



# 14918 - Definitive measurement of WASP-17b's water abundance in preparation for JWST

Cycle: 24, Proposal Category: GO

(JWST Initiative)

(Availability Mode: SUPPORTED)

## INVESTIGATORS

<i>Name</i>	<i>Institution</i>	<i>E-Mail</i>
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Dr. Mark Clampin (CoI)	NASA Goddard Space Flight Center	mark.clampin@nasa.gov

## VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) WASP-17	WFC3/IR	5	27-Jun-2017 21:05:04.0	yes
02	(1) WASP-17	WFC3/IR	5	27-Jun-2017 21:05:21.0	yes
32	(1) WASP-17	WFC3/IR	5	27-Jun-2017 21:05:37.0	yes

15 Total Orbits Used

## ABSTRACT

In the advent of JWST, understanding of exoplanet atmospheres with current facilities will be critical. Here we propose to measure the transmission spectrum of WASP-17b in the near-infrared with both spectroscopic grisms of WFC3. Early HST and Spitzer surveys have shown WASP-17b to

have the clearest atmosphere of any measured exoplanet so far, making it an important archetype hot Jupiter. As such, WASP-17b has been selected as a Hot-Jupiter archetype for detailed study through the JWST GTO program (PI Mountain/Clampin). Higher SNR observations, made possible through spatial scanning, of this key target in the HST WFC3 wavelength range will provide critical guidance and a point for comparison for the upcoming JWST observations. The existing WFC3 data was gathered before the revolutionary spacial scanning technique, and thus the observed H<sub>2</sub>O feature suffers from extremely large errors (factor of ~5X higher) even though WASP-17b remains the very best target for transmission spectroscopy. These observations combined with previous and future observations will allow us to accurately constrain the atmospheric water amplitude and abundance to high precision.

### **OBSERVING DESCRIPTION**

Our HST observations will provide a transmission spectrum for the transiting exoplanet WASP-17b across the 0.8-1.7 micron wavelength range. WASP-17b is a highly inflated exoplanet, and is extremely amenable to atmospheric studies with HST and JWST. These observations with the G102 and G141 grisms in spatial scan mode will allow us to constrain the H<sub>2</sub>O abundance for this clear exoplanet atmosphere and better prepare us for the planned JWST GTO observations of this primed exoplanet target.

For both WFC3 observations, we will use the forward spatial scanning mode using SPARS=25 and NSAMP=7. We will read out a subarray size of 256 pixels to reduce overheads.

G102: we will use a scan rate of 0.0264 arcsec/sec, giving individual exposure times of 134 seconds, resulting in a scan across 3.5 arcsec = 27 pixels per exposure. This will leave plenty of room on the detector to estimate the background flux.

G141: we will use a scan rate of 0.033 arcsec/sec, giving individual exposure times of 134 sec, resulting in a scan across 4.422 arcsec = 35 pixels per exposure. This will leave plenty of room on the detector to estimate the background flux.

Proposal 14918 - W17 G141 (01) - Definitive measurement of WASP-17b's water abundance in preparation for JWST

Wed Jun 28 01:05:40 GMT 2017

<b>Visit</b>	<p><b>Proposal 14918, W17 G141 (01), scheduling</b></p> <p><b>Diagnostic Status: No Diagnostics</b></p> <p>Scientific Instruments: WFC3/IR</p> <p>Special Requirements: Period 3.7354822692 D AND ZERO-PHASE HJD2457192.69798</p> <p><i>Comments: WFC3 G141 IR transit of WASP-17. It is essential that the five orbits be scheduled in a continuous block, preferably free of the SAA as the planet has a long transit.</i></p> <p><i>We will use the spatial scanning mode to lengthen the spectrum along the detector during the exposures, to avoid saturation on relatively long exposures. This mode also greatly increases the efficiency.</i></p>															
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Proposal 14918 - W17 G141 (01) - Definitive measurement of WASP-17b's water abundance in preparation for JWST

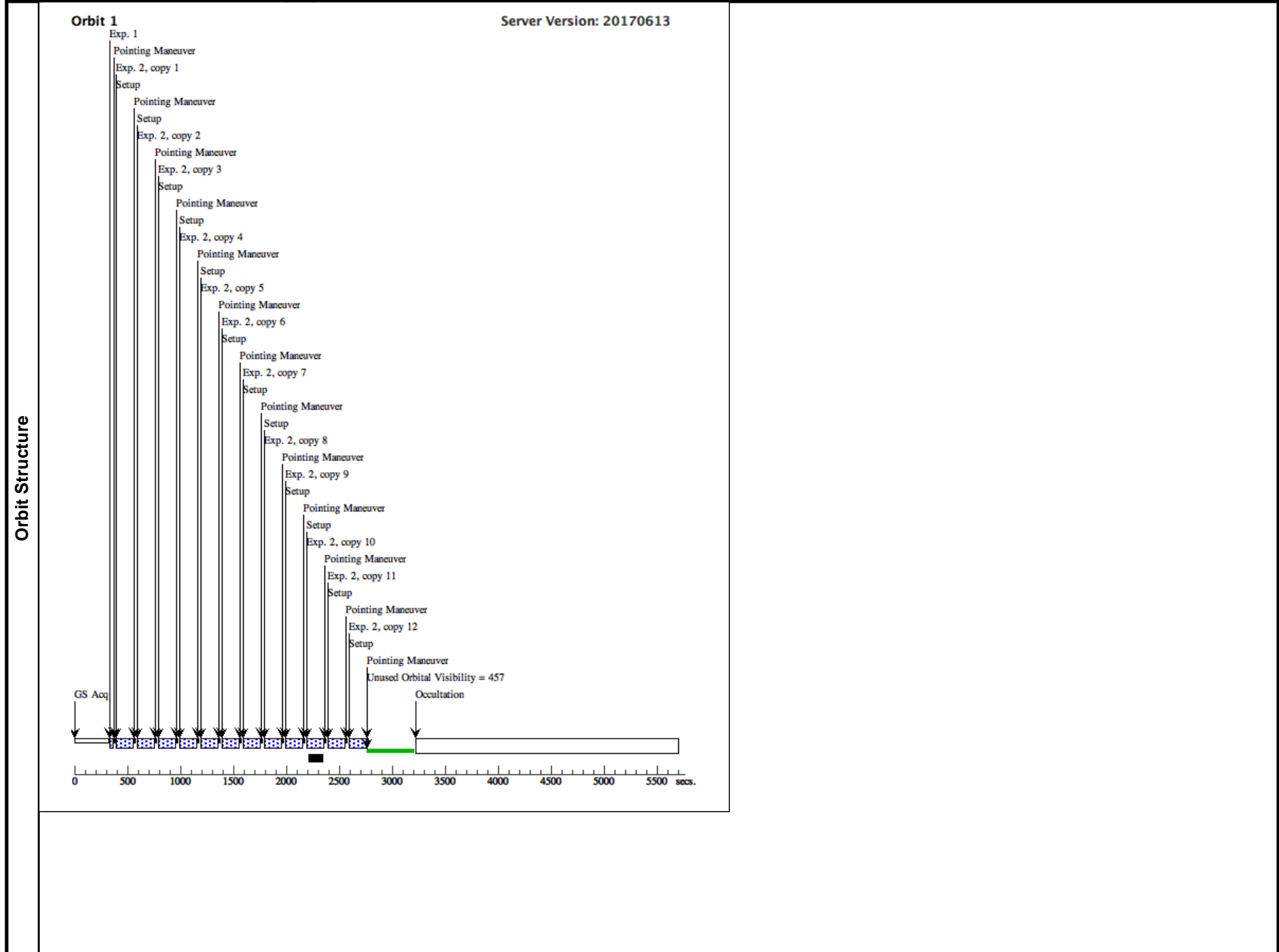
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Exposures	1	W17 F139M O1	(1) WASP-17	WFC3/IR, MULTIACCUM, GRISM256	F139M	SAMP-SEQ=RAPID ; NSAMP=7	PHASE 0.951 TO 0.955 Sequence 1-2 Non-Int in W17 G141 (01)	1.944705 Secs (1.945 Secs) [==>]	[1]
	<i>Comments: Direct image for wavelegnth calibration. Phase constrained so transit occurs between 2nd and 3rd contact.</i>								
	2	W17 G141 O1	(1) WASP-17	WFC3/IR, MULTIACCUM, GRISM256	G141	SAMP-SEQ=SPARS 25; NSAMP=7	SPATIAL SCAN 0.0 33,270.0 Degrees, Forward Sequence 1-2 Non-Int in W17 G141 (01)	134.354049 Secs X 12 (1612.249 Secs) [==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)] [==>(Copy 11)] [==>(Copy 12)]	[1]
3	W17 G141 O2	(1) WASP-17	WFC3/IR, MULTIACCUM, GRISM256	G141	SAMP-SEQ=SPARS 25; NSAMP=7	SPATIAL SCAN 0.0 33,270.0 Degrees, Forward Sequence 3-3 Non-Int in W17 G141 (01)	134.354049 Secs X 15 (2015.311 Secs) [==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)] [==>(Copy 11)] [==>(Copy 12)] [==>(Copy 13)] [==>(Copy 14)] [==>(Copy 15)]	[2]	

