



15864 - Is the Type Ic supernova progenitor WS35 a white dwarf merger product?

Cycle: 27, Proposal Category: GO

(UV Initiative)

(Availability Mode: SUPPORTED)

INVESTIGATORS

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VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) WS35	STIS/CCD STIS/FUV-MAMA STIS/NUV-MAMA	5	13-Feb-2020 13:00:17.0	yes
02	(1) WS35	STIS/CCD STIS/FUV-MAMA STIS/NUV-MAMA	5	13-Feb-2020 13:00:20.0	yes

10 Total Orbits Used

ABSTRACT

We propose to obtain UV spectroscopy of the newly discovered emission-line star WS35 with HST/STIS, which was found in the center of a hydrogen-free infrared nebula. Its pure C-O-Ne surface composition, high surface temperature of 200,000 K, and luminosity of 40,000 L_{sun} imply it is an immediate progenitor of a Type Ic supernova. Optical spectra show an unprecedented, supernova-like velocity of its wind, 16,000 km/s, whose kinetic energy exceeds the photon energy implied by WS35's luminosity. This argues for the wind acceleration being aided by a strong magnetic field and rapid rotation. It has thus been suggested that WS35 is the result of a recent white dwarf merger exceeding the Chandrasekhar mass, caught during its brief evolution from merging to collapse. We aim to test this scenario using FUV/NUV spectra of WS35, which allow to assess the velocity profile of its wind, and thereby provide a clear diagnostic of the wind acceleration mechanism. The proposed STIS UV spectroscopy will further provide unprecedented information on the surface chemical composition that is not accessible from optical spectra, in particular on the so far unconstrained neon abundance. Our results will shed new light on the formation of Type Ic supernovae, constrain the origin of Type Ia supernovae in the double-degenerate scenario, and test the idea of generating strong magnetic fields in stellar mergers.

OBSERVING DESCRIPTION

We obtained 10 HST orbits to observe WS35 with STIS long-slit spectroscopy in low spectral resolution with gratings G140L in the FUV and G230L in the NUV. We wish to obtain data for two epochs to constrain a possible binary nature of the object. We further use time-tag mode to get information on short-term variability, which could be of the order of 5min.

