



14099 - Measuring the L-T transition for a warm Saturn exoplanet

Cycle: 23, Proposal Category: GO

(Availability Mode: AVAILABLE)

INVESTIGATORS

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VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) HAT-P-18 WAVE	STIS/CCD	5	04-Nov-2015 21:23:37.0	yes
02	(1) HAT-P-18 CCDFLAT NONE WAVE	STIS/CCD	5	04-Nov-2015 21:23:45.0	yes
04	(1) HAT-P-18	WFC3/IR	5	04-Nov-2015 21:24:13.0	yes

15 Total Orbits Used

ABSTRACT

The transition between the L and T spectral types in brown dwarf atmospheres is characterized by a reduction in high-altitude clouds and the appearance of strong methane absorption features. Models predict a similar transition for gas giant atmospheres across the 1000-1200K temperature range. To test this theory, we propose to measure a detailed transmission spectrum for HAT-P-18b, a Saturn-mass gas giant with an equilibrium temperature of 850K. Our observations will span the 0.3-1.7 micron wavelength range, with the near-infrared wavelengths allowing us to constrain methane and water absorption, and the optical wavelengths probing alkali metal absorption and high-altitude Rayleigh scattering processes. The latter will allow us to determine whether or not aerosols are present in the atmosphere, addressing a key aspect of atmospheric chemistry. Our measurement for HAT-P-18b will be the first transmission spectrum for a gas giant exoplanet that is predicted to have a high methane abundance, providing unique constraints for atmosphere models.

OBSERVING DESCRIPTION

Our observations will provide a near-UV to near-IR transmission spectrum for the transiting exoplanet HAT-P-18b. This broad wavelength coverage will allow us to distinguish between fundamental atmosphere classes: the optical provides leverage for distinguishing between a clear or hazy atmosphere, while the near-IR contains strong CH₄ absorption features, providing leverage for distinguishing between an L-type or T-type atmosphere.

****Observing strategy:**

To construct the transmission spectrum, we require three separate transits of HAT-P-18b to be observed: one with STIS+G430L, one with STIS+G750L, and one with WFC3+G141. All observations will be made in spectroscopic mode. The STIS observations will be made with the 52x2 arcsec slit to avoid slit losses.

At the start of each visit, a standard phase-constrained acquisition image will be taken. The subsequent observing strategy consists of taking repeated exposures for the duration of the transit, plus time before and after the transit to establish the out-of-transit baseline flux.

For the STIS observations, we will use stare mode with individual exposure times of 278 sec. We will read out a subarray size of 128 pixels to reduce

Proposal 14099 (STScI Edit Number: 3, Created: Wednesday, November 4, 2015 9:24:19 PM EST) - Overview

overheads (SIZEAXIS2=128). As in our previous Programs 12473 and 11740, we will override the default wavelength calibrations at the end of each exposure (WAVECAL=NO). This is done to improve the duty cycle and to allow the thermal breathing trends to be removed, which is only possible with an un-interrupted photometric time series. A separate wavelength calibration image will be taken at the end of each visit. For the G750L visit, we will also take fringe flat images to correct for fringing longward of ~ 720 nm.

For the WFC3 observations, we will use the forward spatial scanning mode with a scan rate of 0.03 arcsec/sec and individual exposure times of 90 sec (i.e. resulting in a scan across 2.7 arcsec = 20 pixels per exposure). This will leave plenty of room on the detector to estimate the background flux. We will read out a subarray size of 256 pixels to reduce overheads.

Each transit of HAT-P-18b lasts for 163 minutes. We therefore require 5 consecutive HST orbits per visit. The first orbit serves to allow the telescope to thermally relax into its new pointing position, which is a standard procedure adopted by all HST exoplanet observations. The remaining orbits will provide good phase coverage during transit, and allow for sufficient time (2 HST orbits - one before transit, one after transit) to establish the baseline stellar flux necessary to accurately measure the transit depth. For both STIS and WFC3, we will obtain 10 exposures per HST orbit, coming to a total of 20 in-transit and 20 out-of-transit exposures for each instrument.

**Signal-to-noise estimates:

HAT-P-18b is a bright target in the conventional sense ($V=12.8$ mag, $J=10.8$ mag). The challenge with exoplanet transmission spectroscopy is to measure subtle changes in the system brightness across different wavelength channels, via variations in the transit depth. We used the HST Exposure Time Calculator with the observational setup described above to calculate the following SNRs per resolution element per exposure: 112 for G430L, 246 for G750L, and 1059 for G141. Note that the G141 flux will be spread over ~ 20 pixel columns due to the spatial scan (see above), so there is no risk of saturation.

Proposal 14099 - HAT-P-18 G430L (01) - Measuring the L-T transition for a warm Saturn exoplanet

Thu Nov 05 02:24:19 GMT 2015

Visit	<p>Proposal 14099, HAT-P-18 G430L (01), scheduled</p> <p>Diagnostic Status: No Diagnostics</p> <p>Scientific Instruments: STIS/CCD</p> <p>Special Requirements: Period 5.50802888 D AND ZERO-PHASE HJD2456461.06765758</p> <p><i>Comments: HAT-P-18. G430L. STIS visit. It is essential that the five orbits be scheduled in a continuous block, free of the SAA.</i></p>					
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes
(1)		HAT-P-18 Alt Name1: GSC-02594-00646	RA: 17 05 23.1510 (256.3464625d) Dec: +33 00 44.97 (33.01249d) Equinox: J2000	Proper Motion RA: -21.3 mas/yr Proper Motion Dec: -44.2 mas/yr Epoch of Position: 2000	V=12.8	Reference Frame: ICRS
<p><i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i></p>						

Proposal 14099 - HAT-P-18 G430L (01) - Measuring the L-T transition for a warm Saturn exoplanet

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
1	ACQ, phase constrained (STIS.ta.733341)	(1) HAT-P-18	STIS/CCD, ACQ, F28X50LP	MIRROR	GAIN=4	PHASE 0.9655 TO 0.9680		1.5 Secs (1.5 Secs) [==>]	[1]
2	HAT-P-18 G430L Orbit 1 (STIS.sp.733350)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G430L 4300 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO			120 Secs (120 Secs) [==>]	[1]
<i>Comments: Short exposure to minimize instrument systematic of first exposure of orbit. Note that it is longer than first exposures in subsequent orbits, due to acquisition overhead and to ensure all time is used.</i>									
3	HAT-P-18 G430L Orbit 1 (STIS.sp.733928)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G430L 4300 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO			278 Secs X 8 (2224 Secs) [==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)]	[1]
4	HAT-P-18 G430L Orbit 2 (STIS.sp.733352)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G430L 4300 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO			1 Secs (1 Secs) [==>]	[2]
<i>Comments: Short exposure to minimize instrument systematic of first exposure in each orbit.</i>									
5	HAT-P-18 G430L Orbit 2 (STIS.sp.733928)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G430L 4300 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO			278 Secs X 10 (2780 Secs) [==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)]	[2]
6	HAT-P-18 G430L Orbit 3 (STIS.sp.733352)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G430L 4300 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO			1 Secs (1 Secs) [==>]	[3]
<i>Comments: Short exposure to minimize instrument systematic of first exposure in each orbit.</i>									

