



17401 - Accreting white dwarfs as probes of compact binary evolution

Cycle: 31, Proposal Category: GO

(UV Initiative, Treasury)

(Availability Mode: AVAILABLE)

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Proposal 17401 (STScI Edit Number: 8, Created: Wednesday, November 19, 2025, 12:00:24PM Eastern Standard Time) - Overview

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Dr. Knox S. Long (CoI)	Eureka Scientific Inc.

VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) 2MASS-J23375921+4308509	COS/FUV	2	19-Nov-2025 12:00:16.0	yes
02	(2) 1RXS-J022917.1-395851	COS/FUV	2	19-Nov-2025 12:00:17.0	yes
03	(3) GD-1662	COS/FUV	4	19-Nov-2025 12:00:17.0	yes
04	(4) HS-0506+7725	COS/FUV	3	19-Nov-2025 12:00:18.0	yes
05	(5) V-V794-AQL	COS/FUV	3	19-Nov-2025 12:00:18.0	yes
06	(6) V-V442-OPH	COS/FUV	2	19-Nov-2025 12:00:19.0	yes
38	(38) SDSS-J152509.57+360054.5	COS/FUV COS/NUV	5	19-Nov-2025 12:00:20.0	yes
43	(43) WD-J104019.50-495129.7	COS/FUV COS/NUV	3	19-Nov-2025 12:00:21.0	yes
4G	(43) WD-J104019.50-495129.7	S/C	1	19-Nov-2025 12:00:21.0	yes
4H	(44) WD-J104019.50-4951-SAFE-TARGET	COS/FUV COS/NUV	3	19-Nov-2025 12:00:22.0	yes
20	(20) V-LL-LYR	COS/FUV	3	19-Nov-2025 12:00:23.0	yes
2A	(20) V-LL-LYR	S/C	1	19-Nov-2025 12:00:23.0	yes

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
2B	(21) V-LL-LYR-SAFE-TGT	COS/FUV	3	19-Nov-2025 12:00:23.0	yes

35 Total Orbits Used

ABSTRACT

In the last 20 years, the study of compact interacting binaries has led to two major breakthroughs in astrophysics: the discovery of dark energy and the first detection of gravitational waves. Although binaries are critically important to probe the properties of the Universe and to test fundamental physical theories, our understanding of their evolution and final fate is still far from being complete.

Accreting white dwarfs are ideal laboratories in which to test the models of compact binary evolution. We here propose a COS Treasury program specifically designed to explore those regions of the parameter space that have been previously poorly studied and where major discrepancies between the theory and observations are found. Combining the high-quality ultraviolet data with the parallaxes from Gaia, we will accurately measure effective temperatures, masses and accretion rates for 43 accreting white dwarfs (the minimum number required to homogeneously sample the entire physical parameter space spanned by this diverse population) thereby testing the mechanisms of angular momentum loss which drive the evolution of all kinds of binaries. The white dwarf masses are a key ingredient in the pathway toward Supernova Type Ia explosions and by obtaining their accurate values, we will constrain both the single and the double-detonation scenarios. Finally, only the ultraviolet allows the detection of the nitrogen and carbon resonance lines. From their abundances, we will establish the formation channel for the most compact systems, which will later be used to verify the performance of the space-mission LISA and calibrate the detector for future gravitational wave source discoveries.

OBSERVING DESCRIPTION

The aim of this program is to obtain far-ultraviolet COS/G140L and STIS/E140M spectroscopy of 43 accreting white dwarfs (WDs). While past HST observations have studied in detail systems in the range $70 \text{ min} < P_{\text{orb}} < 150 \text{ min}$, very few observations have been obtained of long- ($P_{\text{orb}} > 150 \text{ min}$) and short- ($P_{\text{orb}} < 60 \text{ min}$) period accreting WDs, which are therefore the targets of this proposal. We have screened the literature for systems that are bright enough to be observed in 3-5 HST orbits each. The set of targets was chosen to ensure a homogeneous sampling across the P_{orb} range and to maximise the chances of detecting the white dwarf in the UV, and avoiding frequently outbursting systems.

The actual observations are relatively straight-forward: target acquisition, depending on the target brightness, will be done with NUV imaging, or in dispersed light using the G140L grating. The coordinates and proper motions are drawn from Gaia EDR3, hence we do not precede the observations with a search pattern. Depending on the target brightness, the COS G140L spectroscopy will extend over 2, 3, 4, 5 or 6 orbits. We will make use of all 4 FP-POS positions.

Our targets can be divided in two sub-classes according to their observational properties:

- 1) Dwarf novae: systems characterised by quiescent states of low-mass transfer rates onto the WD, interrupted by bright disc outbursts, during which the binary typically brightens by 2-5 magnitudes. In order to ensure that HST spectroscopy is obtained during quiescence, when the WD is the dominant emission source and safe to be observed, we will obtain intensive continuous ground-based monitoring up until 24 h prior to the HST observations.
- 2) Nova-likes: systems in the period range $180 \text{ min} < P_{\text{orb}} < 240 \text{ min}$ are typically characterised by high average accretion rates which keep the disc in a stable hot state, equivalent to a permanent outburst. Occasionally the accretion rate drastically drops ("low state"), unveiling the WD. Our sample contains nine systems falling in this category. We will photometrically monitor them from the ground and the HST observations will be triggered via a non-disruptive ToO only once a low state is detected. We expect to detect ~ 3 systems in low state per year and the observations of these targets will be carried out over the over the next three cycles, in order to observe nine of these systems by the end of the program.

Several of the targets are eclipsing (GS Pav, V1239 Her, 2MASS J15445360+2553488, YZ LMi, ZTFJ040749.30-000716.67, ASASSN-14cn, ZTF-J2252-0519). We have included their ephemeris in order for the target acquisition to be scheduled such that it avoids the eclipse (in a similar strategy as that of HST program 12870).

Impact of reduced-gyro operations:

For the dwarf nova targets, we do not expect major impacts on our program in case of reduced-gyro operations. Prolonged acquisition procedures will not significantly affect the SNR we can achieve, while a possible extension of our program into the next cycle will not preclude our science goals, nor affect in any way the feasibility of our observations.

Instead, with time critical observations becoming much more difficult to schedule, reduced-gyro operations could potentially affect the observations of the nova-like targets, for which long-term non-disruptive ToOs are necessary. The additional nova-likes targets we have include can possibility mitigate the limitations related to reduced-gyro operations by offering (i) alternative pointings that can compensate for the limited sky accessibility, (ii) alternative observing windows that could be more easily included in the schedule.

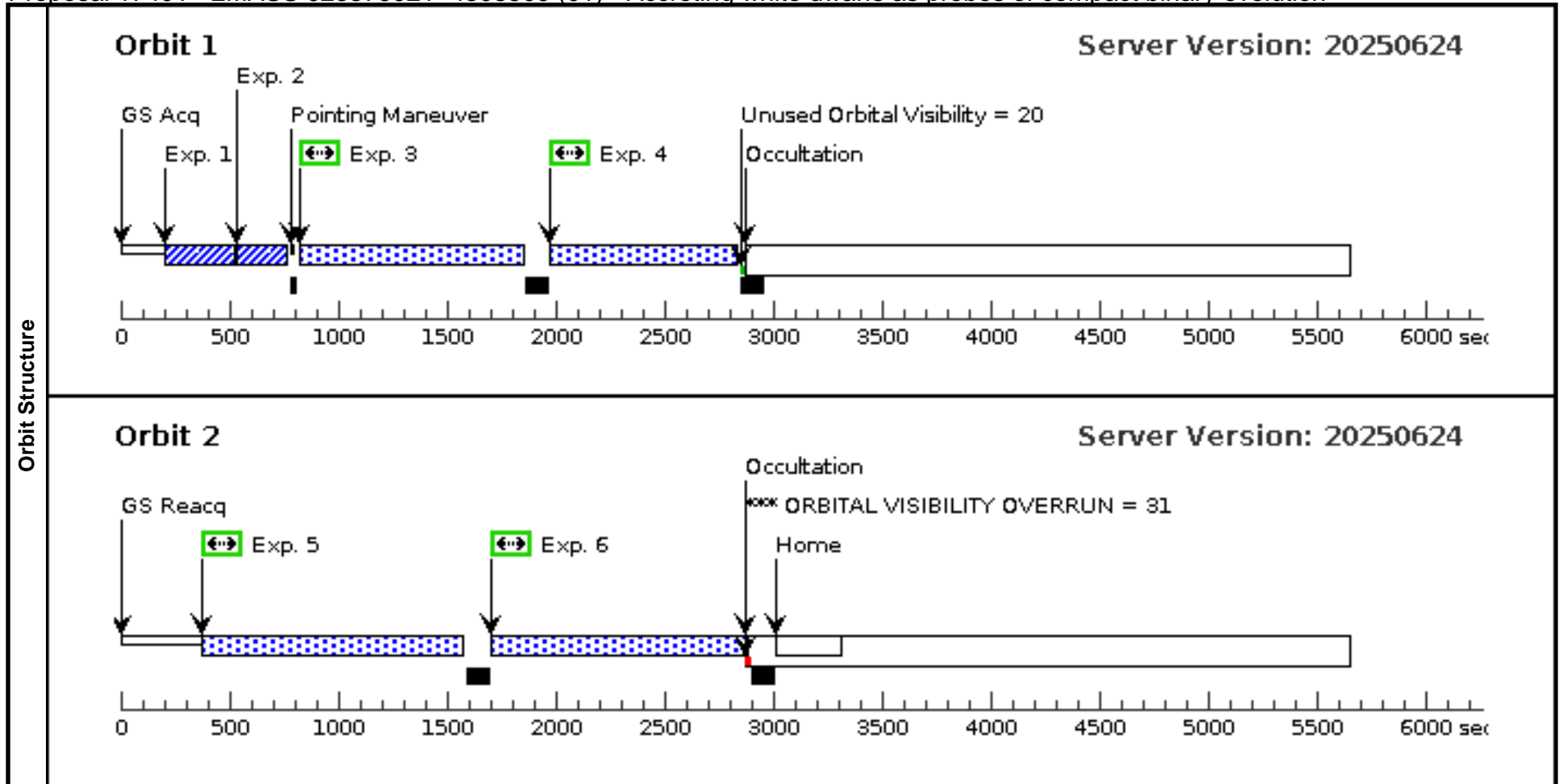
Proposal 17401 - 2MASS-J23375921+4308509 (01) - Accreting white dwarfs as probes of compact binary evolution