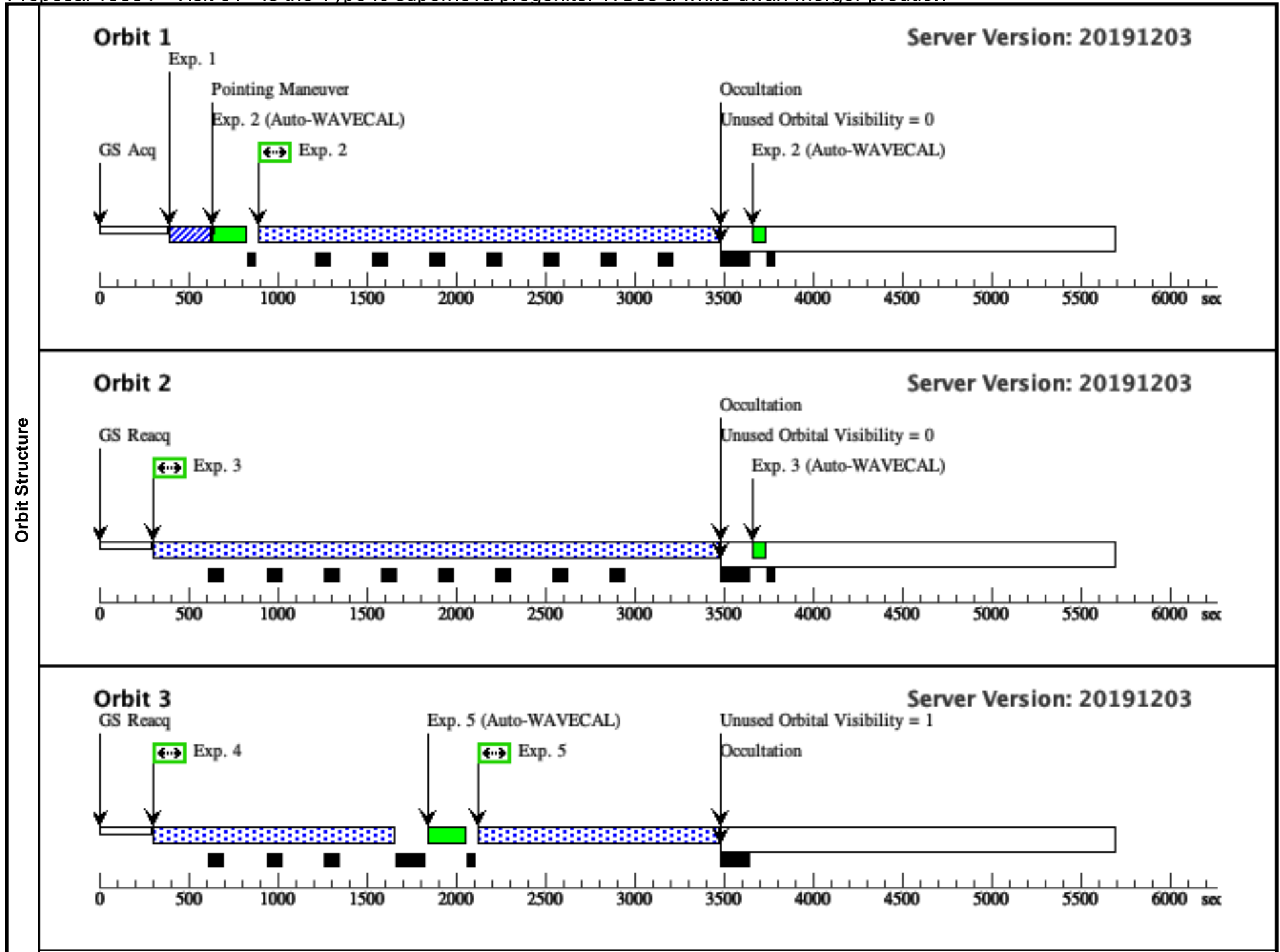
We split our observing time in two visits with 5 orbits each. With an integration time of 7051s per visit in the NUV and 7645s per visit in the FUV we obtain a minimum S/N of 20 per visit in the relevant wavelength ranges (~1600-3000Å in the NUV, and ~1250-1600Å in the FUV, cf. ETC runs STIS.sp.1368104 and STIS.sp.1368105). For time-tag mode we chose buffer times that are roughly a factor three lower than the ETC recommendations to be safe against uncertainties in the source flux and the interstellar extinction (which is included in our uploaded model spectrum). We adopt the standard 52"x0.2" slit.

