



15835 - Near-Ultraviolet Follow-up of the X-ray-detected Warm-Hot Intergalactic Medium Toward 1ES 1553+113

Cycle: 27, Proposal Category: GO

(UV Initiative)

(Availability Mode: SUPPORTED)

INVESTIGATORS

<i>Name</i>	<i>Institution</i>	<i>E-Mail</i>
Dr. Sowgat Muzahid (PI) (ESA Member) (Contact)	Leibniz-Institut fur Astrophysik Potsdam (AIP)	sowgatm@gmail.com
Dr. Sean Johnson (CoI) (AdminUSPI) (Contact)	University of Michigan	sdj@astro.princeton.edu
Prof. Joop Schaye (CoI) (ESA Member)	Universiteit Leiden	schaye@strw.leidenuniv.nl
Prof. Jane C. Charlton (CoI)	The Pennsylvania State University	charlton@astro.psu.edu
Prof. Hsiao-Wen Chen (CoI)	University of Chicago	hchen@oddjob.uchicago.edu
Dr. Anand Narayanan (CoI)	Indian Institute of Space Science and Technology	anand@iist.ac.in
Ms. Nastasha Wijers (CoI) (ESA Member)	Universiteit Leiden	wijers@strw.leidenuniv.nl
Dr. John S. Mulchaey (CoI)	Carnegie Institution of Washington	mulchaey@obs.carnegiescience.edu

VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) QSO-B1553+113	COS/FUV COS/NUV	2	30-Jul-2020 15:00:14.0	yes
02	(1) QSO-B1553+113	COS/FUV COS/NUV	2	30-Jul-2020 15:00:15.0	yes
52	(1) QSO-B1553+113	COS/FUV COS/NUV	1	30-Jul-2020 15:00:16.0	yes

5 Total Orbits Used

ABSTRACT

One of the strengths of modern cosmology is that the cosmic baryon density predicted by Big Bang nucleosynthesis precisely matches the measured value from the cosmic microwave background. At high redshift the vast majority of these baryons are found in the photoionized intergalactic medium. However, cosmological simulations of structure formation predict that a substantial portion of these baryons will be shock heated into the elusive warm-hot intergalactic medium (WHIM) as the universe evolves. Confirmation of these predictions through observations of the WHIM and measurements of its physical state provide a sensitive test of our understanding of the growth of structure in the universe. Recently, Nicastro et al. (2018) reported the first detections of two intervening OVII absorption systems in the X-ray spectrum of the blazar 1ES 1553+113. These systems are thought to be arising from the WHIM with a temperature of $\sim 10^6$ K. Nevertheless, the redshift of the blazar is not known with adequate precision, which makes the origin(s) of the OVII bearing gas and its contribution to the cosmic baryon density highly uncertain. Here we propose to obtain NUV spectra of the blazar using the COS G185M grating. The proposed spectra will enable us to detect intervening Lyman alpha absorbers with $\log N(\text{HI})/\text{cm}^{-2} > 13.0$ from $z = 0.45 - 0.7$. The stringent lower (upper) limit on the blazar's redshift to be obtained from the presence (absence) of any such absorber will be crucial to validate or rule out the claim of potentially solving the so called "missing baryons problem".

OBSERVING DESCRIPTION

In order to place stringent constraints on the blazar (1ES1553+113) redshift and determine whether the O VII absorber at $z = 0.4339$ is in the blazar's host environment, we will obtain NUV COS spectra to search for intervening IGM HI Ly α lines at $z > 0.41$. Contiguous spectral coverage in the NUV from 1767 to 2068 Angstrom with a S/N per resolution element of 12 will enable us to detect intervening IGM Ly α absorption systems with $\log N(\text{HI}) > 13$ from $z = 0.45$ to 0.7 . Based on the expected number of such HI systems per unit redshift from Danforth et al. (2016), we expect to detect 10 Ly α lines per $dz = 0.1$ interval. Non-detection of Ly α at $z > 0.45$ in combination with our deep and highly complete galaxy redshift survey in the field would indicate that the blazar is in the same environment as the WHIM candidate at $z=0.4339$.

