



## 16253 - Cycles and the Seven Dwarfs

Cycle: 28, Proposal Category: GO

(Availability Mode: SUPPORTED)

### INVESTIGATORS

<i>Name</i>	<i>Institution</i>	<i>E-Mail</i>
<b>Dr. Thomas R. Ayres (PI) (Contact)</b>	<b>University of Colorado at Boulder</b>	<b>thomas.ayres@colorado.edu</b>

### VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
10	(1) HD-131156A (2) HD-131156B	COS/FUV COS/NUV	2	10-Dec-2020 14:01:43.0	yes
11	(3) HD-165341A (4) HD-165341B	COS/FUV COS/NUV	2	10-Dec-2020 14:01:46.0	yes

4 Total Orbits Used

### ABSTRACT

Somewhat erratic, unpredictable long-term starspot cycles of late-type stars are relevant for understanding the Sun's analogous high-energy modulations, in turn crucial to Earth-impacting "Space Weather" (SW). Similarly, stellar SW counterparts can have equivalent, or more severe, adverse effects on their exoplanets. The solar decadal oscillations are symptomatic of a deep-seated magnetic pump -- the Dynamo -- whose internal workings remain elusive. Key question: is Sun's 11-year cycle normal, or instead a transition state? A 3-year addition to earlier long-term joint Chandra/HST programs (previously involving Alpha Cen [G2V+K1V] and Procyon [F5IV]), will be carried out for two new targets, the nearby binaries Xi Bootis (G8V+K4) and 70 Ophiuchi (K0V+K5V).

The HST part of the program will measure FUV subcoronal (T~100,000 K) emission lines, like Si IV 140 nm, C IV 155 nm, and density-sensitive O

IV 140 nm, as well as important chromospheric ( $T \sim 10,000$  K) features, like O I 130 nm, C II 133 nm, and C I 165 nm, in both pairs of stars, with one COS visit per system each year. Analyses of these emissions, including dynamical information encoded in Doppler shifts and distortions of the line shapes, can provide insight concerning the thermal/pressure structure of the anomalously hot outer layers of these stars; and how these structural properties change with the ebb and flow of a magnetic activity cycle.

## **OBSERVING DESCRIPTION**

### **\*\*Impact of Reduced Gyro Mode\*\***

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The two targets, Xi Boo and 70 Oph, would be minimally affected by RGM, because there are no explicit time or orientation constraints on the observations: they can be carried out anytime during Cycle 28 when the targets are available. Further, the Visits are only 2 orbits, minimizing the possibility of re-acquisition failures. However, if target acquisition times are increased by RGM, then one or more of the exposures would have to be shortened. This can be done with minimal loss of science.  
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### Phase II Observing Description (for normal 3-gyro operations)

Observational goal is to obtain FUV moderate-resolution spectra of both components of Xi Boo and 70 Oph using Cosmic Origins Spectrograph (COS). The primaries, Xi Boo A and 70 Oph A, are nearly identical in FUV flux levels, about a tenth of well-studied (by STIS) Alpha Cen A. Normally, high-sensitivity COS would be off-limits for these primaries, because their strong Lyman-alpha features would violate FUV detector bright limits. However, the COS G130M-1222 setting places Lyman-alpha in the gap between detector sides A and B, so can be used for cool stars, as long as their other strong emissions (e.g., C II 133 nm) are "safe" (as is the case for Xi Boo A and 70 Oph A). When combined with COS G160M-1533, the spectral range 107-171 nm can be covered, with only two small gaps: 120.7-122.3 nm (Lyman-alpha) and 151.5-153.3 nm (devoid of important lines). COS's 15 km/s resolution is fully adequate for the relatively broad FUV emission features of these cool stars, especially the "hot" lines like Si III 120.6 nm ( $T \sim 30,000$  K), Si IV 139.3 nm (80,000 K), and C IV 154.8 nm (100,000 K). The COS FUV channel is so sensitive that a mere 1000 s exposure obtains  $S/N \sim 100$  (line peak) at Si III, 80 at Si IV, and 65 at C IV, for both primary stars. Each grating setting requires 4 "FP-Splits" to mitigate fixed pattern noise.