Wed Nov 19 17:00:24 GMT 2025

Visit	<p>Proposal 17401, 2MASS-J23375921+4308509 (01), implementation</p> <p>Diagnostic Status: Warning</p> <p>Scientific Instruments: COS/FUV</p> <p>Special Requirements: SCHED 100%; BETWEEN 15-JUN-2021:00:00:00 AND 15-JAN-2022:00:00:00; BETWEEN 15-JUN-2022:00:00:00 AND 15-JAN-2023:00:00:00; BETWEEN 15-JUN-2023:00:00:00 AND 15-JAN-2024:00:00:00; BETWEEN 15-JUN-2024:00:00:00 AND 15-JAN-2025:00:00:00; ON HOLD ; TOO RESPONSE TIME 21.0D</p> <p><i>On Hold Comments: LONG-TERM TOO</i></p> <p><i>This is one of the targets in which the white dwarf is visible only during a low state. We will trigger the ToO when the occurrence of a low state is detected by our ground-based monitoring.</i></p>					
	<p>(2MASS-J23375921+4308509 (01)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p>					
Diagnostics						
Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
	(1)	2MASS-J23375921+4308509	RA: 23 37 59.2215 (354.4967563d) Dec: +43 08 50.89 (43.14747d) Equinox: J2000	Proper Motion RA: 5.05944115231642 mas/yr Proper Motion Dec: -8.709179843108805 mas/yr Parallax: 0.0011740599949903849" Epoch of Position: 2016.0	V=14.31+/-0.08 NUV = 17.49, Gaia EDR3 1925596502097065344, G=16.2, parallax = 1.17(4) mas.	Reference Frame: ICRS
<p><i>Comments:</i> <i>Category=STAR</i> <i>Description=[INTERACTING BINARY, IRREGULAR VARIABLE, NOVA-LIKE]</i> <i>Extended=NO</i></p>						

Proposal 17401 - 2MASS-J23375921+4308509 (01) - Accreting white dwarfs as probes of compact binary evolution

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
Exposures	1	ACQ/PEAK XD (1527541)	(1) 2MASS-J233759 21+4308509	COS/FUV, ACQ/PEAKXD, PSA 1105 A	G140L	LIFETIME-POS=LP 4		17 Secs (17 Secs) [==>]	[1]	
	<i>Comments: I ran a WD spectrum with Teff=50000K, scaled to the distance to the system, through the ETC. The exposure time allows for the system fading 2 more magnitudes.</i>									
	2	ACQ/PEAK D (1527541)	(1) 2MASS-J233759 21+4308509	COS/FUV, ACQ/PEAKD, PSA	G140L 1105 A	CENTER=FLUX-W T-FLR; NUM-POS=5; STEP-SIZE=0.9; LIFETIME-POS=L P4			17 Secs (17 Secs) [==>]	[1]
	<i>Comments: I ran a WD spectrum with Teff=50000K, scaled to the distance to the system, through the ETC. The exposure time allows for the system fading 2 more magnitudes.</i>									
	3	FP-POS=1 (1527545)	(1) 2MASS-J233759 21+4308509	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=1; BUFFER-TIME=87 2; LIFETIME-POS=L P3			872 Secs (872 Secs) [==>]	[1]
	<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>									
4	FP-POS=2 (1527545)	(1) 2MASS-J233759 21+4308509	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=2; BUFFER-TIME=87 2; LIFETIME-POS=L P3			806 Secs (806 Secs) [==>]	[1]	
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>										
5	FP-POS=3 (1527545)	(1) 2MASS-J233759 21+4308509	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=3; BUFFER-TIME=11 46; LIFETIME-POS=L P3			1146 Secs (1146 Secs) [==>]	[2]	
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>										
6	FP-POS=4 (1527545)	(1) 2MASS-J233759 21+4308509	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=4; BUFFER-TIME=11 46; LIFETIME-POS=L P3			1130 Secs (1130 Secs) [==>]	[2]	
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>										



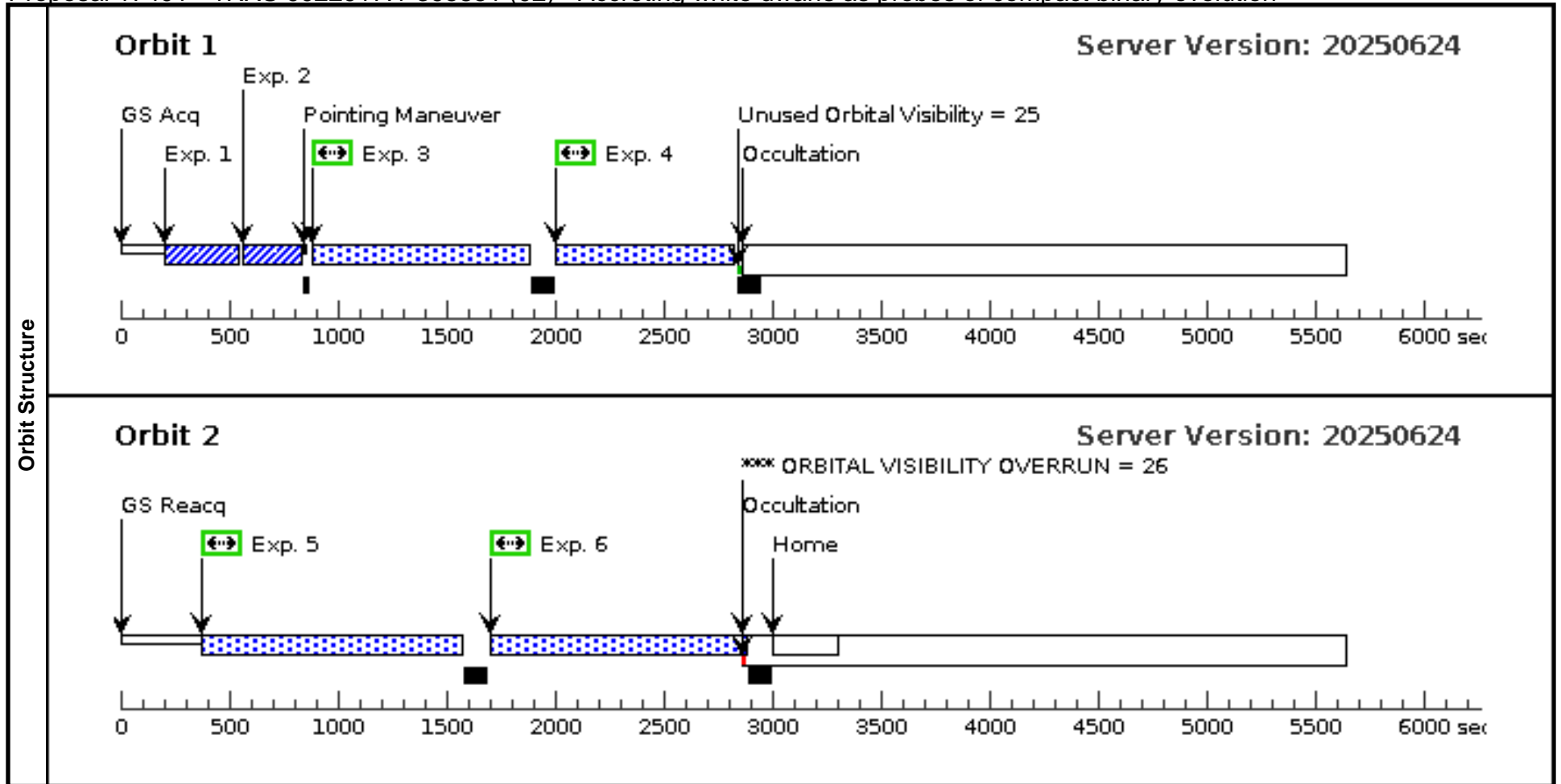
Proposal 17401 - 1RXS-J022917.1-395851 (02) - Accreting white dwarfs as probes of compact binary evolution

Wed Nov 19 17:00:24 GMT 2025

Visit	<p>Proposal 17401, 1RXS-J022917.1-395851 (02), implementation</p> <p>Diagnostic Status: Warning</p> <p>Scientific Instruments: COS/FUV</p> <p>Special Requirements: SCHED 100%; BETWEEN 30-JUN-2021:00:00:00 AND 22-FEB-2022:00:00:00; BETWEEN 30-JUN-2022:00:00:00 AND 22-FEB-2023:00:00:00; BETWEEN 30-JUN-2023:00:00:00 AND 22-FEB-2024:00:00:00; BETWEEN 30-JUN-2024:00:00:00 AND 22-FEB-2025:00:00:00; ON HOLD ; TOO RESPONSE TIME 21.0D</p> <p><i>On Hold Comments: LONG-TERM TOO</i></p> <p><i>This is one of the targets in which the white dwarf is visible only during a low state. We will trigger the ToO when the occurrence of a low state is detected by our ground-based monitoring.</i></p>					
	<p>(1RXS-J022917.1-395851 (02)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p>					
Diagnostics						
Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
	(2)	1RXS-J022917.1-395851	RA: 02 29 17.0025 (37.3208438d) Dec: -39 59 1.61 (-39.98378d) Equinox: J2000	Proper Motion RA: 42.614043637291694 mas/yr Proper Motion Dec: 2.7212311407274634 mas/yr Parallax: 0.0020367016033755827" Epoch of Position: 2016.0	V=13.52+/-0.05 FUV = 13.4, Gaia EDR3 4951395327129718 528, G=13.5, parallax = 2.04(1) mas.	Reference Frame: ICRS
<p><i>Comments:</i> <i>Category=STAR</i> <i>Description=[INTERACTING BINARY, IRREGULAR VARIABLE, NOVA-LIKE]</i> <i>Extended=NO</i></p>						

