



15884 - Confirmation of an Astrometrically Detected Exoplanet Candidate Orbiting the closest Brown Dwarfs

Cycle: 27, Proposal Category: GO
(Availability Mode: SUPPORTED)

INVESTIGATORS

<i>Name</i>	<i>Institution</i>	<i>E-Mail</i>
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Dr. Mattia Libralato (CoI) (AdminUSPI)	Space Telescope Science Institute - ESA - JWST	libra@stsci.edu

VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) LUHMAN-16-SFT	WFC3/UVIS	1	11-Oct-2021 19:00:16.0	yes
02	(3) LUHMAN16AB-SFT	WFC3/UVIS	1	11-Oct-2021 19:00:17.0	yes

2 Total Orbits Used

ABSTRACT

At a distance of slightly less than 2 pc WISE 1049-5319 (or Luhman 16 AB) is the closest system to Earth after those of alpha Centauri and Barnard's Star.

The analysis of astrometric measurements collected in 12 HST epochs (between 2014.64 and 2016.76) revealed a weak astrometric signature indicating the presence of an exoplanet in the Luh16AB system. This candidate was reported in 2017. A single additional epoch ---collected in August 2018--- and re-analysis of the data with more advanced methods increased the significance level of the planet candidate.

Here we propose to use HST in the same observing mode to collect data at two new epochs to confirm or disprove the astrometric signature that suggest the presence of a sub-Neptune-mass exoplanet orbiting one component of the closest known brown dwarf system. If confirmed, it would be the first astrometrically detected exoplanet around a brown dwarf and the third closest known exoplanet; a major discovery which only requires a relatively modest investment of two additional HST orbits.

The proposed epochs will extend the temporal coverage of the astrometric monitoring to two planet orbital periods (based on the current orbit fits). Coverage over two periods is required to: (a) confirm that our model of the planetary system is predictive, and (b) verify and refine uncertainty estimates. Furthermore, should the orbit be non-circular, these additional measurements will determine this and significantly improve the orbital semi-major axis measurement.

