



15476 - The Formation History of Milky Way Satellite Canes Venatici I

Cycle: 25, Proposal Category: GO

(Availability Mode: SUPPORTED)

INVESTIGATORS

<i>Name</i>	<i>Institution</i>	<i>E-Mail</i>
Dr. Daniel R. Weisz (PI) (Contact)	University of California - Berkeley	dan.weisz@berkeley.edu
Dr. Nitya Kallivayalil (CoI)	The University of Virginia	nityajk@gmail.com
Prof. Michael Boylan-Kolchin (CoI)	University of Texas at Austin	mbk@astro.as.utexas.edu
Ms. Tom Zick (CoI)	University of California - Berkeley	tzick@berkeley.edu
Mr. Sean Linden (CoI)	The University of Virginia	stl7ey@virginia.edu
Paul Zivick (CoI)	The University of Virginia	pzivick@gmail.com

VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) CVN-I-DSPH-1 ANY	ACS/WFC WFC3/UVIS	2	07-Jun-2018 14:00:18.0	yes
02	(1) CVN-I-DSPH-1 ANY	ACS/WFC WFC3/UVIS	2	07-Jun-2018 14:00:21.0	yes
03	(2) CVN-I-DSPH-2 ANY	ACS/WFC WFC3/UVIS	2	07-Jun-2018 14:00:23.0	yes
04	(2) CVN-I-DSPH-2 ANY	ACS/WFC WFC3/UVIS	2	07-Jun-2018 14:00:24.0	yes

8 Total Orbits Used

ABSTRACT

We propose deep ACS/WFC and WFC3/UVIS F606W and F814W imaging of Milky Way (MW) satellite Canes Venatici I (CVn I) to measure its star formation history (SFH) and establish first epoch imaging for its proper motion (PM) determination. CVn I ($M_v = -7.9$) is a distant (~ 220 kpc) MW satellite that straddles the boundary between ultra-faint and classical dwarf galaxy. It is among the most diffuse, faint satellites in the local Universe ($r_h > 500$ pc, $\mu = 28.1$ mag/arcsec²), but whether its low density is intrinsic or environmentally driven remains unknown.

The SFH and PM motion measurements enable a wide range of immediate and long term science including (1) determining if the timing and intensity of star formation can explain CVn I's low density through dark matter core creation; (2) examining how reionization impacts galaxy formation in this transitional mass regime; (3) checking if the sub-Solar stellar IMF follows the trends established in other ultra-faint galaxies; (4) determining if CVn I's orbital history includes a close pericentric passage (< 20 kpc) to the MW's disk, which would have tidally disturbed CVn I; (5) providing one of the most precise measurements of the total mass of the MW's halo. Furthermore, this program will put CVn I on same observational footing as all other MW satellites.

Only HST can achieve any of these science goals. However, all archival HST observations of CVn I are inadequate. Maximizing the time baseline for PM measurement is scientifically imperative, but it will be 2 cycles until normal channels for small programs are restored. Waiting will substantially compromise the precision of the PM determination.