Proposal 14918 - W17 G141 (01) - Definitive measurement of WASP-17b's water abundance in preparation for JWST

4	W17 G141 O3	(1) WASP-17	WFC3/IR, MULTIACCUM, GRISM256	G141	SAMP-SEQ=SPARS 25; NSAMP=7	SPATIAL SCAN 0.0 33,270.0 Degrees, Forward	Sequence 4-4 Non-In t in W17 G141 (01)	134.354049 Secs X 15 (2015.311 Secs)	[==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)] [==>(Copy 11)] [==>(Copy 12)] [==>(Copy 13)] [==>(Copy 14)] [==>(Copy 15)]	[3]
5	W17 G141 O4	(1) WASP-17	WFC3/IR, MULTIACCUM, GRISM256	G141	SAMP-SEQ=SPARS 25; NSAMP=7	SPATIAL SCAN 0.0 33,270.0 Degrees, Forward	Sequence 5-5 Non-In t in W17 G141 (01)	134.354049 Secs X 15 (2015.311 Secs)	[==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)] [==>(Copy 11)] [==>(Copy 12)] [==>(Copy 13)] [==>(Copy 14)] [==>(Copy 15)]	[4]

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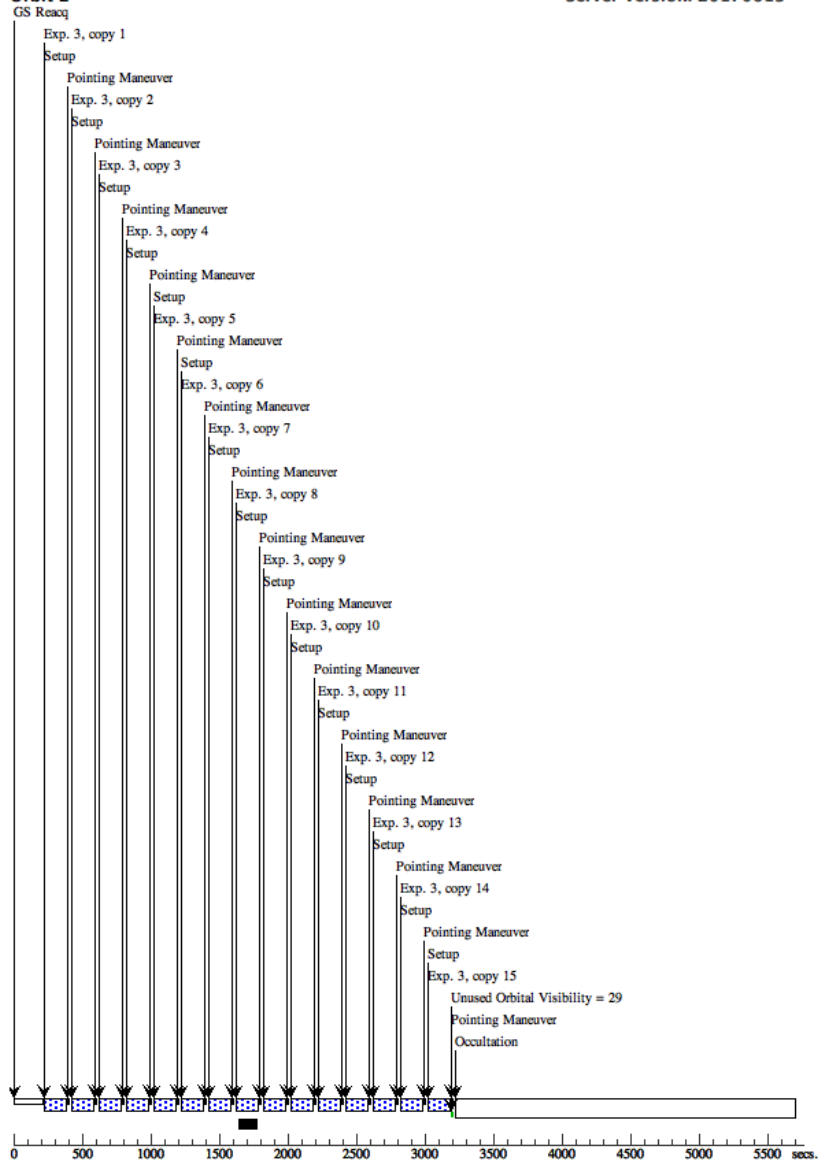
6	W17 G141 O5	(1) WASP-17	WFC3/IR, MULTIACCUM, GRISM256	G141	SAMP-SEQ=SPARS 25; NSAMP=7	SPATIAL SCAN 0.0 33,270.0 Degrees, Forward	Sequence 6-6 Non-In t in W17 G141 (01) 134.354049 Secs X 13 (1746.603 Secs)	[5]
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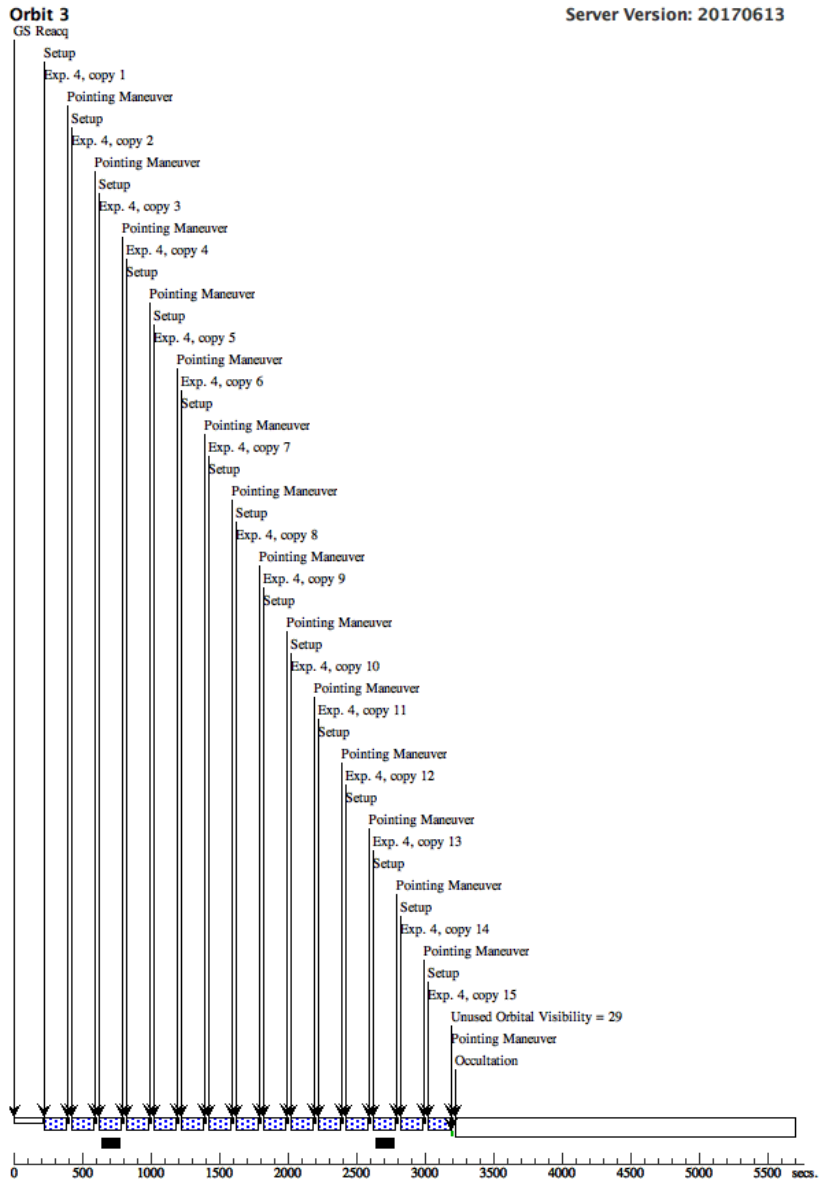
Proposal 14918 - W17 G141 (01) - Definitive measurement of WASP-17b's water abundance in preparation for JWST

