



15412 - Observing the final part of the beta Pic Hill Sphere Transit in the far-UV

Cycle: 25, Proposal Category: GO

(UV Initiative)

(Availability Mode: AVAILABLE)

INVESTIGATORS

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VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) -BET-PIC	COS/FUV	3	04-Jan-2018 21:00:36.0	yes
02	(1) -BET-PIC	COS/FUV	3	04-Jan-2018 21:00:39.0	yes

6 Total Orbits Used

ABSTRACT

The Hill sphere of the directly-imaged planet Beta Pic b is currently transiting its host-star, Beta Pic. During this event, material in the Hill sphere (e.g. left-over debris from planet formation, rings, gas torus around satellites, etc.) could transit the star. With Cycle 24 HST/COS observations which probed the first half of the Hill sphere transit from April to August 2017, we detected a significant flux decrease in the blue and red wings of the HI Ly-alpha emission line and the 1334.53 CII absorption line. This is consistent with the first ever detection of an extended envelope of hydrogen and carbon surrounding Beta Pic b. With the upcoming Director's Discretionary Program #15396 we will observe the second half of the Hill sphere transit.

We propose to monitor the final part of the Hill sphere transit (the expected trailing Hill sphere gas) and post-transit baseline (for Hill sphere transit verification), using the same instrument setup and observation strategy. As with the previous 2017 observations, we will again use the "Airglow Virtual Motion" technique, which is an extremely powerful technique to observe Lyman-alpha with COS. This is a very unique opportunity to monitor the transit of the environment of a ~20 million years old young planet that has been directly imaged orbiting a 4th magnitude star; the next such transit will not happen again for another 22 years. The aims for this proposal are to 1) obtain a post-transit baseline for transit verification 2) probe the shape and the dynamics of the gas trailing the Hill sphere, and 3) probe the composition of the gas trailing the Hill sphere.

OBSERVING DESCRIPTION

We have a total of two visits scheduled. Each visit consists of three orbits. Each visit has an identical observing strategy.

For each visit:

Orbit 1: Observe target centered in the aperture

Orbit 2: The target (Beta Pictoris) is to be offset -1.1 arcseconds in the dispersion direction

Orbit 3: The target (Beta Pictoris) is to be offset +1.1 arcseconds in the dispersion direction

The 1.1 arcsecond offsets shifts are important for the detection of hydrogen, nitrogen and oxygen which are all affected by airglow contamination. An increase in the off-axis target distance along the dispersion axis effectively shifts the line emission feature in wavelength relative to the airglow emission.