Proposal 17401 - 1RXS-J022917.1-395851 (02) - Accreting white dwarfs as probes of compact binary evolution

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
Exposures	1	ACQ/PEAK XD (1527549)	(2) 1RXS-J022917.1 -395851	COS/FUV, ACQ/PEAKXD, PSA 1105 A	G140L	LIFETIME-POS=LP 4		25 Secs (25 Secs) [==>]	[1]	
	<i>Comments: I ran a WD spectrum with $T_{\text{eff}}=50000\text{K}$, scaled to the distance to the system, through the ETC. The exposure time allows for the system fading 2 more magnitudes.</i>									
	2	ACQ/PEAK D (1527549)	(2) 1RXS-J022917.1 -395851	COS/FUV, ACQ/PEAKD, PSA 1105 A	G140L	CENTER=FLUX-W T-FLR; NUM-POS=5; STEP-SIZE=0.9; LIFETIME-POS=L P4		25 Secs (25 Secs) [==>]	[1]	
	<i>Comments: I ran a WD spectrum with $T_{\text{eff}}=50000\text{K}$, scaled to the distance to the system, through the ETC. The exposure time allows for the system fading 2 more magnitudes.</i>									
	3	FP-POS=1 (1527551)	(2) 1RXS-J022917.1 -395851	COS/FUV, TIME-TAG, PSA 800 A	G140L	FLASH=YES; FP-POS=1; BUFFER-TIME=83 9; LIFETIME-POS=L P3		839 Secs (839 Secs) [==>]	[1]	
	<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>									
4	FP-POS=2 (1527551)	(2) 1RXS-J022917.1 -395851	COS/FUV, TIME-TAG, PSA 800 A	G140L	FLASH=YES; FP-POS=2; BUFFER-TIME=83 9; LIFETIME-POS=L P3		768 Secs (768 Secs) [==>]	[1]		
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>										
5	FP-POS=3 (1527551)	(2) 1RXS-J022917.1 -395851	COS/FUV, TIME-TAG, PSA 800 A	G140L	FLASH=YES; FP-POS=3; BUFFER-TIME=11 45; LIFETIME-POS=L P3		1145 Secs (1145 Secs) [==>]	[2]		
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>										
6	FP-POS=4 (1527551)	(2) 1RXS-J022917.1 -395851	COS/FUV, TIME-TAG, PSA 800 A	G140L	FLASH=YES; FP-POS=4; BUFFER-TIME=11 45; LIFETIME-POS=L P3		1124 Secs (1124 Secs) [==>]	[2]		
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>										



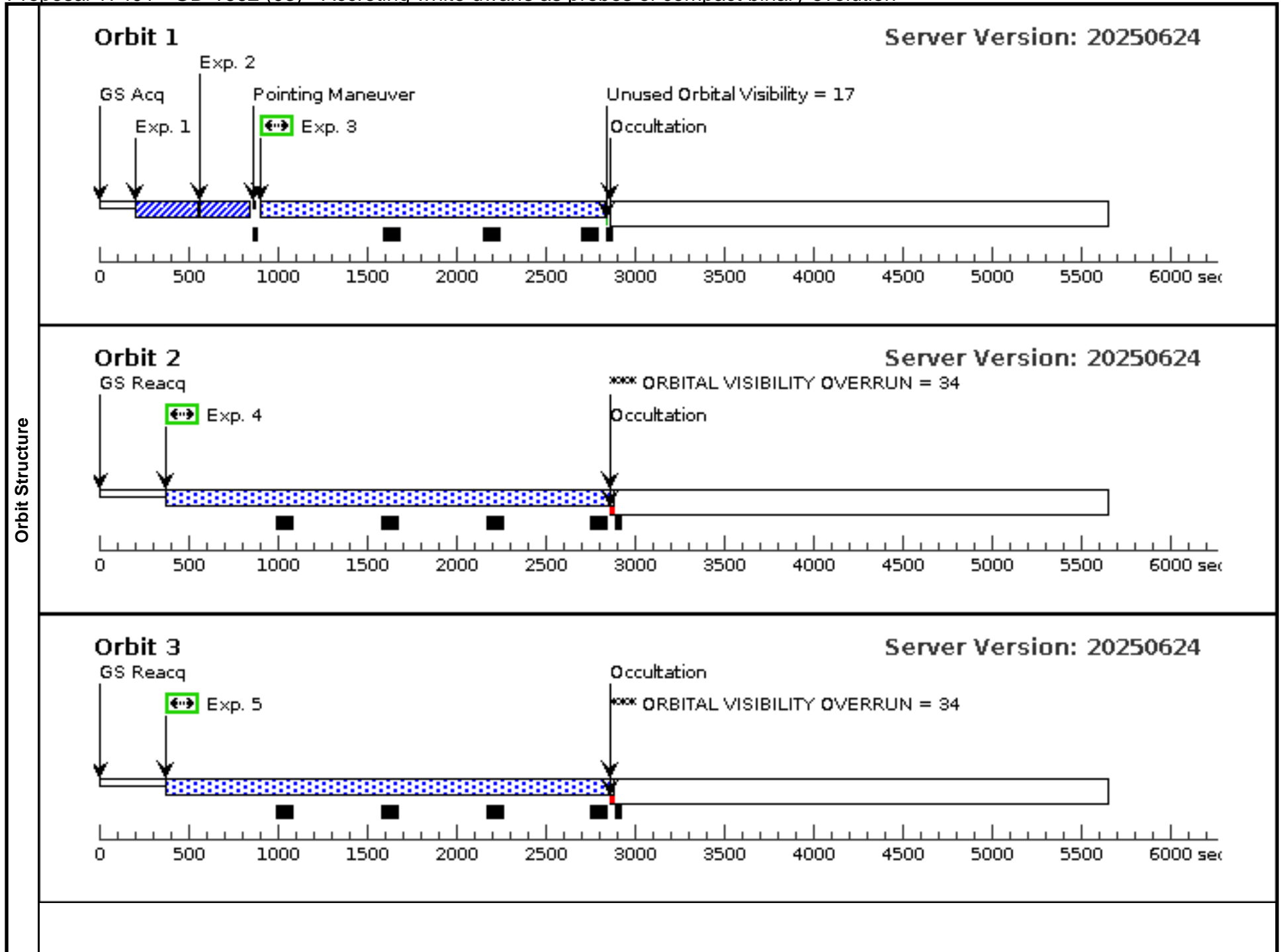
Proposal 17401 - GD-1662 (03) - Accreting white dwarfs as probes of compact binary evolution

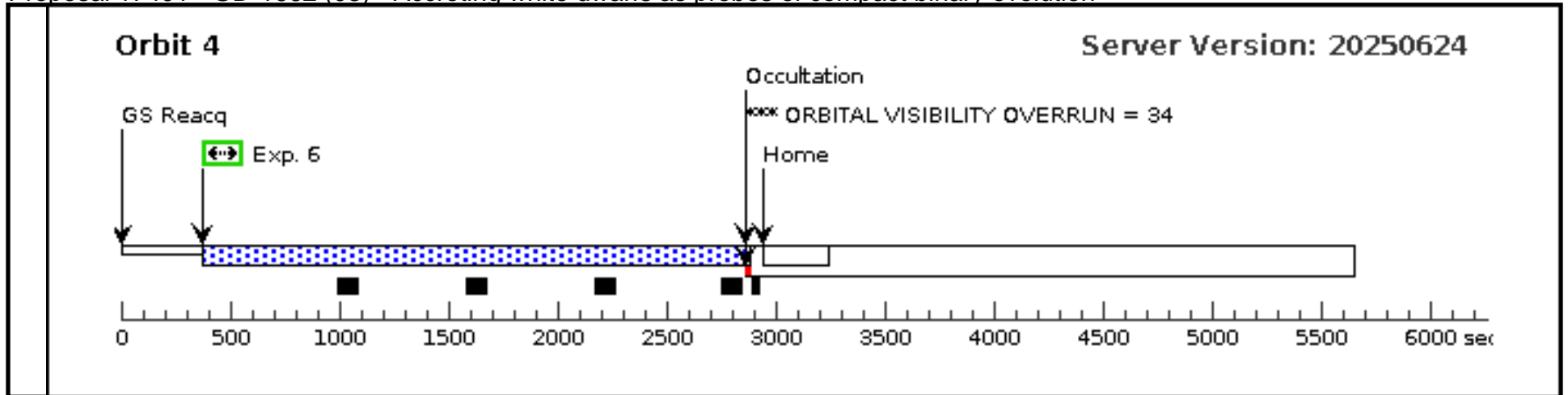
Wed Nov 19 17:00:24 GMT 2025

Visit	Proposal 17401, GD-1662 (03), implementation Diagnostic Status: Warning Scientific Instruments: COS/FUV Special Requirements: SCHED 100%; BETWEEN 23-OCT-2023:00:00:00 AND 25-NOV-2023:00:00:00; ON HOLD ; TOO RESPONSE TIME 21.0D <i>On Hold Comments: LONG-TERM TOO</i> <i>This is one of the targets in which the white dwarf is visible only during a low state. We will trigger the ToO when the occurrence of a low state is detected by our ground-based monitoring.</i>																	
	Diagnosics (GD-1662 (03)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN (GD-1662 (03)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN (GD-1662 (03)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN																	
Fixed Targets	<table border="1"> <thead> <tr> <th>#</th> <th>Name</th> <th>Target Coordinates</th> <th>Targ. Coord. Corrections</th> <th>Fluxes</th> <th>Miscellaneous</th> </tr> </thead> <tbody> <tr> <td>(3)</td> <td>GD-1662</td> <td>RA: 23 29 0.4543 (352.2518929d) Dec: -29 46 46.08 (-29.77947d) Equinox: J2000</td> <td>Proper Motion RA: -13.847933168928266 mas/yr Proper Motion Dec: -6.005677448492444 mas/yr Parallax: 0.001640177333531042" Epoch of Position: 2016.0</td> <td>V=13.2+/-0.3 FUV=14.0, NUV=13.8, Gaia EDR3 2329317895999827840, G=13.6, parallax = 1.64(2) mas.</td> <td>Reference Frame: ICRS</td> </tr> </tbody> </table>						#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous	(3)	GD-1662	RA: 23 29 0.4543 (352.2518929d) Dec: -29 46 46.08 (-29.77947d) Equinox: J2000	Proper Motion RA: -13.847933168928266 mas/yr Proper Motion Dec: -6.005677448492444 mas/yr Parallax: 0.001640177333531042" Epoch of Position: 2016.0	V=13.2+/-0.3 FUV=14.0, NUV=13.8, Gaia EDR3 2329317895999827840, G=13.6, parallax = 1.64(2) mas.	Reference Frame: ICRS
	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous												
(3)	GD-1662	RA: 23 29 0.4543 (352.2518929d) Dec: -29 46 46.08 (-29.77947d) Equinox: J2000	Proper Motion RA: -13.847933168928266 mas/yr Proper Motion Dec: -6.005677448492444 mas/yr Parallax: 0.001640177333531042" Epoch of Position: 2016.0	V=13.2+/-0.3 FUV=14.0, NUV=13.8, Gaia EDR3 2329317895999827840, G=13.6, parallax = 1.64(2) mas.	Reference Frame: ICRS													
<i>Comments:</i> Category=STAR Description=[INTERACTING BINARY, IRREGULAR VARIABLE, NOVA-LIKE] Extended=NO																		