Based on the COS ETC, obtaining the desired spectral coverage and S/N will require observations with COS G185M gratings with central wavelength settings of 1882, 1913, 1941, and 1953, each observed for one orbit for a total of four orbits of COS NUV spectroscopy. We will split each central wavelength setting into four FP-POS positions as recommended to minimize flat fielding errors. Moreover, as per the instructions after the gyro failure, we will complete the total allocated 4 orbits in two visits.

Proposal 15835 (STScI Edit Number: 0, Created: Thursday, July 30, 2020 at 2:00:17 PM Eastern Standard Time) - Overview

Since the co-ordinates of the target are obtained from GAIA with an accuracy better than 0.02", we will directly use PEAKXD followed by PEAKD for acquisition. For the flux observed in the existing COS spectrum ($F(1300) = 1.5 \times 10^{-14}$), ACQ/IMAGE is a viable option. However, it exceeds the bright limit if the flux is increased by a factor of 2.5 (1 mag) owing to unexpected flare. Therefore, we adopt the spectroscopic acquisition for instruments' safety. We used the existing archival COS spectrum of the blazar for calculating the exposure time for spectroscopic acquisition (COS.sa.1366483).

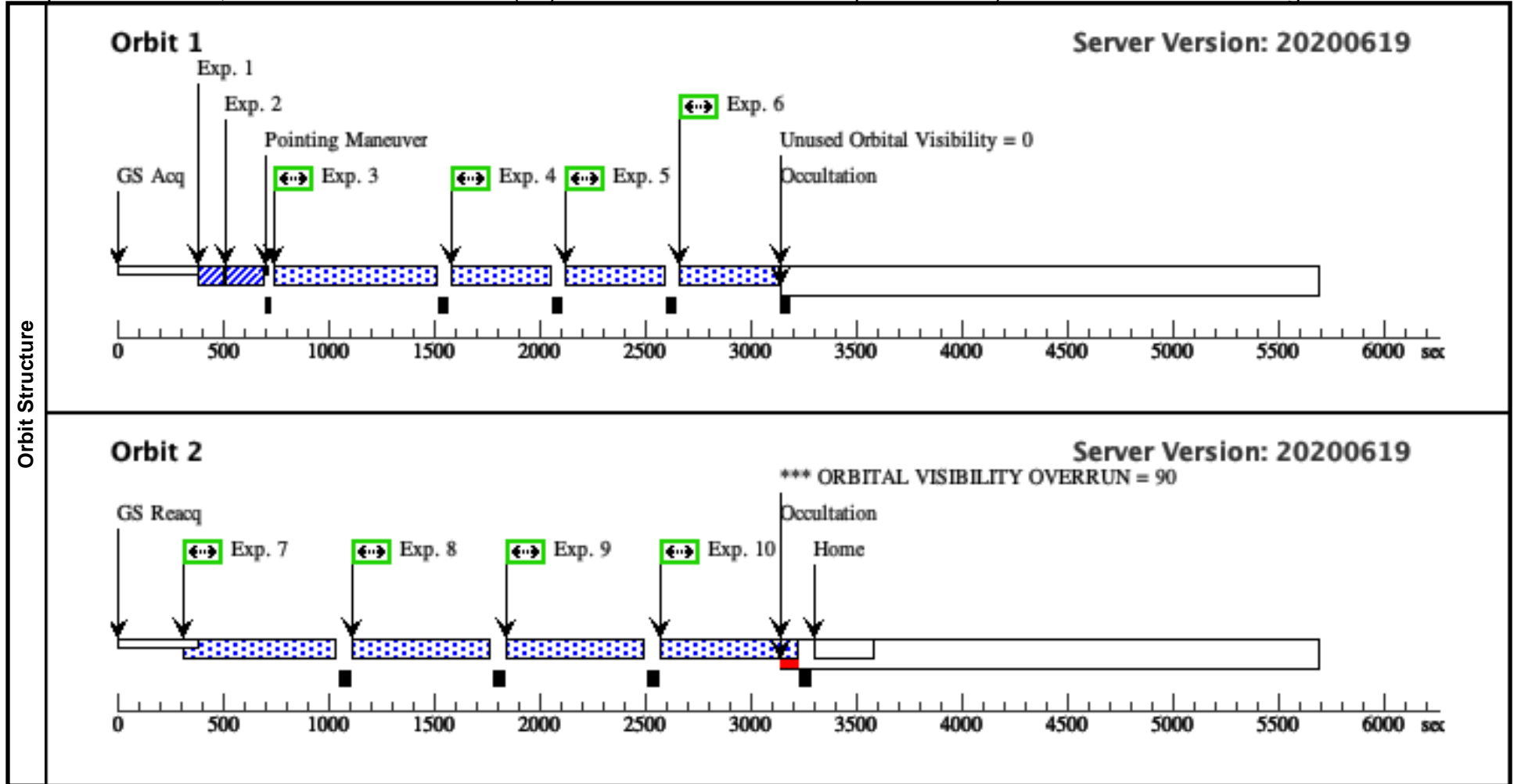
Impact of reduced-gyro operations: There are no significant impact of the reduced-gyro mode (RGM) operations except for an increased target acquisition time of 2 minutes.