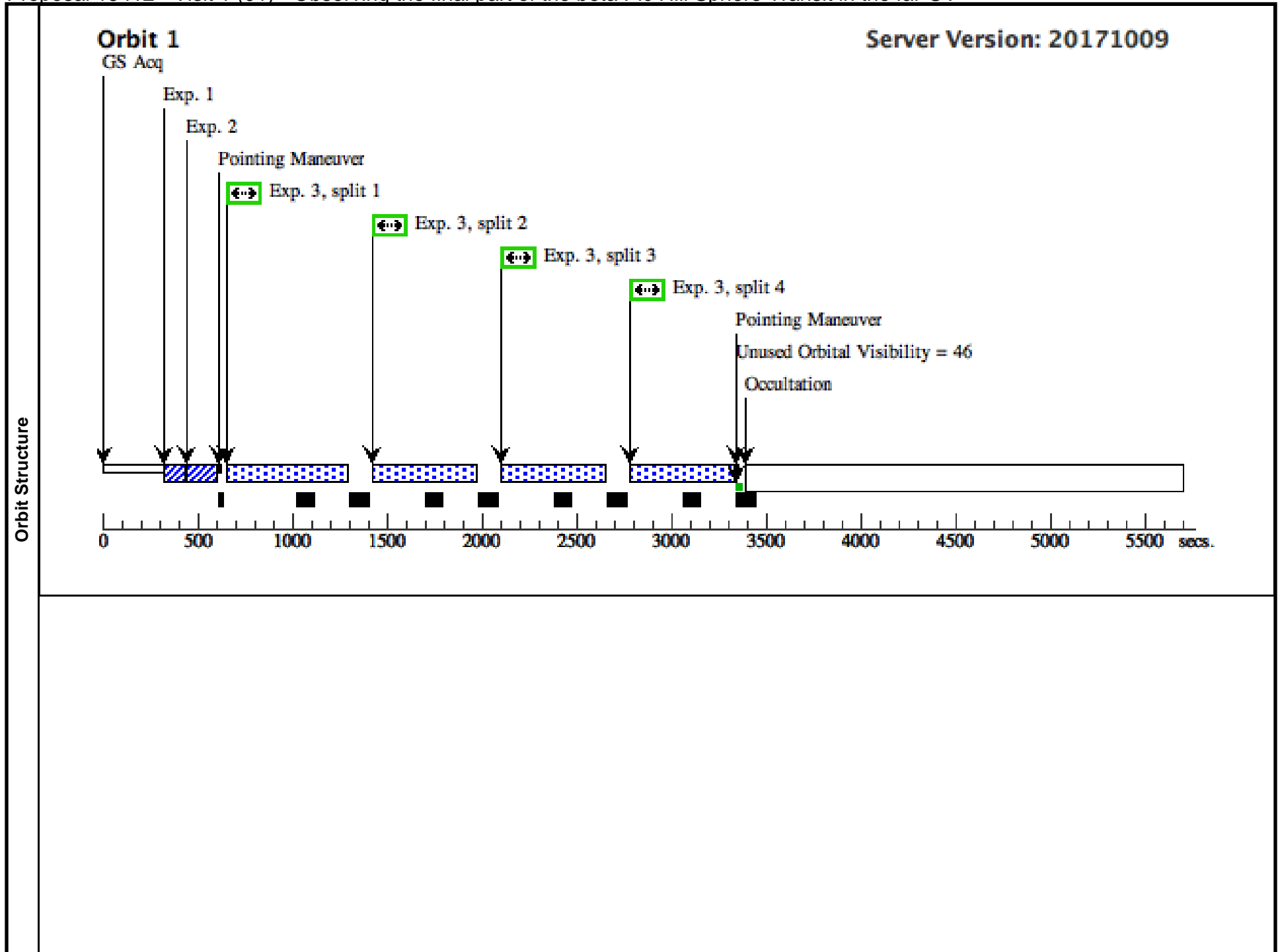
Proposal 15412 - Visit 1 (01) - Observing the final part of the beta Pic Hill Sphere Transit in the far-UV

Fri Jan 05 02:00:40 GMT 2018

Visit	<p>Proposal 15412, Visit 1 (01), implementation</p> <p>Diagnostic Status: Warning</p> <p>Scientific Instruments: COS/FUV</p> <p>Special Requirements: BETWEEN 04-MAR-2018:00:00:00 AND 01-APR-2018:23:59:00</p> <p><i>Comments: Orbit 1: Observe target centered in the aperture</i> <i>Orbit 2: The target (Beta Pictoris) is to be offset -1.1 arcseconds in the dispersion direction</i> <i>Orbit 3: The target (Beta Pictoris) is to be offset +1.1 arcseconds in the dispersion direction</i></p>																	
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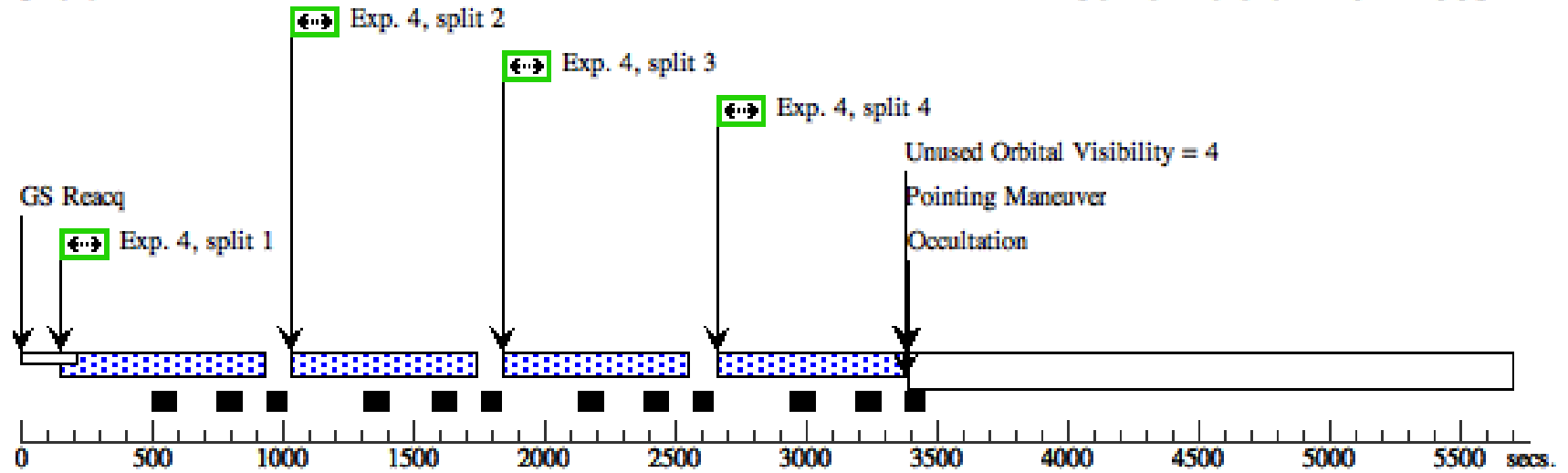
Proposal 15412 - Visit 1 (01) - Observing the final part of the beta Pic Hill Sphere Transit in the far-UV