Proposal 17401 - GD-1662 (03) - Accreting white dwarfs as probes of compact binary evolution

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
Exposures	1	ACQ/PEAK XD (1527561)	(3) GD-1662	COS/FUV, ACQ/PEAKXD, PSA	G140L 1105 A	LIFETIME-POS=LP 4		27 Secs (27 Secs) [==>]	[1]
	<i>Comments: I ran a WD spectrum with Teff=50000K, scaled to FUV=16.6 (as measured from IUE data obtained in low state), through the ETC. The exposure time allows for the system fading another magnitude.</i>								
	2	ACQ/PEAK D (1527561)	(3) GD-1662	COS/FUV, ACQ/PEAKD, PSA	G140L 1105 A	CENTER=FLUX-W T-FLR; NUM-POS=5; STEP-SIZE=0.9; LIFETIME-POS=L P4		27 Secs (27 Secs) [==>]	[1]
	<i>Comments: I ran a WD spectrum with Teff=50000K, scaled to FUV=16.6 (as measured from IUE data obtained in low state), through the ETC. The exposure time allows for the system fading another magnitude.</i>								
	3	FP-POS=1 (1527563)	(3) GD-1662	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=1; BUFFER-TIME=55 5; LIFETIME-POS=L P3		1776 Secs (1776 Secs) [==>]	[1]
	<i>Comments: Buffer-time set conservatively short to avoid overflow, in case the system is observed in a brighter state than the low state minimum.</i>								
4	FP-POS=2 (1527563)	(3) GD-1662	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=2; BUFFER-TIME=58 6; LIFETIME-POS=L P3		2454 Secs (2454 Secs) [==>]	[2]	
<i>Comments: Buffer-time set conservatively short to avoid overflow, in case the system is observed in a brighter state than the low state minimum.</i>									
5	FP-POS=3 (1527563)	(3) GD-1662	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=3; BUFFER-TIME=58 6; LIFETIME-POS=L P3		2454 Secs (2454 Secs) [==>]	[3]	
<i>Comments: Buffer-time set conservatively short to avoid overflow, in case the system is observed in a brighter state than the low state minimum.</i>									
6	FP-POS=4 (1527563)	(3) GD-1662	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=4; BUFFER-TIME=58 6; LIFETIME-POS=L P3		2454 Secs (2454 Secs) [==>]	[4]	
<i>Comments: Buffer-time set conservatively short to avoid overflow, in case the system is observed in a brighter state than the low state minimum.</i>									





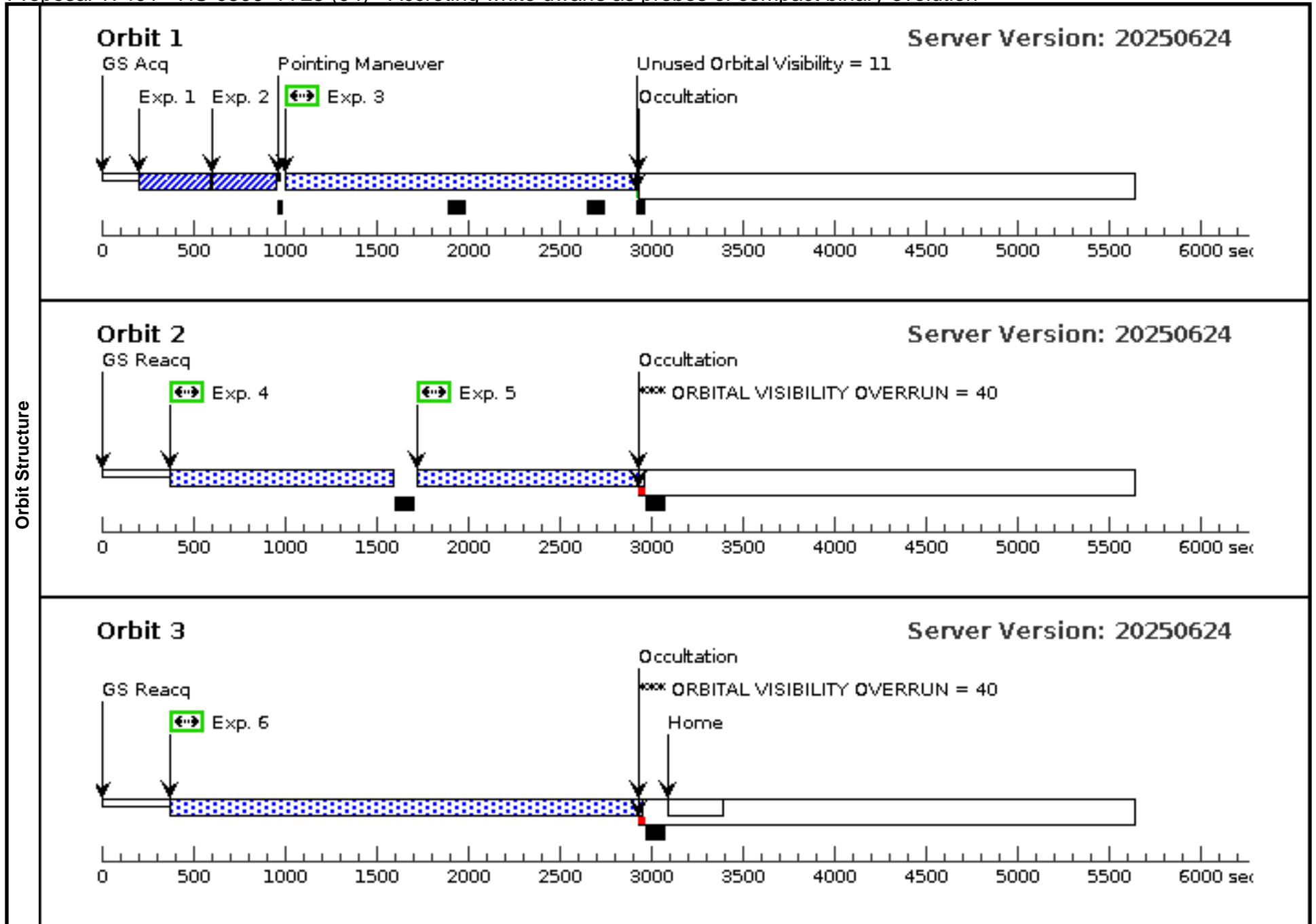
Proposal 17401 - HS-0506+7725 (04) - Accreting white dwarfs as probes of compact binary evolution

Wed Nov 19 17:00:24 GMT 2025

Visit	<p>Proposal 17401, HS-0506+7725 (04), implementation</p> <p>Diagnostic Status: Warning</p> <p>Scientific Instruments: COS/FUV</p> <p>Special Requirements: SCHED 100%; BETWEEN 01-SEP-2021:00:00:00 AND 05-APR-2022:00:00:00; BETWEEN 01-SEP-2022:00:00:00 AND 05-APR-2023:00:00:00; BETWEEN 01-SEP-2023:00:00:00 AND 05-APR-2024:00:00:00; BETWEEN 01-SEP-2024:00:00:00 AND 05-APR-2025:00:00:00; ON HOLD ; TOO RESPONSE TIME 21.0D</p> <p><i>On Hold Comments: LONG-TERM TOO</i></p> <p><i>This is one of the targets in which the white dwarf is visible only during a low state. We will trigger the ToO when the occurrence of a low state is detected by our ground-based monitoring.</i></p>					
	<p>(HS-0506+7725 (04)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p> <p>(HS-0506+7725 (04)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p>					
Diagnosics						
Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
	(4)	HS-0506+7725	RA: 05 13 36.5757 (78.4023988d) Dec: +77 28 42.67 (77.47852d) Equinox: J2000	Proper Motion RA: 1.2821538884881634 mas/yr Proper Motion Dec: -6.849241478209648 mas/yr Parallax: 0.001085954010123805" Epoch of Position: 2016.0	V=14.7 FUV=16.5, NUV=16.5, Gaia EDR3 5521586977692176 64, G=15.04, parallax = 1.09(2) mas.	Reference Frame: ICRS
<p><i>Comments:</i> <i>Category=STAR</i> <i>Description=[INTERACTING BINARY, IRREGULAR VARIABLE, NOVA-LIKE]</i> <i>Extended=NO</i></p>						

Proposal 17401 - HS-0506+7725 (04) - Accreting white dwarfs as probes of compact binary evolution