Proposal 15835 - QSO-B1553+113-VISIT-01 (01) - Near-Ultraviolet Follow-up of the X-ray-detected Warm-Hot Intergalactic Medium T...

Visit	Proposal 15835, QSO-B1553+113-VISIT-01 (01), completed Thu Jul 30 19:00:17 GMT 2020 Diagnostic Status: Warning Scientific Instruments: COS/FUV, COS/NUV Special Requirements: (none)																
	Diagnostics (QSO-B1553+113-VISIT-01 (01)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN																
Fixed Targets	<table border="1"> <thead> <tr> <th>#</th> <th>Name</th> <th>Target Coordinates</th> <th>Targ. Coord. Corrections</th> <th>Fluxes</th> <th>Miscellaneous</th> </tr> </thead> <tbody> <tr> <td>(1)</td> <td>QSO-B1553+113</td> <td> RA: 15 55 43.0440 (238.9293500d) Dec: +11 11 24.37 (11.19010d) Equinox: J2000 </td> <td></td> <td> V=14.57 G=14, FUV=16.3, NUV=15.6, F(1300) = 1.5e-14 from HST/C OS </td> <td>Reference Frame: ICRS</td> </tr> </tbody> </table>					#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous	(1)	QSO-B1553+113	RA: 15 55 43.0440 (238.9293500d) Dec: +11 11 24.37 (11.19010d) Equinox: J2000		V=14.57 G=14, FUV=16.3, NUV=15.6, F(1300) = 1.5e-14 from HST/C OS	Reference Frame: ICRS
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Comments: Co-ordinates are from GAIA Category=GALAXY Description=[BL LAC] Extended=NO																	

Proposal 15835 - QSO-B1553+113-VISIT-01 (01) - Near-Ultraviolet Follow-up of the X-ray-detected Warm-Hot Intergalactic Medium T...

