



# 15448 - Constraining the emergent EUV ionizing emission in the reawakening monster in Mrk 590

Cycle: 25, Proposal Category: GO/DD

(Availability Mode: SUPPORTED)

## INVESTIGATORS

<i>Name</i>	<i>Institution</i>	<i>E-Mail</i>
<b>Dr. Marianne Vestergaard (PI) (ESA Member) (Contact)</b>	<b>University of Copenhagen, Niels Bohr Institute</b>	<b>vester@dark-cosmology.dk</b>
Dr. Sandra Raimundo (CoI) (ESA Member)	University of Copenhagen, Niels Bohr Institute	s.raimundo@dark-cosmology.dk
Prof. Bradley M Peterson (CoI) (AdminUSPI)	The Ohio State University	peterson.12@osu.edu
Dr. Gisella De Rosa (CoI)	Space Telescope Science Institute	gderosa@stsci.edu
Daniel P. Lawther (CoI) (ESA Member)	University of Copenhagen, Niels Bohr Institute	danielpe@dark-cosmology.dk
Dr. Jun Yi (Kevin) Koay (CoI)	Academia Sinica, Institute of Astronomy and Astrophysics	jykoay@asiaa.sinica.edu.tw

## VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) NGC-863	COS/FUV COS/NUV	1	22-Jan-2018 20:08:35.0	yes
02	(1) NGC-863	COS/FUV COS/NUV	1	22-Jan-2018 20:08:36.0	yes
S3	(1) NGC-863	COS/FUV COS/NUV	1	22-Jan-2018 20:08:37.0	yes

3 Total Orbits Used

## ABSTRACT

After a 10-year hiatus, Mrk 590 is rekindling its nuclear activity! This offers a very rare opportunity to document the onset of AGN activity as it occurs that can lead to a better understanding of the long-standing issue of how AGNs are fuelled and how AGN work. We have monitored Mrk 590 ~bi-daily with Swift since late August 2017 in response to sudden and strong X-ray flares that continue. Early December we triggered our HST/COS ToO program of 2 single-orbit visits and additional ToO programs with VLT/Xshooter and Gemini that run for a month and spectroscopically cover from 3200Å to 2.2micron. We now observe with Swift daily for at least another month. Our recent VLT/MUSE, X-shooter and HST/COS data show a strengthening of the UV-optical-IR broad lines. In particular, we see appearance of HeI IR lines that indicate a strengthening of the EUV continuum; the continuum is still low in the UV-optical and non-existing in the near-IR. We kindly ask for two additional single-orbit visits in order to follow the strengthening of the broad component of the UV Helium, Lyman-alpha and Carbon lines in order to better constrain the build-up of the ionizing EUV continuum source. The HST/COS data are pivotal, since we cannot obtain these constraints on the far-UV and EUV emission any other way.

An extensive and ancillary database is being acquired on Mrk 590 from radio through X-ray wavelengths - also during this current brief monitoring campaign. These new and existing HST/COS data will contribute with unique information to the legacy database on this intriguing object, Mrk 590, that will undoubtedly spawn new insight on AGN and their fuelling.

## OBSERVING DESCRIPTION

This is an extension of proposal 15404. The source has already been successfully observed.

Same observing strategy applies.

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We propose to obtain COS low resolution FUV observations of the Seyfert 1 galaxy Mrk 590 to determine if the AGN thermal emission and broad emission lines have re-appeared in response to the significant increase in X-ray and far-UV emission as detected by Swift when this program is triggered. Using the G140L grating @1105 Å we will be able to cover the Ly $\alpha$ 1215, NV1240, Si IV + O IV1400, C IV1549, HeII1640 and CIII]1909 broad emission lines in a single exposure. The grating resolution is adequate to resolve the typical AGN narrow emission components; the width of the [O III]5007 narrow line is FWHM~400 km s<sup>-1</sup>.