Orbit 2

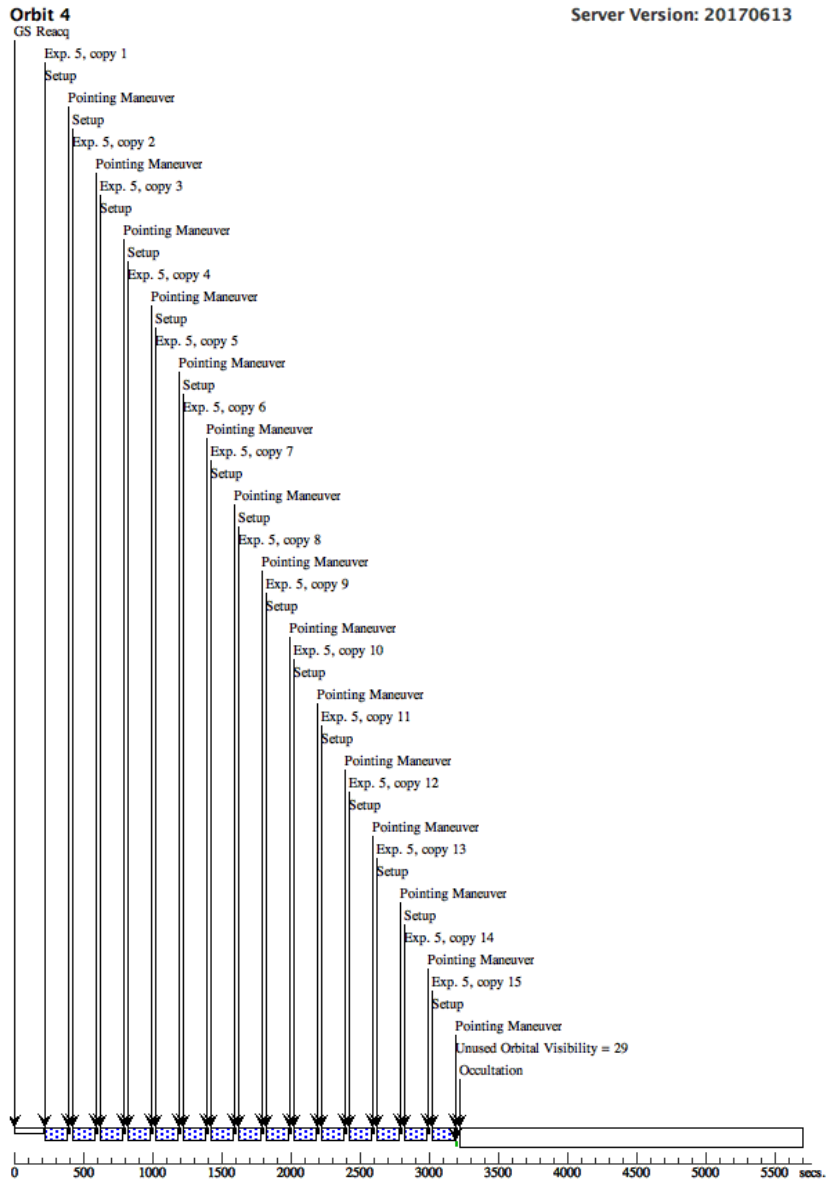
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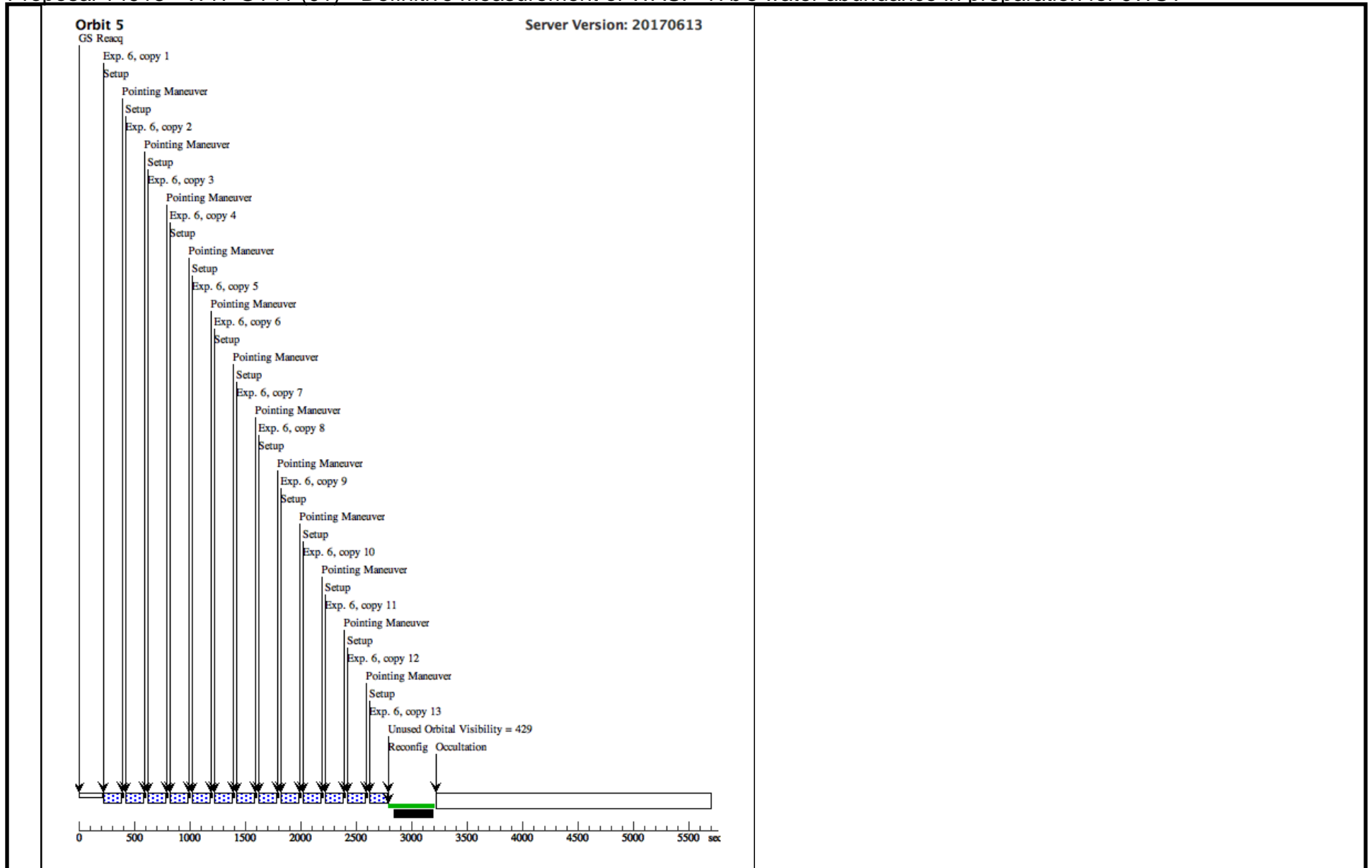