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
Exposures	1	ACQ-PEAK XD (COS.sa.136 6483)	(1) QSO-B1553+113	COS/FUV, ACQ/PEAKXD, PSA	G130M 1291 A	SEGMENT=BOTH; NUM-POS=3; STEP-SIZE=1.3; CENTER=FLUX-W T		5 Secs (5 Secs) [==>]	[1]
	2	ACQ-PEAK D (COS.sa.136 6483)	(1) QSO-B1553+113	COS/FUV, ACQ/PEAKD, PSA	G130M 1291 A	NUM-POS=5; STEP-SIZE=0.9; SEGMENT=BOTH; CENTER=FLUX-W T-FLR		5 Secs (5 Secs) [==>]	[1]
	3	TIME-TAG- G185M-188 2-FP-1 (COS.sp.136 6492)	(1) QSO-B1553+113	COS/NUV, TIME-TAG, PSA	G185M 1882 A	FP-POS=1; BUFFER-TIME=15 61		456 Secs (456 Secs) [==>]	[1]
	4	TIME-TAG- G185M-188 2-FP-2 (COS.sp.136 6492)	(1) QSO-B1553+113	COS/NUV, TIME-TAG, PSA	G185M 1882 A	FP-POS=2; BUFFER-TIME=15 61		454 Secs (454 Secs) [==>]	[1]
	5	TIME-TAG- G185M-188 2-FP-3 (COS.sp.136 6492)	(1) QSO-B1553+113	COS/NUV, TIME-TAG, PSA	G185M 1882 A	FP-POS=3; BUFFER-TIME=15 61		454 Secs (454 Secs) [==>]	[1]
	6	TIME-TAG- G185M-188 2-FP-4 (COS.sp.136 6492)	(1) QSO-B1553+113	COS/NUV, TIME-TAG, PSA	G185M 1882 A	FP-POS=4; BUFFER-TIME=15 61		454 Secs (454 Secs) [==>]	[1]
	7	TIME-TAG- G185M-191 3-FP-1 (COS.sp.136 6492)	(1) QSO-B1553+113	COS/NUV, TIME-TAG, PSA	G185M 1913 A	FP-POS=1; BUFFER-TIME=15 61		633 Secs (633 Secs) [==>]	[2]
	8	TIME-TAG- G185M-191 3-FP-2 (COS.sp.136 6492)	(1) QSO-B1553+113	COS/NUV, TIME-TAG, PSA	G185M 1913 A	FP-POS=2; BUFFER-TIME=15 61		632 Secs (632 Secs) [==>]	[2]
	9	TIME-TAG- G185M-191 3-FP-3 (COS.sp.136 6492)	(1) QSO-B1553+113	COS/NUV, TIME-TAG, PSA	G185M 1913 A	FP-POS=3; BUFFER-TIME=15 61		632 Secs (632 Secs) [==>]	[2]
	10	TIME-TAG- G185M-191 3-FP-4 (COS.sp.136 6492)	(1) QSO-B1553+113	COS/NUV, TIME-TAG, PSA	G185M 1913 A	FP-POS=4; BUFFER-TIME=15 61		632 Secs (632 Secs) [==>]	[2]

