



14598 - The Incredibly Long-Lived SN 2005ip

Cycle: 24, Proposal Category: GO

(UV Initiative)

(Availability Mode: SUPPORTED)

INVESTIGATORS

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VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) SN-2005IP	STIS/CCD STIS/FUV-MAMA STIS/NUV-MAMA	5	29-Jul-2016 13:50:13.0	yes

5 Total Orbits Used

ABSTRACT

Type II_n supernovae (SNe II_n) are defined by their relatively narrow spectral line features associated with a dense circumstellar medium (CSM) formed by the progenitor star. The nature of the progenitor and mass loss remains relatively unknown. Shock interaction with the dense CSM can

often result in significant UV emission for several years post-explosion, thereby probing the CSM characteristics, progenitor mass loss history and, ultimately, the progenitor itself. The Type IIn SN 2005ip proves to be one of the most interesting and well-studied targets within this subclass. Compared to all other supernovae, SN 2005ip is the most luminous for its age. Now more than 11 years post-explosion, the SN has released $>10^{51}$ erg throughout its lifetime as the forward shock continues to collide with a dense CSM. Here we propose HST/STIS-MAMA UV observations of SN 2005ip to investigate the massive CSM. When accounting for the shock travel time, these observations will probe material lost from the progenitor more than 1000 years prior to the explosion. We already have a single HST/STIS spectrum of SN 2005ip from 2014, which was obtained while the shock was still within a higher mass regime. With just 5 orbits, a second spectrum will allow us to directly trace the evolution of the CSM and produce new constraints on the pre-SN mass-loss history. Coinciding with Cycle 24's UV Initiative, this program offers new insight regarding both the progenitor and explosion characteristics of the SN IIn subclass.

OBSERVING DESCRIPTION

We propose to obtain a single STIS-MAMA spectrum of SN 2005ip in both the G140 and G230L bandpasses. The goal is to detect UV signatures of CSM interaction to (1) constrain the CSM characteristics, including the geometry and composition, and (2) confirm emission from CSM interaction as the heating source in the warm-dust model presented by Fox et al. (2011). We choose STIS over COS because, while COS is optimized for isolated faint point-sources, STIS offers spatially resolved spectra that exploit the intrinsically high resolution of HST over a large spectral range (accommodating both the C III] 1909 and Mg II 2800 lines). The long slit length will allow us to subtract contamination from zodiacal light, earthshine, the host galaxy, and H II regions.

Proposal 14598 - STIS-SN2005ip (01) - The Incredibly Long-Lived SN 2005ip

Fri Jul 29 17:50:15 GMT 2016

Visit	Proposal 14598, STIS-SN2005ip (01), implementation Diagnostic Status: No Diagnostics Scientific Instruments: STIS/NUV-MAMA, STIS/CCD, STIS/FUV-MAMA Special Requirements: (none) <i>Comments: Normal ACO.</i>					
	Patterns	#	Primary Pattern	Secondary Pattern	Exposures	
	(1)	Pattern Type=STIS-ALONG-SLIT Coordinate Frame=POS-TARG Purpose=DITHER Pattern Orientation=90.0 Number Of Points=4 Angle Between Sides= Point Spacing=1 Center Pattern=true Line Spacing=		(2), (3), (4), (5), (6)		
Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
	(1)	SN-2005IP	RA: 09 32 6.4200 (143.0267500d) Dec: +08 26 44.40 (8.44567d) Equinox: J2000		V=20.5	Reference Frame: ICRS
	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i>					