The pair of FUV settings plus FP-SPLITS can be done in a single orbit for the primary of each system, including the initial Guide Star and NUV

Proposal 16253 (STScI Edit Number: 4, Created: Thursday, December 10, 2020 at 2:01:47 PM Eastern Standard Time) - Overview  
imaging target acquisitions.

At the beginning of the second orbit, the secondary star would be captured by a small offset maneuver and NUV imaging acquisition. The same dual-setting FUV sequence then would be carried out, extending over the remaining unocculted part of the second orbit.

The COS visits would be once a year for each system (4 HST orbits per year, total): experience with Alpha Cen AB has shown that the FUV variability, while significant, is not as pronounced as that of the X-rays (which are observed twice a year for each system). Helpfully, key coronal forbidden lines Fe XII 134.9 nm (T~2 MK) and Fe XXI 135.4 nm (10 MK) are available in the FUV to assess the stellar coronal states at the times of the COS pointings. For this reason, the HST and Chandra observations need not be simultaneous. In fact, it is better to have them separated, since the FUV coronal forbidden lines would then provide an additional point on the high-energy timeline. Given the very deep COS exposures proposed here (especially at 134.9 nm and 135.4 nm, which receive double coverage in the 2 nm overlap between G130M and G160M), these coronal forbidden lines should be easily detectable in the primary stars (based on previous STIS experience), and perhaps even in the UV-fainter secondaries.

