motions were determined from the best-fit orbital track of the binary, based on five years of HRC-I measurements, including the best-fit linear proper motions applied to the center-of-mass, and refer specifically to the A component. The coords (advanced to epoch 2010.0) and PMs differ slightly from those cited in SIMBAD, since the latter values were measured in a different epoch when the relative orbital motion was different.

#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
(2)	HD128621-1	Offset from HD128620 by	Proper Motion RA: 0.00 sec of time/yr	V=1.33+/-0.1	Offset Position (HD128621-1)
	Alt Name1: ALP-CEN-B	RA Offset: -0.7075 Secs	Proper Motion Dec: 0.00 arcsec/yr		Reference Frame: ICRS
		Dec Offset: -1.011 Arcsec	Parallax: 0.00"		
			Epoch of Position: 2010		

Exposures

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time/[Actual Dur.]	Orbit
1	(STIS.ta.235462)	(1) HD128620	STIS/CCD, ACQ, F25ND5	MIRROR		GS ACQ SCENARIO BASE1B3		0.1 Secs *[==>]*	 ***[1]***

Comments: Castelli and Kurucz Models:G2V;Time to Saturation (for a single exposure) = 2.3 seconds ; Optimum SNR = 179

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time/[Actual Dur.]	Orbit
2	(STIS.sp.235468)	(1) HD128620	STIS/FUV-MAMA, ACCUM, 0.2X0.2	E140M 1425 A				1425 Secs X 3 *[==>(Copy 1)]* *[==>(Copy 2)]* *[==>(Copy 3)]*	 ***[1]***

Comments: input=StarCAT ETC file for ALP-CEN-A; Exposure time= 4.3 ks at wavelength 1393.80 Å gives: SNR = 58 (per resol); Brightest Pixel (at 1215.40 Å)= 1.8 cps; peak S/N~200 at Lyman-alpha

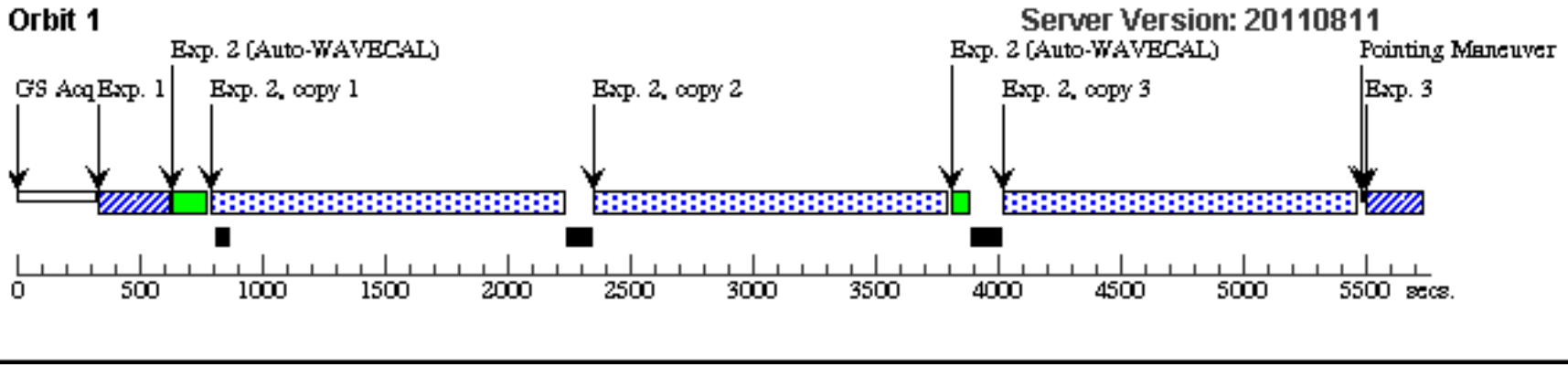
#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time/[Actual Dur.]	Orbit
3	(STIS.sp.235634)	(2) HD128621-1	STIS/CCD, ACQ/PEAK, 0.3X0.05ND	G430L 4300 A				0.1 Secs *[==>]*	 ***[1]***

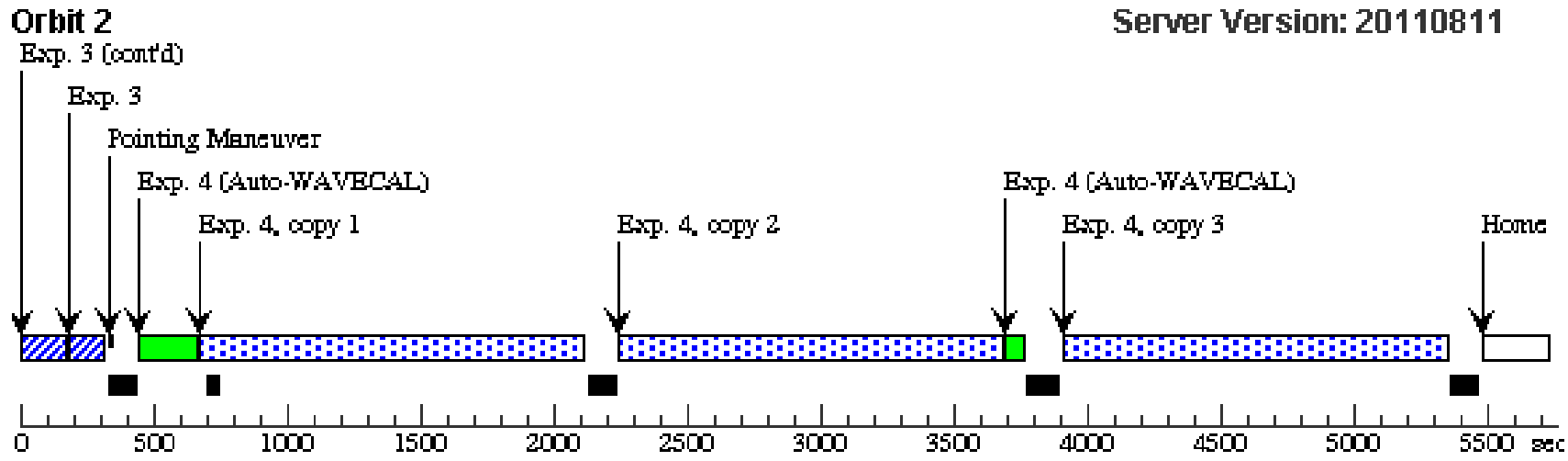
Comments: dispersed light peak-up assuming Castelli-Kurucz Models K0V, V=1.33 mag: at wavelength 5000 Å in 0.1 s, SNR ~ 15 (per resol); Time to Saturation (for a single exposure) ~ 60 seconds. ETC calculation done with 52x0.05 slit, then scaled to 0.3x0.05ND aperture by factor 0.001 (ND=3), as verified by comparing throughputs indicated by separate ETC TARG ACQ runs (without disperser) with ND ap and clear apertures.