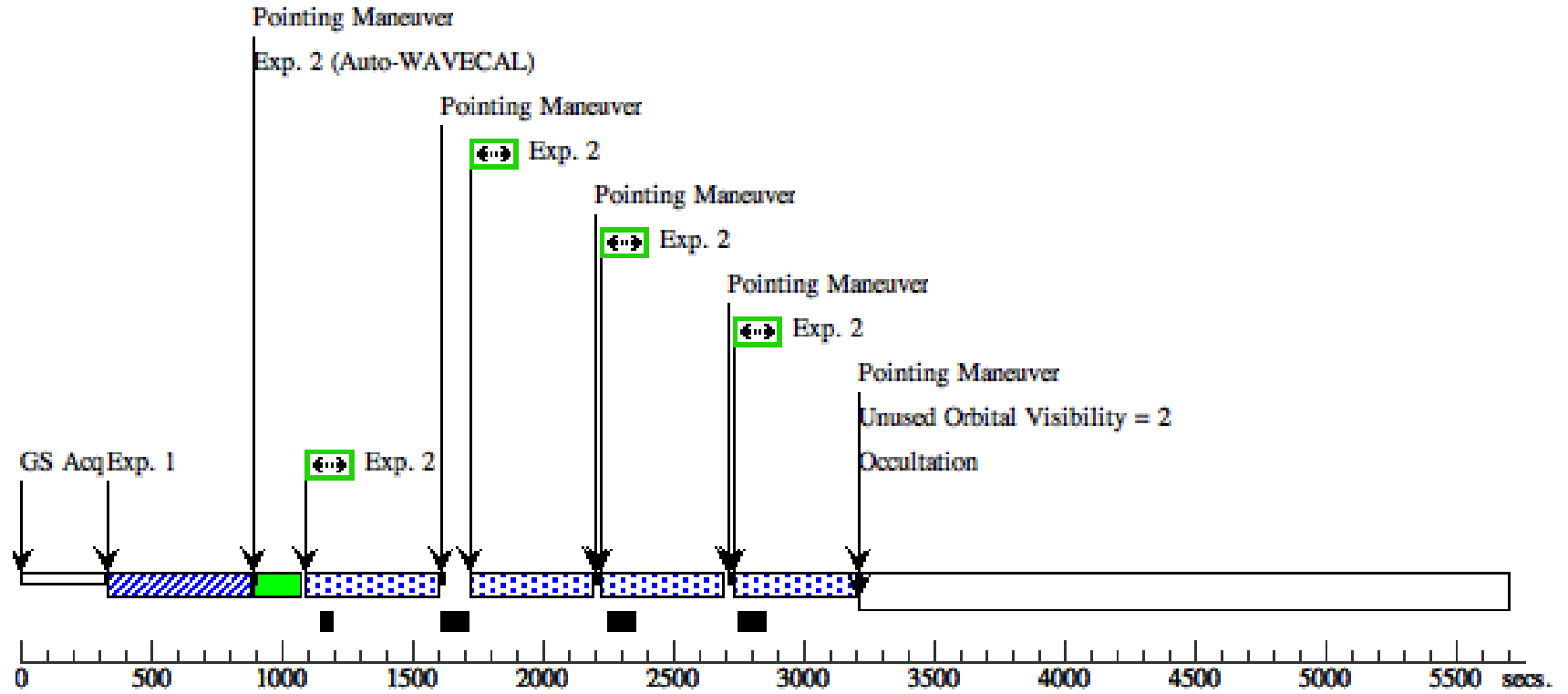
Proposal 14598 - STIS-SN2005ip (01) - The Incredibly Long-Lived SN 2005ip

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
Exposures	1	ACQ (STIS.ta.828 022)	(1) SN-2005IP	STIS/CCD, ACQ, F28X50LP	MIRROR			80 Secs (80 Secs) [==>]	[1]	
	<i>Comments: Exposure time set for 20121115 spectrum of 05ip scaled to V mag = 20.5. Plus an extra 20 seconds to account for potential fading.</i>									
	2	STIS/FUV E xp1 (STIS.sp.50 8430)	(1) SN-2005IP	STIS/FUV-MAMA, ACCUM, 52X0.5	G140L 1425 A			Pattern 1, Exps 2-2 i n STIS-SN2005ip (0 1) (1)	500 Secs (1812 Secs) [==>453.0 Secs (Pattern 1)] [==>453.0 Secs (Pattern 2)] [==>453.0 Secs (Pattern 3)] [==>453.0 Secs (Pattern 4)]	[1]
	3	STIS/FUV E xp2 (STIS.sp.50 8430)	(1) SN-2005IP	STIS/FUV-MAMA, ACCUM, 52X0.5	G140L 1425 A			Pattern 1, Exps 3-3 i n STIS-SN2005ip (0 1) (1)	530 Secs (2536 Secs) [==>634.0 Secs (Pattern 1)] [==>634.0 Secs (Pattern 2)] [==>634.0 Secs (Pattern 3)] [==>634.0 Secs (Pattern 4)]	[2]
	4	STIS/FUV E xp3 (STIS.sp.50 8430)	(1) SN-2005IP	STIS/FUV-MAMA, ACCUM, 52X0.5	G140L 1425 A			Pattern 1, Exps 4-4 i n STIS-SN2005ip (0 1) (1)	530 Secs (2536 Secs) [==>634.0 Secs (Pattern 1)] [==>634.0 Secs (Pattern 2)] [==>634.0 Secs (Pattern 3)] [==>634.0 Secs (Pattern 4)]	[3]
	5	STIS/NUV Exp1 (STIS.sp.51 7595)	(1) SN-2005IP	STIS/NUV-MAMA, ACCUM, 52X0.5	G230L 2376 A			Pattern 1, Exps 5-5 i n STIS-SN2005ip (0 1) (1)	650 Secs (2700 Secs) [==>675.0 Secs (Pattern 1)] [==>675.0 Secs (Pattern 2)] [==>675.0 Secs (Pattern 3)] [==>675.0 Secs (Pattern 4)]	[4]
6	STIS/NUV Exp2 (STIS.sp.51 7595)	(1) SN-2005IP	STIS/NUV-MAMA, ACCUM, 52X0.5	G230L 2376 A			Pattern 1, Exps 6-6 i n STIS-SN2005ip (0 1) (1)	650 Secs (2796 Secs) [==>699.0 Secs (Pattern 1)] [==>699.0 Secs (Pattern 2)] [==>699.0 Secs (Pattern 3)] [==>699.0 Secs (Pattern 4)]	[5]	