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
Exposures	1	ACQ/PEAK XD (1527733)	(4) HS-0506+7725	COS/FUV, ACQ/PEAKXD, PSA	G140L 1105 A	LIFETIME-POS=LP 4		40 Secs (40 Secs) [==>]	[1]	
	<i>Comments: Exposure time based on a 50000K WD model, scaled to the distance to the system, but conservatively allowing for the target fading two more magnitudes.</i>									
	2	ACQ/PEAK D (1527733)	(4) HS-0506+7725	COS/FUV, ACQ/PEAKD, PSA	G140L 1105 A	CENTER=FLUX-W T-FLR; NUM-POS=5; STEP-SIZE=0.9; LIFETIME-POS=L P4		40 Secs (40 Secs) [==>]	[1]	
	<i>Comments: Exposure time based on a 50000K WD model, scaled to the distance to the system, but conservatively allowing for the target fading two more magnitudes.</i>									
	3	FP-POS=1 (1527737)	(4) HS-0506+7725	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=1; BUFFER-TIME=75 2; LIFETIME-POS=L P3		1750 Secs (1750 Secs) [==>]	[1]	
	<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>									
4	FP-POS=2 (1527737)	(4) HS-0506+7725	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=2; BUFFER-TIME=11 85; LIFETIME-POS=L P3		1165 Secs (1165 Secs) [==>]	[2]		
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>										
5	FP-POS=3 (1527737)	(4) HS-0506+7725	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=3; BUFFER-TIME=11 86; LIFETIME-POS=L P3		1186 Secs (1186 Secs) [==>]	[2]		
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>										
6	FP-POS=4 (1527737)	(4) HS-0506+7725	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=4; BUFFER-TIME=25 52; LIFETIME-POS=L P3		2532 Secs (2532 Secs) [==>]	[3]		
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>										



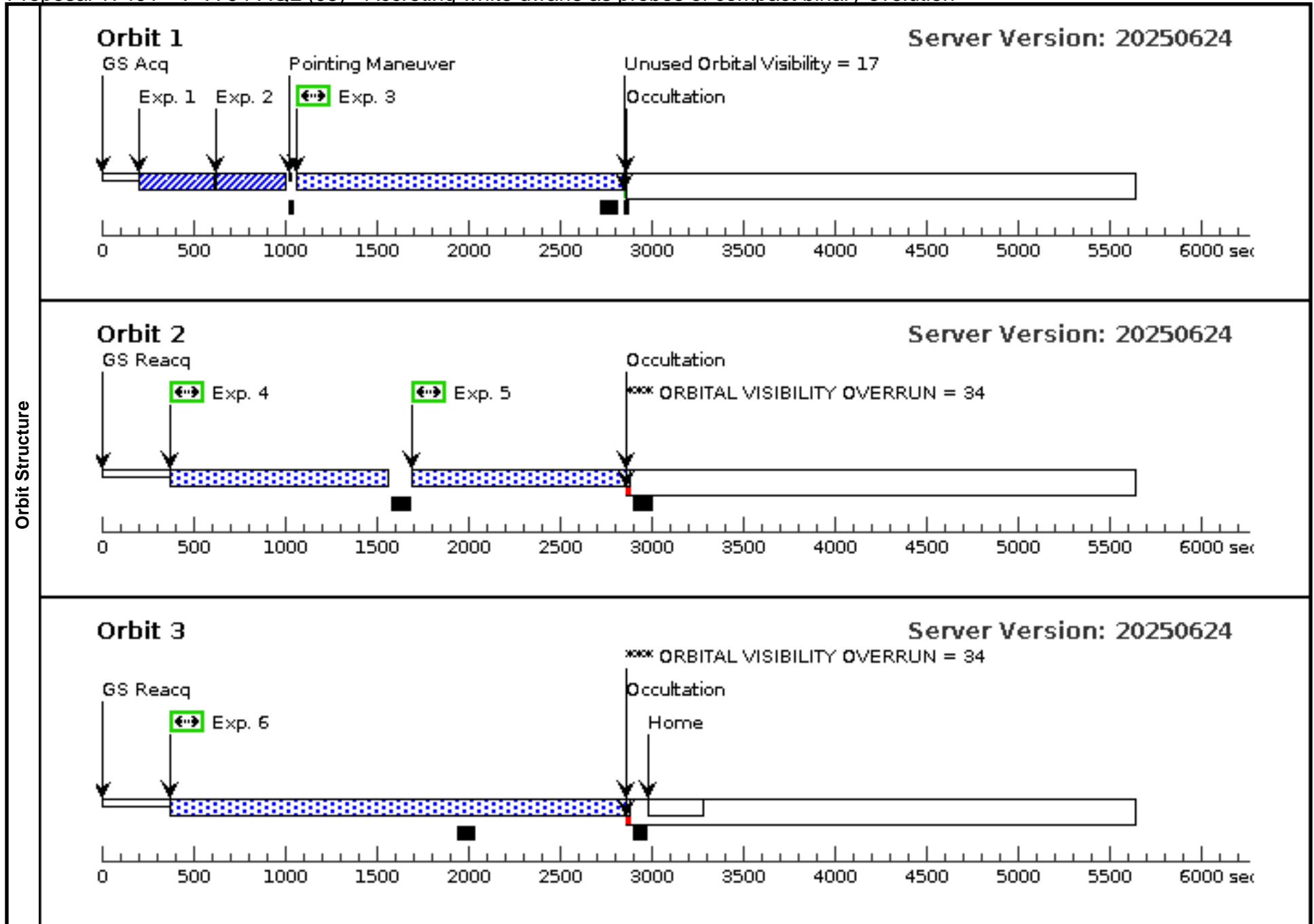
Proposal 17401 - V-V794-AQL (05) - Accreting white dwarfs as probes of compact binary evolution

Wed Nov 19 17:00:24 GMT 2025

Visit	<p>Proposal 17401, V-V794-AQL (05), implementation</p> <p>Diagnostic Status: Warning</p> <p>Scientific Instruments: COS/FUV</p> <p>Special Requirements: SCHED 100%; BETWEEN 18-APR-2021:00:00:00 AND 08-NOV-2021:00:00:00; BETWEEN 18-APR-2022:00:00:00 AND 08-NOV-2022:00:00:00; BETWEEN 18-APR-2023:00:00:00 AND 08-NOV-2023:00:00:00; BETWEEN 18-APR-2024:00:00:00 AND 08-NOV-2024:00:00:00; ON HOLD ; TOO RESPONSE TIME 21.0D</p> <p><i>On Hold Comments: LONG-TERM TOO</i></p> <p><i>This is one of the targets in which the white dwarf is visible only during a low state. We will trigger the ToO when the occurrence of a low state is detected by our ground-based monitoring.</i></p>					
	<p>(V-V794-AQL (05)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p> <p>(V-V794-AQL (05)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p>					
Diagnosics						
Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
	(5)	V-V794-AQL	RA: 20 17 33.8730 (304.3911375d) Dec: -03 39 51.19 (-3.66422d) Equinox: J2000	Proper Motion RA: -40.13430224001505 mas/yr Proper Motion Dec: -14.632375673806653 mas/yr Parallax: 0.0015277963634299084" Epoch of Position: 2016.0	V=14.7 FUV = 16.2, Gaia EDR3 4222299752184927 616, G=14.8, parallax = 1.53(3) mas.	Reference Frame: ICRS
<p><i>Comments:</i> <i>Category=STAR</i> <i>Description=[INTERACTING BINARY, IRREGULAR VARIABLE, NOVA-LIKE]</i> <i>Extended=NO</i></p>						

Proposal 17401 - V-V794-AQL (05) - Accreting white dwarfs as probes of compact binary evolution

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
Exposures	1	ACQ/PEAK XD (1527706)	(5) V-V794-AQL	COS/FUV, ACQ/PEAKXD, PSA	G140L 1105 A	LIFETIME-POS=LP 4		47 Secs (47 Secs) [==>]	[1]
	<i>Comments: I ran a WD spectrum with $T_{\text{eff}}=50000\text{K}$, scaled to the distance to the system, through the ETC. The exposure time allows for the system fading 1 more magnitudes.</i>								
	2	ACQ/PEAK D (1527706)	(5) V-V794-AQL	COS/FUV, ACQ/PEAKD, PSA	G140L 1105 A	CENTER=FLUX-W T-FLR; NUM-POS=5; STEP-SIZE=0.9; LIFETIME-POS=L P4		47 Secs (47 Secs) [==>]	[1]
	<i>Comments: I ran a WD spectrum with $T_{\text{eff}}=50000\text{K}$, scaled to the distance to the system, through the ETC. The exposure time allows for the system fading 1 more magnitudes.</i>								
	3	FP-POS=1 (1527716)	(5) V-V794-AQL	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=1; BUFFER-TIME=15 27; LIFETIME-POS=L P3		1619 Secs (1619 Secs) [==>]	[1]
	<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>								
4	FP-POS=2 (1527716)	(5) V-V794-AQL	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=2; BUFFER-TIME=11 41; LIFETIME-POS=L P3		1141 Secs (1141 Secs) [==>]	[2]	
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>									
5	FP-POS=3 (1527716)	(5) V-V794-AQL	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=3; BUFFER-TIME=11 40; LIFETIME-POS=L P3		1135 Secs (1135 Secs) [==>]	[2]	
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>									
6	FP-POS=4 (1527716)	(5) V-V794-AQL	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=4; BUFFER-TIME=15 27; LIFETIME-POS=L P3		2457 Secs (2457 Secs) [==>]	[3]	
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>									