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
Exposures	1	Positioning (COS.sa.729 672)	(1) -BET-PIC	COS/FUV, ACQ/PEAKXD, PSA	G130M 1291 A	LIFETIME-POS=LP 4; SEGMENT=BOTH	Sequence 1-3 Non-Int in Visit 1 (01)	1.4 Secs (1.4 Secs) [==>]	[1]
	2	Max-Light (COS.sa.729 672)	(1) -BET-PIC	COS/FUV, ACQ/PEAKD, PSA	G130M 1291 A	CENTER=DEF; NUM-POS=5; STEP-SIZE=0.9; LIFETIME-POS=L P4	Sequence 1-3 Non-Int in Visit 1 (01)	1.4 Secs (1.4 Secs) [==>]	[1]
	3	Center (COS.sp.848 182)	(1) -BET-PIC	COS/FUV, TIME-TAG, PSA	G130M 1327 A	BUFFER-TIME=25 0; FLASH=YES; FP-POS=ALL; LIFETIME-POS=L P3	Sequence 1-3 Non-Int in Visit 1 (01)	534 Secs (1996 Secs) [==>499.0 Secs (Split 1)] [==>499.0 Secs (Split 2)] [==>499.0 Secs (Split 3)] [==>499.0 Secs (Split 4)]	[1]
	<i>Comments: The strength of the Stellar Lyman-alpha line has been set to an conservative upper limit to calculate the Buffer-time and check the Bright Objects Limit (FWHM=5 Angstrom, Flux=1e-11).</i>								
	4	Blueward-1. 1arcsec (COS.sp.848 182)	(1) -BET-PIC	COS/FUV, TIME-TAG, PSA	G130M 1327 A	BUFFER-TIME=25 0; FLASH=YES; FP-POS=ALL; LIFETIME-POS=L P3	POS TARG -1.1,null Sequence 4-4 Non-Int in Visit 1 (01)	534 Secs (2644 Secs) [==>661.0 Secs (Split 1)] [==>661.0 Secs (Split 2)] [==>661.0 Secs (Split 3)] [==>661.0 Secs (Split 4)]	[2]
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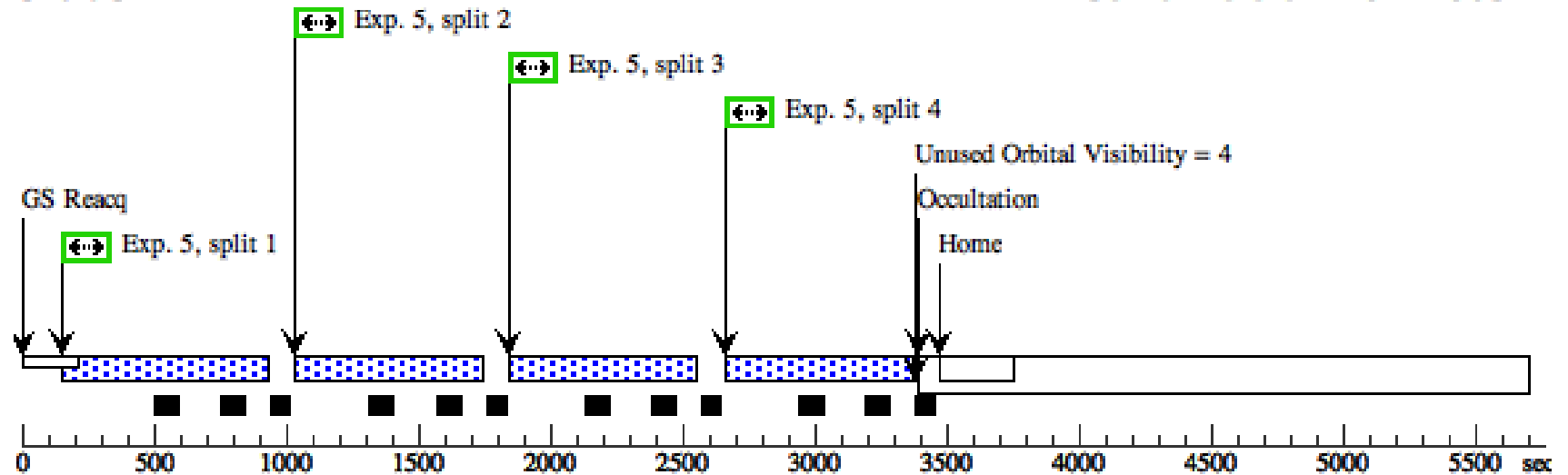
Orbit 2