Orbit 1

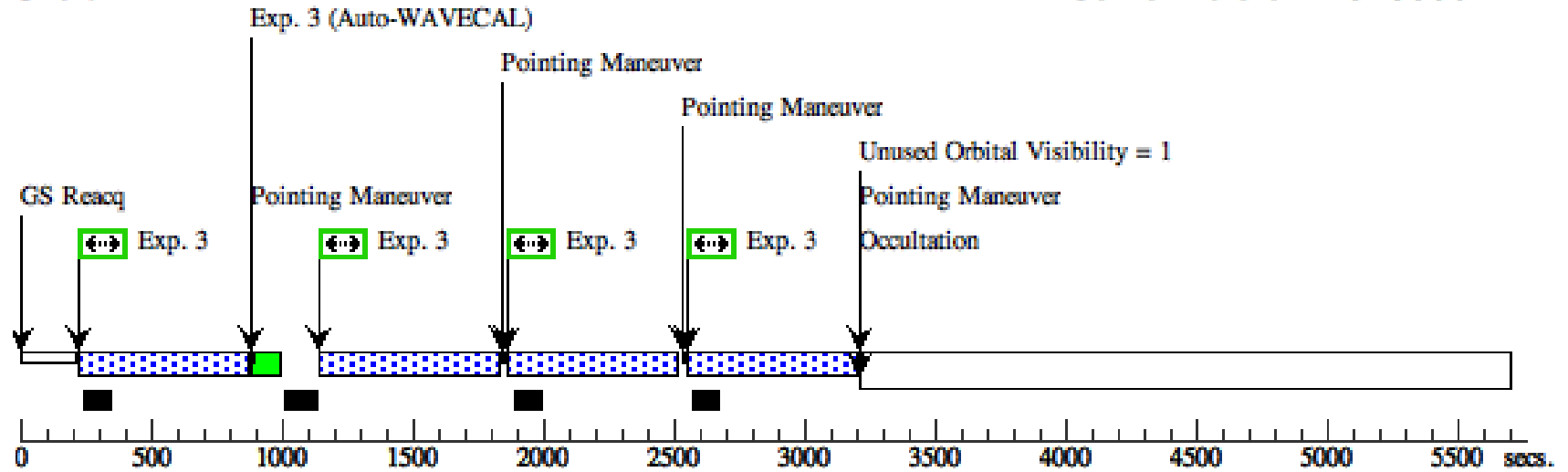
Server Version: 20160601

Orbit Structure



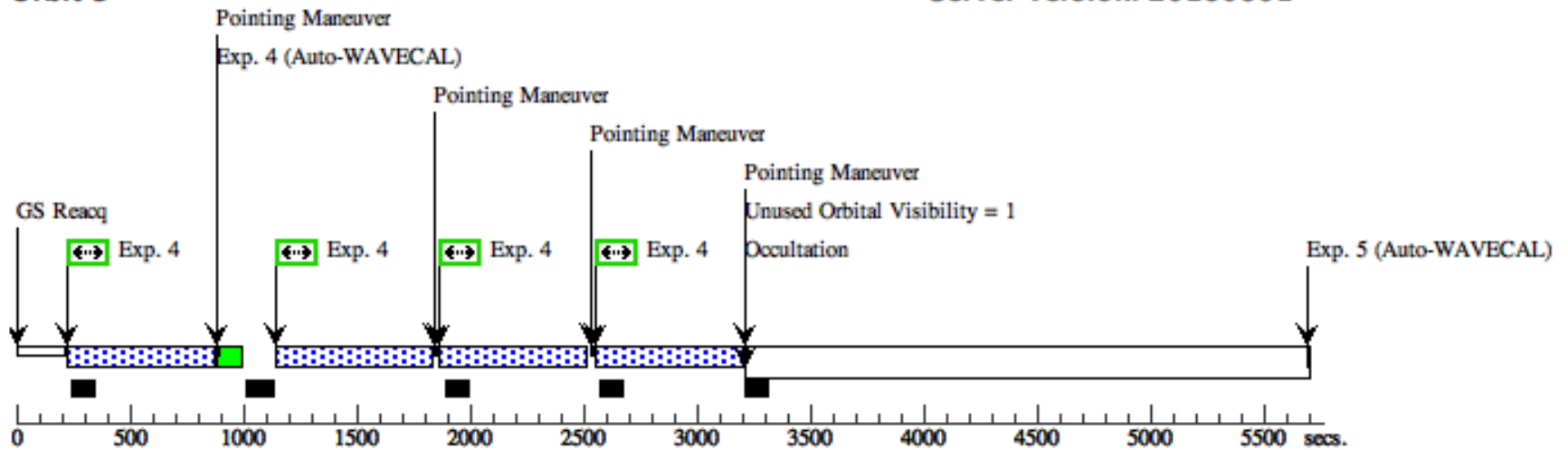
Orbit 2

Server Version: 20160601