Proposal 14918 - W17 G141 (01) - Definitive measurement of WASP-17b's water abundance in preparation for JWST



Proposal 14918 - W17 G141 (01) - Definitive measurement of WASP-17b's water abundance in preparation for JWST





Proposal 14918 - W17 G102 (02) - Definitive measurement of WASP-17b's water abundance in preparation for JWST

Wed Jun 28 01:05:41 GMT 2017

<b>Visit</b>	<p><b>Proposal 14918, W17 G102 (02), failed</b></p> <p><b>Diagnostic Status: No Diagnostics</b></p> <p>Scientific Instruments: WFC3/IR</p> <p>Special Requirements: Period 3.7354822692 D AND ZERO-PHASE HJD2457192.69798</p> <p><i>Comments: WFC3 G102 IR transit of WASP-17. It is essential that the five orbits be scheduled in a continuous block, preferably free of the SAA as the planet has a long transit.</i></p> <p><i>We will use the spatial scanning mode to lengthen the spectrum along the detector during the exposures, to avoid saturation on relatively long exposures. This mode also greatly increases the efficiency.</i></p>												
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Proposal 14918 - W17 G102 (02) - Definitive measurement of WASP-17b's water abundance in preparation for JWST

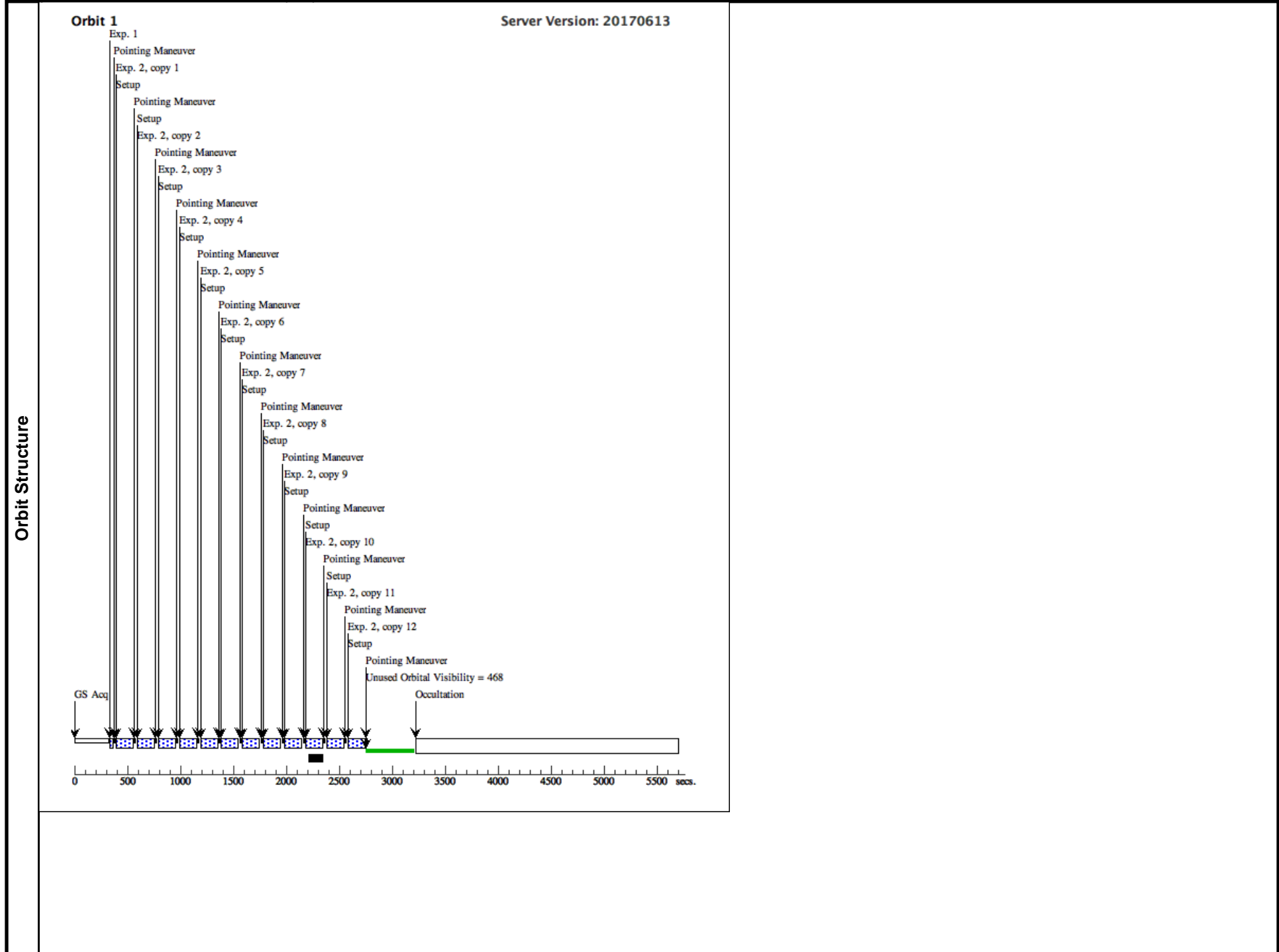
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Proposal 14918 - W17 G102 (02) - Definitive measurement of WASP-17b's water abundance in preparation for JWST