OBSERVING DESCRIPTION

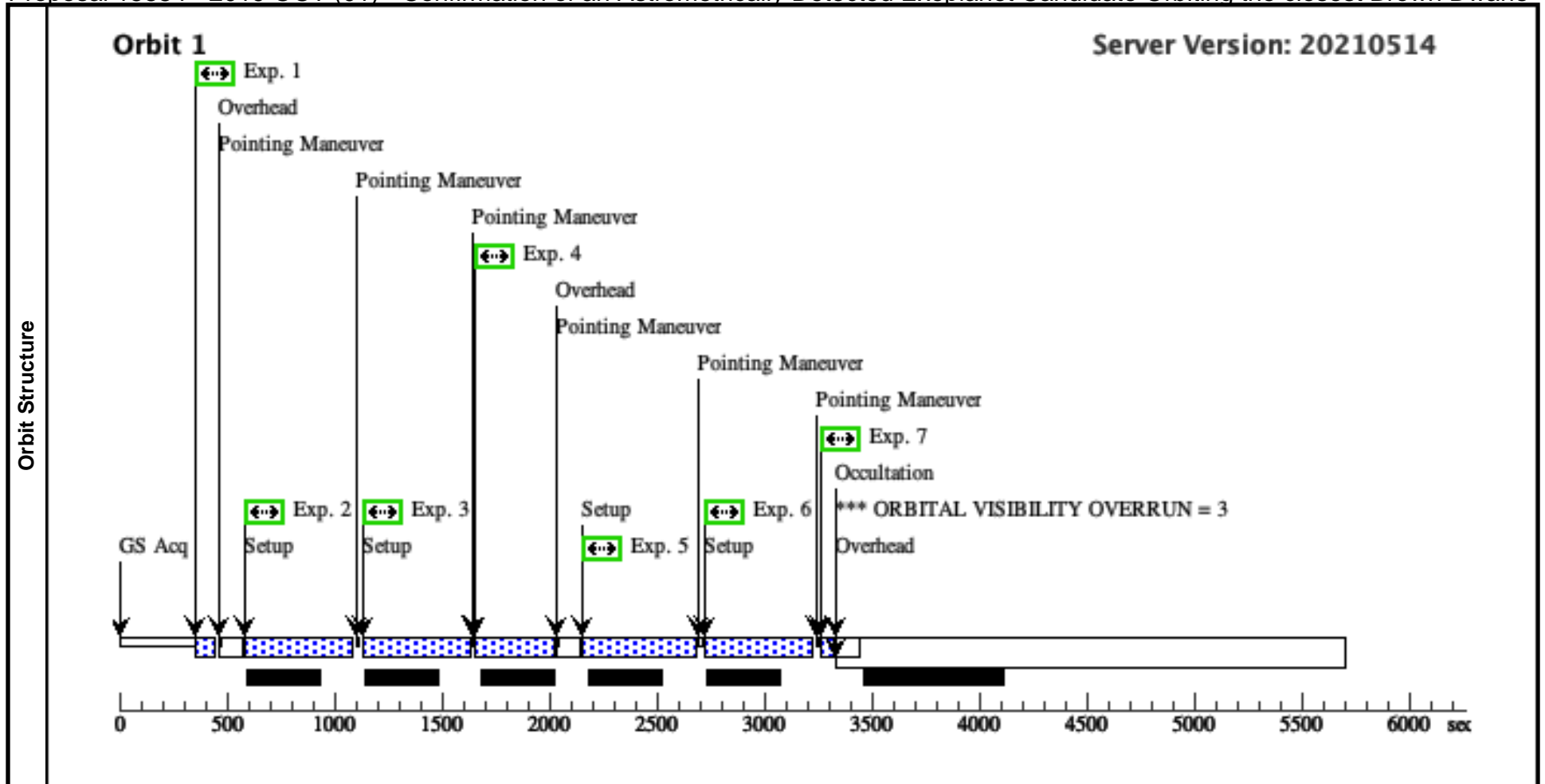
Each epoch contains three images in point source imaging mode, and four in spatial-scanning mode with WFC3/UVIS.

In addition to the images taken in spatial scan mode and to the short two point-source images at the beginning and at the end of the single orbit-visits --- as requested at the phase I stage --- it turned out possible to fit in the orbit one extra point source images of 348s in F606W. The color information will be a precious to discriminate the nature (and so the distance) of the reference sources in the field (for example to separate slow close by red dwarf from far away giants)

Proposal 15884 - 2019 OCT (01) - Confirmation of an Astrometrically Detected Exoplanet Candidate Orbiting the closest Brown Dwarfs

Mon Oct 11 23:00:18 GMT 2021

Visit	<p>Proposal 15884, 2019 OCT (01), completed</p> <p>Diagnostic Status: Warning</p> <p>Scientific Instruments: WFC3/UVIS</p> <p>Special Requirements: ORIENT 225D TO 260 D; BETWEEN 19-OCT-2019 AND 25-OCT-2019</p> <p><i>Comments: Epoch 2019 OCT. This is the epoch is taken at the ricciolo of the cycloid apparent motion. The epoch must be around 2019-10-20 +/- 15 days</i></p>																																																																																				
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Proposal 15884 - 2021 NOV (02) - Confirmation of an Astrometrically Detected Exoplanet Candidate Orbiting the closest Brown Dwarfs

Mon Oct 11 23:00:18 GMT 2021

Visit	Proposal 15884, 2021 NOV (02), implementation Diagnostic Status: Warning Scientific Instruments: WFC3/UVIS Special Requirements: ORIENT 345D TO 35 D; AFTER 01-NOV-2021:00:00:00 <i>Comments: Epoch 2021 NOV</i> <i>This is the epoch is taken at the ricciolo of the cycloid apparent motion.</i> <i>The epoch must be after 2021-11-01</i>																																																																																				
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