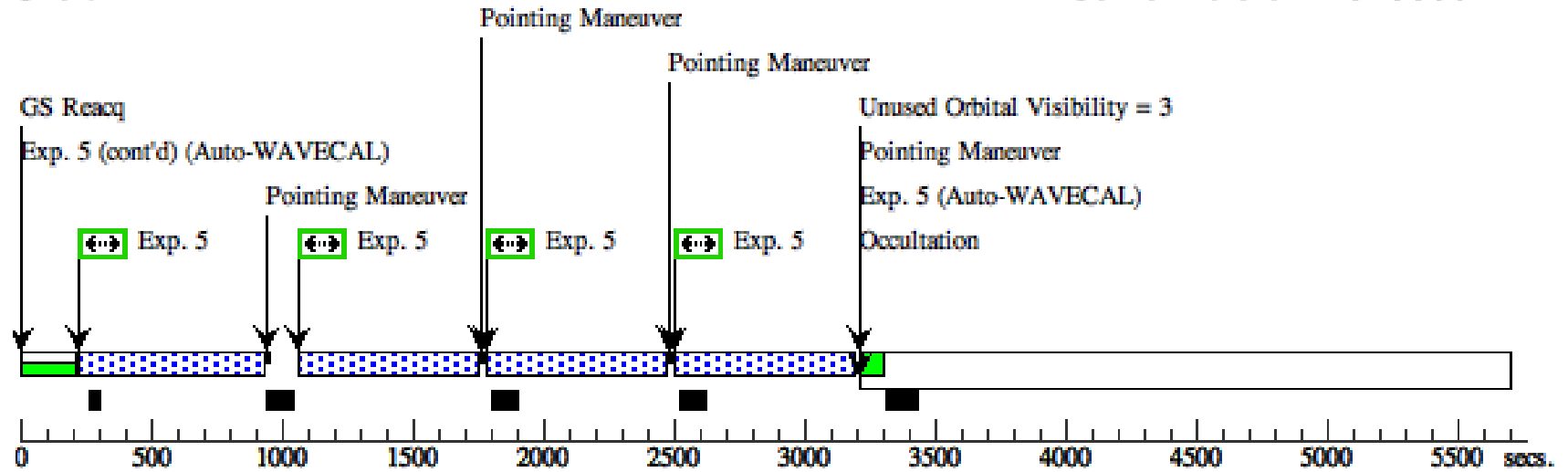
Orbit 3

Server Version: 20160601



Orbit 4

Server Version: 20160601



Orbit 5

Server Version: 20160601