Exposures

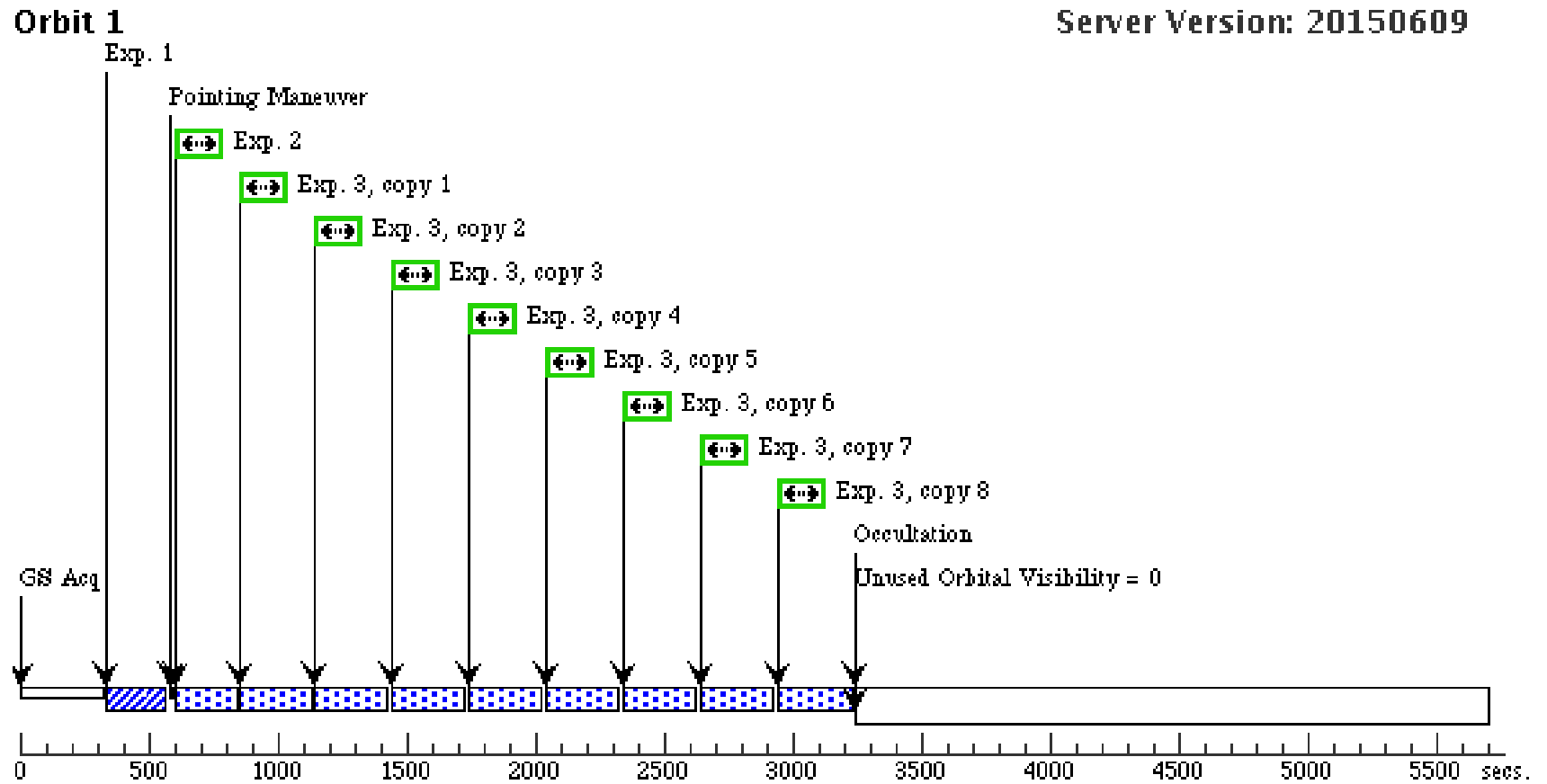
Proposal 14099 - HAT-P-18 G430L (01) - Measuring the L-T transition for a warm Saturn exoplanet

7	HAT-P-18 G430L Orbit 3 (STIS.sp.73 3928)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G430L 4300 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO	278 Secs X 10 (2780 Secs) [==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)]	[3]
8	HAT-P-18 G430L Orbit 4 (STIS.sp.73 3352)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G430L 4300 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO	1 Secs (1 Secs) [==>]	[4]
<i>Comments: Short exposure to minimize instrument systematic of first exposure in each orbit.</i>							
9	HAT-P-18 G430L Orbit 4 (STIS.sp.73 3928)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G430L 4300 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO	278 Secs X 10 (2780 Secs) [==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)]	[4]
10	HAT-P-18 G430L Orbit 5 (STIS.sp.73 3352)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G430L 4300 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO	1 Secs (1 Secs) [==>]	[5]
<i>Comments: Short exposure to minimize instrument systematic of first exposure in each orbit.</i>							
11	HAT-P-18 G430L Orbit 5 (STIS.sp.73 3928)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G430L 4300 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO	278 Secs X 10 (2780 Secs) [==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)]	[5]

Proposal 14099 - HAT-P-18 G430L (01) - Measuring the L-T transition for a warm Saturn exoplanet

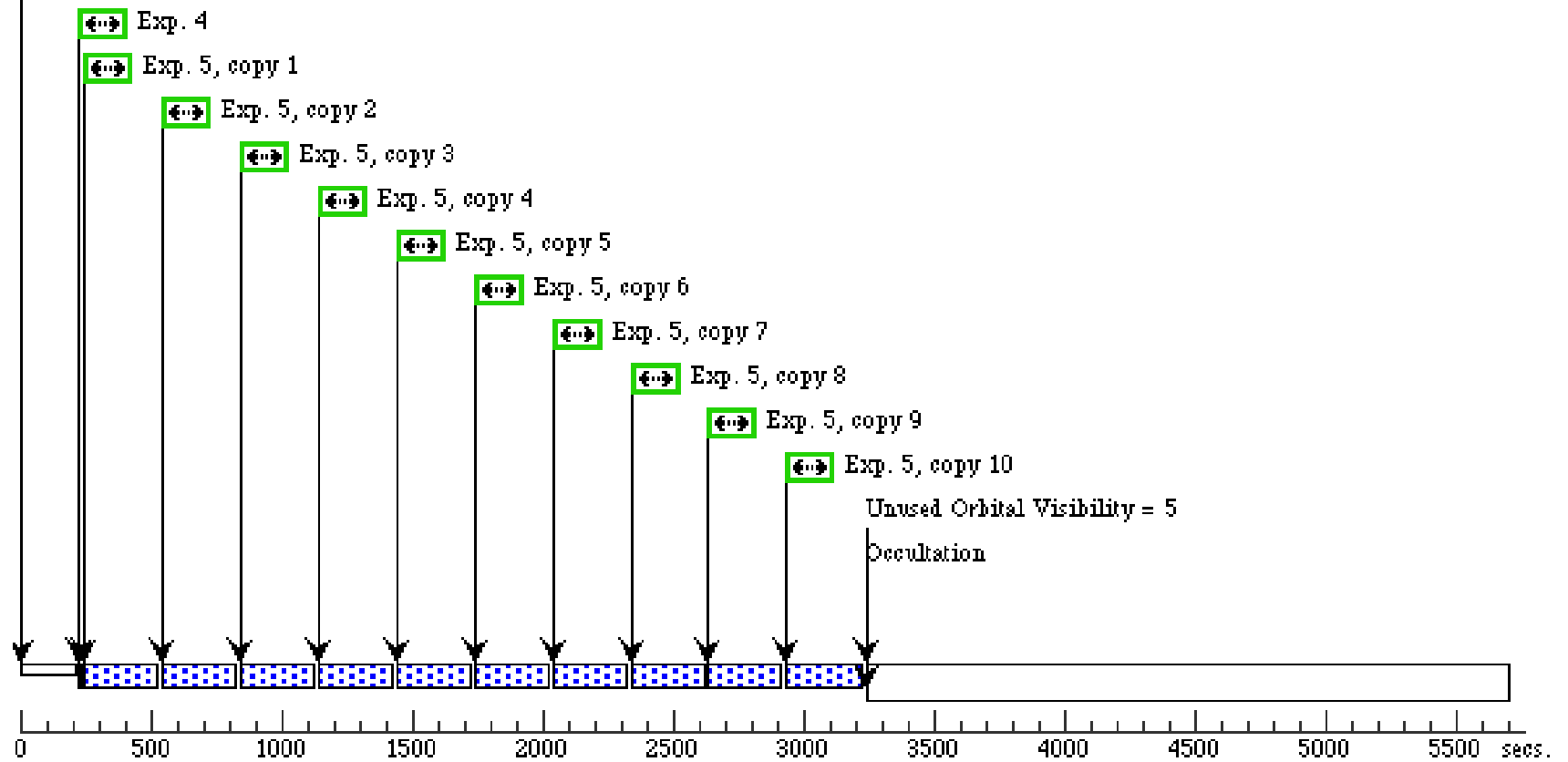
	12 WAVE	WAVE	STIS/CCD, ACCUM, 52X0.2	G430L 4300 A	[==>]	[5]
<i>Comments: Explicit WAVECAL, auto-waves disabled</i>						

