



14100 - A Direct Imaging Experiment to Determine the Origin of H₂ Emission from M dwarf Exoplanetary Systems

Cycle: 23, Proposal Category: GO

(UV Initiative)

(Availability Mode: SUPPORTED)

INVESTIGATORS

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VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) GJ832	ACS/SBC	4	07-Jul-2015 21:05:01.0	yes
02	(1) GJ832	ACS/SBC	4	07-Jul-2015 21:05:05.0	yes

8 Total Orbits Used

ABSTRACT

Recent HST-COS observations of M dwarfs hosting exoplanetary systems have detected photo-excited H₂ emission in these environments, in contrast to previous HST-STIS and IUE observations of non-planet hosting M dwarfs where no H₂ emission is observed. Three plausible origins for this emission include: 1) starspots or the stellar photosphere, 2) the heated atmospheres of orbiting planets, or 3) a second generation circumstellar disk produced by atmospheric mass-loss from the planets in these systems. While direct imaging of exoplanets is extremely challenging in the

visible/IR owing to the large star/planet contrast ratio, we will take advantage of the favorable star/planet contrast ratio in the FUV to ascertain the origin of the H₂ emission. This proposal will use multi-band FUV imaging with ACS/SBC to directly image one well-positioned M dwarf exoplanetary system to measure the distribution of the H₂ gas: GJ 832, at $d = 4.9$ pc with a Jovian-mass exoplanet at 0.7 arcseconds separation. Using a combination of F122M, F140LP, and F165LP imaging, we can isolate the relative spatial distributions of Lyman-alpha, C IV, and H₂ providing a stellar reference (C IV from the stellar atmosphere) and two tracers that may originate from exoplanets or their evaporated atmospheres (Lyman-alpha and H₂). A direct H₂ image of GJ 832b or a circumstellar disk would demonstrate a new observational tool for the detection of planetary systems orbiting low-mass stars and possibly a new means of measuring atmospheric mass loss rates.

OBSERVING DESCRIPTION

The H₂ Lyman band systems are photoexcited ("pumped") by stellar Ly photons and have transitions that span the FUV bandpass from 1150 - 1650 Ang, but the brightest lines reside between 1430 - 1510 Ang. Therefore, in order to image the spatial origin of the H₂ in the GJ 832 system, we wish to isolate this spectral region. In addition, recent work has explored the spectral characteristics of H I Ly emission from hot Jupiters (Menager et al. 2013), produced mainly by resonant scattering of the stellar emission line. To fully understand the H₂ excitation conditions, we wish to know the spatial distribution of both the Ly pumping radiation and the H₂ emission, therefore we also require Ly imaging of the system. Finally, as a control, we wish to isolate one spectral tracer that has no exoplanetary contribution: C IV emission that is only produced in $T \sim 10^5$ K gas in the stellar atmosphere.

All three of these emissions can be isolated using filters on the ACS/SBC. F122M is a dedicated LyA filter. Combining the long-pass FUV filters on ACS is a well-characterized technique for creating custom filter bandpasses for FUV imaging (e.g., Hayes et al. 2013; Fleming et al. 2015), and we propose to apply this technique to isolate H₂ and C IV emissions from the GJ 832 system.

Taking the average GJ 832 FUV spectrum, and using the STScI on-line exposure time

calculator, we find that a $S/N = 100$ per angular resolution element F122M Ly image of GJ 832 can be acquired in $T_{exp} = 4500$ sec. The remaining ACS/SBC filters are long-pass and include a significant red leak, so a filter difference technique is required to isolate a specific FUV bandpass. Using the filter difference $D150 = F140LP - F165LP$, we can quantify the spectral contributions to this effective narrow-band image from our existing COS spectra: Si IV (9.5 %), C IV (34.9 %), He II (21.1 %), and the 1430 - 1520 Ang H₂ region (14.6 %). The D150 band covers both hot gas and H₂ emission, however the hot gas emission (Si IV, C IV, and HeII) arises exclusively in the stellar atmosphere and therefore any detection at the position of the planet in D150 will be attributable to H₂.

