



# 17107 - A very low mass, highly irradiated survivor of the common envelope phase

Cycle: 30, Proposal Category: GO

(Availability Mode: SUPPORTED)

## INVESTIGATORS

<i>Name</i>	<i>Institution</i>
<b>Prof. Boris T. Gaensicke (PI) (ESA Member) (Contact)</b>	<b>The University of Warwick</b>
Prof. Tom R. Marsh (CoI) (ESA Member)	The University of Warwick
Dr. Nicole Reindl (CoI) (ESA Member)	Universitat Potsdam
Dr. Veronika Schaffenroth (CoI) (ESA Member)	Thüringer Landessternwarte Tautenburg (TLS)

## VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) WDJ200328.97+083847.22	STIS/CCD STIS/FUV-MAMA	3	14-Jun-2024 12:00:16.0	yes
02	(1) WDJ200328.97+083847.22	STIS/CCD STIS/FUV-MAMA	3	14-Jun-2024 12:00:17.0	yes
52	(1) WDJ200328.97+083847.22	STIS/CCD STIS/FUV-MAMA	3	14-Jun-2024 12:00:18.0	yes

9 Total Orbits Used

## ABSTRACT

The common envelope (CE) phase is a key driver in the evolution of close binary stars, causing their periods to shrink dramatically, or the two stars to merge altogether. An important parameter for the survival of many systems, including planets enveloped by host giant branch stars, is the minimum companion mass that can survive a CE. No certain survivor of a CE has been found with a mass less than 50 Jupiter masses ( $M_J$ ), well above the lowest theoretical prediction of 5  $M_J$ . With the theory itself uncertain, it is important to establish firmer observational limits. A recently

discovered hot white dwarf / brown dwarf binary with a period of 2.4 hours promises to do just this. Optical data of this system indicate a mass for the brown dwarf lying between 11 and 44  $M_J$ . Better constraints require a measurement of the temperature of the hot white dwarf, which translates to a better mass for the brown dwarf through a combination of distance, model atmosphere, gravitational red-shift and dynamical constraints. Preliminary indications from optical spectra suggest a white dwarf temperature of order 100,000 K and a brown dwarf mass of order 20  $M_J$ , substantially reducing the current gap between theory and observation. However, optical spectra have relatively few and weak features compared to the FUV, and suffer significant contamination from light emitted by the heated face of the brown dwarf in this compact system. The aim of this proposal is to acquire FUV spectra to measure the white dwarf's effective temperature and to obtain improved values for the dynamical and red-shift constraints upon which the mass estimates hinge.