OBSERVING DESCRIPTION

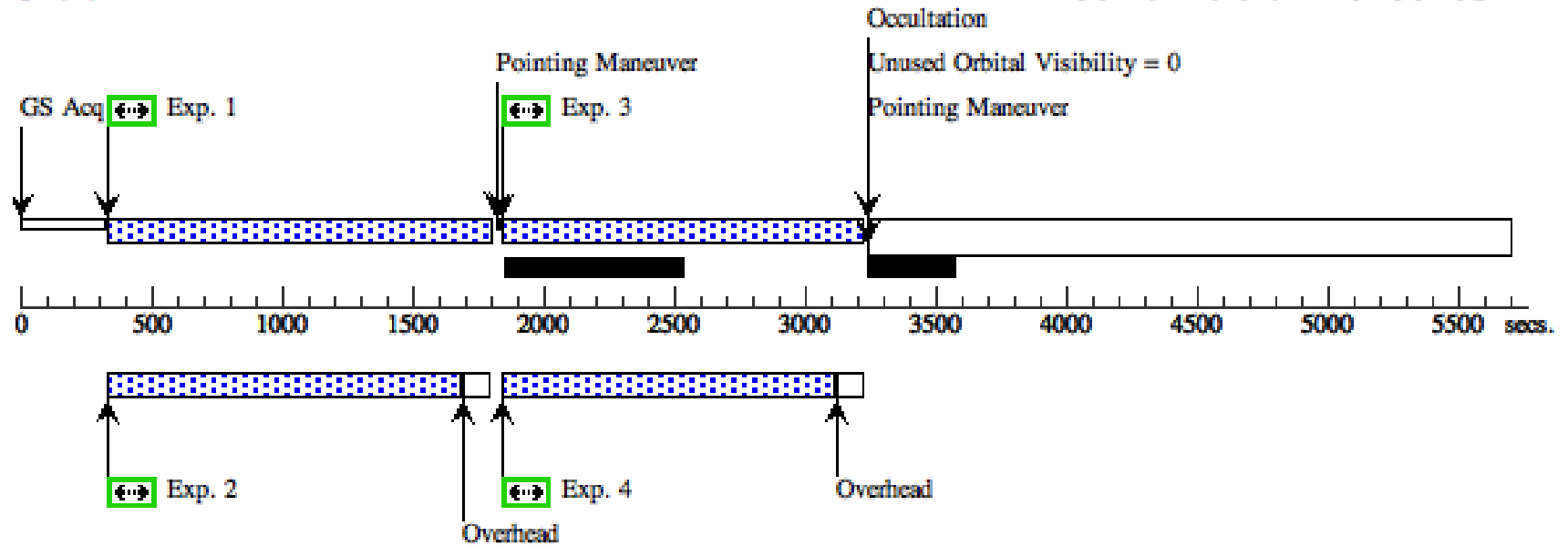
We will image CVn I using the F606W and F814W filters of ACS/WFC and WFC3/UVIS. Our imaging will reach high S/N at the oldest main sequence turnoff in multiple fields, allowing us to measure the detailed star formation history (SFH) of CVn I across all cosmic time. Our imaging is well-dithered, and thus establishes a first epoch proper motion baseline, and is deep enough to explore the sub-Solar stellar initial mass function.

Our observing plan generally follows that specified in the GO-14734, a Treasury survey of Milky Way satellites, which did not include CVn I. Our program consists of two adjacent ACS pointings located within the half-light radius. Each pointing is allocated 4 orbits, two in F606W and two in F814W, spread out over two visits. Each orbit is executed as a single filter and divided into long exposures (~ 1100 - 1300 sec per filter) in order to minimize overheads. For a single orbit in each filter, we add a short integration (60 sec) to mitigate any potential saturation issues with bright stars. We use POS-TARG commands to implement the 4 point dither pattern specified by the ACS users handbook. This dither pattern satisfies PSF sampling requirements for consistent astrometric fidelity with observations of other MW satellites.