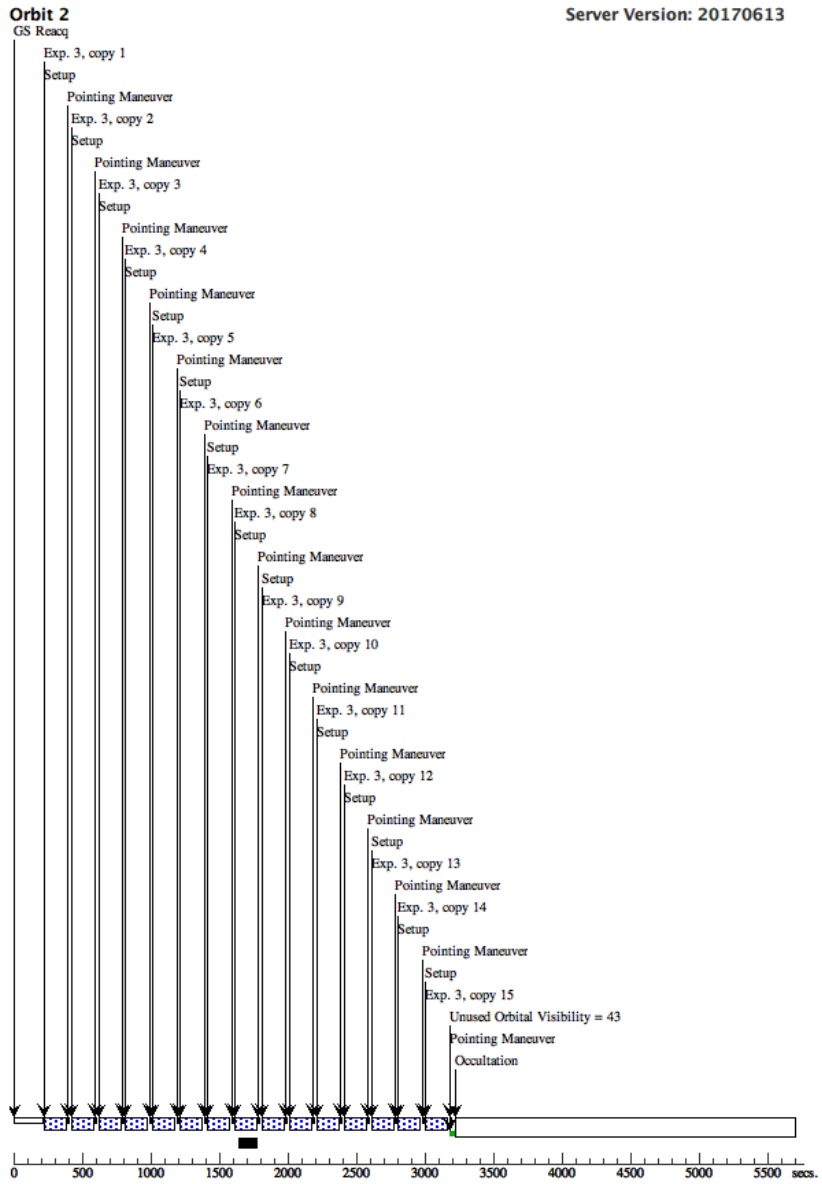
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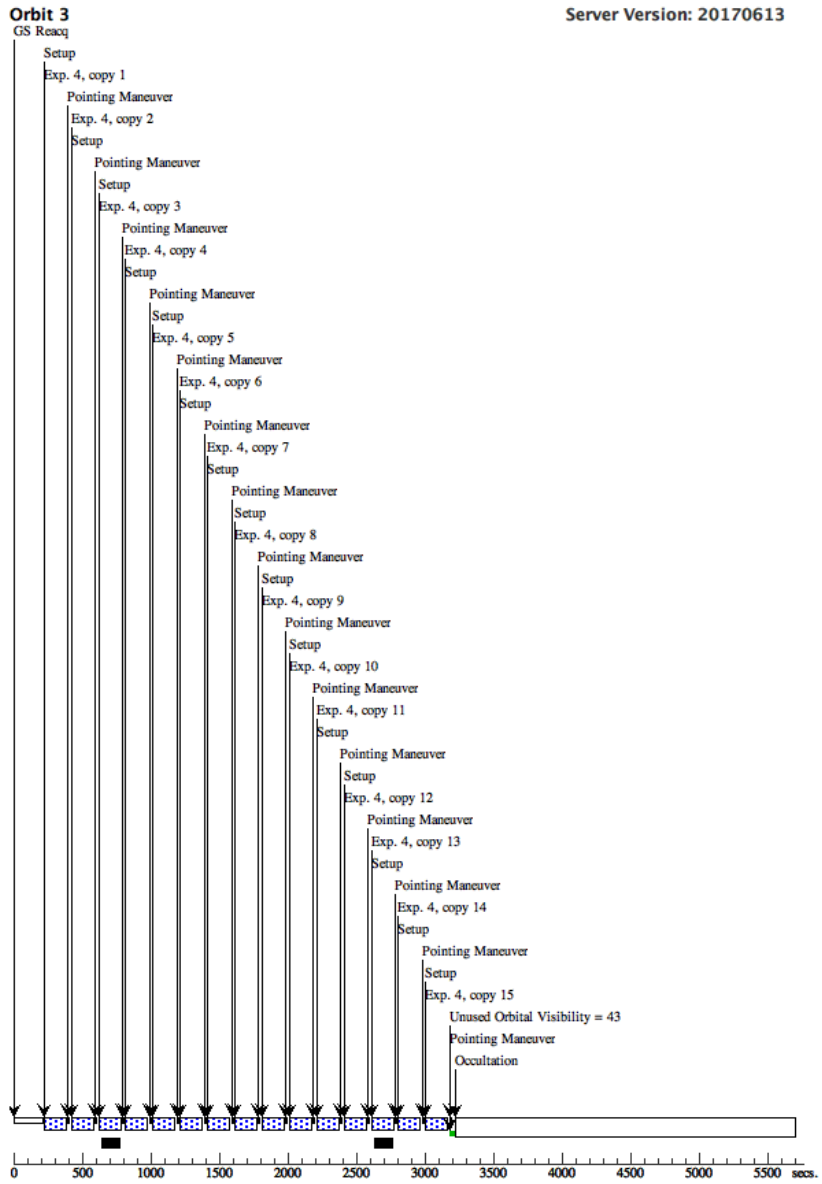
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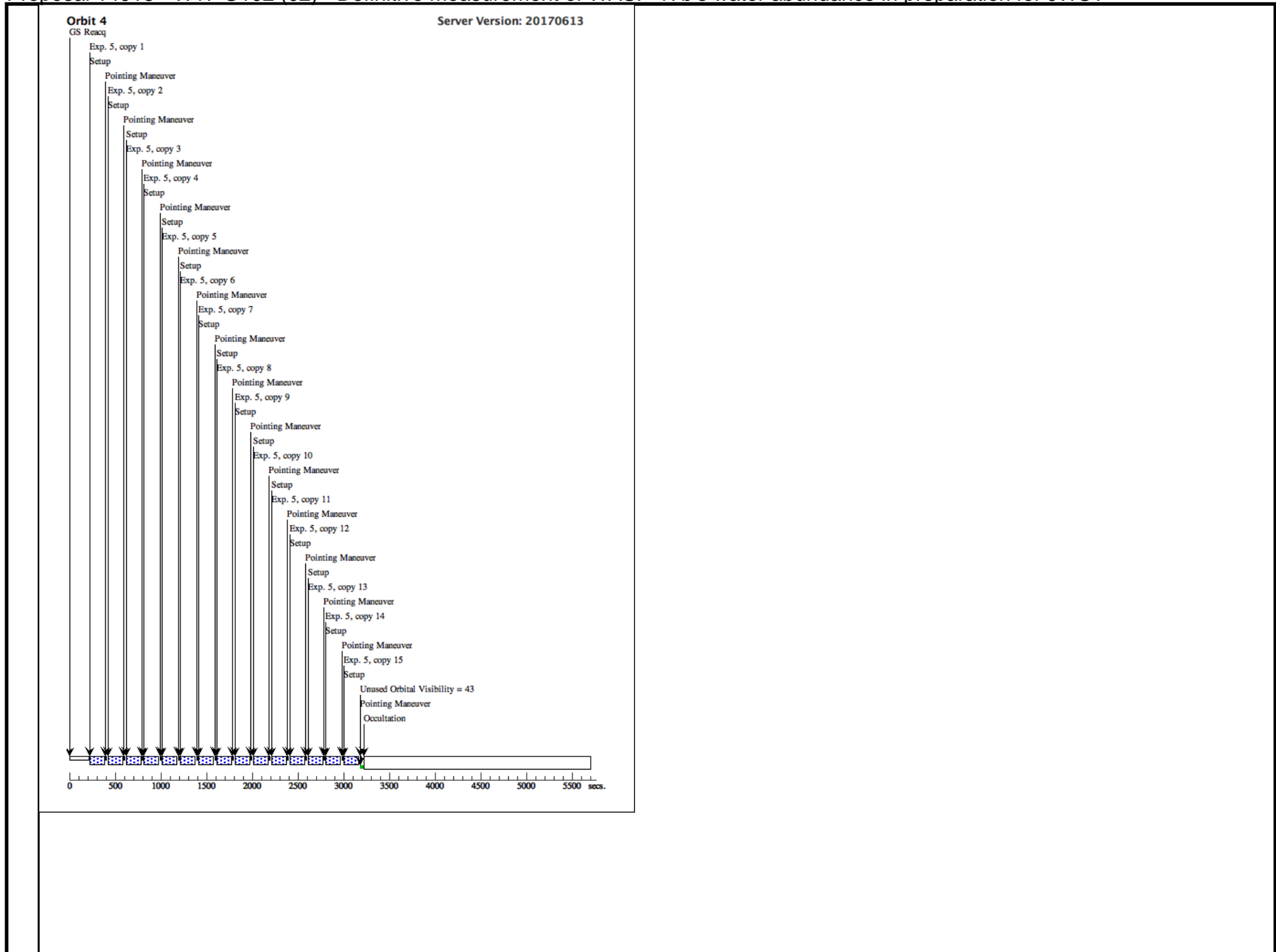
Proposal 14918 - W17 G102 (02) - Definitive measurement of WASP-17b's water abundance in preparation for JWST



