



13840 - The Smith Cloud: Galactic or Extragalactic?

Cycle: 22, Proposal Category: GO

(UV Initiative)

(Availability Mode: SUPPORTED)

INVESTIGATORS

<i>Name</i>	<i>Institution</i>	<i>E-Mail</i>
Dr. Andrew J. Fox (PI) (ESA Member) (Contact)	Space Telescope Science Institute - ESA	afox@stsci.edu
Dr. Kenneth Sembach (CoI) (AdminUSPI)	Space Telescope Science Institute	sembach@stsci.edu
Dr. Felix J. Lockman (CoI)	National Science Foundation	jlockman@nrao.edu
Dr. Bart P. Wakker (CoI)	University of Wisconsin - Madison	wakker@astro.wisc.edu
Prof. Brice Menard (CoI)	The Johns Hopkins University	menard@jhu.edu
Dr. Nicolas Lehner (CoI)	University of Notre Dame	nlehner@nd.edu
Dr. Alex S Hill (CoI)	CSIRO, Australia Telescope National Facility	alex.hill@csiro.au
Dr. Kathleen A Barger (CoI)	University of Notre Dame	kbargers@nd.edu
Prof. Fabian Heitsch (CoI)	University of North Carolina at Chapel Hill	fheitsch@email.unc.edu
David Vincent Stark (CoI)	University of North Carolina at Chapel Hill	dstark@email.unc.edu
Dr. Mubdi Rahman (CoI)	The Johns Hopkins University	mubdi@pha.jhu.edu
Dr. Robert A. Benjamin (CoI)	University of Wisconsin - Whitewater	benjamir@uww.edu

VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) PG2112+059	COS/FUV	3	15-Sep-2014 21:03:57.0	yes
02	(2) RXJ2043.1+0324	COS/FUV	3	15-Sep-2014 21:03:59.0	yes
03	(3) RXJ2139.7+0246	COS/FUV	3	15-Sep-2014 21:04:01.0	yes

9 Total Orbits Used

ABSTRACT

The Milky Way is surrounded by a population of gaseous high-velocity clouds (HVCs), which play an important role in bringing fresh fuel for star formation into the disk. The Smith Cloud is an HVC in an advanced state of accretion, only ~ 30 Myr away from impacting the Galactic plane. It is unique among HVCs in having a known distance (12.4 kpc) and a well-constrained 3D velocity (300 km/s), but we do not yet know its metallicity, which would distinguish between Galactic (metal-enriched) and extragalactic (metal-poor) origins. Here we propose a program to measure the Smith Cloud's metallicity using HST/COS spectra of three AGN lying behind the Cloud and its wake. Combining the UV metal-line (O I and S II) column densities with existing H I column densities will lead to precise abundance measurements for the Smith Cloud, directly constraining its origin. In addition, the relative strength of depleted (Fe II, Si II) to undepleted (S II) UV absorption lines will be used to infer its dust content. Finally, we will search for variation in the ionization level and kinematics between the three sight lines, which would indicate a disruptive encounter with the hot Galactic corona.

OBSERVING DESCRIPTION

This is a nine-orbit COS/G130M spectroscopic program to study the chemical and physical properties of the Smith Cloud, a high-velocity cloud (HVC) falling onto the Galactic disk. The program observes three QSOs, with a single three-orbit visit on each QSO. Each target is fairly faint in the UV, with F_{1300} between 0.5×10^{-14} and 0.8×10^{-14} erg/cm²/s⁻¹/Å⁻¹, so there are no bright-object concerns (all three fields pass BOT checks with both GSC II and GALEX catalogs). All four FP-POS positions are used within each visit with the 1291 CENWAVE setting.

Proposal 13840 - PG2112+059 (01) - The Smith Cloud: Galactic or Extragalactic?

Tue Sep 16 01:04:02 GMT 2014

Visit	Proposal 13840, PG2112+059 (01), implementation Diagnostic Status: No Diagnostics Scientific Instruments: COS/FUV Special Requirements: (none)									
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous			
	(1)	PG2112+059	RA: 21 14 52.5871 (318.7191129d) Dec: +06 07 42.33 (6.12842d) Equinox: J2000		V=15.77 F_1300=0.75e-14 (cgs)	Reference Frame: ICRS				
	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i>									
Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	ACQ/SEAR CH (615809)	(1) PG2112+059	COS/FUV, ACQ/SEARCH, PSA	G130M 1291 A	SCAN-SIZE=2; CENTER=FLUX-W T; STEP-SIZE=1.767			8 Secs (8 Secs) [==>]	[1]
	2	ACQ/PEAK XD (615809)	(1) PG2112+059	COS/FUV, ACQ/PEAKXD, PSA	G130M 1291 A				8 Secs (8 Secs) [==>]	[1]
	3	ACQ/PEAK D (615809)	(1) PG2112+059	COS/FUV, ACQ/PEAKD, PSA	G130M 1291 A	NUM-POS=5.0; STEP-SIZE=0.9; CENTER=FLUX-W T-FLR			8 Secs (8 Secs) [==>]	[1]
	4	PG2112+05 9-FPPOS1 (615812)	(1) PG2112+059	COS/FUV, TIME-TAG, PSA	G130M 1291 A	FP-POS=1; BUFFER-TIME=22 05			2205 Secs (2205 Secs) [==>]	[1]
	5	PG2112+05 9-FPPOS2 (615812)	(1) PG2112+059	COS/FUV, TIME-TAG, PSA	G130M 1291 A	FP-POS=2; BUFFER-TIME=24 24			2935 Secs (2935 Secs) [==>]	[2]
	6	PG2112+05 9-FPPOS3 (615812)	(1) PG2112+059	COS/FUV, TIME-TAG, PSA	G130M 1291 A	FP-POS=3; BUFFER-TIME=13 75			1375 Secs (1375 Secs) [==>]	[3]
	7	PG2112+05 9-FPPOS4 (615812)	(1) PG2112+059	COS/FUV, TIME-TAG, PSA	G130M 1291 A	FP-POS=4; BUFFER-TIME=13 75			1375 Secs (1375 Secs) [==>]	[3]