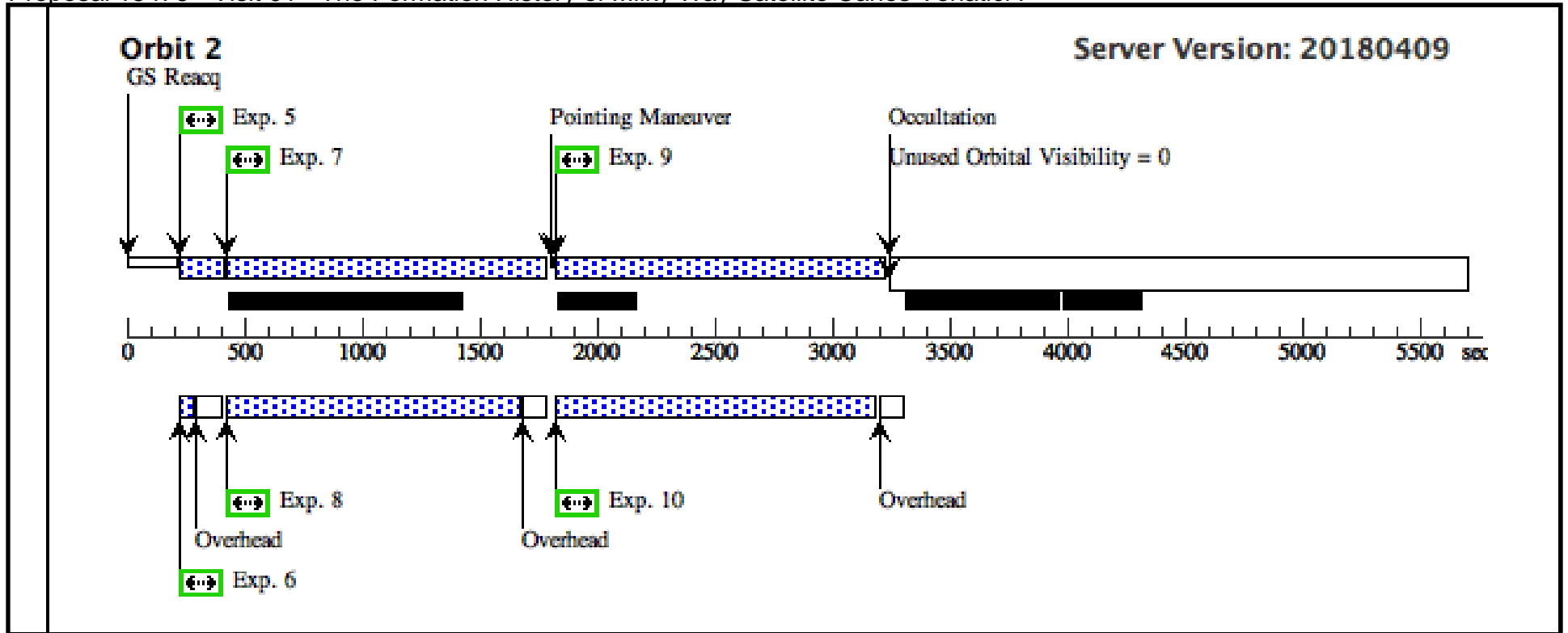
WFC3/UVIS observations are taken in parallel mode, and follow a similar observing strategy to ACS. For the 60 sec UVIS exposures, we use FLASH=10. To maximize schedulability of this program, we only constrain the roll angle to avoid bright foreground stars that may negatively affect the parallels. However, we do require that different visits for the same