Orbit Structure



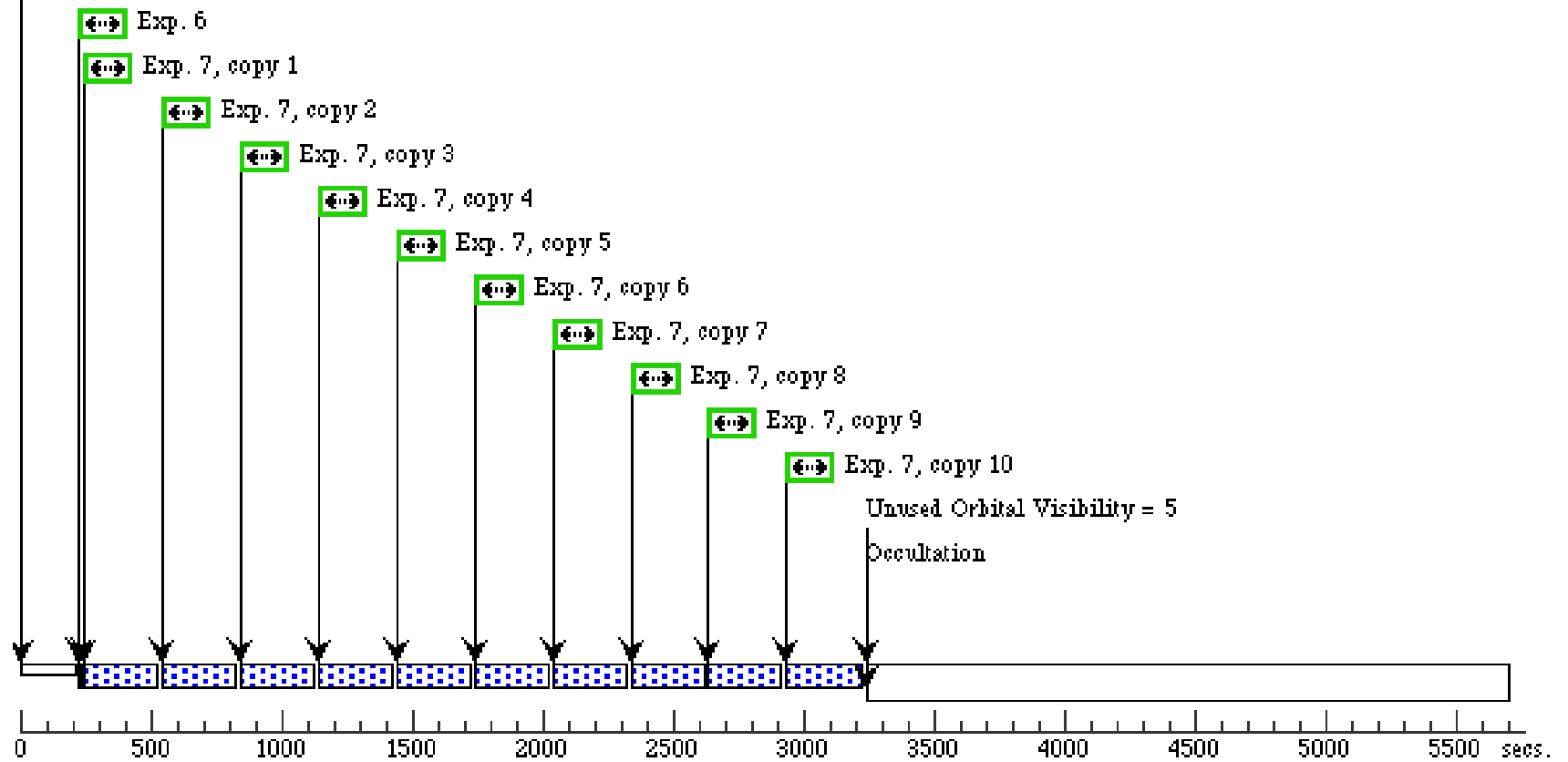
Orbit 2

GS Reacq



Orbit 3

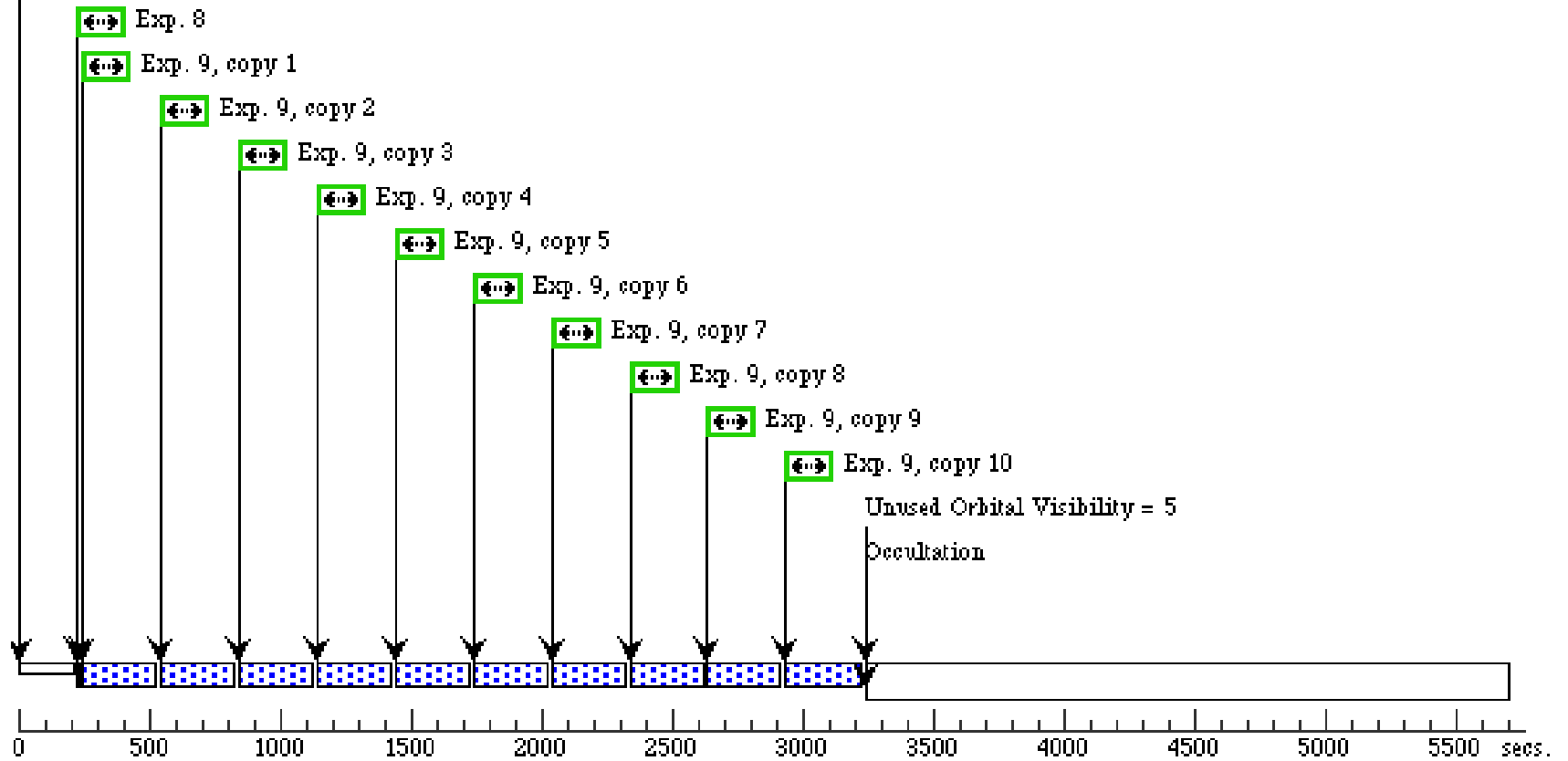
GS Reacq



Orbit 4

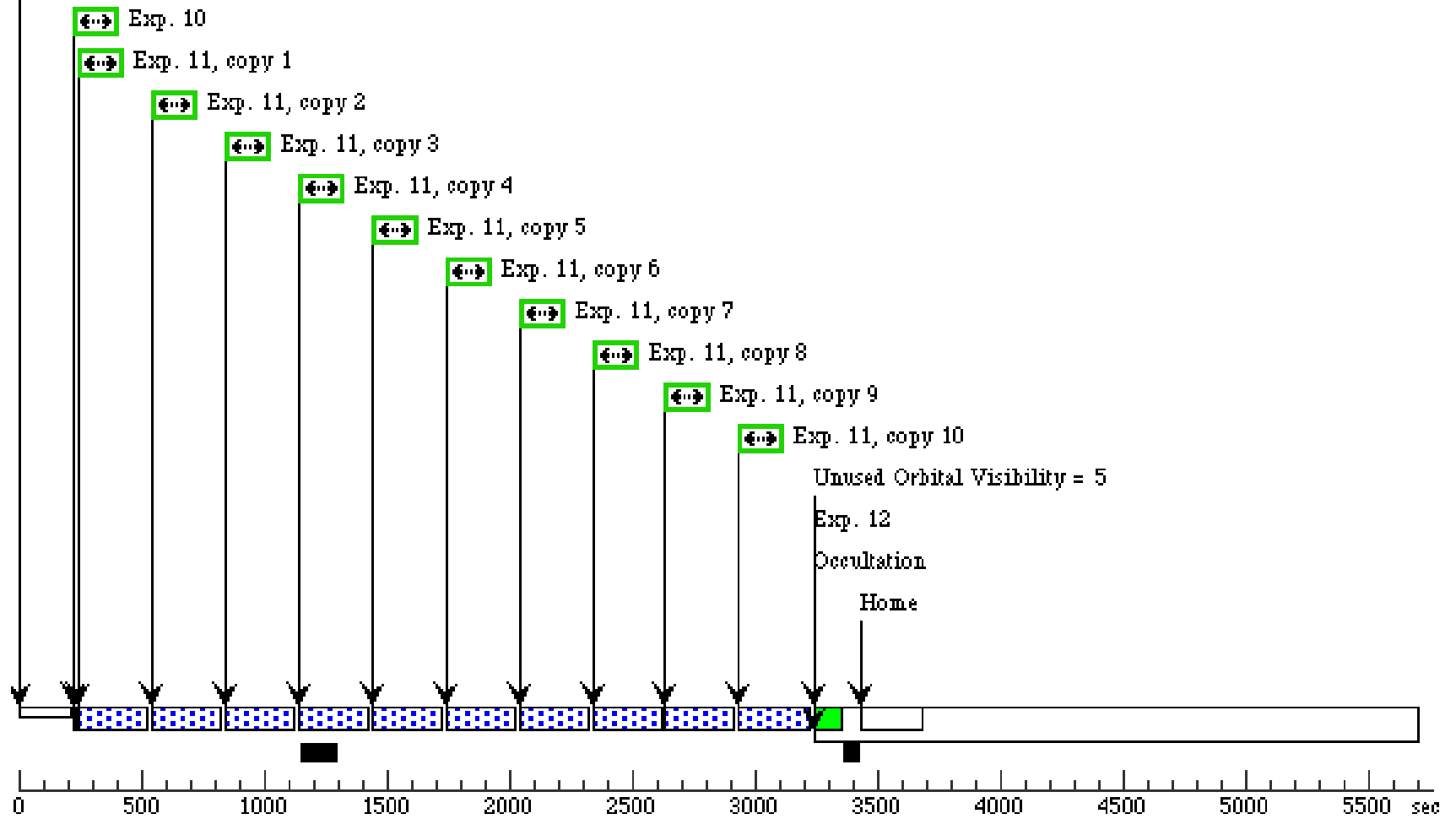
Server Version: 20150609

GS Reacq



Orbit 5

GS Reacq



Proposal 14099 - HAT-P-18 G750L (02) - Measuring the L-T transition for a warm Saturn exoplanet

Thu Nov 05 02:24:20 GMT 2015

Visit	Proposal 14099, HAT-P-18 G750L (02), implementation Diagnostic Status: No Diagnostics Scientific Instruments: STIS/CCD Special Requirements: SCHED 80%; Period 5.50802888 D AND ZERO-PHASE HJD2456461.06765758 <i>Comments: HAT-P-18. G750L. STIS visit. It is essential that the five orbits be scheduled in a continuous block, free of the SAA.</i>					
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes
(1)		HAT-P-18 Alt Name1: GSC-02594-00646	RA: 17 05 23.1510 (256.3464625d) Dec: +33 00 44.97 (33.01249d) Equinox: J2000	Proper Motion RA: -21.3 mas/yr Proper Motion Dec: -44.2 mas/yr Epoch of Position: 2000	V=12.8	Reference Frame: ICRS
<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i>						

Proposal 14099 - HAT-P-18 G750L (02) - Measuring the L-T transition for a warm Saturn exoplanet

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
1	ACQ, phase constrained (STIS.ta.733 341)	(1) HAT-P-18	STIS/CCD, ACQ, F28X50LP	MIRROR	GAIN=4	PHASE 0.9655 TO 0.9680	Sequence 1-3 Non-Int in HAT-P-18 G750 L (02)	1.5 Secs (1.5 Secs) [==>]	[1]
2	HAT-P-18 G750L Orbit 1	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G750L 7751 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO		Sequence 1-3 Non-Int in HAT-P-18 G750 L (02)	100 Secs (100 Secs) [==>]	[1]