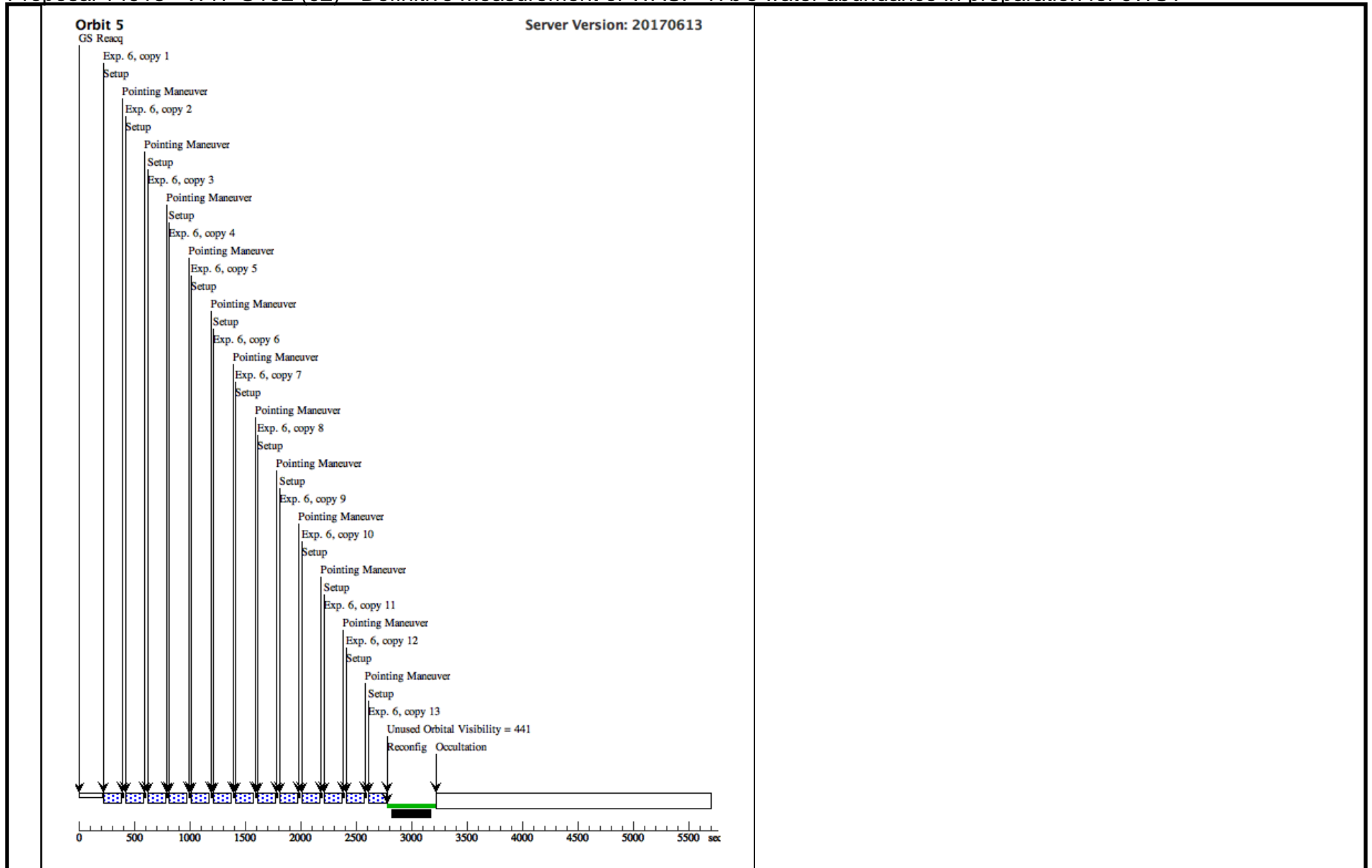
Proposal 14918 - W17 G102 (02) - Definitive measurement of WASP-17b's water abundance in preparation for JWST



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Proposal 14918 - W17 G102 (02) - Definitive measurement of WASP-17b's water abundance in preparation for JWST



Proposal 14918 - W17 G102 (32) - Definitive measurement of WASP-17b's water abundance in preparation for JWST