pointings have the same ORIENTATION. Given the minimal roll angle constraints and the diffuse nature of CVn I, there is a >50% chance that one of the UVIS fields will sample stars within the galaxy's half-light radius.

Orbit 1

Server Version: 20180409



Orbit Structure



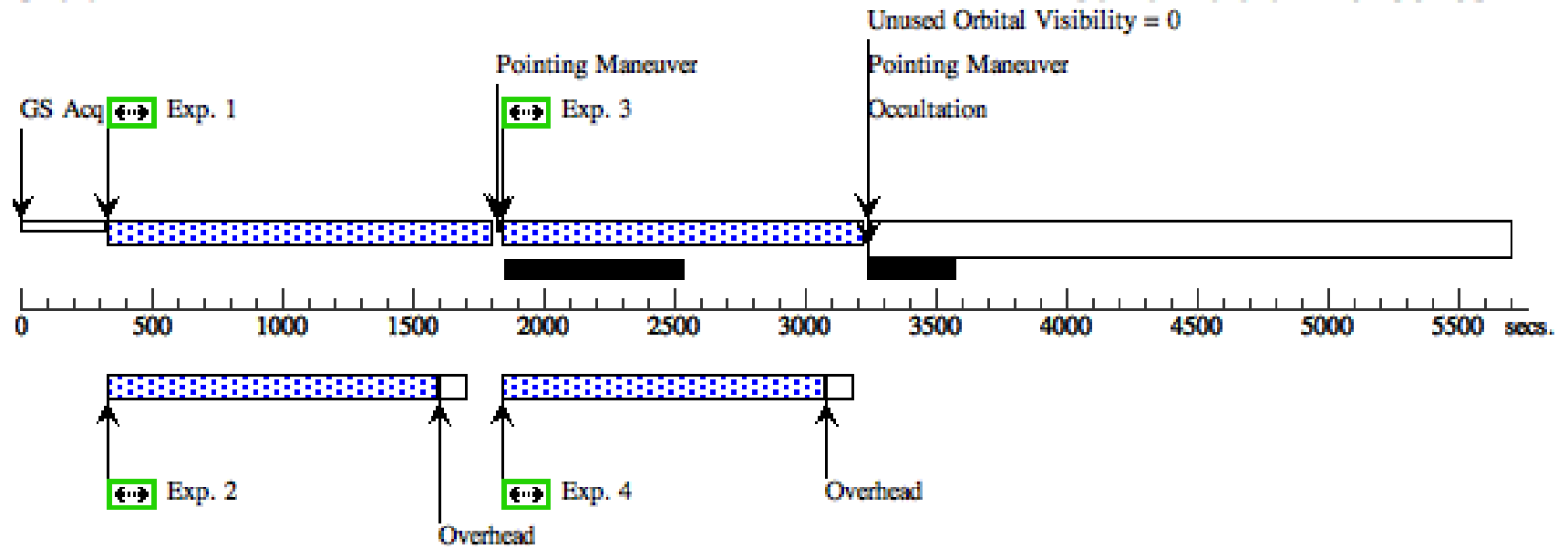
Proposal 15476 - Visit 02 - The Formation History of Milky Way Satellite Canes Venatici I