<i>Comments: Short exposure to minimize instrument systematic of first exposure of orbit. Note that it is longer than first exposures in subsequent orbits, due to acquisition overhead and to ensure all time is used.</i>									
3	HAT-P-18 G750L Orbit 1 (STIS.sp.73 3930)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G750L 7751 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO		Sequence 1-3 Non-Int in HAT-P-18 G750 L (02)	250 Secs X 8 (2000 Secs) [==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)]	[1]
4	HAT-P-18 G750L Orbit 2	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G750L 7751 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO		Sequence 4-5 Non-Int in HAT-P-18 G750 L (02)	1 Secs (1 Secs) [==>]	[2]
<i>Comments: Short exposure to minimize instrument systematic of first exposure in each orbit.</i>									
5	HAT-P-18 G750L Orbit 2 (STIS.sp.73 3930)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G750L 7751 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO		Sequence 4-5 Non-Int in HAT-P-18 G750 L (02)	250 Secs X 10 (2500 Secs) [==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)]	[2]
6	HAT-P-18 G750L Orbit 3	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G750L 7751 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO		Sequence 6-7 Non-Int in HAT-P-18 G750 L (02)	1 Secs (1 Secs) [==>]	[3]
<i>Comments: Short exposure to minimize instrument systematic of first exposure in each orbit.</i>									

Exposures

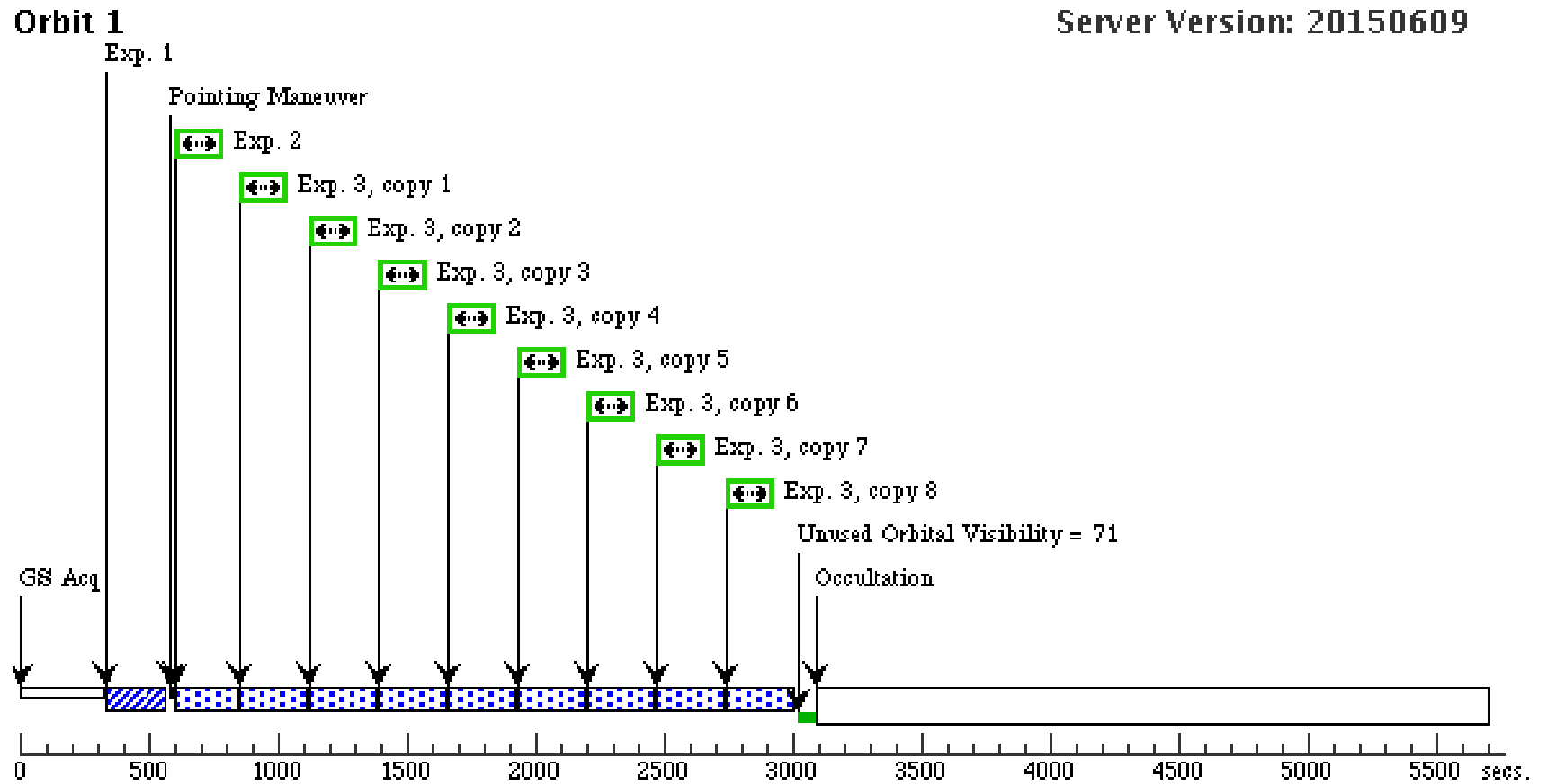
Proposal 14099 - HAT-P-18 G750L (02) - Measuring the L-T transition for a warm Saturn exoplanet