Server Version: 20171009



Orbit 3

Server Version: 20171009



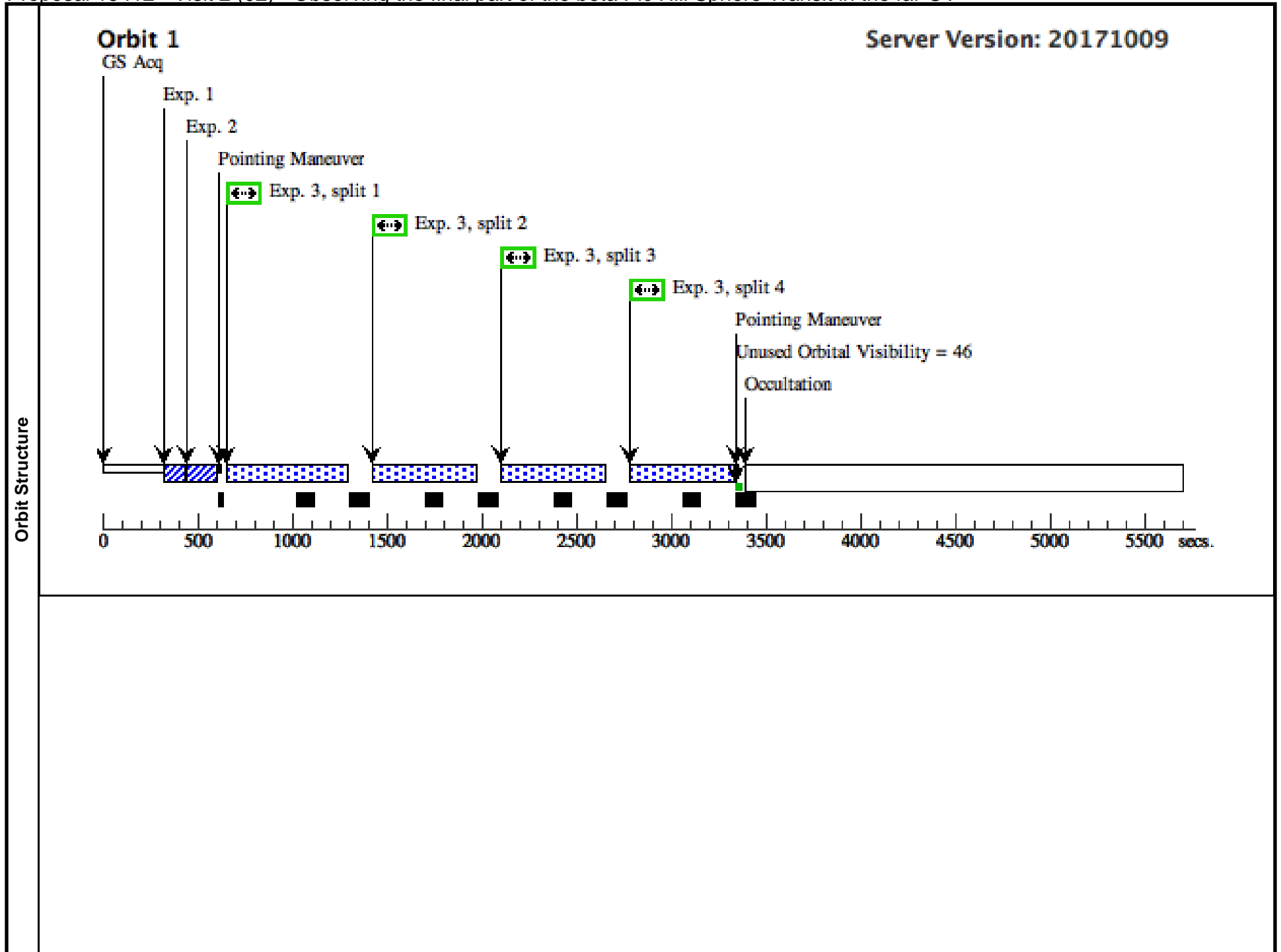
Proposal 15412 - Visit 2 (02) - Observing the final part of the beta Pic Hill Sphere Transit in the far-UV