As explained in the Phase I, the target has been closely monitored with SWIFT, and in the past 80 days UVW2 fluxes (centroid wavelength 1928)

Proposal 15448 (STScI Edit Number: 1, Created: Monday, January 22, 2018 8:08:38 PM EST) - Overview

have varied from a minimum of  $\sim 2.2 \times 10^{-15}$  erg s<sup>-1</sup> cm<sup>-2</sup> Ang<sup>-1</sup> to a max of  $\sim 4.0$  erg s<sup>-1</sup> cm<sup>-2</sup> Ang<sup>-1</sup>.

Note: the target has already been successfully observed in the same configurations during faint state (PID 13518)

ACQUISITION: The target has precise 2MASS coordinates (ICRS reproduced with an error  $< 0.1''$ ). We plan to use ACQ/IMG NUV MIRRORB and an exp time of 140s. For our ETC calculations we assume a flat continuum in fl normalized to half of the min observed UVW2 flux  $f_1 \sim 1.1 \times 10^{-15}$  erg s<sup>-1</sup> cm<sup>-2</sup> Ang<sup>-1</sup> @1928 Ang (<http://etc.stsci.edu/etc/results/COS.ta.1030198/>). This will ensure acquisition in case of "faint" state. The source will not pose a threat to the detector even if it ends up being 5 times more luminous than the max detected in the past 80 days (<http://etc.stsci.edu/etc/results/COS.ta.1030199/>).

SCIENCE: We use the G140L grating at 1105Ang with FP\_POS=ALL. For ETC calculations we assume a flat continuum normalized to the mean observed UVW2 flux in the past 80 days (<http://etc.stsci.edu/etc/results/COS.sp.1030204/>). Buffer time is set to 2/3 of the ETC buffer time. The source will not pose a threat to the detector even if it ends up being 5 times more luminous than the past 80 days (<http://etc.stsci.edu/etc/results/COS.sp.1030205/>).

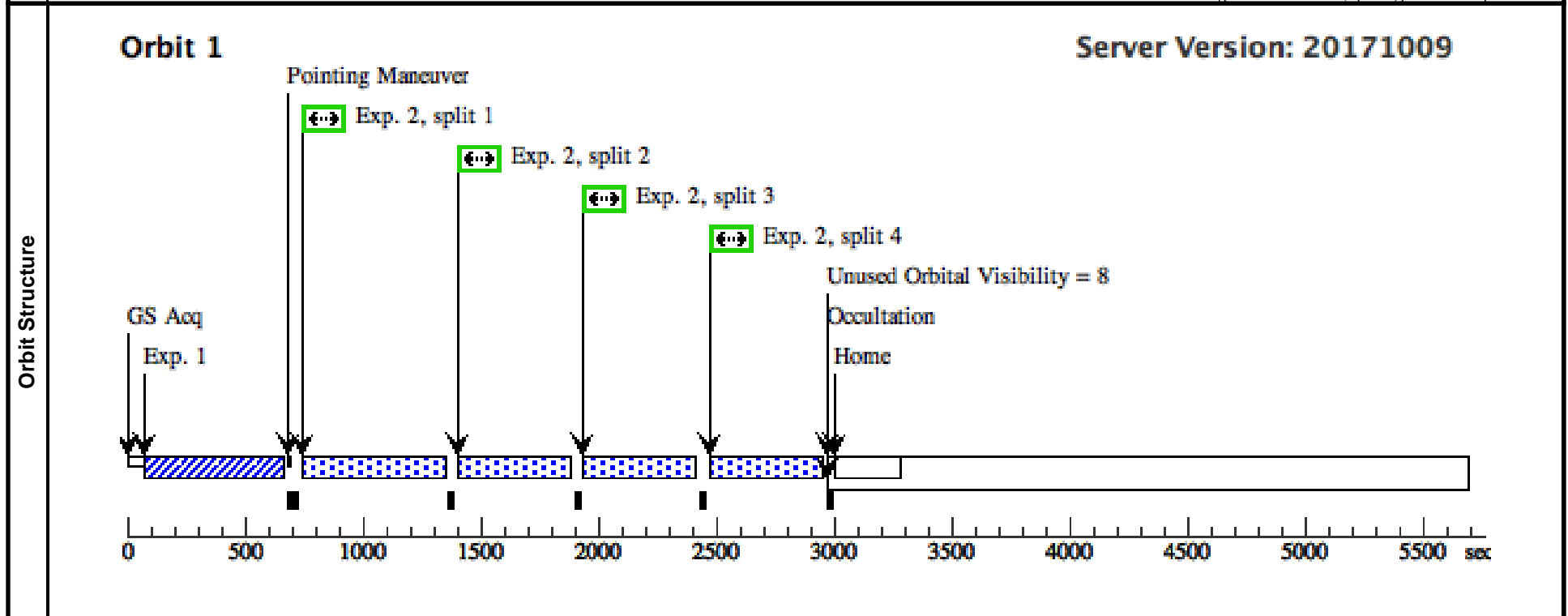
Proposal 15448 - Visit 01 - Constraining the emergent EUV ionizing emission in the reawakening monster in Mrk 590