7	HAT-P-18 G750L Orbit 3 (STIS.sp.73 3930)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G750L 7751 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO	Sequence 6-7 Non-Int in HAT-P-18 G750 L (02)	250 Secs X 10 (2500 Secs) [==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)]	[3]	
8	HAT-P-18 G750L Orbit 4	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G750L 7751 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO	Sequence 8-9 Non-Int in HAT-P-18 G750 L (02)	1 Secs (1 Secs) [==>]	[4]	
<i>Comments: Short exposure to minimize instrument systematic of first exposure in each orbit.</i>									
9	HAT-P-18 G750L Orbit 4 (STIS.sp.73 3930)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G750L 7751 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO	Sequence 8-9 Non-Int in HAT-P-18 G750 L (02)	250 Secs X 10 (2500 Secs) [==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)]	[4]	
10	HAT-P-18 G750L Orbit 5	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G750L 7751 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO	Sequence 10-11 Non-Int in HAT-P-18 G750L (02)	1 Secs (1 Secs) [==>]	[5]	
<i>Comments: Short exposure to minimize instrument systematic of first exposure in each orbit.</i>									
11	HAT-P-18 G750L Orbit 5 (STIS.sp.73 3930)	(1) HAT-P-18	STIS/CCD, ACCUM, 52X2	G750L 7751 A	CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0; WAVECAL=NO	Sequence 10-11 Non-Int in HAT-P-18 G750L (02)	250 Secs X 10 (2500 Secs) [==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)]	[5]	

Proposal 14099 - HAT-P-18 G750L (02) - Measuring the L-T transition for a warm Saturn exoplanet

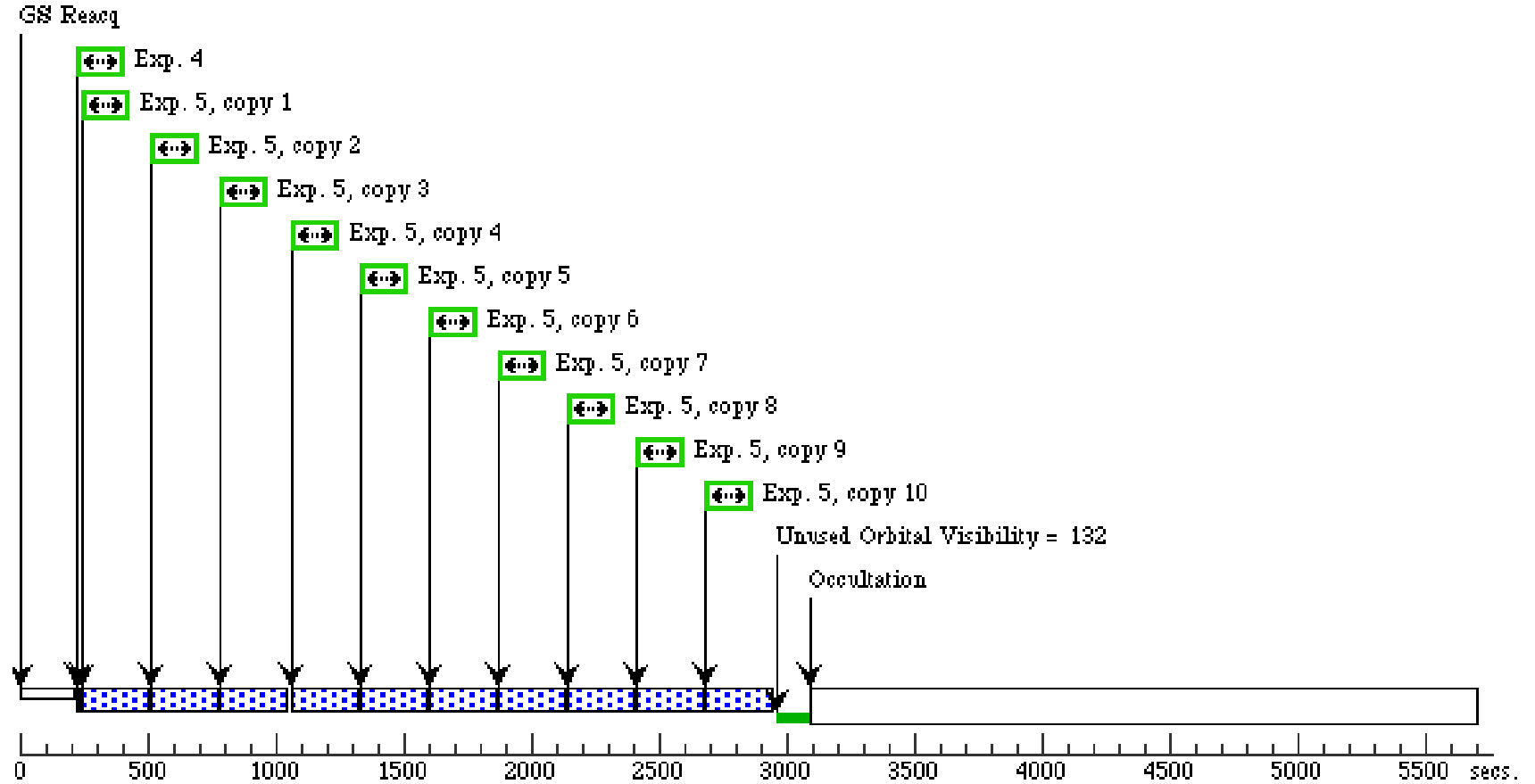
12	WAVE	WAVE	STIS/CCD, ACCUM, 52X0.2	G750L 7751 A		[==>]	[5]
<i>Comments: Explicit WAVECAL, auto-waves disabled</i>							
13	Flat	CCDFLAT	STIS/CCD, ACCUM, 0.3X0.09	G750L 7751 A		[==>(Copy 1)] [==>(Copy 2)]	[5]
14	Tungsten La mp	NONE	STIS/CCD, ACCUM, 0.3X0.09	G750L 7751 A	LAMP=TUNGSTE N; CR-SPLIT=NO; GAIN=4; SIZEAXIS2=128.0	240 Secs X 4 (960 Secs) [==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)]	[5]