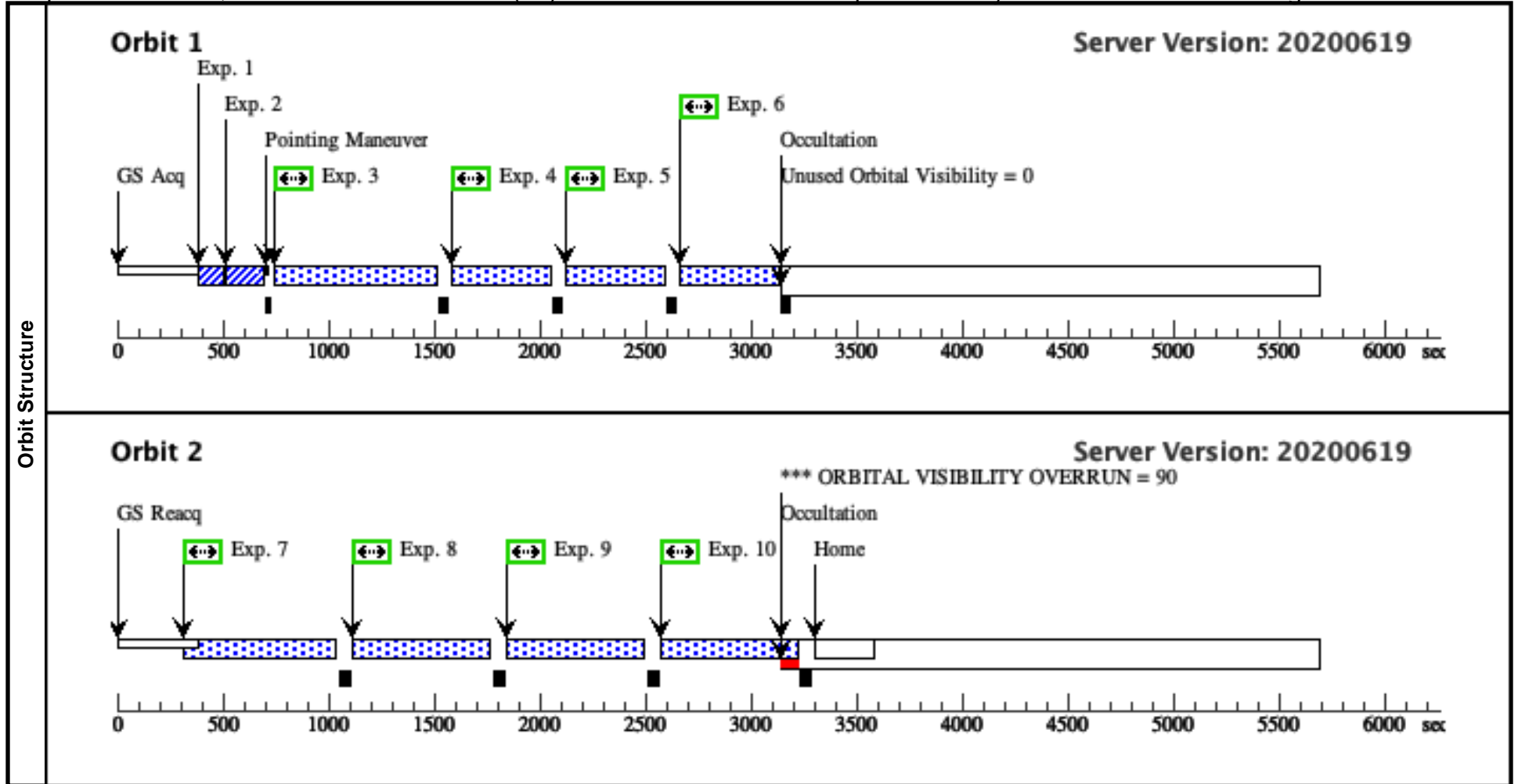


Proposal 15835 - QSO-B1553+113-VISIT-02 (02) - Near-Ultraviolet Follow-up of the X-ray-detected Warm-Hot Intergalactic Medium T...

Visit	Proposal 15835, QSO-B1553+113-VISIT-02 (02), failed Thu Jul 30 19:00:17 GMT 2020 Diagnostic Status: Warning Scientific Instruments: COS/FUV, COS/NUV Special Requirements: (none)																																					
	Diagnostics	(QSO-B1553+113-VISIT-02 (02)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN																																				
Fixed Targets		<table border="1" style="width: 100%; border-collapse: collapse;"> <thead> <tr> <th data-bbox="136 389 241 414">#</th> <th data-bbox="241 389 472 414">Name</th> <th data-bbox="472 389 892 414">Target Coordinates</th> <th data-bbox="892 389 1312 414">Targ. Coord. Corrections</th> <th data-bbox="1312 389 1606 414">Fluxes</th> <th data-bbox="1606 389 2005 414">Miscellaneous</th> </tr> </thead> <tbody> <tr> <td data-bbox="136 414 241 690" rowspan="2">(1)</td> <td data-bbox="241 414 472 690" rowspan="2">QSO-B1553+113</td> <td data-bbox="472 414 892 446">RA: 15 55 43.0440 (238.9293500d)</td> <td data-bbox="892 414 1312 446"></td> <td data-bbox="1312 414 1606 446">V=14.57</td> <td data-bbox="1606 414 2005 446">Reference Frame: ICRS</td> </tr> <tr> <td data-bbox="472 446 892 479">Dec: +11 11 24.37 (11.19010d)</td> <td data-bbox="892 446 1312 479"></td> <td data-bbox="1312 446 1606 479">G=14,</td> <td data-bbox="1606 446 2005 479"></td> </tr> <tr> <td data-bbox="472 479 892 511">Equinox: J2000</td> <td data-bbox="892 479 1312 511"></td> <td data-bbox="1312 479 1606 511">FUV=16.3,</td> <td data-bbox="1606 479 2005 511"></td> </tr> <tr> <td data-bbox="472 511 892 544"></td> <td data-bbox="892 511 1312 544"></td> <td data-bbox="1312 511 1606 544">NUV=15.6,</td> <td data-bbox="1606 511 2005 544"></td> </tr> <tr> <td data-bbox="472 544 892 576"></td> <td data-bbox="892 544 1312 576"></td> <td data-bbox="1312 544 1606 576">F(1300) = 1.5e-14 from HST/C</td> <td data-bbox="1606 544 2005 576"></td> </tr> <tr> <td data-bbox="472 576 892 609"></td> <td data-bbox="892 576 1312 609"></td> <td data-bbox="1312 576 1606 609">OS</td> <td data-bbox="1606 576 2005 609"></td> </tr> </tbody> </table>	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous	(1)	QSO-B1553+113	RA: 15 55 43.0440 (238.9293500d)		V=14.57	Reference Frame: ICRS	Dec: +11 11 24.37 (11.19010d)		G=14,		Equinox: J2000		FUV=16.3,				NUV=15.6,				F(1300) = 1.5e-14 from HST/C				OS					
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Proposal 15835 - QSO-B1553+113-VISIT-02 (02) - Near-Ultraviolet Follow-up of the X-ray-detected Warm-Hot Intergalactic Medium T...