# Proposal 16253 - Visit 10 - Cycles and the Seven Dwarfs

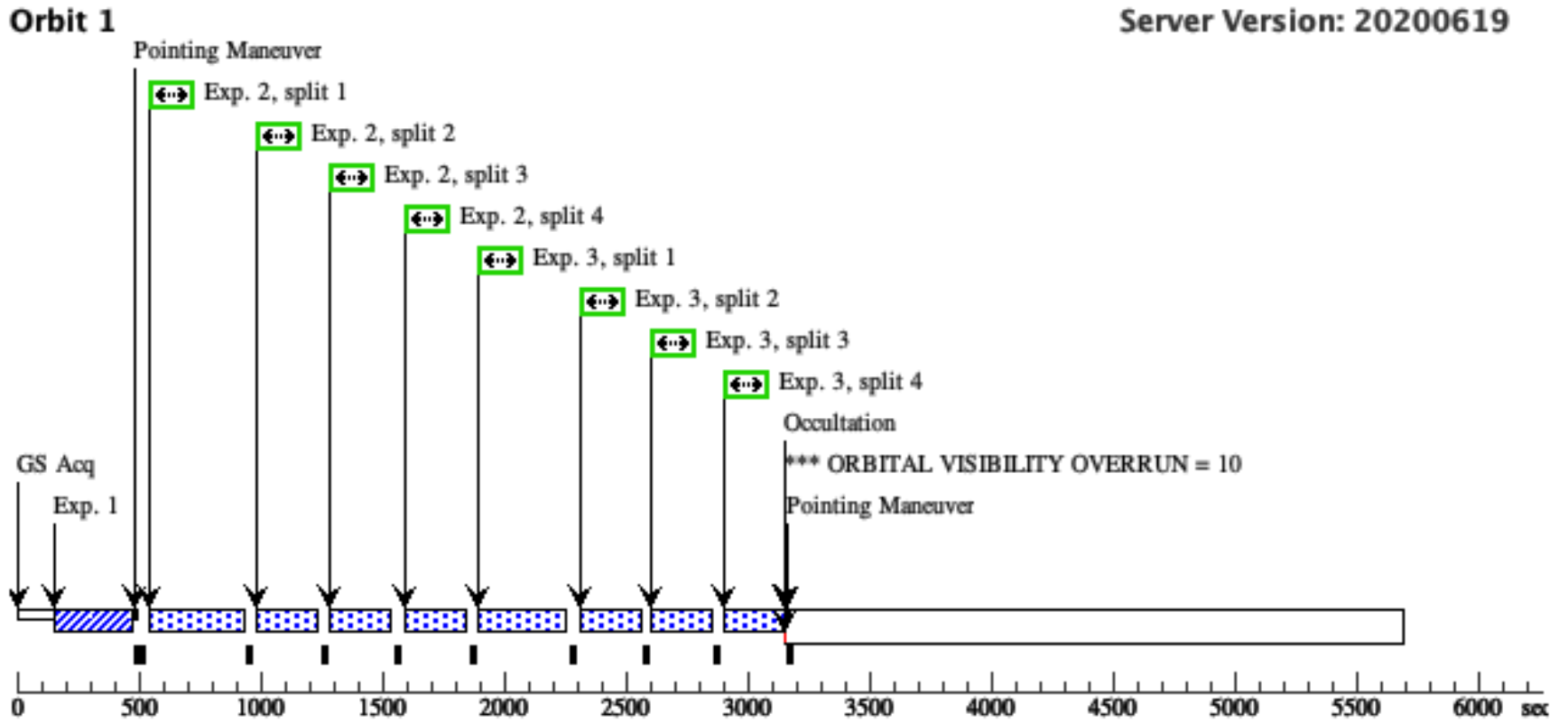
Thu Dec 10 19:01:47 GMT 2020

<b>Visit</b>	<b>Proposal 16253, Visit 10, implementation</b> <b>Diagnostic Status: Warning</b> Scientific Instruments: COS/FUV, COS/NUV Special Requirements: ORIENT 131D TO 11 D																																		
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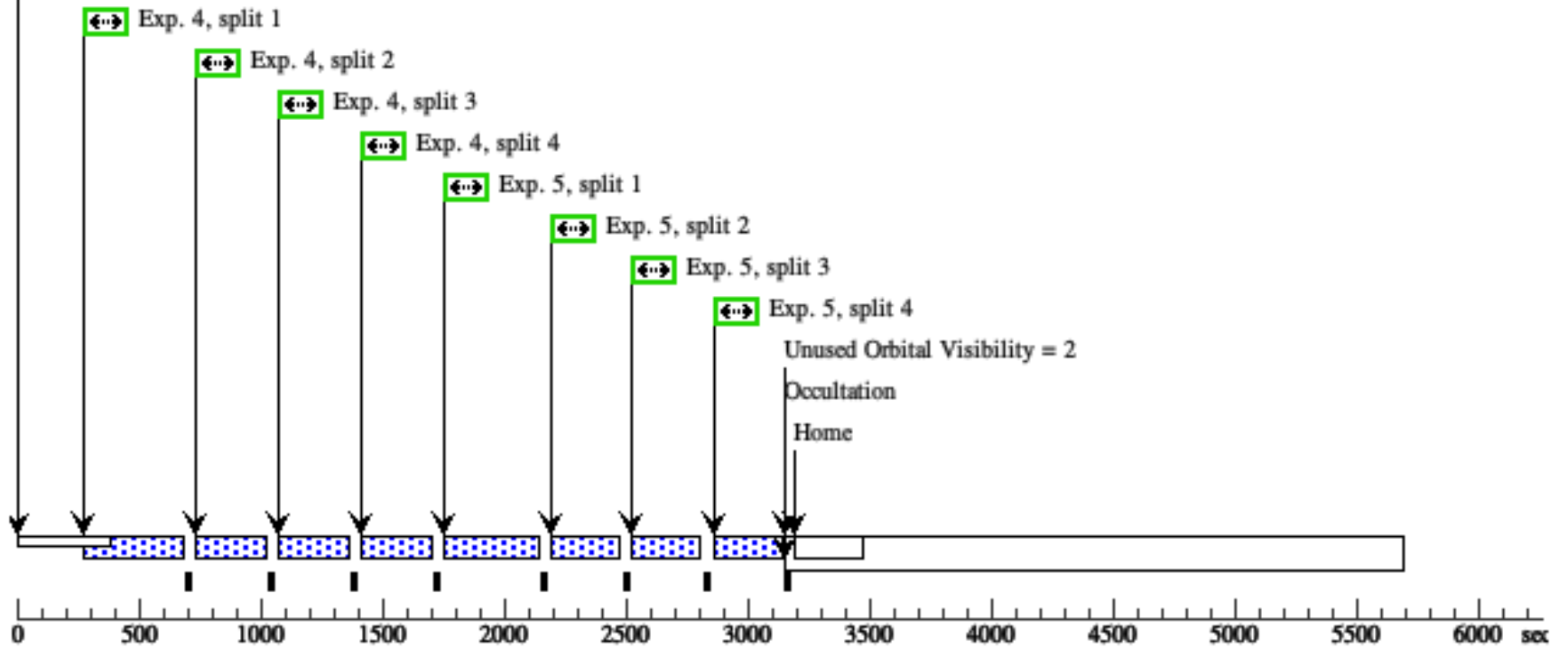
Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
	1	(COS.ta.136 7079)	(1) HD-131156A	COS/NUV, ACQ/IMAGE, BOA	MIRRORB			GS ACQ SCENARI O BASE1BE		10 Secs (10 Secs) [==>]	[1]
	<i>Comments: Castelli-Kurucz Models G8V, renormalized to vegamag = 4.68 in filter Johnson/V</i>										
	2	(COS.sp.136 7281)	(1) HD-131156A	COS/FUV, TIME-TAG, PSA	G130M 1222 A		BUFFER-TIME=24 00; FP-POS=ALL			200 Secs (800 Secs) [==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[1]