This spectral assignment will be verified using our existing spectroscopy and potential Ly detection at the planetary position. We note that while the H₂

region contains the highest density of photoexcited H₂ lines per wavelength, this 14.6% is contaminated by weak S I and N I emission lines from the lower chromosphere of the parent star. The S I and N I make up approximately one-third of the 1430 - 1510 Ang flux.

Factoring in the contribution from the neutral metal lines, and taking the count outputs from the on-line ACS exposure time calculator, we estimate that if the entire H₂ signal originates from GJ 832b, this signal will be detected at $S/N(H_2) = (X*0.146*(S/N140)^2) / \text{SQRT}((S/N140)^2 + (S/N165)^2)$, where 0.146 is the 14.6% H₂ contribution, X is the 1430 - 1510 Ang flux contributed by H₂ fluorescence ($X \sim 2/3$, measured from the COS data, see previous paragraph), and S/N140 and S/N165 are the F140LP and F165LP signal-to-noise ratios, respectively. Therefore with $S/N = 100$ in both the F140LP and F165LP images, we expect to detect GJ 832b in H₂ to a 7- significance if all of the molecular emission originates on the planet. Similarly, if the molecular emission originates in a second generation disk, we will detect the total emission from the disk using photometric annuli, although the significance will be reduced to 3 - 5 sigma, depending on the angular extent of the disk. Using the observed GJ 832 spectrum and the on-line ACS exposure time calculator, we estimate that $S/N = 100$ can be achieved in 4500 s with F140LP and 9200s with F165LP.

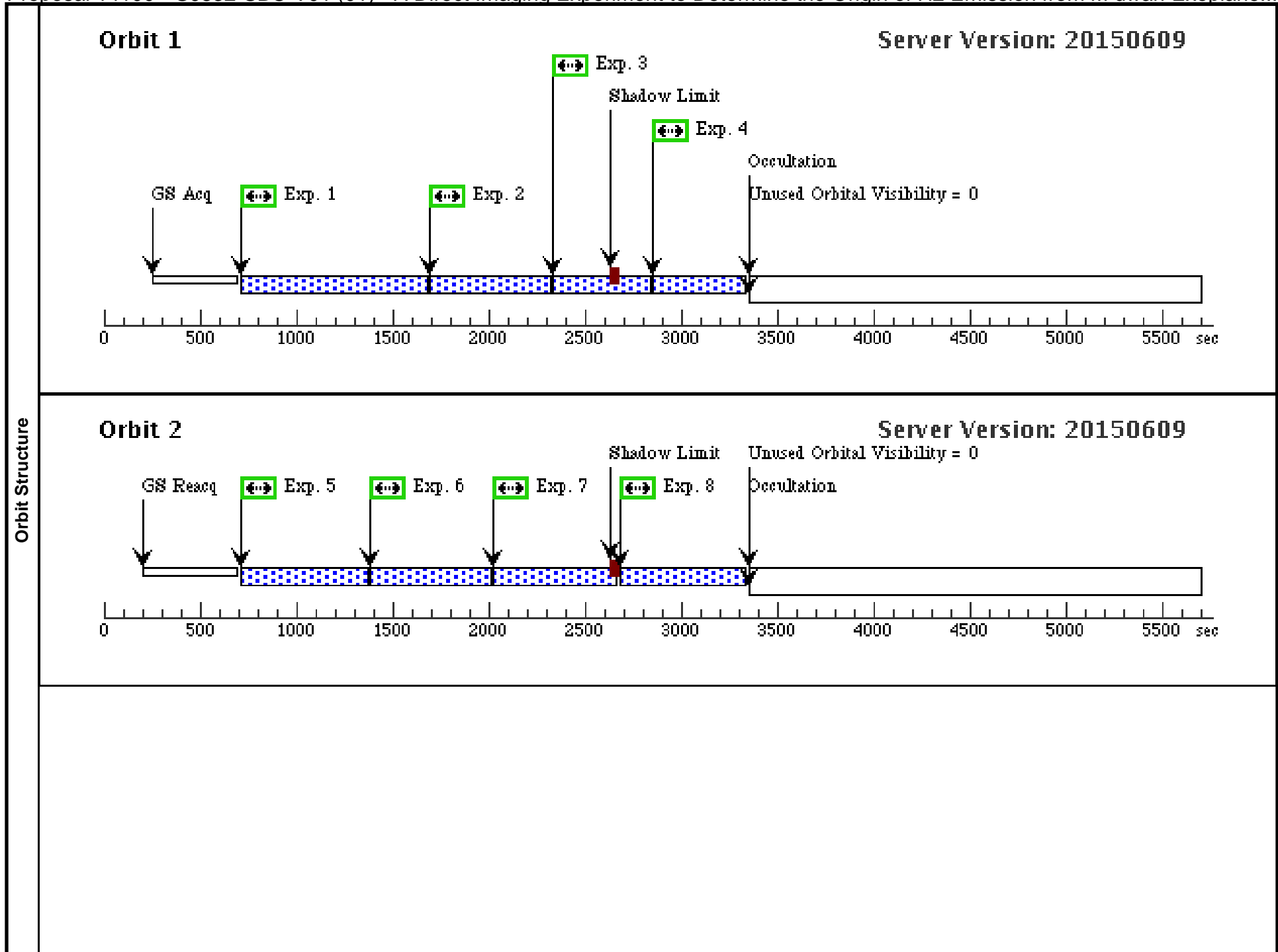
Proposal 14100 - GJ832-SBC-V01 (01) - A Direct Imaging Experiment to Determine the Origin of H2 Emission from M dwarf Exoplanet...