Thu Jun 07 18:00:26 GMT 2018

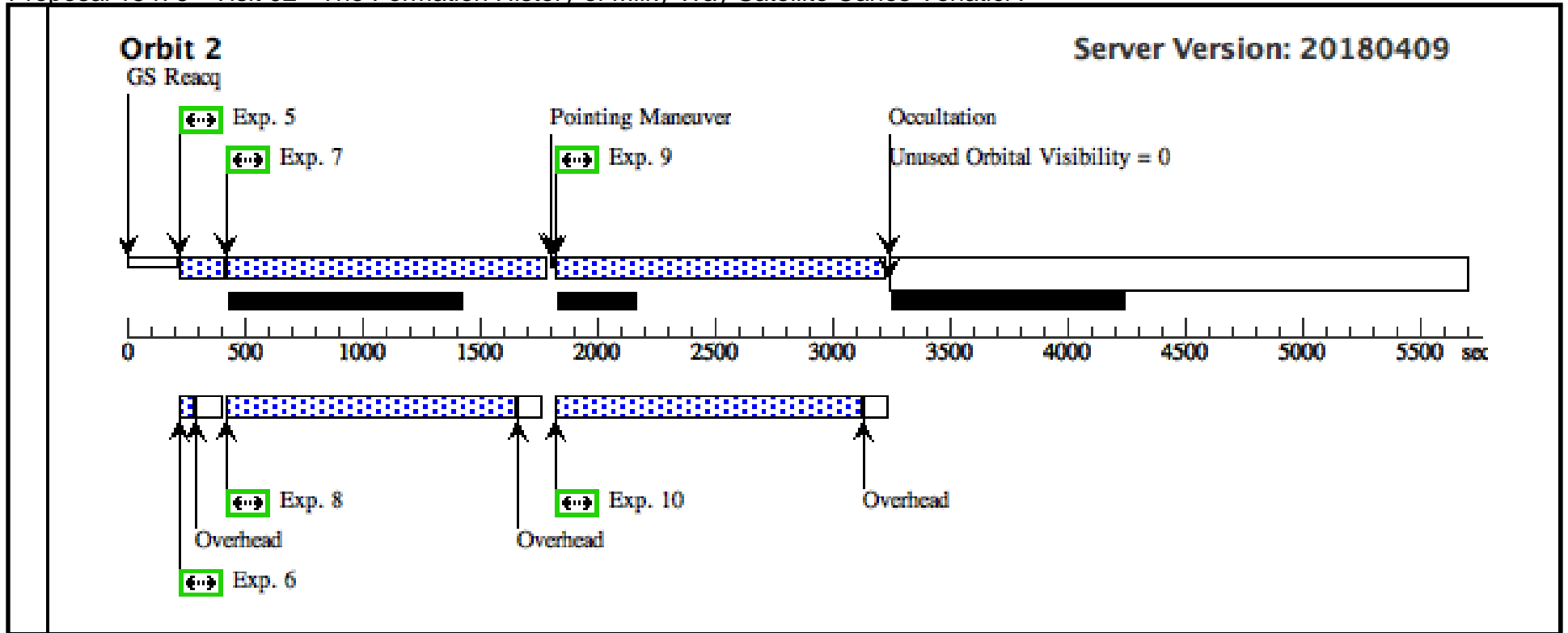
Visit	Proposal 15476, Visit 02, implementation Diagnostic Status: No Diagnostics Scientific Instruments: WFC3/UVIS, ACS/WFC Special Requirements: SAME ORIENT AS 01									
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous			
	(1)	CVN-I-DSPH-1	RA: 13 28 18.2939 (202.0762246d) Dec: +33 35 1.34 (33.58371d) Equinox: J2000		V=19+/-1	Reference Frame: NED				
	<i>Comments: Faintest targets have V~28-29. A few foreground stars with V~13</i> Category=GALAXY Description=[DWARF SPHEROIDAL]									
Exposures	#	Label	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	(1) CVN-I-DSPH-1	ACS/WFC, ACCUM, WFC	F814W		POS TARG 0.0,0.0	Prime + Parallel Group 1-2 in Visit 02	1269 Secs (1269 Secs)	[1]	
	2	ANY	WFC3/UVIS, ACCUM, UVIS	F814W			Prime + Parallel Group 1-2 in Visit 02	1229 Secs (1229 Secs)	[1]	
	3	(1) CVN-I-DSPH-1	ACS/WFC, ACCUM, WFC	F814W		POS TARG 0.148,0.086	Prime + Parallel Group 3-4 in Visit 02	1260 Secs (1260 Secs)	[1]	
	4	ANY	WFC3/UVIS, ACCUM, UVIS	F814W			Prime + Parallel Group 3-4 in Visit 02	1229 Secs (1229 Secs)	[1]	
	5	(1) CVN-I-DSPH-1	ACS/WFC, ACCUM, WFC	F814W		POS TARG 0.222,0.240	Prime + Parallel Group 5-6 in Visit 02	60 Secs (60 Secs)	[2]	
	6	ANY	WFC3/UVIS, ACCUM, UVIS	F814W	FLASH=10		Prime + Parallel Group 5-6 in Visit 02	60 Secs (60 Secs)	[2]	
	7	(1) CVN-I-DSPH-1	ACS/WFC, ACCUM, WFC	F814W		POS TARG 0.222,0.240	Prime + Parallel Group 7-8 in Visit 02	1240 Secs (1240 Secs)	[2]	
	8	ANY	WFC3/UVIS, ACCUM, UVIS	F814W			Prime + Parallel Group 7-8 in Visit 02	1230 Secs (1230 Secs)	[2]	
	9	(1) CVN-I-DSPH-1	ACS/WFC, ACCUM, WFC	F814W		POS TARG 0.074,0.154	Prime + Parallel Group 9-10 in Visit 02	1280 Secs (1280 Secs)	[2]	
10	ANY	WFC3/UVIS, ACCUM, UVIS	F814W			Prime + Parallel Group 9-10 in Visit 02	1300 Secs (1300 Secs)	[2]		