Fri Jan 05 02:00:41 GMT 2018

Visit	<p>Proposal 15412, Visit 2 (02), implementation</p> <p>Diagnostic Status: Warning</p> <p>Scientific Instruments: COS/FUV</p> <p>Special Requirements: BETWEEN 05-MAY-2018:00:00:00 AND 15-JUL-2018:23:59:00</p> <p><i>Comments: Orbit 1: Observe target centered in the aperture</i> <i>Orbit 2: The target (Beta Pictoris) is to be offset -1.1 arcseconds in the dispersion direction</i> <i>Orbit 3: The target (Beta Pictoris) is to be offset +1.1 arcseconds in the dispersion direction</i></p>																	
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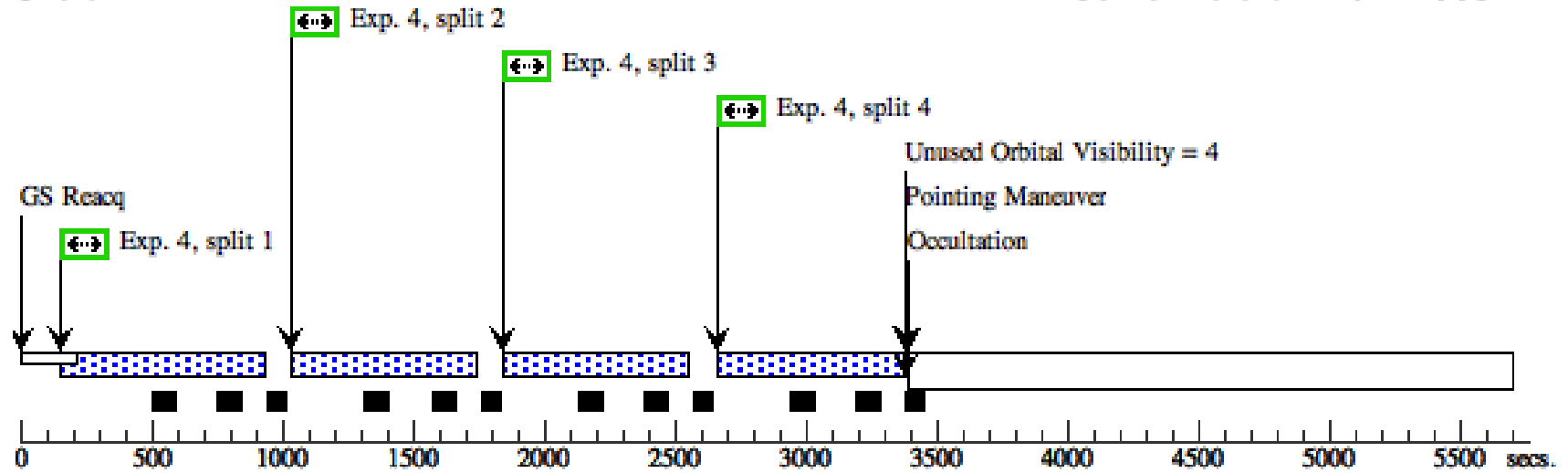
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Orbit 2

Server Version: 20171009



Orbit 3

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