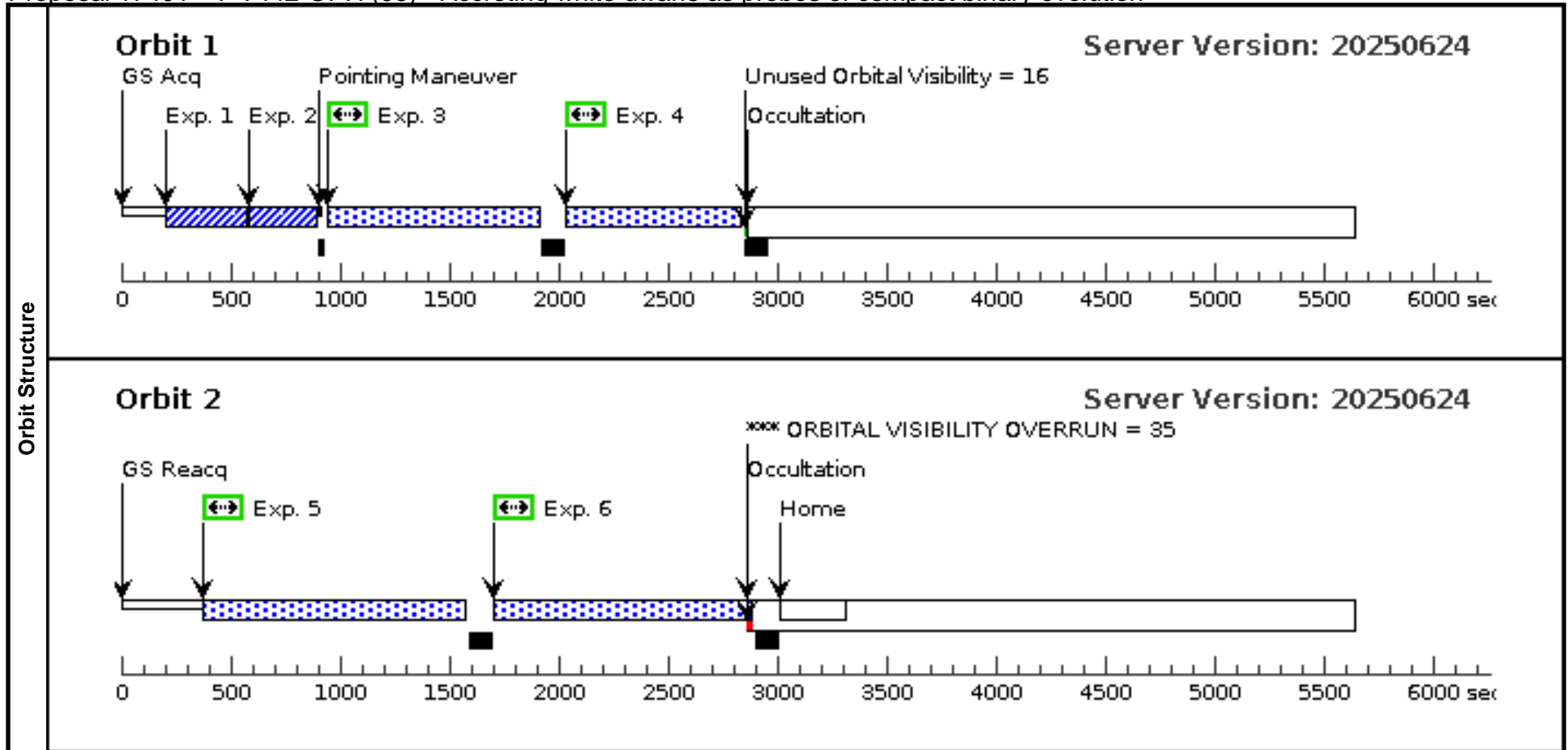
Proposal 17401 - V-V442-OPH (06) - Accreting white dwarfs as probes of compact binary evolution

Wed Nov 19 17:00:24 GMT 2025

Visit	<p>Proposal 17401, V-V442-OPH (06), implementation</p> <p>Diagnostic Status: Warning</p> <p>Scientific Instruments: COS/FUV</p> <p>Special Requirements: SCHED 100%; BETWEEN 07-MAR-2022:00:00:00 AND 10-OCT-2022:00:00:00; BETWEEN 07-MAR-2023:00:00:00 AND 10-OCT-2023:00:00:00; BETWEEN 07-MAR-2024:00:00:00 AND 10-OCT-2024:00:00:00; BETWEEN 07-MAR-2025:00:00:00 AND 10-OCT-2025:00:00:00; ON HOLD ; TOO RESPONSE TIME 21.0D</p> <p><i>On Hold Comments: LONG-TERM TOO</i></p> <p><i>This is one of the targets in which the white dwarf is visible only during a low state. We will trigger the ToO when the occurrence of a low state is detected by our ground-based monitoring.</i></p>																	
	<p>(V-V442-OPH (06)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p>																	
Diagnostics																		
Fixed Targets	<table border="1"> <thead> <tr> <th>#</th> <th>Name</th> <th>Target Coordinates</th> <th>Targ. Coord. Corrections</th> <th>Fluxes</th> <th>Miscellaneous</th> </tr> </thead> <tbody> <tr> <td>(6)</td> <td>V-V442-OPH</td> <td>RA: 17 32 15.1318 (263.0630492d) Dec: -16 15 22.01 (-16.25611d) Equinox: J2000</td> <td>Proper Motion RA: -3.9615056735586376 mas/yr Proper Motion Dec: 3.5013410220099415 mas/yr Parallax: 0.002070837745656075" Epoch of Position: 2016.0</td> <td>V=13.4 FUV=16.7, NUV=16.5, G=13.9, Gaia EDR3 4124654392314770 048, G=parallax=2.07(1) mas.</td> <td>Reference Frame: ICRS</td> </tr> </tbody> </table>	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous	(6)	V-V442-OPH	RA: 17 32 15.1318 (263.0630492d) Dec: -16 15 22.01 (-16.25611d) Equinox: J2000	Proper Motion RA: -3.9615056735586376 mas/yr Proper Motion Dec: 3.5013410220099415 mas/yr Parallax: 0.002070837745656075" Epoch of Position: 2016.0	V=13.4 FUV=16.7, NUV=16.5, G=13.9, Gaia EDR3 4124654392314770 048, G=parallax=2.07(1) mas.	Reference Frame: ICRS					
	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous												
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<p><i>Comments:</i> <i>Category=STAR</i> <i>Description=[INTERACTING BINARY, IRREGULAR VARIABLE, NOVA-LIKE]</i> <i>Extended=NO</i></p>																		

Proposal 17401 - V-V442-OPH (06) - Accreting white dwarfs as probes of compact binary evolution

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
Exposures	1	ACQ/PEAK XD (1527475)	(6) V-V442-OPH	COS/FUV, ACQ/PEAKXD, PSA	G140L 1105 A	LIFETIME-POS=LP 4		32 Secs (32 Secs) [==>]	[1]	
	<i>Comments: I ran a WD spectrum with $T_{\text{eff}}=50000\text{K}$, scaled to $FUV=16.8$ (as measured from <i>Galex</i> photometry in low state), through the ETC. The exposure time allows for the system fading another magnitude.</i>									
	2	ACQ/PEAK D (1527475)	(6) V-V442-OPH	COS/FUV, ACQ/PEAKD, PSA	G140L 1105 A	CENTER=FLUX-W T-FLR; NUM-POS=5; STEP-SIZE=0.9; LIFETIME-POS=L P4		32 Secs (32 Secs) [==>]	[1]	
	<i>Comments: I ran a WD spectrum with $T_{\text{eff}}=50000\text{K}$, scaled to $FUV=16.8$ (as measured from <i>Galex</i> photometry in low state), through the ETC. The exposure time allows for the system fading another magnitude.</i>									
	3	FP-POS=1 (1527650)	(6) V-V442-OPH	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=1; BUFFER-TIME=81 2; LIFETIME-POS=L P3		812 Secs (812 Secs) [==>]	[1]	
	<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>									
4	FP-POS=2 (1527650)	(6) V-V442-OPH	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=2; BUFFER-TIME=81 2; LIFETIME-POS=L P3		746 Secs (746 Secs) [==>]	[1]		
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>										
5	FP-POS=3 (1527650)	(6) V-V442-OPH	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=3; BUFFER-TIME=11 46; LIFETIME-POS=L P3		1146 Secs (1146 Secs) [==>]	[2]		
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>										
6	FP-POS=4 (1527650)	(6) V-V442-OPH	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=4; BUFFER-TIME=11 46; LIFETIME-POS=L P3		1130 Secs (1130 Secs) [==>]	[2]		
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the low state).</i>										



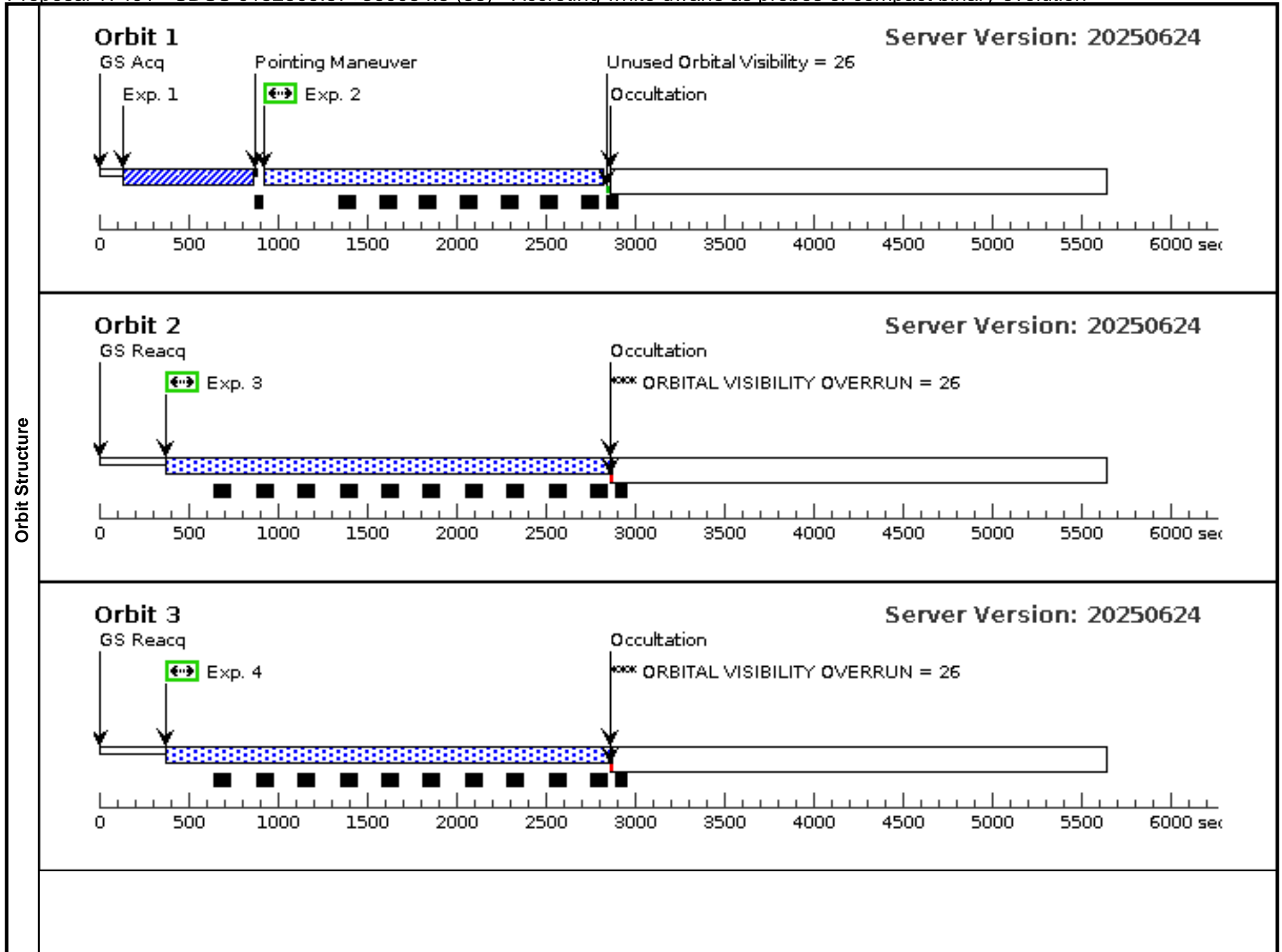
Proposal 17401 - SDSS-J152509.57+360054.5 (38) - Accreting white dwarfs as probes of compact binary evolution