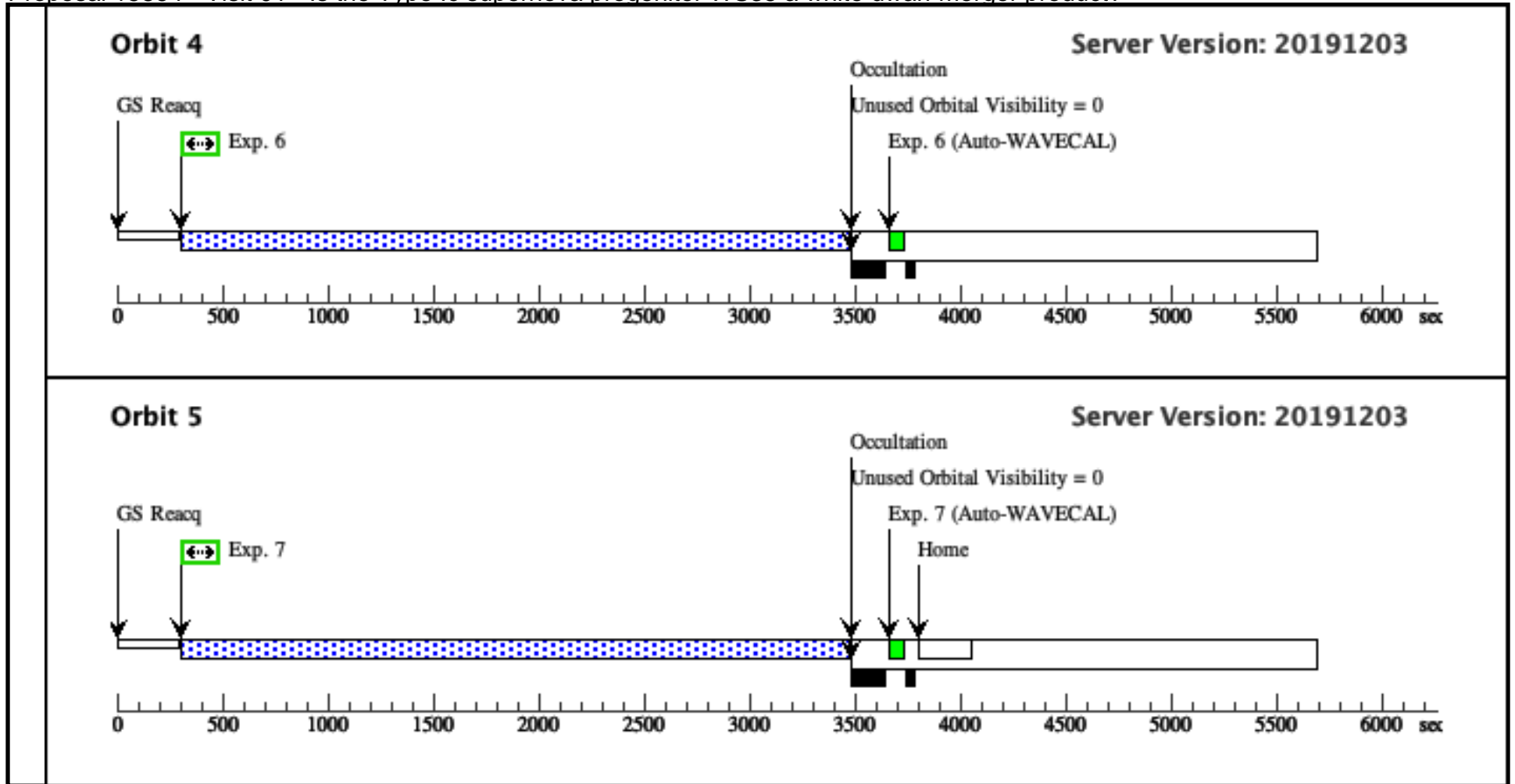
For target acquisition we use a standard point source acquisition.

Proposal 15864 - Visit 01 - Is the Type Ic supernova progenitor WS35 a white dwarf merger product?