Orbit Structure



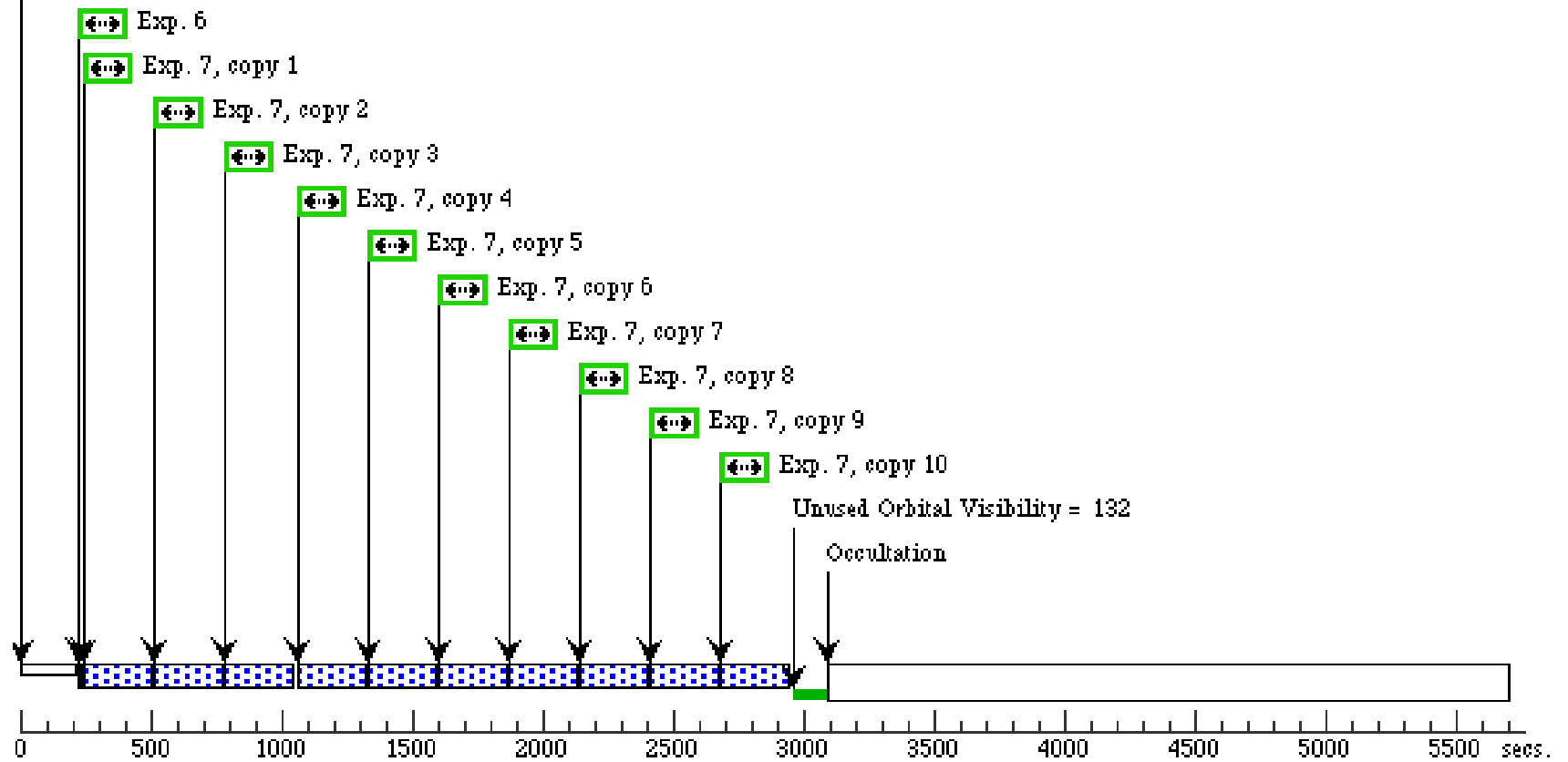
Orbit 2

Server Version: 20150609



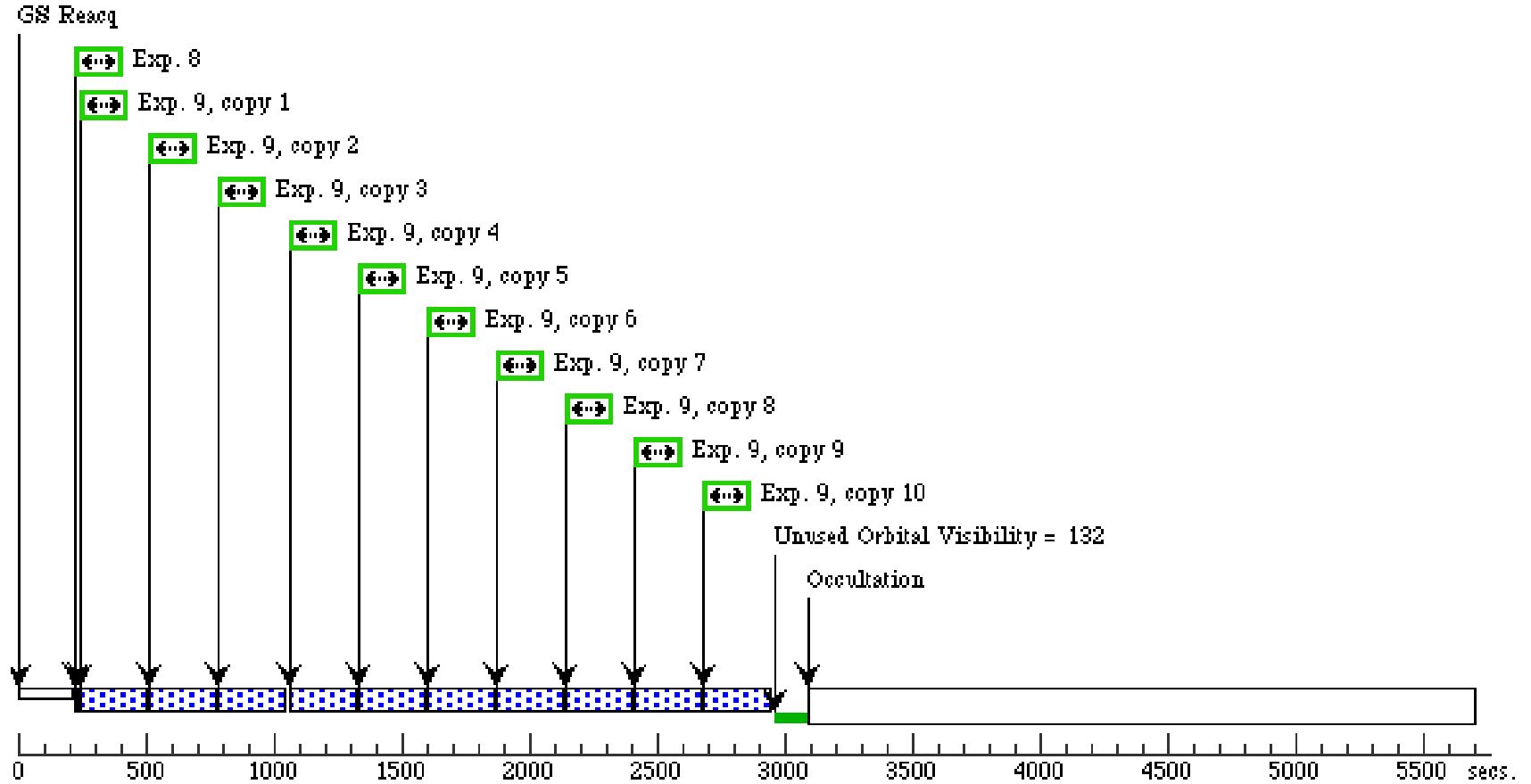
Orbit 3

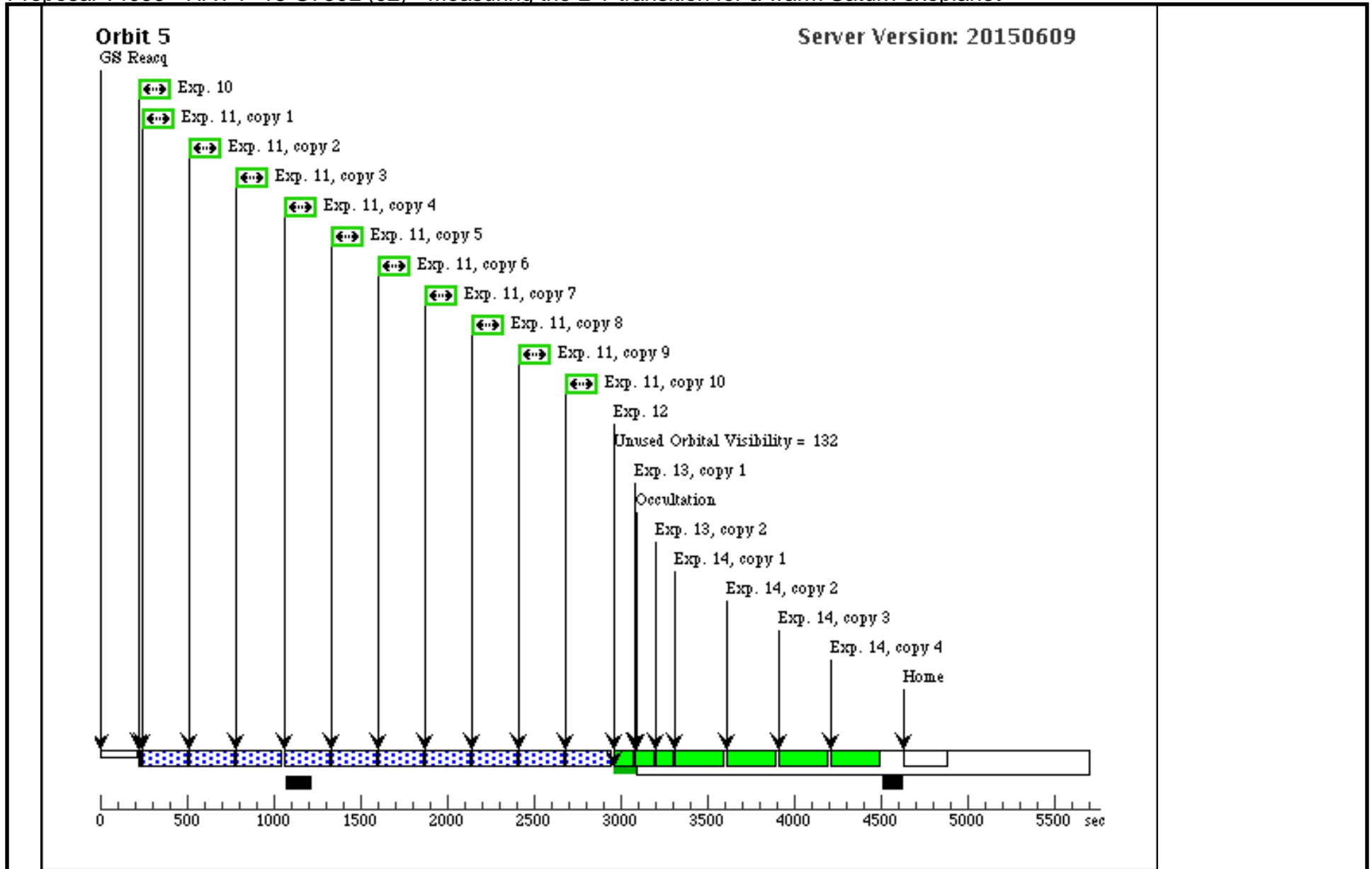
GS Reacq



Orbit 4

Server Version: 20150609





Proposal 14099 - HAT-P-18 WFC3 (04) - Measuring the L-T transition for a warm Saturn exoplanet