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time/[Actual Dur.]	Orbit
4	(STIS.sp.235468)	(2) HD128621-1	STIS/FUV-MAMA, ACCUM, 0.2X0.2	E140M 1425 A				1425 Secs X 3 *[==>(Copy 1)]* *[==>(Copy 2)]* *[==>(Copy 3)]*	 ***[2]***

Comments: ETC run is for ALP-CEN-A, which has very similar spectrum to ALP-CEN-B in observed flux densities

Orbit Structure





Visit

Proposal 12758, Visit 02, implementation Thu Sep 08 01:29:32 GMT 2011

Diagnostic Status: No Diagnostics

Scientific Instruments: STIS/CCD, STIS/FUV-MAMA

Special Requirements: CVZ; BETWEEN 06-OCT-2012:12:00:00 AND 08-OCT-2012:12:00:00

Fixed Targets

#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
(1)	HD128620	RA: 14 39 31.2880 (219.8803667d)	Proper Motion RA: -3.80 arcsec/yr	V=-0.1+/-0.1	Reference Frame: ICRS
	Alt Name1: ALP-CEN-A	Dec: -60 49 59.06 (-60.83307d)	Proper Motion Dec: +0.35 arcsec/yr	FUV peak fluxes are 5E-12 per Angstrom (excluding Lyman-alpha, which is ~ 8E-11)	
		Equinox: J2000	Parallax: 0.747"		
			Epoch of Position: 2010.0		
			Radial Velocity: 24 km/sec		

Comments: The target coords for epoch 2009.95 were measured directly from a high-precision Chandra HRC-I image, and include the parallactic shift at that time. The effective proper motions were determined from the best-fit orbital track of the binary, based on five years of HRC-I measurements, including the best-fit linear proper motions applied to the center-of-mass, and refer specifically to the A component. The coords (advanced to epoch 2010.0) and PMs differ slightly from those cited in SIMBAD, since the latter values were measured in a different epoch when the relative orbital motion was different.

(3)	HD128621-2	Offset from HD128620 by	Proper Motion RA: 0.00 sec of time/yr	V=1.33+/-0.1	Offset Position (HD128621-2)
	Alt Name1: ALP-CEN-B	RA Offset: -0.6702 Secs	Proper Motion Dec: 0.00 arcsec/yr		Reference Frame: ICRS
		Dec Offset: -0.506 Arcsec	Parallax: 0.00"		
			Epoch of Position: 2010		

Exposures

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time/[Actual Dur.]	Orbit
1	(STIS.ta.235462)	(1) HD128620	STIS/CCD, ACQ, F25ND5	MIRROR		GS ACQ SCENARIO BASE1B3		0.1 Secs	
								[==>]	***[1]***

Comments: Castelli and Kurucz Models:G2V;Time to Saturation (for a single exposure) = 2.3 seconds ; Optimum SNR = 179

2	(STIS.sp.235468)	(1) HD128620	STIS/FUV-MAMA, ACCUM, 0.2X0.2	E140M 1425 A				1425 Secs X 3	
								[==>(Copy 1)]	
								[==>(Copy 2)]	***[1]***
								[==>(Copy 3)]	

Comments: input=StarCAT ETC file for ALP-CEN-A; Exposure time= 4.3 ks at wavelength 1393.80 Å gives: SNR = 58 (per resol); Brightest Pixel (at 1215.40 Å)= 1.8 cps; peak S/N~200 at Lyman-alpha

3	(STIS.sp.235634)	(3) HD128621-2	STIS/CCD, ACQ/PEAK, 0.3X0.05ND	G430L 4300 A				0.1 Secs	
								[==>]	***[1]***

Comments: dispersed light peak-up assuming Castelli-Kurucz Models K0V, V=1.33 mag: at wavelength 5000 Å in 0.1 s, SNR ~ 15 (per resol); Time to Saturation (for a single exposure) ~ 60 seconds. ETC calculation done with 52x0.05 slit, then scaled to 0.3x0.05ND aperture by factor 0.001 (ND=3), as verified by comparing throughputs indicated by separate ETC TARG ACQ runs (without disperser) with ND ap and clear apertures.

4	(STIS.sp.235468)	(3) HD128621-2	STIS/FUV-MAMA, ACCUM, 0.2X0.2	E140M 1425 A				1425 Secs X 3	
								[==>(Copy 1)]	
								[==>(Copy 2)]	***[2]***
								[==>(Copy 3)]	

Comments: ETC run is for ALP-CEN-A, which has very similar spectrum to ALP-CEN-B in observed flux densities

Orbit Structure