	<i>Comments: h_hd131156_uvsum_1m_51163_etc.dat from StarCAT (STIS E140M-1425)</i>										
	3	(COS.sp.136 7282)	(1) HD-131156A	COS/FUV, TIME-TAG, PSA	G160M 1533 A		BUFFER-TIME=40 00; FP-POS=ALL			193 Secs (772 Secs) [==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[1]
<i>Comments: h_hd131156_uvsum_1m_51163_etc.dat from StarCAT (STIS E140M-1425)</i>											
4	(COS.sp.136 7319)	(2) HD-131156B	COS/FUV, TIME-TAG, PSA	G130M 1222 A		BUFFER-TIME=93 00; FP-POS=ALL			235 Secs (940 Secs) [==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[2]	
<i>Comments: B fluxes scaled by 0.14 from A [h_hd131156_uvsum_1m_51163_etc.dat from StarCAT (STIS E140M-1425)]</i>											
5	(COS.sp.136 7320)	(2) HD-131156B	COS/FUV, TIME-TAG, PSA	G160M 1533 A		BUFFER-TIME=13 400; FP-POS=ALL			227 Secs (908 Secs) [==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[2]	
<i>Comments: B fluxes scaled by 0.14 from A [h_hd131156_uvsum_1m_51163_etc.dat from StarCAT (STIS E140M-1425)]</i>											