Wed Jun 28 01:05:42 GMT 2017

<b>Visit</b>	<p><b>Proposal 14918, W17 G102 (32)</b></p> <p><b>Diagnostic Status: No Diagnostics</b></p> <p>Scientific Instruments: WFC3/IR</p> <p>Special Requirements: Period 3.7354822692 D AND ZERO-PHASE HJD2457192.69798</p> <p><i>Comments: WFC3 G102 IR transit of WASP-17. It is essential that the five orbits be scheduled in a continuous block, preferably free of the SAA as the planet has a long transit.</i></p> <p><i>We will use the spatial scanning mode to lengthen the spectrum along the detector during the exposures, to avoid saturation on relatively long exposures. This mode also greatly increases the efficiency.</i></p>															
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2	W17 G102 O1	(1) WASP-17	WFC3/IR, MULTIACCUM, GRISM256	G102	SAMP-SEQ=SPARS25; NSAMP=7	SPATIAL SCAN 0.0 264,270.0 Degrees, Forward	Sequence 1-2 Non-Int in W17 G102 (32)	134.354049 Secs X 12 (1612.249 Secs)	
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3	W17 G102 O2	(1) WASP-17	WFC3/IR, MULTIACCUM, GRISM256	G102	SAMP-SEQ=SPARS25; NSAMP=7	SPATIAL SCAN 0.0 264,270.0 Degrees, Forward	Sequence 3-3 Non-Int in W17 G102 (32)	134.354049 Secs X 15 (2015.311 Secs)	
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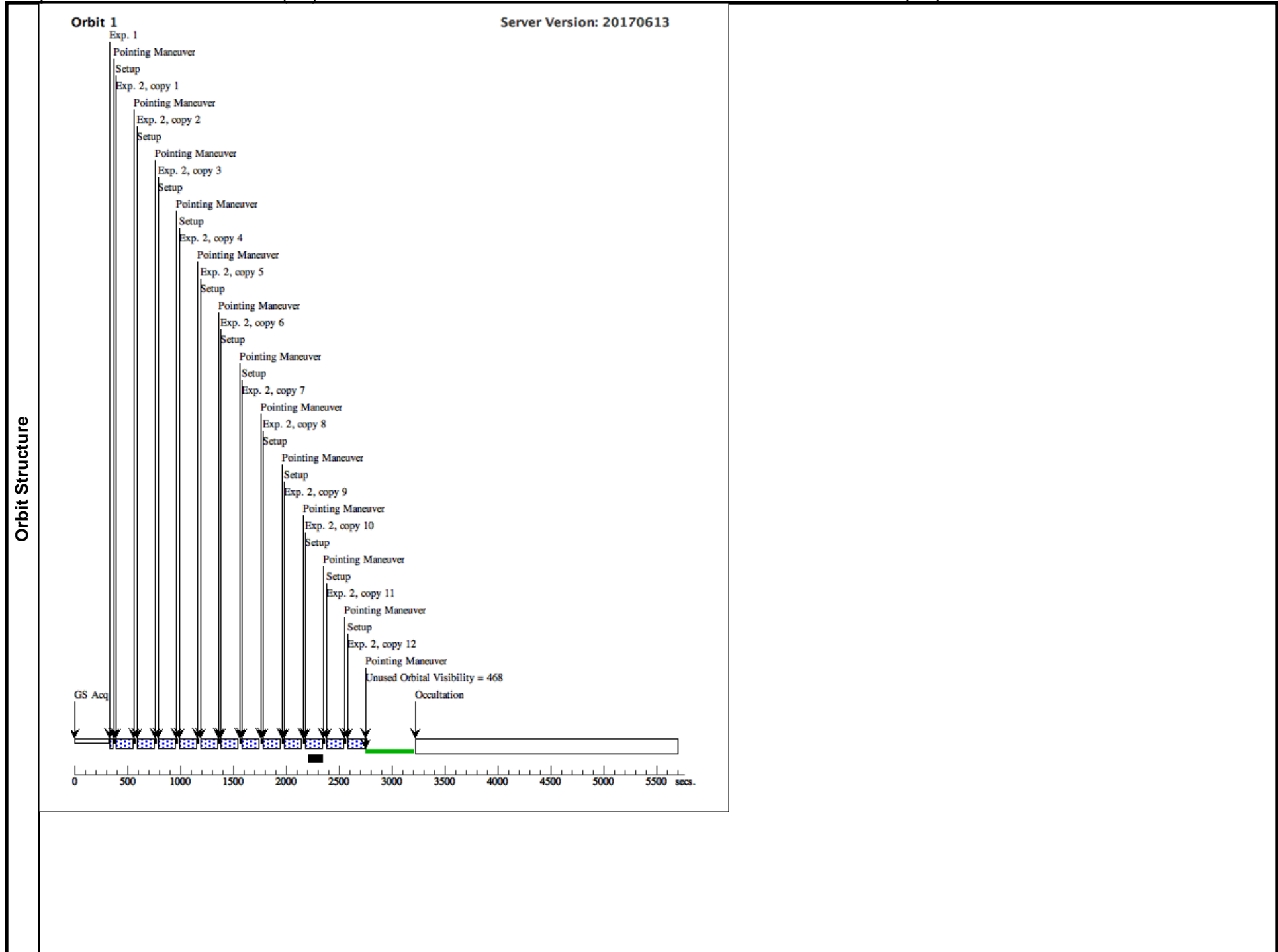
Exposures

Proposal 14918 - W17 G102 (32) - Definitive measurement of WASP-17b's water abundance in preparation for JWST

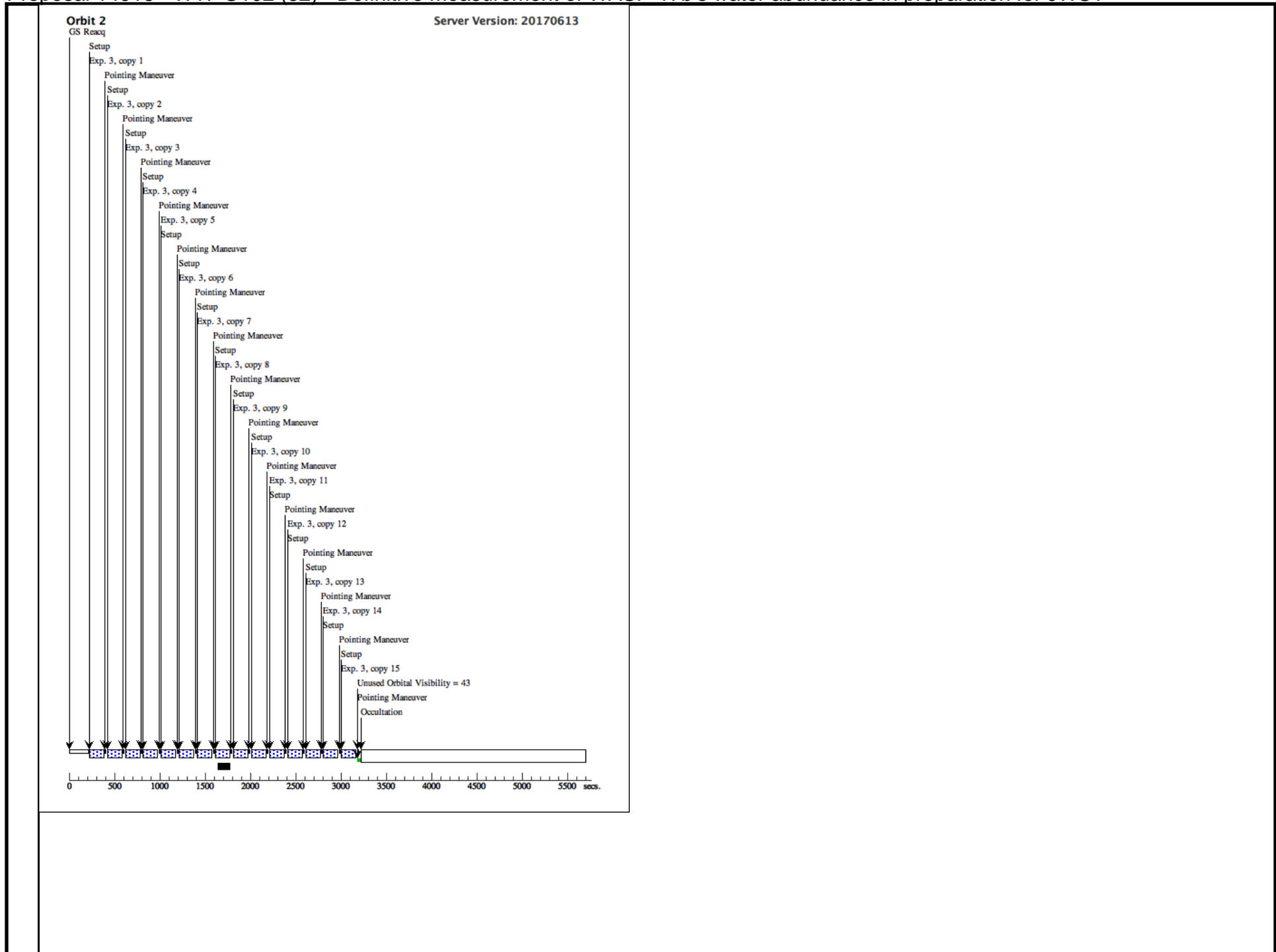
4	W17 G102 O3	(1) WASP-17	WFC3/IR, MULTIACCUM, GRISM256	G102	SAMP-SEQ=SPARS 25; NSAMP=7	SPATIAL SCAN 0.0 264,270.0 Degrees, Forward	Sequence 4-4 Non-In t in W17 G102 (32)	134.354049 Secs X 15 (2015.311 Secs)	[==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)] [==>(Copy 11)] [==>(Copy 12)] [==>(Copy 13)] [==>(Copy 14)] [==>(Copy 15)]	[3]
5	W17 G102 O4	(1) WASP-17	WFC3/IR, MULTIACCUM, GRISM256	G102	SAMP-SEQ=SPARS 25; NSAMP=7	SPATIAL SCAN 0.0 264,270.0 Degrees, Forward	Sequence 5-5 Non-In t in W17 G102 (32)	134.354049 Secs X 15 (2015.311 Secs)	[==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)] [==>(Copy 11)] [==>(Copy 12)] [==>(Copy 13)] [==>(Copy 14)] [==>(Copy 15)]	[4]