Thu Nov 05 02:24:20 GMT 2015

Visit	<p>Proposal 14099, HAT-P-18 WFC3 (04), completed</p> <p>Diagnostic Status: No Diagnostics</p> <p>Scientific Instruments: WFC3/IR</p> <p>Special Requirements: Period 5.50802888 D AND ZERO-PHASE HJD2456461.06765758</p> <p><i>Comments: HAT-P-18. G141. WFC3 visit. It is essential that the five orbits be scheduled in a continuous block, free of the SAA.</i></p> <p><i>We will use the spatial scanning mode to lengthen the spectrum along the slit during the exposures, to avoid saturation on relatively long exposures. This mode greatly increases the efficiency.</i></p>					
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes
(1)		HAT-P-18	RA: 17 05 23.1510 (256.3464625d)	Proper Motion RA: -21.3 mas/yr	V=12.8	Reference Frame: ICRS
		Alt Name1: GSC-02594-00646	Dec: +33 00 44.97 (33.01249d) Equinox: J2000	Proper Motion Dec: -44.2 mas/yr Epoch of Position: 2000		
	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i>					

Proposal 14099 - HAT-P-18 WFC3 (04) - Measuring the L-T transition for a warm Saturn exoplanet

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
1	HAT-P-18 WFC3, phase constrained (WFC3IR.im.733974)	(1) HAT-P-18	WFC3/IR, MULTIACCUM, GRISM256	F130N	SAMP-SEQ=RAPID ; NSAMP=7	PHASE 0.9655 TO 0.9680; GS ACQ SCENARIO BASE1B3		1.944705 Secs (1.945 Secs) [==>]	[1]
<p>Comments: Direct image for wavelength calibration. Phase constrained.</p>									

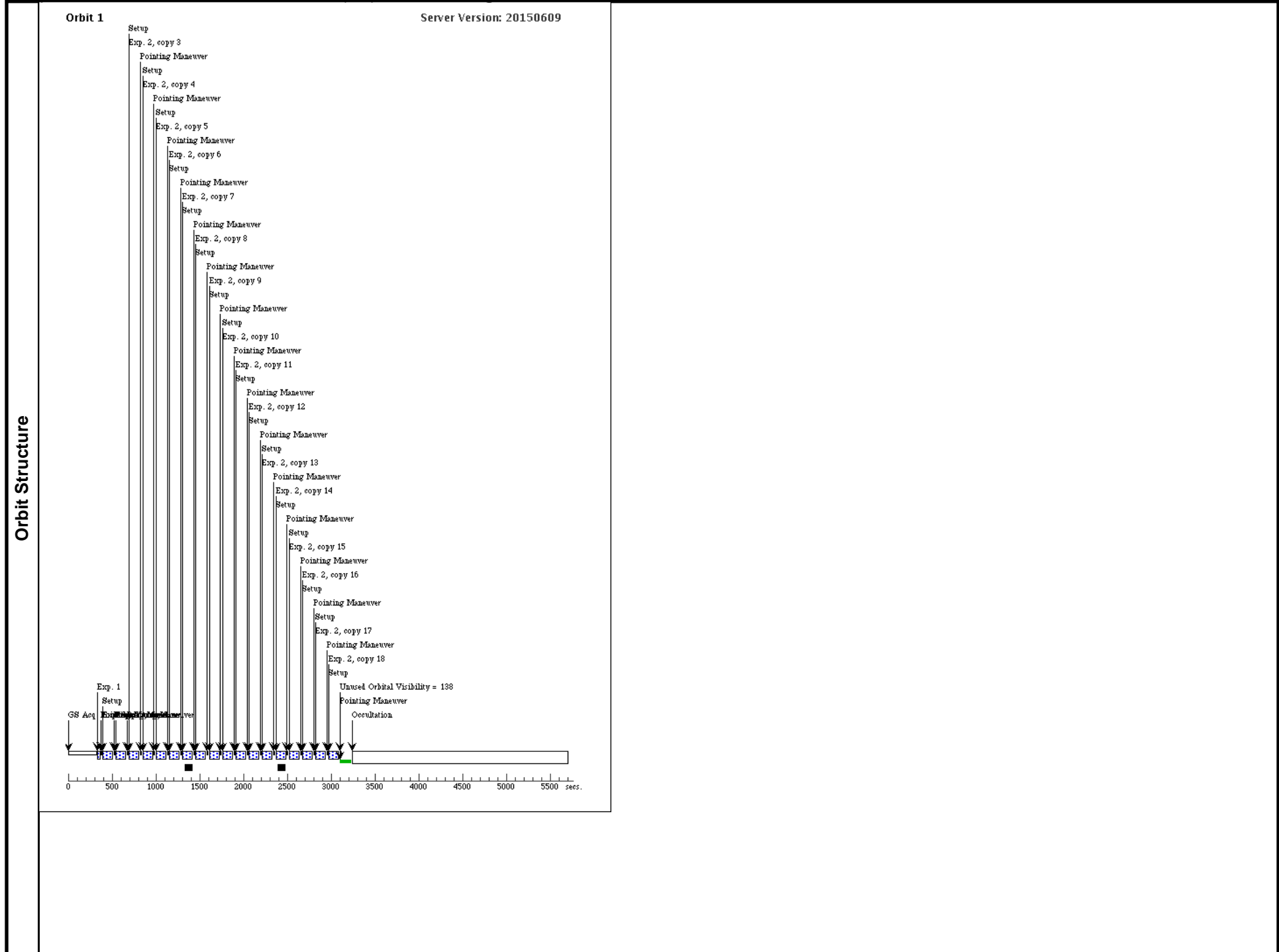
Exposures

	<p>[==>(Copy 39)] [==>(Copy 40)] [==>(Copy 41)] [==>(Copy 42)] [==>(Copy 43)] [==>(Copy 44)] [==>(Copy 45)] [==>(Copy 46)] [==>(Copy 47)] [==>(Copy 48)] [==>(Copy 49)] [==>(Copy 50)] [==>(Copy 51)] [==>(Copy 52)] [==>(Copy 53)] [==>(Copy 54)] [==>(Copy 55)] [==>(Copy 56)] [==>(Copy 57)] [==>(Copy 58)]</p>	<p>[3]</p>
	<p>[==>(Copy 59)] [==>(Copy 60)] [==>(Copy 61)] [==>(Copy 62)] [==>(Copy 63)] [==>(Copy 64)] [==>(Copy 65)] [==>(Copy 66)] [==>(Copy 67)] [==>(Copy 68)] [==>(Copy 69)] [==>(Copy 70)] [==>(Copy 71)] [==>(Copy 72)] [==>(Copy 73)] [==>(Copy 74)] [==>(Copy 75)] [==>(Copy 76)] [==>(Copy 77)] [==>(Copy 78)]</p>	<p>[4]</p>

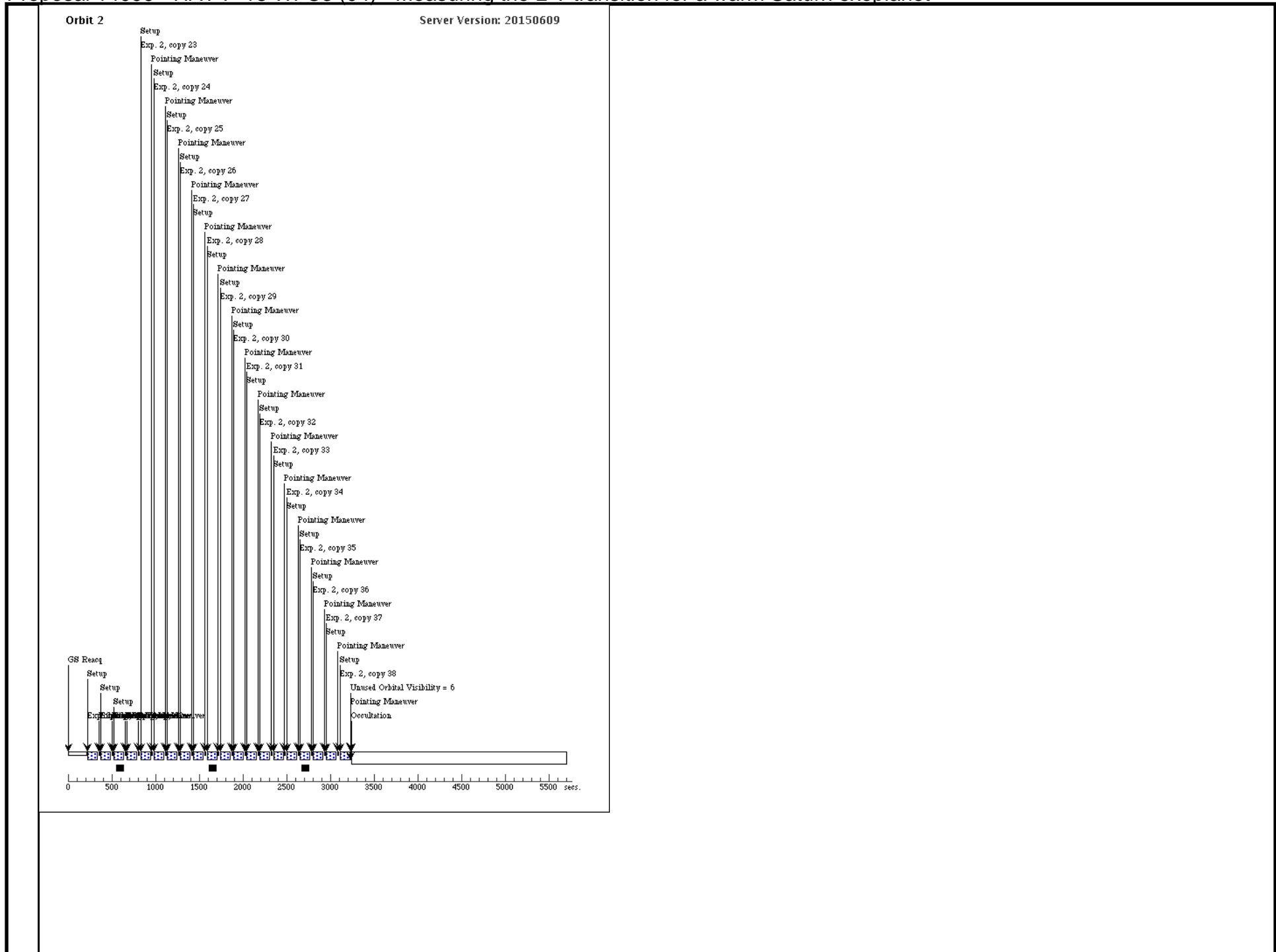
Proposal 14099 - HAT-P-18 WFC3 (04) - Measuring the L-T transition for a warm Saturn exoplanet