Orbit Structure



### Orbit 2

GS Reacq



# Proposal 16253 - Visit 11 - Cycles and the Seven Dwarfs

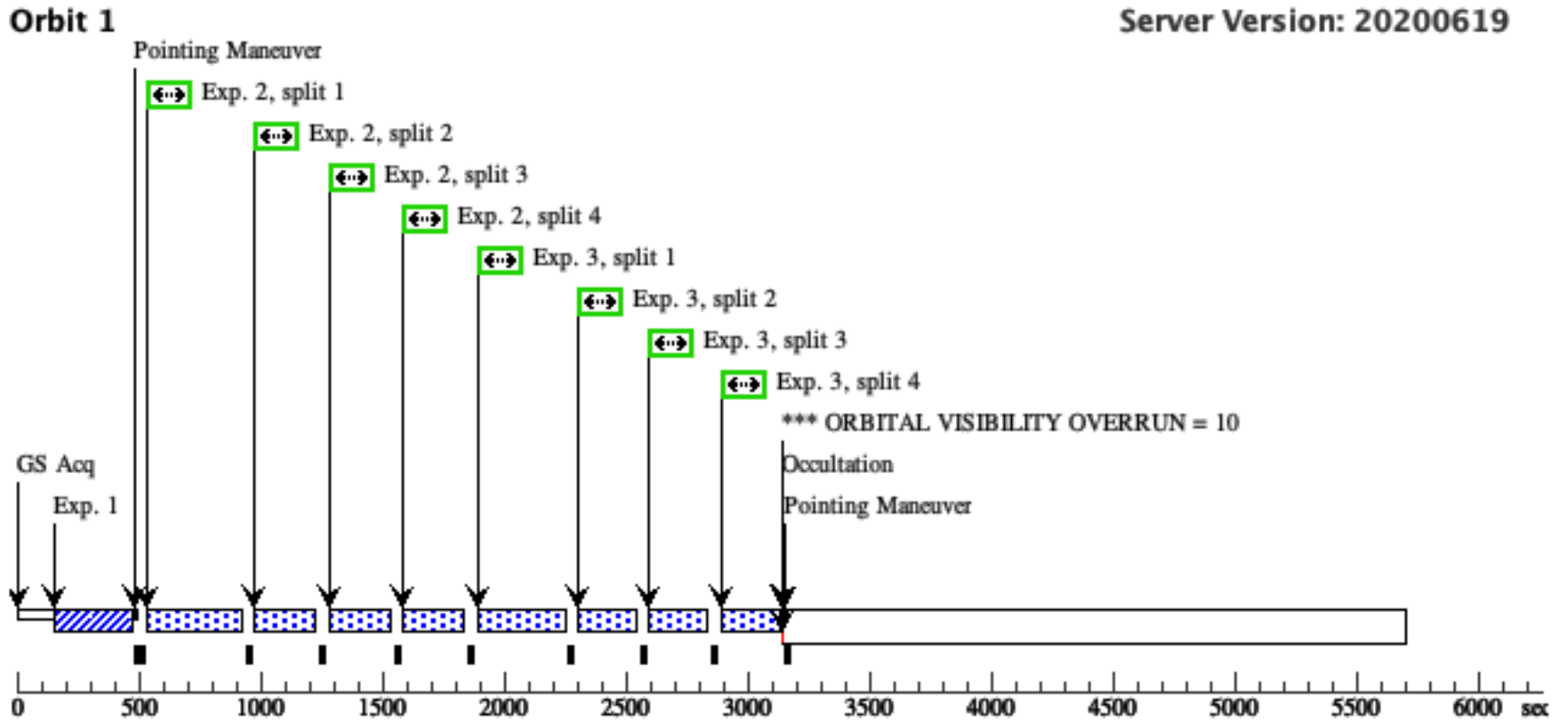
Thu Dec 10 19:01:47 GMT 2020

<b>Visit</b>	<b>Proposal 16253, Visit 11, completed</b> <b>Diagnostic Status: Warning</b> Scientific Instruments: COS/FUV, COS/NUV Special Requirements: ORIENT 316D TO 196 D																																		
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	1	(COS.ta.136 7302)	(3) HD-165341A	COS/NUV, ACQ/IMAGE, BOA	MIRRORB			GS ACQ SCENARI O BASE1BE		7 Secs (7 Secs) [==>]	[1]
	<i>Comments: Castelli-Kurucz Models K0V, renormalized to vegamag = 4.22 in filter Johnson/V (from BSC and WDS)</i>										
	2	(COS.sp.136 7312)	(3) HD-165341A	COS/FUV, TIME-TAG, PSA	G130M 1222 A	BUFFER-TIME=28 00; FP-POS=ALL				200 Secs (800 Secs) [==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[1]
	3	(COS.sp.136 7313)	(3) HD-165341A	COS/FUV, TIME-TAG, PSA	G160M 1533 A	BUFFER-TIME=43 00; FP-POS=ALL				191 Secs (764 Secs) [==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[1]
4	(COS.sp.136 7325)	(4) HD-165341B	COS/FUV, TIME-TAG, PSA	G130M 1222 A	BUFFER-TIME=92 00; FP-POS=ALL				235 Secs (940 Secs) [==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[2]	
5	(COS.sp.136 7326)	(4) HD-165341B	COS/FUV, TIME-TAG, PSA	G160M 1533 A	BUFFER-TIME=12 700; FP-POS=ALL				224 Secs (896 Secs) [==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[2]	

Orbit Structure



**Orbit 2**

GS Reacq