### **OBSERVING DESCRIPTION**

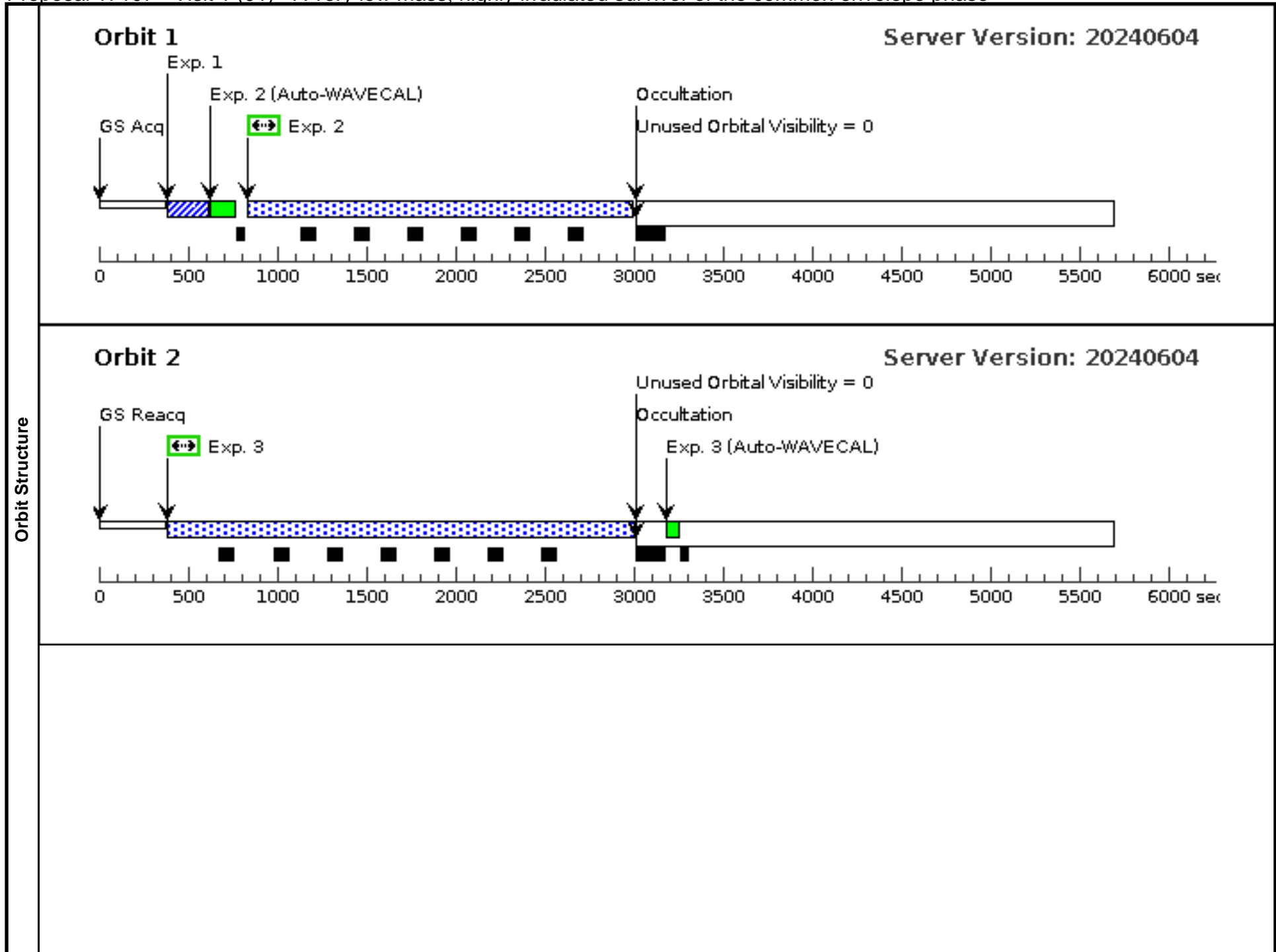
The goal of this proposal is to obtain far-ultraviolet, orbital phase-resolved, high-resolution spectroscopy of WD 2003+0838 to measure the effective temperature and orbital velocity of the white dwarf, in order to determine its mass and hence the mass of its brown dwarf companion, and to look for evidence of mass loss and/or accretion in the system and circumstellar material.

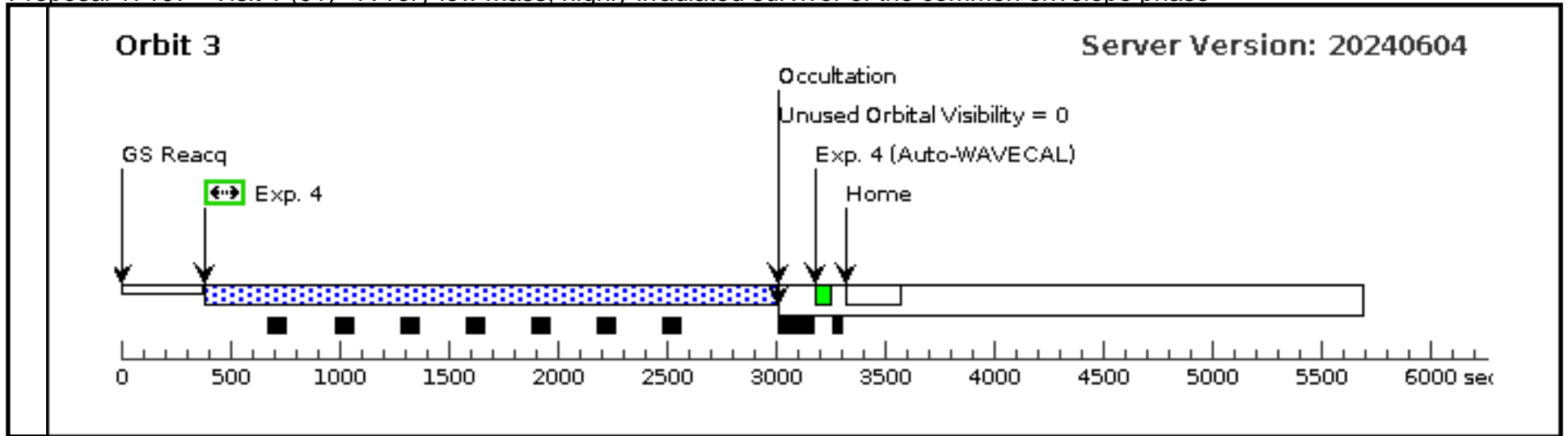
The observations will be split into 2 x 3 orbits, following standard STIS procedures: target acquisition with the F28X50LP long-pass filter, followed by E140M time-tag spectroscopy (such that we can bin the data into arbitrary phase bins during the analysis) using the 0.2"x0.2" aperture. The two visits can be scheduled totally independently.