Thu Feb 13 18:00:21 GMT 2020

Visit	Proposal 15864, Visit 01, implementation Diagnostic Status: No Diagnostics Scientific Instruments: STIS/NUV-MAMA, STIS/CCD, STIS/FUV-MAMA Special Requirements: AFTER 06-SEP-2020:00:00:00																											
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Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit																		
	1	ACQ (STIS.ta.136 5087)	(1) WS35	STIS/CCD, ACQ, F28X50LP	MIRROR	ACQTYPE=POINT			1 Secs (1 Secs) [==>]	[1]																		
	2	STIS/NUV (STIS.sp.13 68104)	(1) WS35	STIS/NUV-MAMA, TIME-TAG, 52X0.2	G230L 2376 A	BUFFER-TIME=32 0			1000 Secs (2562 Secs) [==>2562.0 Secs]	[1]																		
	3	STIS/NUV (STIS.sp.13 68104)	(1) WS35	STIS/NUV-MAMA, TIME-TAG, 52X0.2	G230L 2376 A	BUFFER-TIME=32 0			1000 Secs (3156 Secs) [==>3156.0 Secs]	[2]																		
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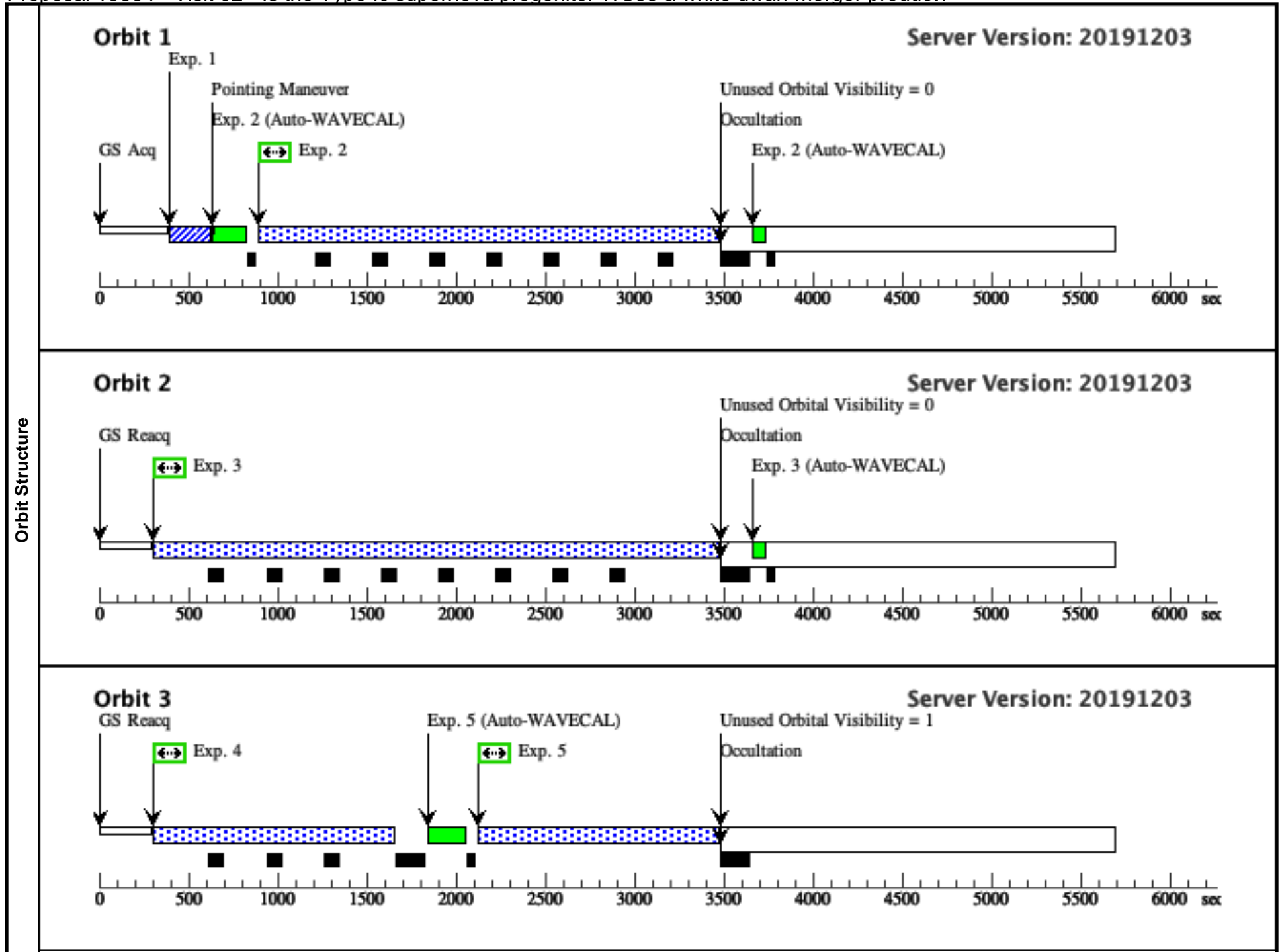




Proposal 15864 - Visit 02 - Is the Type Ic supernova progenitor WS35 a white dwarf merger product?

Thu Feb 13 18:00:22 GMT 2020

Visit	Proposal 15864, Visit 02, implementation Diagnostic Status: No Diagnostics Scientific Instruments: STIS/NUV-MAMA, STIS/CCD, STIS/FUV-MAMA Special Requirements: AFTER 06-SEP-2020:00:00:00																											
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Orbit Structure