Proposal 14918 - W17 G102 (32) - Definitive measurement of WASP-17b's water abundance in preparation for JWST

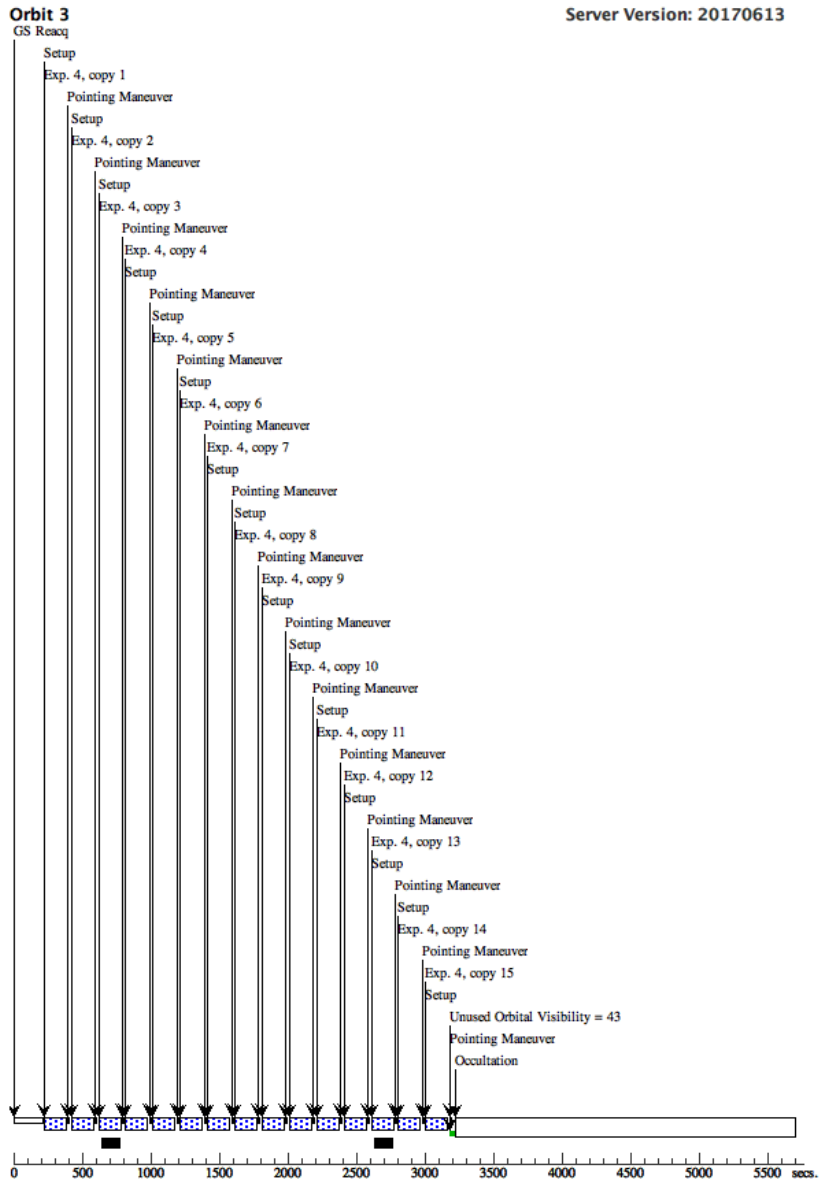
6	W17 G102 O5	(1) WASP-17	WFC3/IR, MULTIACCUM, GRISM256	G102	SAMP-SEQ=SPARS 25; NSAMP=7	SPATIAL SCAN 0.0 264,270.0 Degrees, Forward	Sequence 6-6 Non-In t in W17 G102 (32) 134.354049 Secs X 13 (1746.603 Secs)	[5]
[==>(Copy 1)] [==>(Copy 2)] [==>(Copy 3)] [==>(Copy 4)] [==>(Copy 5)] [==>(Copy 6)] [==>(Copy 7)] [==>(Copy 8)] [==>(Copy 9)] [==>(Copy 10)] [==>(Copy 11)] [==>(Copy 12)] [==>(Copy 13)]								



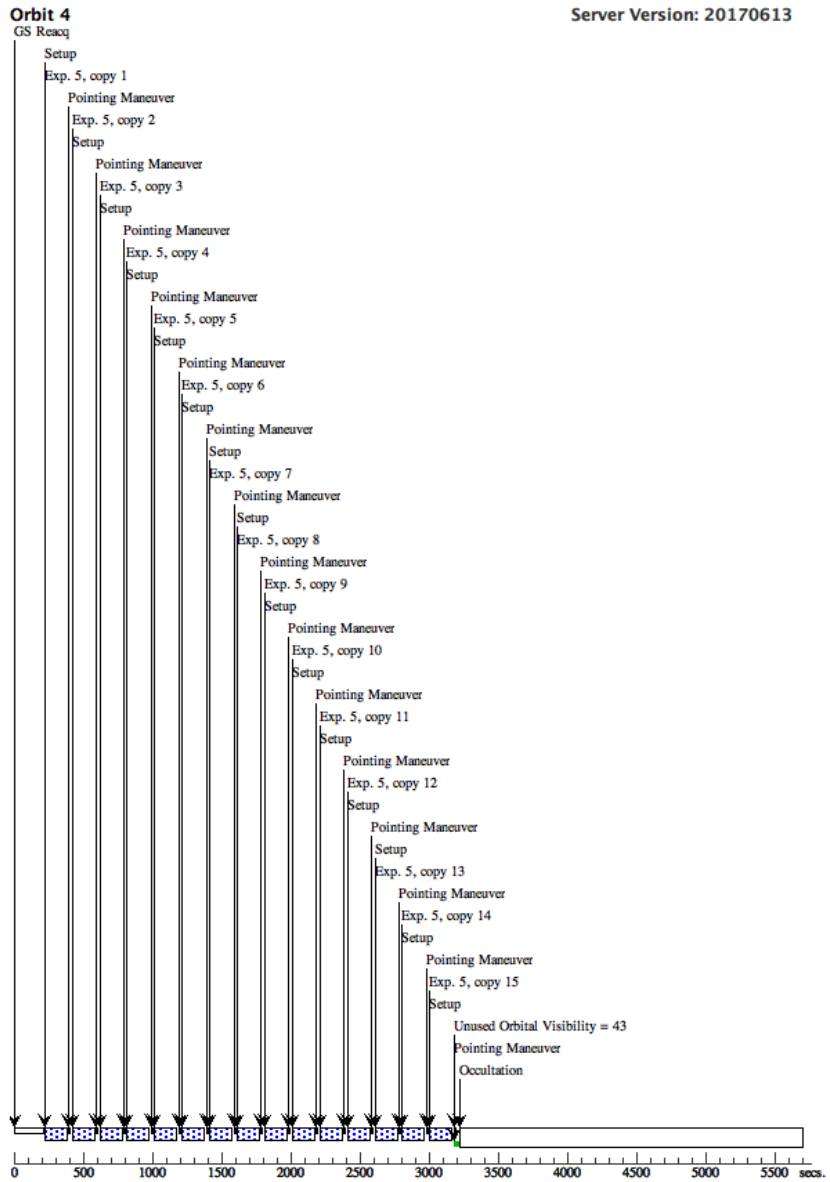
Proposal 14918 - W17 G102 (32) - Definitive measurement of WASP-17b's water abundance in preparation for JWST



Proposal 14918 - W17 G102 (32) - Definitive measurement of WASP-17b's water abundance in preparation for JWST



Proposal 14918 - W17 G102 (32) - Definitive measurement of WASP-17b's water abundance in preparation for JWST



Proposal 14918 - W17 G102 (32) - Definitive measurement of WASP-17b's water abundance in preparation for JWST