Proposal 17107 - Visit 1 (01) - A very low mass, highly irradiated survivor of the common envelope phase

Fri Jun 14 16:00:19 GMT 2024

Visit	<b>Proposal 17107, Visit 1 (01), implementation</b> <b>Diagnostic Status: No Diagnostics</b> Scientific Instruments: STIS/CCD, STIS/FUV-MAMA Special Requirements: (none)																																																																																										
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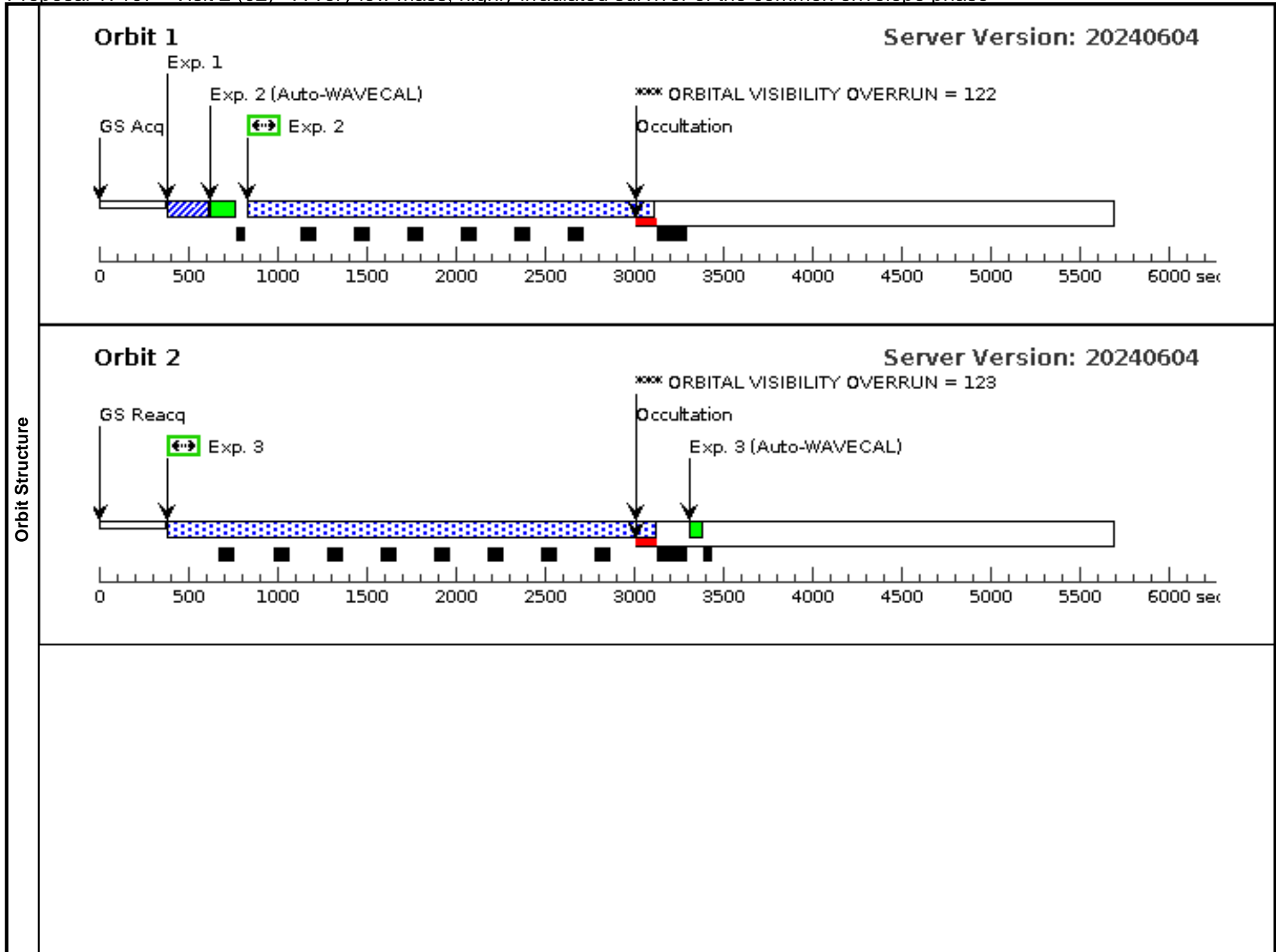