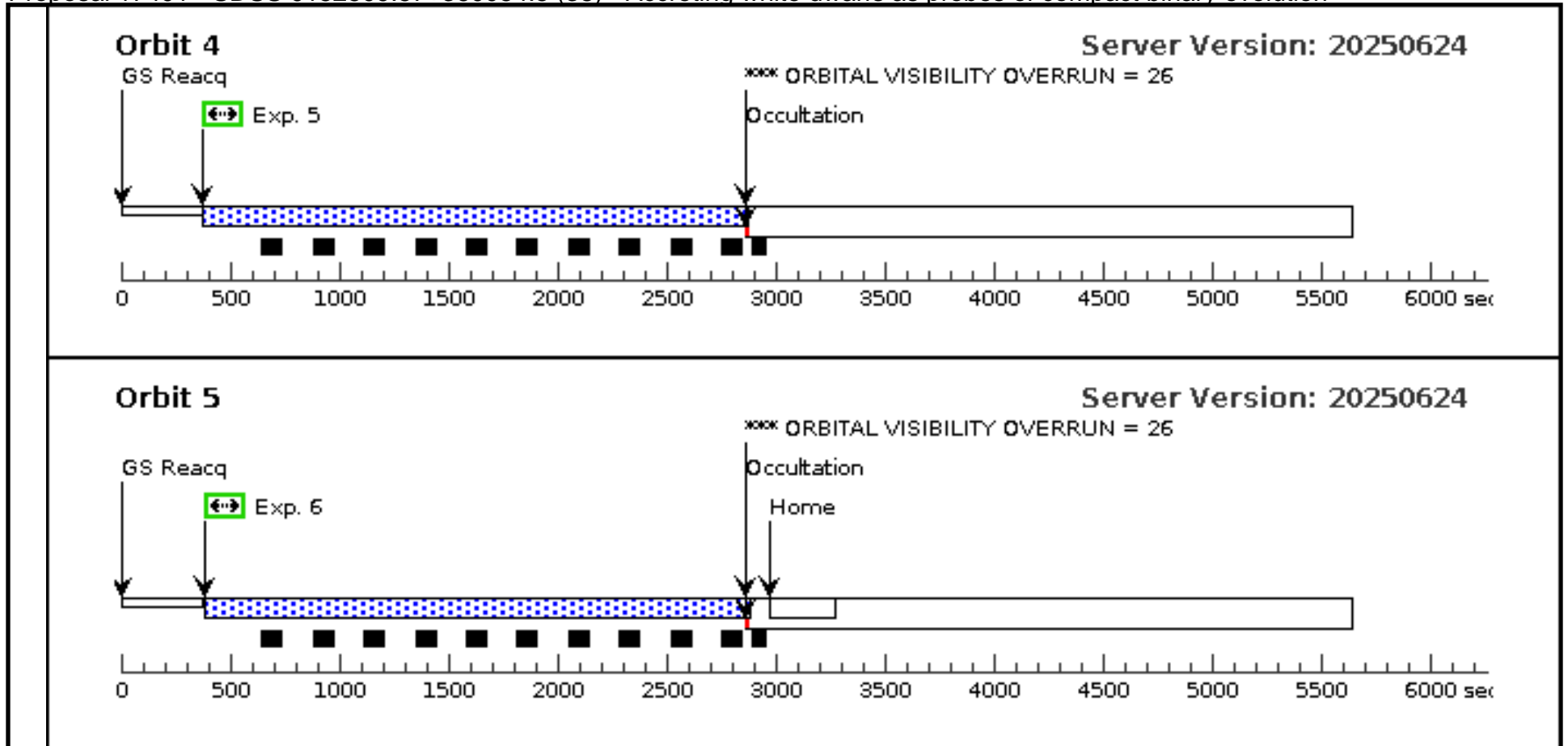
Wed Nov 19 17:00:24 GMT 2025

Visit	Proposal 17401, SDSS-J152509.57+360054.5 (38), implementation Diagnostic Status: Warning Scientific Instruments: COS/FUV, COS/NUV Special Requirements: SCHED 100%; BETWEEN 01-MAR-2025:00:00:00 AND 13-MAR-2025:00:00:00; BETWEEN 19-MAR-2025:00:00:00 AND 10-APR-2025:00:00:00; BETWEEN 18-APR-2025:00:00:00 AND 07-MAY-2025:00:00:00; BETWEEN 16-MAY-2025:00:00:00 AND 05-JUN-2025:00:00:00																	
	Diagnostics	(SDSS-J152509.57+360054.5 (38)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN (SDSS-J152509.57+360054.5 (38)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN (SDSS-J152509.57+360054.5 (38)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN (SDSS-J152509.57+360054.5 (38)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN																
Fixed Targets		<table border="1"> <thead> <tr> <th>#</th> <th>Name</th> <th>Target Coordinates</th> <th>Targ. Coord. Corrections</th> <th>Fluxes</th> <th>Miscellaneous</th> </tr> </thead> <tbody> <tr> <td>(38)</td> <td>SDSS-J152509.57+360054.5</td> <td>RA: 15 25 9.5732 (231.2898883d) Dec: +36 00 54.41 (36.01511d) Equinox: J2000</td> <td>Proper Motion RA: 4.986026977395869 mas/yr Proper Motion Dec: -15.85931643659449 mas/yr Parallax: 0.00196588153504011" Epoch of Position: 2016.0</td> <td>V=19.64 FUV=20.1, NUV=19.5, Gaia EDR3 1375131155313563136, G=19.8, parallax=1.9(2) mas.</td> <td>Reference Frame: ICRS</td> </tr> </tbody> </table>						#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous	(38)	SDSS-J152509.57+360054.5	RA: 15 25 9.5732 (231.2898883d) Dec: +36 00 54.41 (36.01511d) Equinox: J2000	Proper Motion RA: 4.986026977395869 mas/yr Proper Motion Dec: -15.85931643659449 mas/yr Parallax: 0.00196588153504011" Epoch of Position: 2016.0	V=19.64 FUV=20.1, NUV=19.5, Gaia EDR3 1375131155313563136, G=19.8, parallax=1.9(2) mas.
	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous												
(38)	SDSS-J152509.57+360054.5	RA: 15 25 9.5732 (231.2898883d) Dec: +36 00 54.41 (36.01511d) Equinox: J2000	Proper Motion RA: 4.986026977395869 mas/yr Proper Motion Dec: -15.85931643659449 mas/yr Parallax: 0.00196588153504011" Epoch of Position: 2016.0	V=19.64 FUV=20.1, NUV=19.5, Gaia EDR3 1375131155313563136, G=19.8, parallax=1.9(2) mas.	Reference Frame: ICRS													
Comments: No V magnitude available, reporting u-SDSS magnitude instead. Category=STAR Description=[ACCRETION DISK, DWARF NOVA, INTERACTING BINARY, IRREGULAR VARIABLE] Extended=NO																		

Proposal 17401 - SDSS-J152509.57+360054.5 (38) - Accreting white dwarfs as probes of compact binary evolution

Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
	1	ACQ/IMAG E (1529618)	(38) SDSS-J152509. 57+360054.5	COS/NUV, ACQ/IMAGE, PSA	MIRRORB					210 Secs (210 Secs) [==>]	[1]
	2	FP-POS=1 (1529621)	(38) SDSS-J152509. 57+360054.5	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=1; BUFFER-TIME=22 6; LIFETIME-POS=L P3				1693 Secs (1693 Secs) [==>]	[1]
	3	FP-POS=2 (1529621)	(38) SDSS-J152509. 57+360054.5	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=2; BUFFER-TIME=23 4; LIFETIME-POS=L P3				2450 Secs (2450 Secs) [==>]	[2]
	4	FP-POS=3 (1529621)	(38) SDSS-J152509. 57+360054.5	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=3; BUFFER-TIME=23 4; LIFETIME-POS=L P3				2450 Secs (2450 Secs) [==>]	[3]
	5	FP-POS=4 (1529621)	(38) SDSS-J152509. 57+360054.5	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=4; BUFFER-TIME=23 4; LIFETIME-POS=L P3				2450 Secs (2450 Secs) [==>]	[4]
	6	FP-POS=4 (1529621)	(38) SDSS-J152509. 57+360054.5	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=4; BUFFER-TIME=23 4; LIFETIME-POS=L P3				2450 Secs (2450 Secs) [==>]	[5]