Wed Jul 08 01:05:07 GMT 2015

Visit	<p>Proposal 14100, GJ832-SBC-V01 (01)</p> <p>Diagnostic Status: No Diagnostics</p> <p>Scientific Instruments: ACS/SBC</p> <p>Special Requirements: (none)</p> <p><i>Comments: BOT checked and passed. The only flags are when Aladin mistake the target (GJ 832, M1V) for an O5V star. KF - 06/29/15</i></p>					
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes
(1)		GJ832 Alt Name1: HD-204961	RA: 21 33 33.9753 (323.3915638d) Dec: -49 00 32.42 (-49.00901d) Equinox: J2000	Proper Motion RA: -46.05 mas/yr Proper Motion Dec: -817.63 mas/yr Parallax: 0.20187" Epoch of Position: 2000 Radial Velocity: 4.3 km/sec	V=8.672+/-0.1	Reference Frame: ICRS
<p><i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i></p> <p><i>Coord and proper motions - 2007A&A...474..653V</i></p> <p><i>Extended=YES</i></p>						

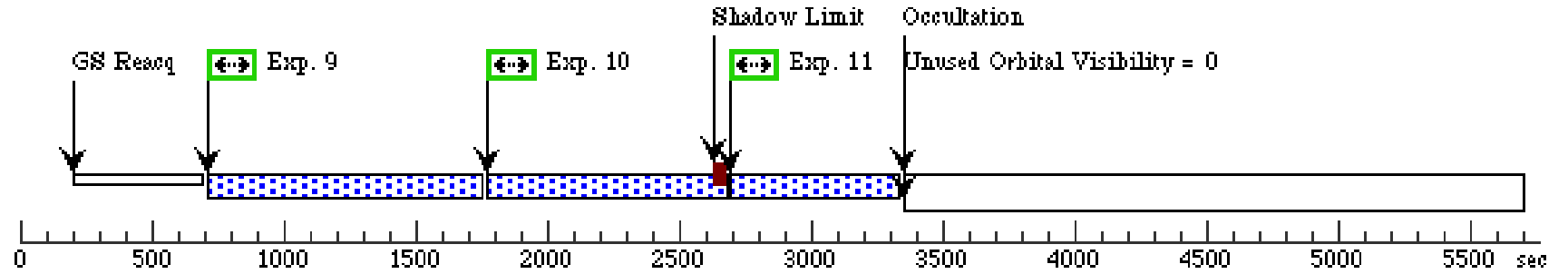
Proposal 14100 - GJ832-SBC-V01 (01) - A Direct Imaging Experiment to Determine the Origin of H2 Emission from M dwarf Exoplanet...

Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
	1	V01-F122M _1 (ACS.im.71 5035)	(1) GJ832	ACS/SBC, ACCUM, SBC	F122M			SHADOW		900 Secs (900 Secs) [==>]	[1]
	2	V01-F122M _2 (ACS.im.71 5035)	(1) GJ832	ACS/SBC, ACCUM, SBC	F122M			SHADOW		600 Secs (600 Secs) [==>]	[1]
	3	V01-F140L P_1 (ACS.im.71 5039)	(1) GJ832	ACS/SBC, ACCUM, SBC	F140LP					450 Secs (450 Secs) [==>]	[1]
	4	V01-F140L P_2 (ACS.im.71 5039)	(1) GJ832	ACS/SBC, ACCUM, SBC	F140LP					453 Secs (453 Secs) [==>]	[1]
	5	V01-F122M _3 (ACS.im.71 5035)	(1) GJ832	ACS/SBC, ACCUM, SBC	F122M			SHADOW		600 Secs (600 Secs) [==>]	[2]
	6	V01-F122M _4 (ACS.im.71 5035)	(1) GJ832	ACS/SBC, ACCUM, SBC	F122M			SHADOW		600 Secs (600 Secs) [==>]	[2]
	7	V01-F165L P_1 (ACS.im.71 5038)	(1) GJ832	ACS/SBC, ACCUM, SBC	F165LP					600 Secs (600 Secs) [==>]	[2]
	8	V01-F165L P_2 (ACS.im.71 5038)	(1) GJ832	ACS/SBC, ACCUM, SBC	F165LP					627 Secs (627 Secs) [==>]	[2]
	9	V01-F122M _5 (ACS.im.71 5035)	(1) GJ832	ACS/SBC, ACCUM, SBC	F122M			SHADOW		1000 Secs (1000 Secs) [==>]	[3]
	10	V01-F140L P_3 (ACS.im.71 5039)	(1) GJ832	ACS/SBC, ACCUM, SBC	F140LP					850 Secs (850 Secs) [==>]	[3]
	11	V01-F165L P_3 (ACS.im.71 5038)	(1) GJ832	ACS/SBC, ACCUM, SBC	F165LP					589 Secs (589 Secs) [==>]	[3]
	12	V01-F122M _6 (ACS.im.71 5035)	(1) GJ832	ACS/SBC, ACCUM, SBC	F122M			SHADOW		1000 Secs (1000 Secs) [==>]	[4]
	13	V01-F140L P_4 (ACS.im.71 5039)	(1) GJ832	ACS/SBC, ACCUM, SBC	F140LP					850 Secs (850 Secs) [==>]	[4]
14	V01-F165L P_4 (ACS.im.71 5038)	(1) GJ832	ACS/SBC, ACCUM, SBC	F165LP					589 Secs (589 Secs) [==>]	[4]	