Tue Jan 23 01:08:38 GMT 2018

<b>Visit</b>	Proposal 15448, Visit 01, completed				
	Diagnostic Status: No Diagnostics				
	Scientific Instruments: COS/FUV, COS/NUV				
	Special Requirements: SCHED 100%				

<b>Fixed Targets</b>	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
	(1)	NGC-863	RA: 02 14 33.5790 (33.6399125d) Dec: -00 46 0.28 (-.76674d) Equinox: J2000		V=13.81	Reference Frame: ICRS
	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i>					
	Category=GALAXY Description=[ACCRETION DISK, BLR, SEYFERT] Extended=NO					

<b>Exposures</b>	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	NUV MIRR OR B ACQ (COS.ta.103 0198)	(1) NGC-863	COS/NUV, ACQ/IMAGE, PSA	MIRRORB				140 Secs (140 Secs) [==>]	[1]
	2	FUV G140L /1105 Scienc e (COS.sp.103 0204)	(1) NGC-863	COS/FUV, TIME-TAG, PSA	G140L 1105 A	FP-POS=ALL; BUFFER-TIME=50 00			474 Secs (1720 Secs) [==>430.0 Secs (Split 1)] [==>430.0 Secs (Split 2)] [==>430.0 Secs (Split 3)] [==>430.0 Secs (Split 4)]	[1]



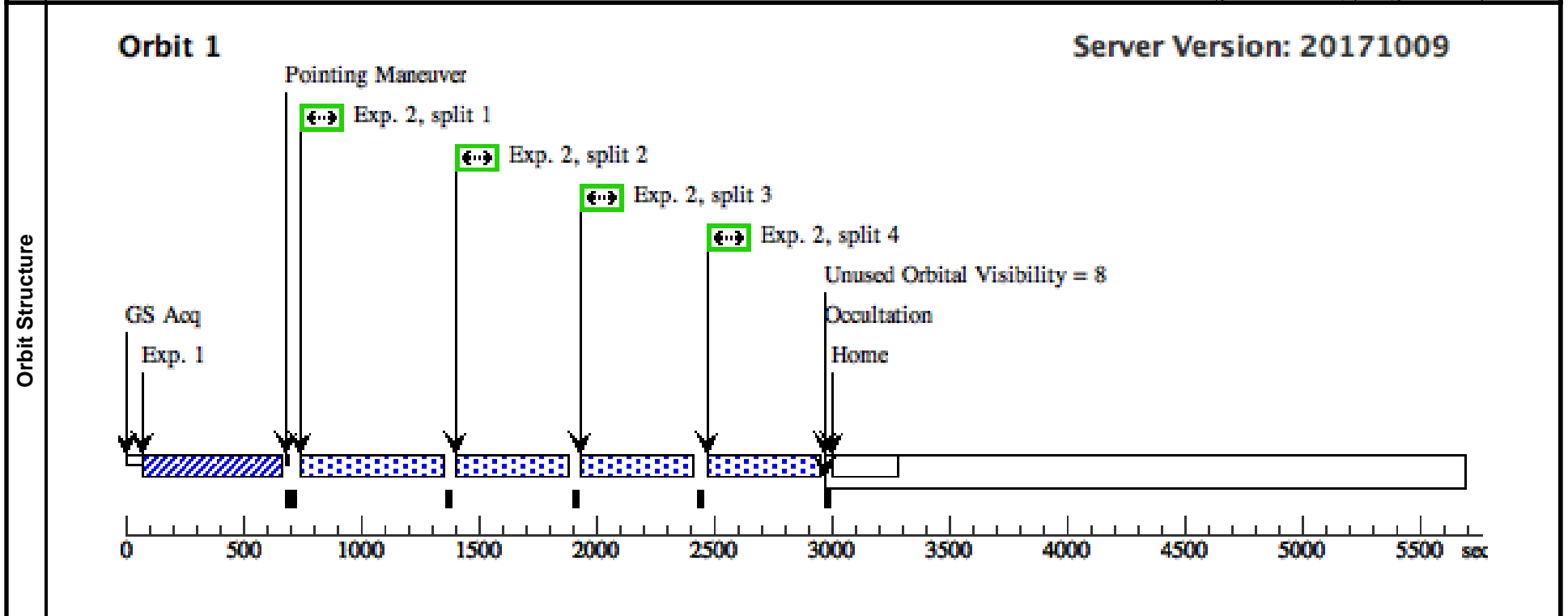
Proposal 15448 - Visit 02 - Constraining the emergent EUV ionizing emission in the reawakening monster in Mrk 590