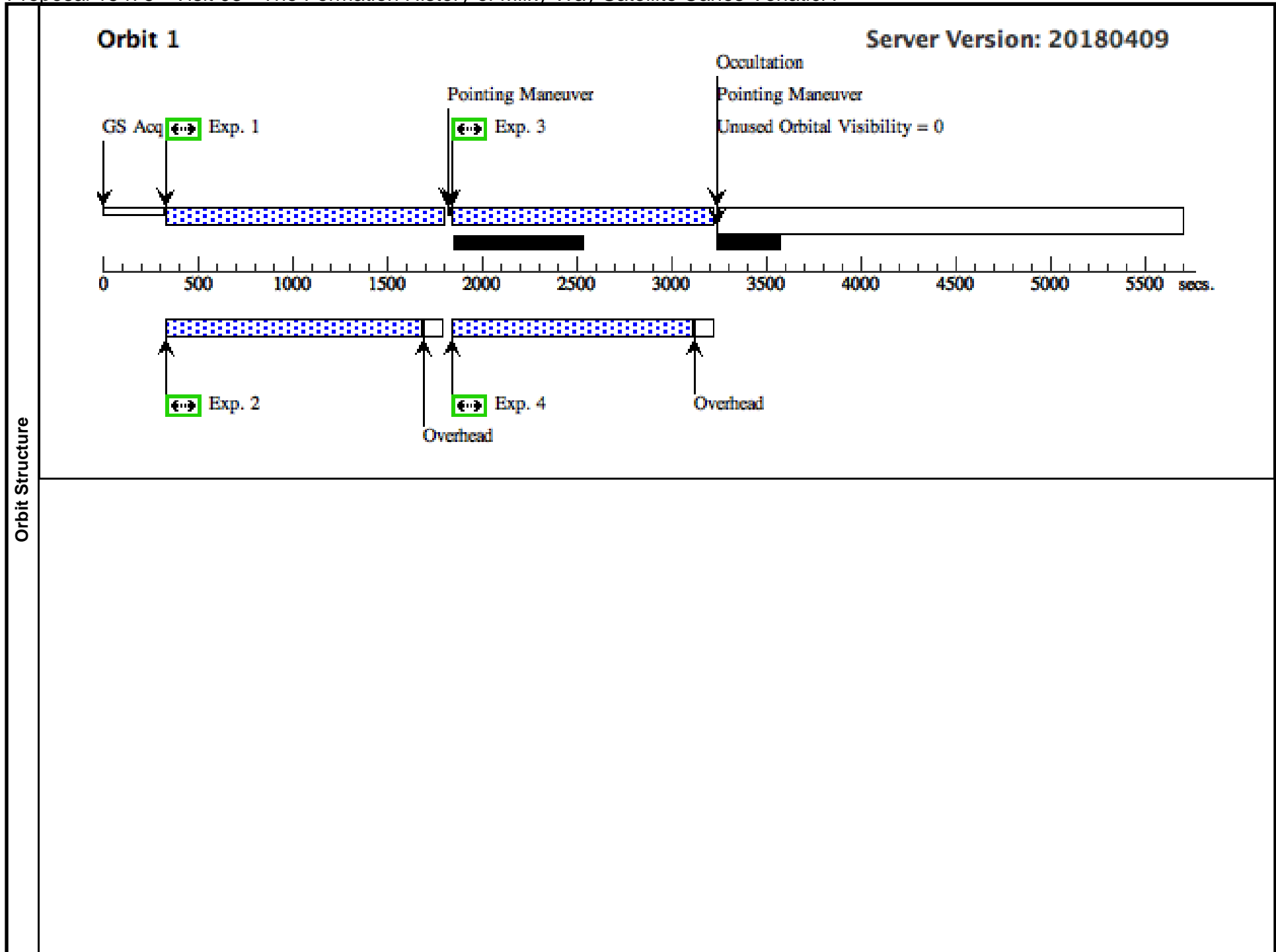
Orbit 1

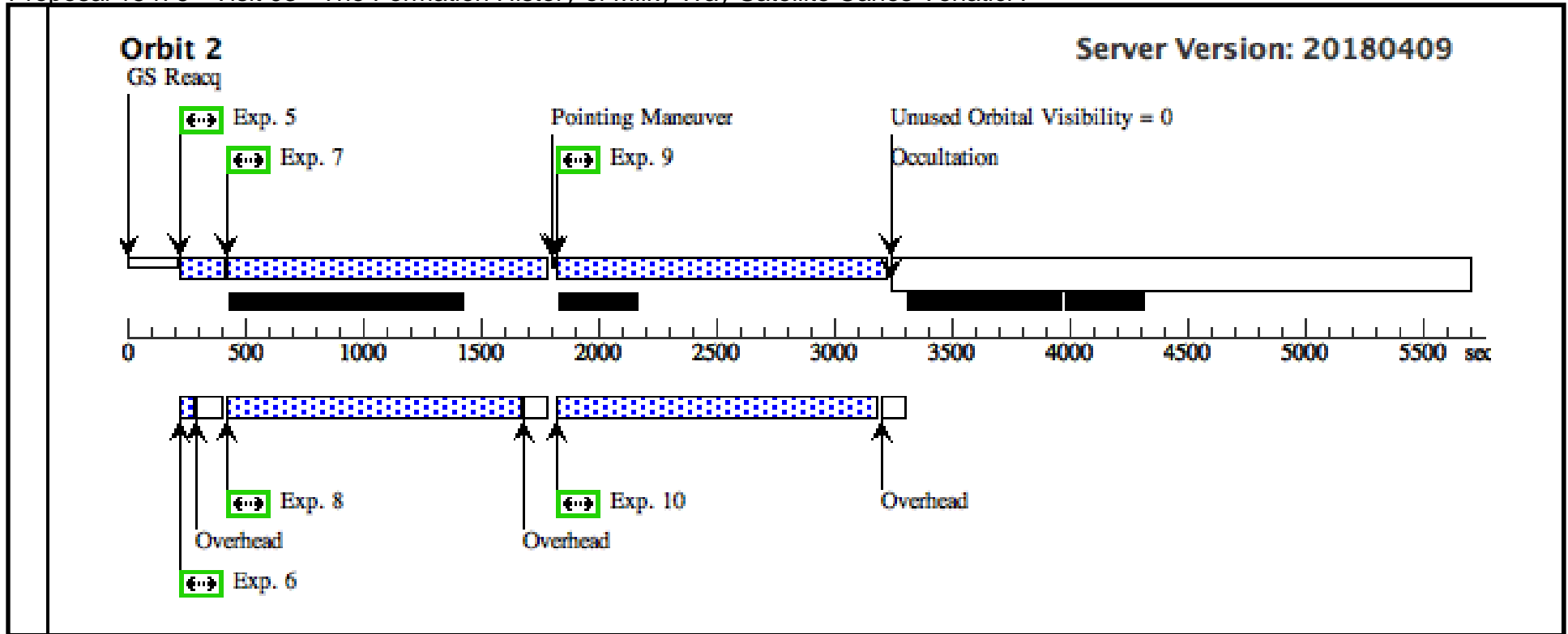
Server Version: 20180409



Orbit Structure







Proposal 15476 - Visit 04 - The Formation History of Milky Way Satellite Canes Venatici I

Thu Jun 07 18:00:26 GMT 2018

Visit	Proposal 15476, Visit 04, implementation Diagnostic Status: No Diagnostics Scientific Instruments: WFC3/UVIS, ACS/WFC Special Requirements: SAME ORIENT AS 03									
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous			
	(2)	CVN-I-DSPH-2	RA: 13 28 2.0601 (202.0085837d) Dec: +33 33 30.88 (33.55858d) Equinox: J2000		V=19+/-1	Reference Frame: NED				
	<i>Comments: Faintest targets have V~28-29. A few foreground stars with V~13</i> Category=GALAXY Description=[DWARF SPHEROIDAL]									
Exposures	#	Label	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	(2)	CVN-I-DSPH-2	ACS/WFC, ACCUM, WFC	F814W		POS TARG 0.0,0.0	Prime + Parallel Group 1-2 in Visit 04	1269 Secs (1269 Secs)	[1]