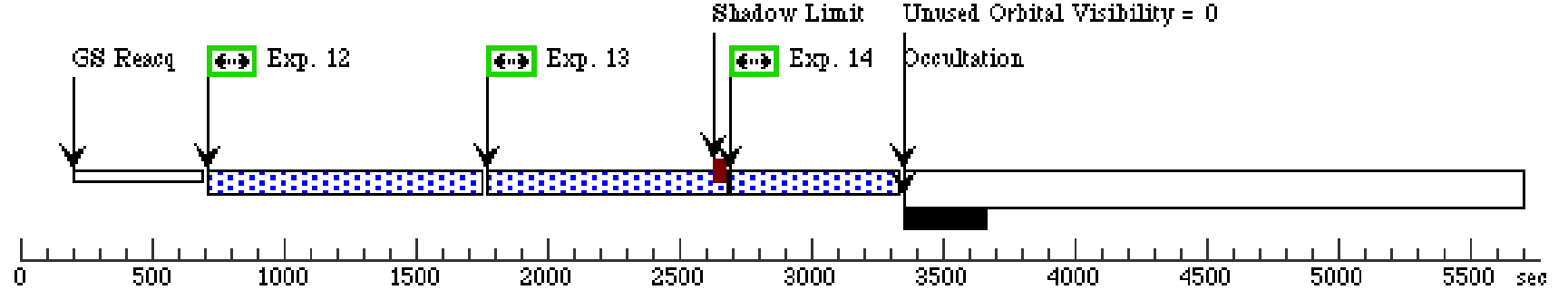
Orbit 3

Server Version: 20150609



Orbit 4

Server Version: 20150609



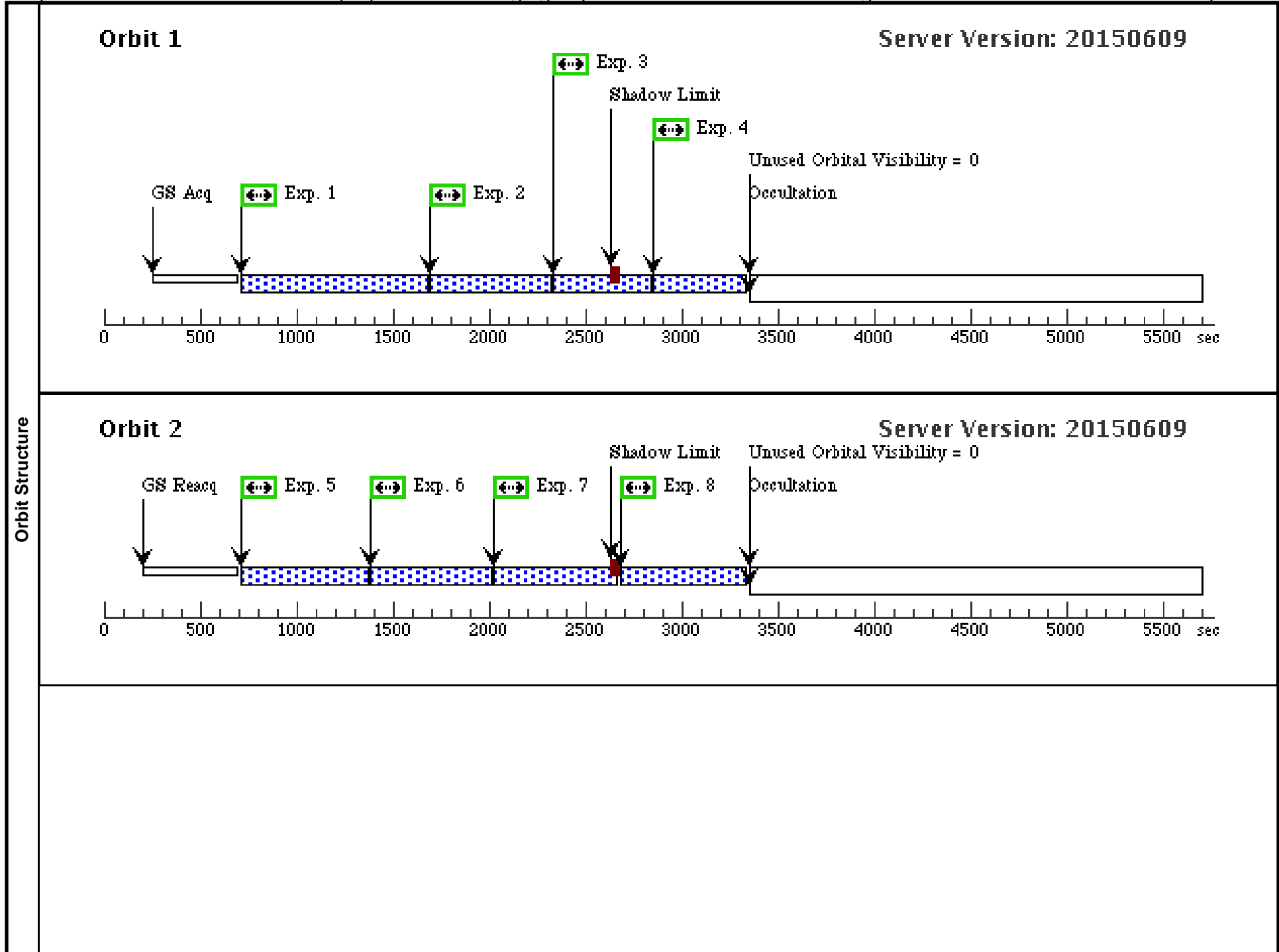
Proposal 14100 - GJ832-SBC-V02 (02) - A Direct Imaging Experiment to Determine the Origin of H2 Emission from M dwarf Exoplanet...