Tue Jan 23 01:08:38 GMT 2018

<b>Visit</b>	Proposal 15448, Visit 02, failed				
	Diagnostic Status: No Diagnostics				
	Scientific Instruments: COS/FUV, COS/NUV				
	Special Requirements: SCHED 100%				

<b>Fixed Targets</b>	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
	(1)	NGC-863	RA: 02 14 33.5790 (33.6399125d) Dec: -00 46 0.28 (-.76674d) Equinox: J2000		V=13.81	Reference Frame: ICRS
	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i>					
	Category=GALAXY Description=[ACCRETION DISK, BLR, SEYFERT] Extended=NO					

<b>Exposures</b>	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	NUV MIRR OR B ACQ (COS.ta.103 0198)	(1) NGC-863	COS/NUV, ACQ/IMAGE, PSA	MIRRORB				140 Secs (140 Secs) [==>]	[1]
	2	FUV G140L /1105 Scienc e (COS.sp.103 0204)	(1) NGC-863	COS/FUV, TIME-TAG, PSA	G140L 1105 A	FP-POS=ALL; BUFFER-TIME=50 00			474 Secs (1720 Secs) [==>430.0 Secs (Split 1)] [==>430.0 Secs (Split 2)] [==>430.0 Secs (Split 3)] [==>430.0 Secs (Split 4)]	[1]



Proposal 15448 - SafingAutoRepeat (S3) - Constraining the emergent EUV ionizing emission in the reawakening monster in Mrk 590

Tue Jan 23 01:08:38 GMT 2018

<b>Visit</b>	Proposal 15448, SafingAutoRepeat (S3)				
	Diagnostic Status: No Diagnostics				
	Scientific Instruments: COS/FUV, COS/NUV				
	Special Requirements: SCHED 100%				

<b>Fixed Targets</b>	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
	(1)	NGC-863	RA: 02 14 33.5790 (33.6399125d) Dec: -00 46 0.28 (-.76674d) Equinox: J2000		V=13.81	Reference Frame: ICRS
	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i>					
	Category=GALAXY Description=[ACCRETION DISK, BLR, SEYFERT] Extended=NO					

<b>Exposures</b>	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	NUV MIRR OR B ACQ (COS.ta.103 0198)	(1) NGC-863	COS/NUV, ACQ/IMAGE, PSA	MIRRORB				140 Secs (140 Secs) [==>]	[1]
	2	FUV G140L /1105 Scienc e (COS.sp.103 0204)	(1) NGC-863	COS/FUV, TIME-TAG, PSA	G140L 1105 A	FP-POS=ALL; BUFFER-TIME=50 00			474 Secs (1720 Secs) [==>430.0 Secs (Split 1)] [==>430.0 Secs (Split 2)] [==>430.0 Secs (Split 3)] [==>430.0 Secs (Split 4)]	[1]