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
Exposures	1	ACQ-PEAK XD (COS.sa.136 6483)	(1) QSO-B1553+113	COS/FUV, ACQ/PEAKXD, PSA	G130M 1291 A	SEGMENT=BOTH; NUM-POS=3; STEP-SIZE=1.3; CENTER=FLUX-W T		5 Secs (5 Secs) [==>]	[1]
	2	ACQ-PEAK D (COS.sa.136 6483)	(1) QSO-B1553+113	COS/FUV, ACQ/PEAKD, PSA	G130M 1291 A	NUM-POS=5; STEP-SIZE=0.9; SEGMENT=BOTH; CENTER=FLUX-W T-FLR		5 Secs (5 Secs) [==>]	[1]
	3	TIME-TAG- G185M-195 3-FP-1 (COS.sp.136 6492)	(1) QSO-B1553+113	COS/NUV, TIME-TAG, PSA	G185M 1953 A	FP-POS=1; BUFFER-TIME=15 61		455 Secs (455 Secs) [==>]	[1]
	4	TIME-TAG- G185M-195 3-FP-2 (COS.sp.136 6492)	(1) QSO-B1553+113	COS/NUV, TIME-TAG, PSA	G185M 1953 A	FP-POS=2; BUFFER-TIME=15 61		455 Secs (455 Secs) [==>]	[1]
	5	TIME-TAG- G185M-195 3-FP-3 (COS.sp.136 6492)	(1) QSO-B1553+113	COS/NUV, TIME-TAG, PSA	G185M 1953 A	FP-POS=3; BUFFER-TIME=15 61		454 Secs (454 Secs) [==>]	[1]
	6	TIME-TAG- G185M-195 3-FP-4 (COS.sp.136 6492)	(1) QSO-B1553+113	COS/NUV, TIME-TAG, PSA	G185M 1953 A	FP-POS=4; BUFFER-TIME=15 61		454 Secs (454 Secs) [==>]	[1]
	7	TIME-TAG- G185M-194 1-FP-1 (COS.sp.136 6492)	(1) QSO-B1553+113	COS/NUV, TIME-TAG, PSA	G185M 1941 A	FP-POS=1; BUFFER-TIME=15 61		633 Secs (633 Secs) [==>]	[2]
	8	TIME-TAG- G185M-194 1-FP-2 (COS.sp.136 6492)	(1) QSO-B1553+113	COS/NUV, TIME-TAG, PSA	G185M 1941 A	FP-POS=2; BUFFER-TIME=15 61		632 Secs (632 Secs) [==>]	[2]
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Proposal 15835 - QSO-B1553+113-VISIT-52 (52) - Near-Ultraviolet Follow-up of the X-ray-detected Warm-Hot Intergalactic Medium T...

Thu Jul 30 19:00:17 GMT 2020

Visit	Proposal 15835, QSO-B1553+113-VISIT-52 (52) Diagnostic Status: No Diagnostics Scientific Instruments: COS/FUV, COS/NUV Special Requirements: (none)									
	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous				
Fixed Targets	(1)	QSO-B1553+113	RA: 15 55 43.0440 (238.9293500d) Dec: +11 11 24.37 (11.19010d) Equinox: J2000		V=14.57 G=14, FUV=16.3, NUV=15.6, F(1300) = 1.5e-14 from HST/C OS	Reference Frame: ICRS				
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