Wed Jul 08 01:05:08 GMT 2015

Visit	<p>Proposal 14100, GJ832-SBC-V02 (02)</p> <p>Diagnostic Status: No Diagnostics</p> <p>Scientific Instruments: ACS/SBC</p> <p>Special Requirements: (none)</p> <p><i>Comments: BOT checked and passed. The only flags are when Aladin mistake the target (GJ 832, M1V) for an O5V star. KF - 06/29/15</i></p>					
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes
(1)		GJ832 Alt Name1: HD-204961	RA: 21 33 33.9753 (323.3915638d) Dec: -49 00 32.42 (-49.00901d) Equinox: J2000	Proper Motion RA: -46.05 mas/yr Proper Motion Dec: -817.63 mas/yr Parallax: 0.20187" Epoch of Position: 2000 Radial Velocity: 4.3 km/sec	V=8.672+/-0.1	Reference Frame: ICRS
<p><i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i></p> <p><i>Coord and proper motions - 2007A&A...474..653V</i></p> <p><i>Extended=YES</i></p>						

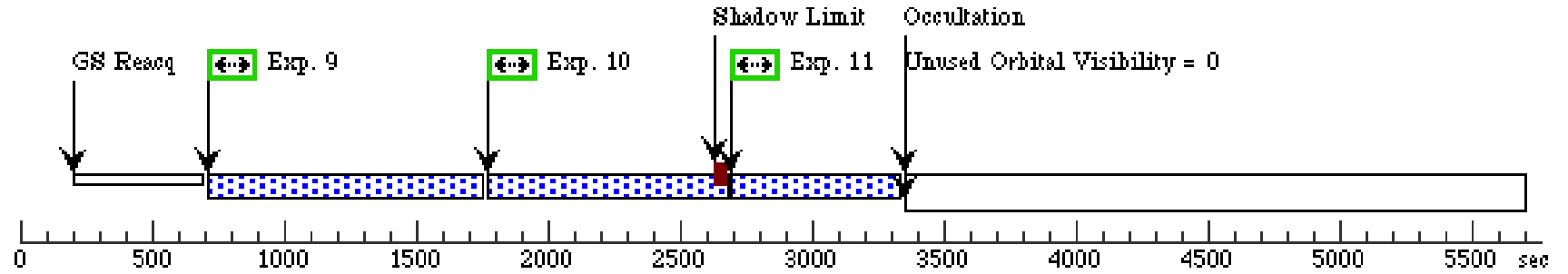
Proposal 14100 - GJ832-SBC-V02 (02) - A Direct Imaging Experiment to Determine the Origin of H2 Emission from M dwarf Exoplanet...

Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
	1	V02-F122M _1 (ACS.im.71 5035)	(1) GJ832	ACS/SBC, ACCUM, SBC	F122M			SHADOW		900 Secs (900 Secs) [==>]	[1]
	2	V02-F122M _2 (ACS.im.71 5035)	(1) GJ832	ACS/SBC, ACCUM, SBC	F122M			SHADOW		600 Secs (600 Secs) [==>]	[1]
	3	V02-F140L P_1 (ACS.im.71 5039)	(1) GJ832	ACS/SBC, ACCUM, SBC	F140LP					450 Secs (450 Secs) [==>]	[1]
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	6	V02-F122M _4 (ACS.im.71 5035)	(1) GJ832	ACS/SBC, ACCUM, SBC	F122M			SHADOW		600 Secs (600 Secs) [==>]	[2]
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	8	V02-F165L P_2 (ACS.im.71 5038)	(1) GJ832	ACS/SBC, ACCUM, SBC	F165LP					627 Secs (627 Secs) [==>]	[2]
	9	V02-F122M _5 (ACS.im.71 5035)	(1) GJ832	ACS/SBC, ACCUM, SBC	F122M			SHADOW		1000 Secs (1000 Secs) [==>]	[3]
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	11	V02-F165L P_3 (ACS.im.71 5038)	(1) GJ832	ACS/SBC, ACCUM, SBC	F165LP					589 Secs (589 Secs) [==>]	[3]
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	13	V02-F140L P_4 (ACS.im.71 5039)	(1) GJ832	ACS/SBC, ACCUM, SBC	F140LP					850 Secs (850 Secs) [==>]	[4]
14	V02-F165L P_4 (ACS.im.71 5038)	(1) GJ832	ACS/SBC, ACCUM, SBC	F165LP					589 Secs (589 Secs) [==>]	[4]	



Orbit 3

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Orbit 4

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