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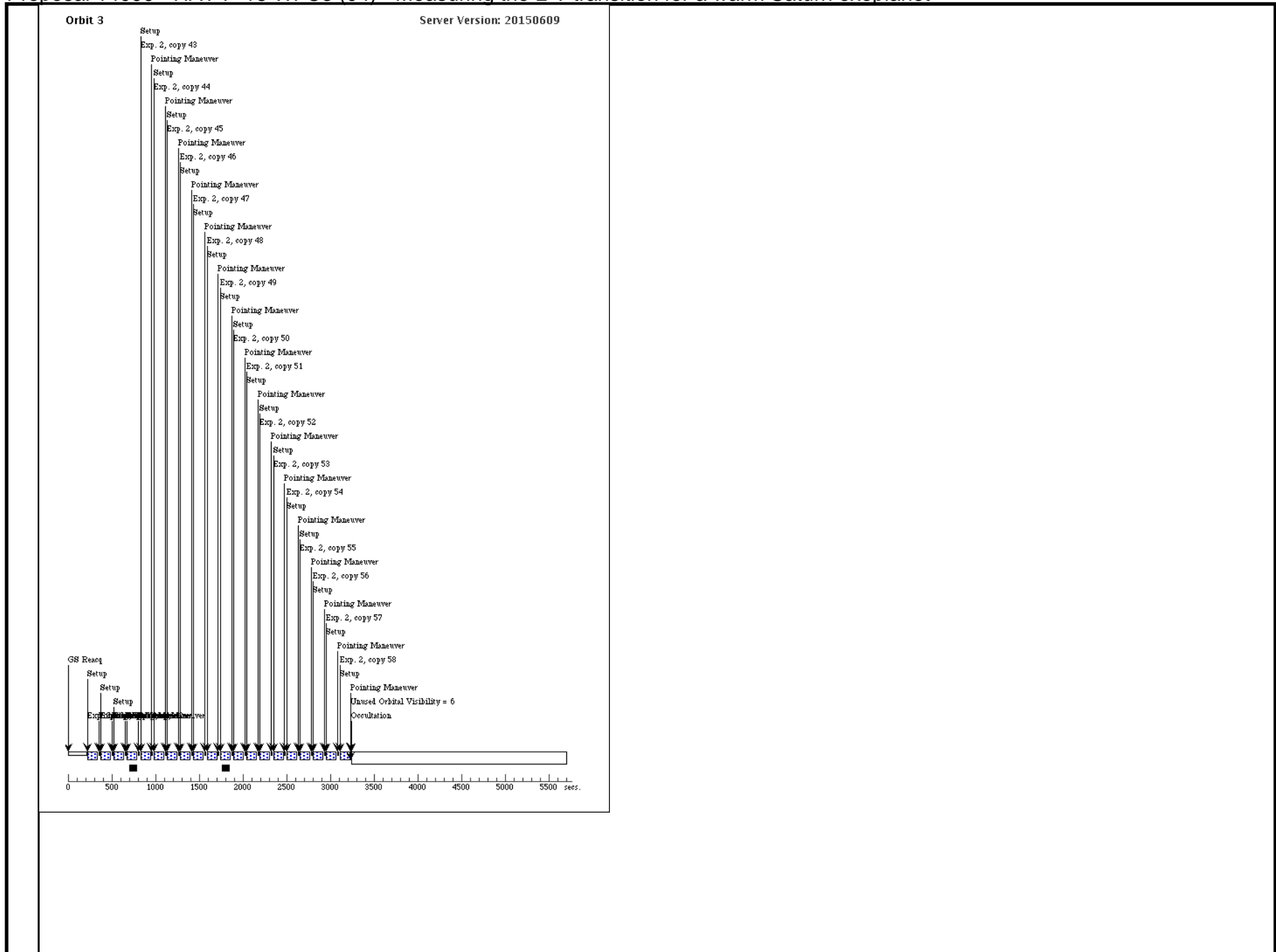
Comments: Spatial scan.



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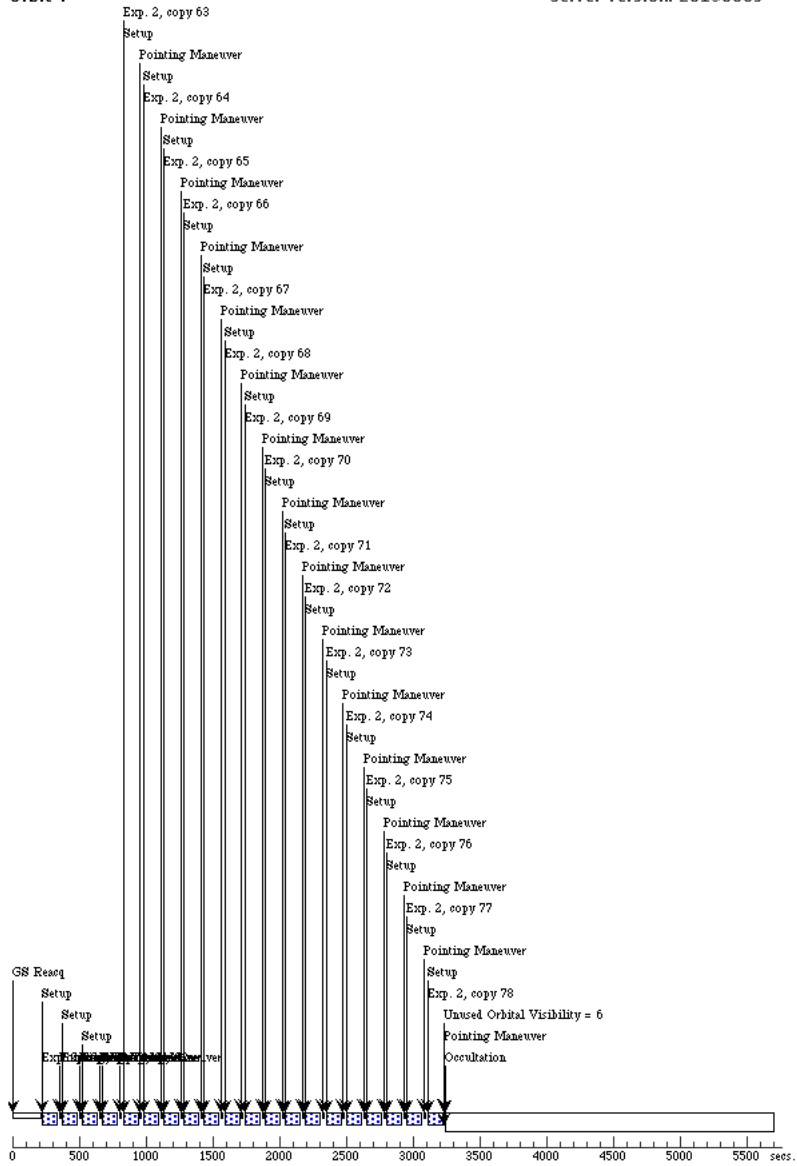
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