	2	ANY		WFC3/UVIS, ACCUM, UVIS	F814W			Prime + Parallel Group 1-2 in Visit 04	1229 Secs (1229 Secs)	[1]
	3	(2)	CVN-I-DSPH-2	ACS/WFC, ACCUM, WFC	F814W		POS TARG 0.148,0.086	Prime + Parallel Group 3-4 in Visit 04	1260 Secs (1260 Secs)	[1]
	4	ANY		WFC3/UVIS, ACCUM, UVIS	F814W			Prime + Parallel Group 3-4 in Visit 04	1229 Secs (1229 Secs)	[1]
	5	(2)	CVN-I-DSPH-2	ACS/WFC, ACCUM, WFC	F814W		POS TARG 0.222,0.240	Prime + Parallel Group 5-6 in Visit 04	60 Secs (60 Secs)	[2]
	6	ANY		WFC3/UVIS, ACCUM, UVIS	F814W	FLASH=10		Prime + Parallel Group 5-6 in Visit 04	60 Secs (60 Secs)	[2]
	7	(2)	CVN-I-DSPH-2	ACS/WFC, ACCUM, WFC	F814W		POS TARG 0.222,0.240	Prime + Parallel Group 7-8 in Visit 04	1240 Secs (1240 Secs)	[2]
	8	ANY		WFC3/UVIS, ACCUM, UVIS	F814W			Prime + Parallel Group 7-8 in Visit 04	1230 Secs (1230 Secs)	[2]
	9	(2)	CVN-I-DSPH-2	ACS/WFC, ACCUM, WFC	F814W		POS TARG 0.074,0.154	Prime + Parallel Group 9-10 in Visit 04	1280 Secs (1280 Secs)	[2]
10	ANY		WFC3/UVIS, ACCUM, UVIS	F814W			Prime + Parallel Group 9-10 in Visit 04	1300 Secs (1300 Secs)	[2]	

