



15488 - UV Spectroscopy of the Nearby Superluminous Supernova SN2018bsz

Cycle: 25, Proposal Category: GO/DD

(Availability Mode: SUPPORTED)

INVESTIGATORS

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VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) SN2018BSZ	COS/FUV COS/NUV STIS/CCD STIS/NUV-MAMA	3	08-Jun-2018 20:03:31.0	yes

3 Total Orbits Used

ABSTRACT

Proposal 15488 (STScI Edit Number: 1, Created: Friday, June 8, 2018 7:03:33 PM EST) - Overview

The latest generation of wide-field time-domain surveys has led to the discovery of rare, extreme transients such as superluminous supernovae (SLSNe). Recently, exotic power sources such as a magnetar central engine have gained traction as a plausible explanation for their enormous luminosities. However, due to their low volumetric rates, only a handful of SLSNe have been discovered at $z < 0.1$ where the effects of the power source can be probed in the greatest detail. A key region of the spectrum for testing the effects of various proposed power sources is the UV, where the spectral energy distributions peak. Fully understanding the UV characteristics and diversity revealed by ground-based observations of high-redshift events and HST observations of low-redshift events requires additional observations. Here we propose to obtain FUV and NUV observations with HST of SN2018bsz at a redshift of $z = 0.026$, the nearest SLSN discovered to date. In addition to one epoch of FUV+NUV observations, we request two additional epochs of NUV observations to monitor the UV evolution. SN2018bsz provides an excellent opportunity to obtain high signal-to-noise UV spectroscopy that can be compared with other events with UV spectra and with radiative transfer models. Characterizing the UV spectra of low- z SLSNe with HST is critical for understanding the rest-frame UV spectra of future events that will be discovered at $z > 1$ with upcoming surveys such as LSST.

OBSERVING DESCRIPTION

We request a single visit with 2 orbits of COS/FUV and 1 orbit of STIS/NUV.

Proposal 15488 - Visit 01 - UV Spectroscopy of the Nearby Superluminous Supernova SN2018bsz

Sat Jun 09 00:03:33 GMT 2018

Visit	Proposal 15488, Visit 01, implementation Diagnostic Status: Warning Scientific Instruments: STIS/NUV-MAMA, STIS/CCD, COS/FUV, COS/NUV Special Requirements: BEFORE 02-JUL-2018:00:00:00									
	Diagnosics (Visit 01) Warning (Orbit Planner): INEFFICIENT ORDERING OF FP-POS POSITIONS									
Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous				
	(1)	SN2018BSZ	RA: 16 09 39.1100 (242.4129583d) Dec: -32 03 45.63 (-32.06268d) Equinox: J2000		V=16.03 U = 15.50, UVW1 = 15.51, UVM2 = 16.05, UVW2 = 16.42	Reference Frame: ICRS				
Comments: Category=EXT-STAR Description=[SUPERNOVA] Extended=NO										
Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	(COS.ta.116 7707)	(1) SN2018BSZ	COS/NUV, ACQ/IMAGE, PSA	MIRRORB				70 Secs (70 Secs) [==>]	[1]
	2	(COS.sp.116 7733)	(1) SN2018BSZ	COS/FUV, TIME-TAG, PSA	G140L 1105 A	BUFFER-TIME=60 00; FP-POS=ALL; FLASH=YES			532 Secs (2128 Secs) [==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[1]
	3	(COS.sp.116 7786)	(1) SN2018BSZ	COS/FUV, TIME-TAG, PSA	G140L 1105 A	BUFFER-TIME=60 00; FP-POS=ALL; FLASH=YES			659 Secs (2636 Secs) [==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[2]
	4		(1) SN2018BSZ	STIS/CCD, ACQ, F28X50LP	MIRROR				15 Secs (15 Secs) [==>]	[3]
	5	(STIS.sp.11 67787)	(1) SN2018BSZ	STIS/NUV-MAMA, ACCUM, 52X0.2	G230L 2376 A				2378 Secs (2378 Secs) [==>]	[3]