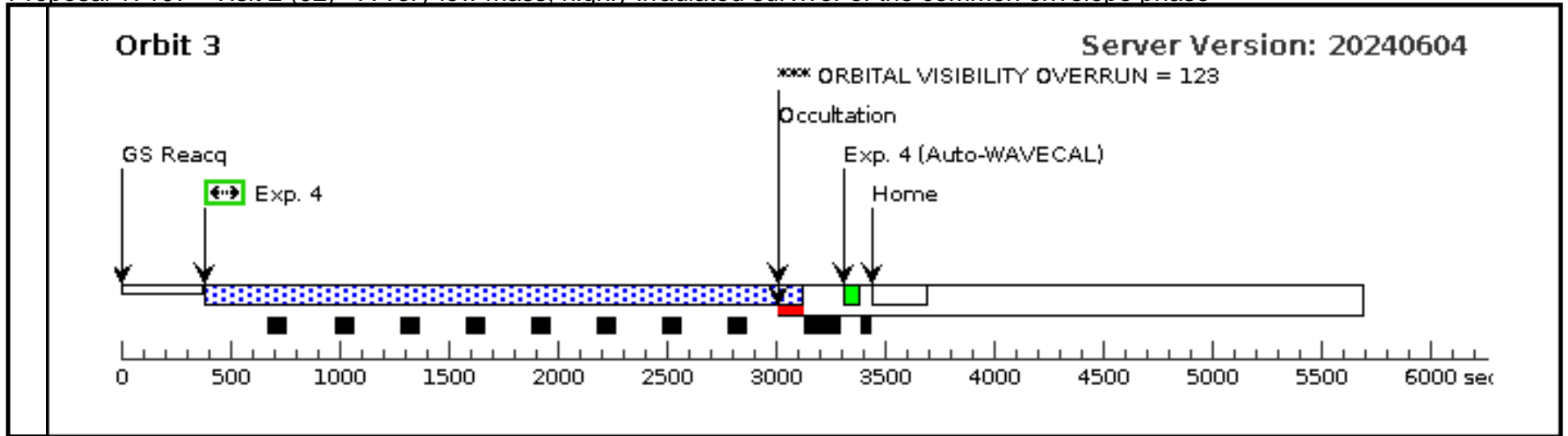


Proposal 17107 - Visit 2 (02) - A very low mass, highly irradiated survivor of the common envelope phase

Fri Jun 14 16:00:19 GMT 2024

<b>Visit</b>	<b>Proposal 17107, Visit 2 (02), failed</b> <b>Diagnostic Status: Warning</b> Scientific Instruments: STIS/CCD, STIS/FUV-MAMA Special Requirements: (none)										
	(Visit 2 (02)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN (Visit 2 (02)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN (Visit 2 (02)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN										
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Proposal 17107 - Visit 52 (52) - A very low mass, highly irradiated survivor of the common envelope phase

Fri Jun 14 16:00:19 GMT 2024

Visit	<b>Proposal 17107, Visit 52 (52), implementation</b> <b>Diagnostic Status: No Diagnostics</b> Scientific Instruments: STIS/CCD, STIS/FUV-MAMA Special Requirements: (none) Comments: HOPR repeat of visit 02.																					
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