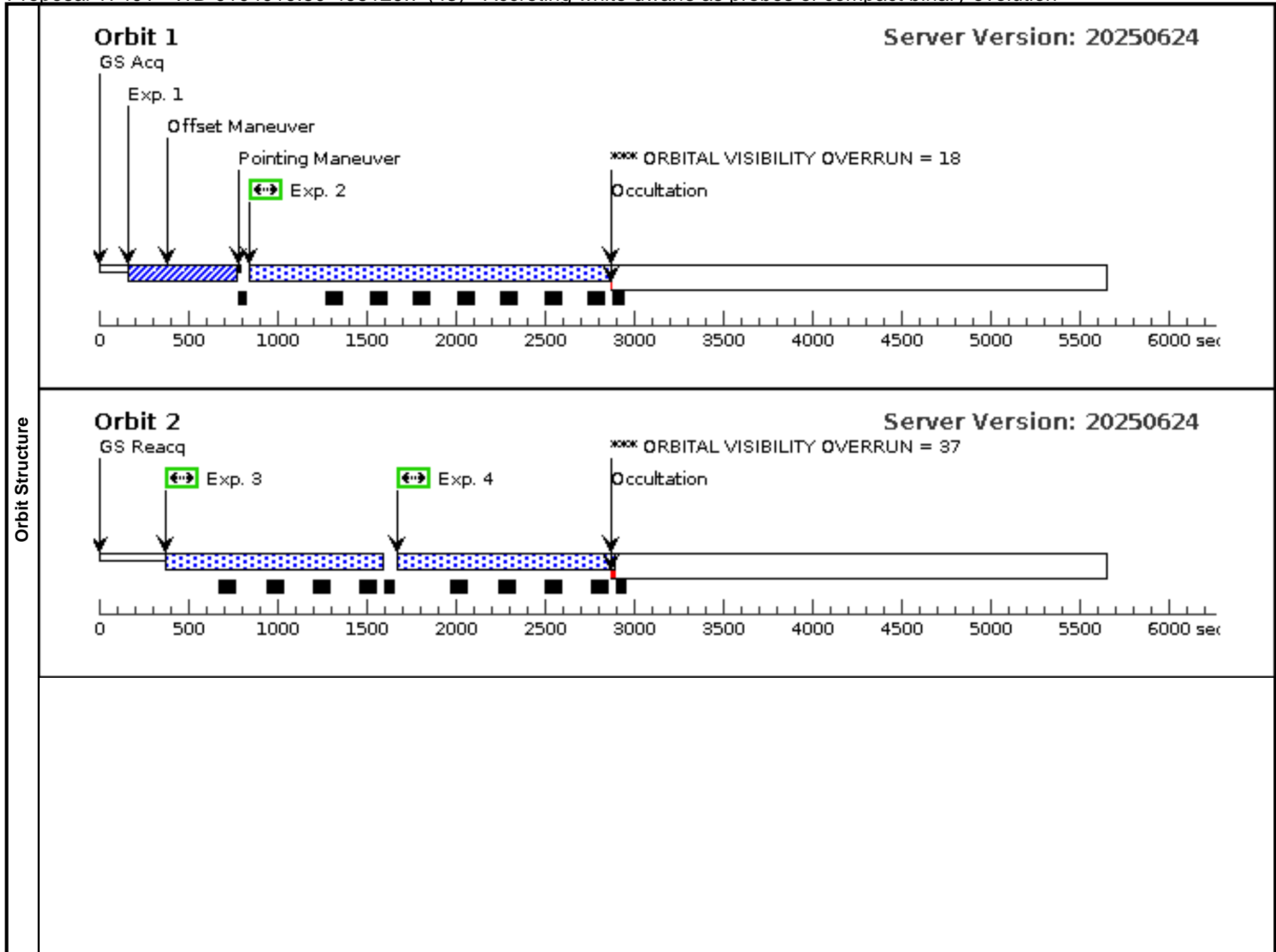
Proposal 17401 - WD-J104019.50-495129.7 (43) - Accreting white dwarfs as probes of compact binary evolution

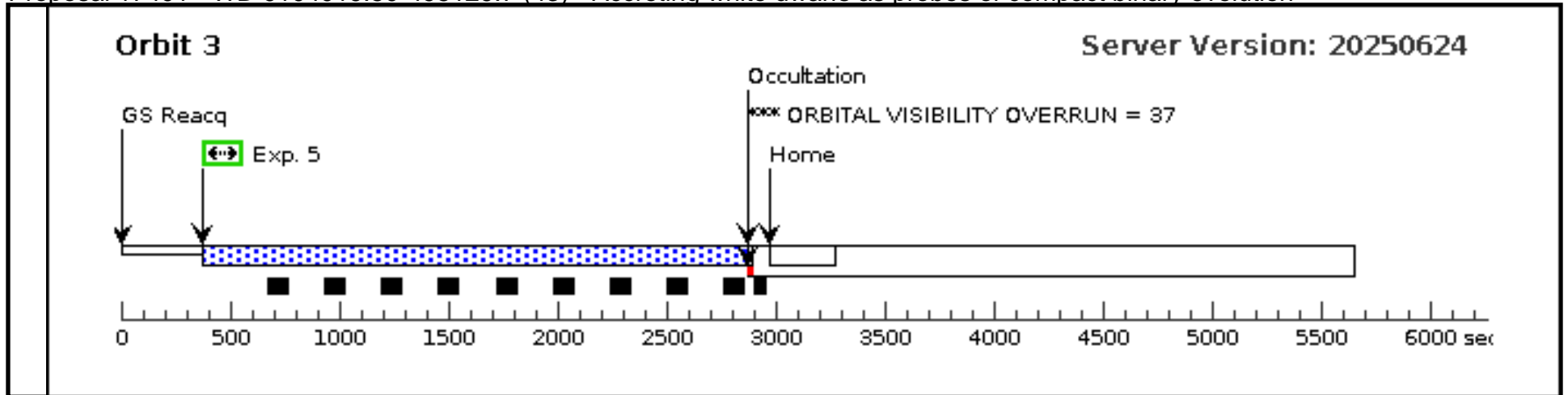
Wed Nov 19 17:00:24 GMT 2025

Visit	Proposal 17401, WD-J104019.50-495129.7 (43), implementation Diagnostic Status: Warning Scientific Instruments: COS/FUV, COS/NUV Special Requirements: SCHED 100%; ORIENT 170D TO 45 D; BETWEEN 01-NOV-2025:00:00:00 AND 01-DEC-2025:00:00:00 <i>Comments: This should be scheduled in the evening local-time on a Thursday or Friday. Flags need to be cleared during the work day. Weekends and holidays are to be avoided since the CS must clear the target within 24 hours of HST execution</i>																	
	Diagnosics (WD-J104019.50-495129.7 (43)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN (WD-J104019.50-495129.7 (43)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN (WD-J104019.50-495129.7 (43)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN																	
Fixed Targets	<table border="1"> <thead> <tr> <th>#</th> <th>Name</th> <th>Target Coordinates</th> <th>Targ. Coord. Corrections</th> <th>Fluxes</th> <th>Miscellaneous</th> </tr> </thead> <tbody> <tr> <td>(43)</td> <td>WD-J104019.50-495129.7</td> <td>RA: 10 40 19.4785 (160.0811604d) Dec: -49 51 29.43 (-49.85818d) Equinox: J2000</td> <td>Proper Motion RA: -11.746504535541535 mas/yr Proper Motion Dec: 16.566512133372953 mas/yr Parallax: 0.003783567563747677" Epoch of Position: 2016.0</td> <td>V=17 Gaia EDR3 5360970516650651136, G=17.1, parallax= 3.78(5) mas.</td> <td>Reference Frame: ICRS</td> </tr> </tbody> </table>						#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous	(43)	WD-J104019.50-495129.7	RA: 10 40 19.4785 (160.0811604d) Dec: -49 51 29.43 (-49.85818d) Equinox: J2000	Proper Motion RA: -11.746504535541535 mas/yr Proper Motion Dec: 16.566512133372953 mas/yr Parallax: 0.003783567563747677" Epoch of Position: 2016.0	V=17 Gaia EDR3 5360970516650651136, G=17.1, parallax= 3.78(5) mas.	Reference Frame: ICRS
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(43)	WD-J104019.50-495129.7	RA: 10 40 19.4785 (160.0811604d) Dec: -49 51 29.43 (-49.85818d) Equinox: J2000	Proper Motion RA: -11.746504535541535 mas/yr Proper Motion Dec: 16.566512133372953 mas/yr Parallax: 0.003783567563747677" Epoch of Position: 2016.0	V=17 Gaia EDR3 5360970516650651136, G=17.1, parallax= 3.78(5) mas.	Reference Frame: ICRS													
<i>Comments:</i> Category=STAR Description=[ACCRETION DISK, DWARF NOVA, INTERACTING BINARY, IRREGULAR VARIABLE] Extended=NO																		

Proposal 17401 - WD-J104019.50-495129.7 (43) - Accreting white dwarfs as probes of compact binary evolution

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
Exposures	1	ACQ/IMAG E (1895722)	(43) WD-J104019.50 -495129.7	COS/NUV, ACQ/IMAGE, PSA	MIRRORB		USE OFFSET V43S AF	150 Secs (150 Secs) [==>]	[1]	
	<i>Comments: Using mirror B to avoid exceeding count rate per pixel in case the system is brighter than expected (but still safe for the detector). The exposure time computed assuming a black body with $T_{\text{eff}}=17000\text{K}$, scaled to $FUV=17$, and has been increase by an amount equal to the duration of the full eclipse (90 seconds). This will allow to reach the required S/N even if an eclipse occurs during the acquisition. 50 seconds more have been added to be conservative.</i>									
	2	FP-POS=1 (1529784)	(43) WD-J104019.50 -495129.7	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=1; BUFFER-TIME=24 5; LIFETIME-POS=L P3	USE OFFSET V43S AF	1827 Secs (1827 Secs) [==>]	[1]	
	<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the quiescent level).</i>									
	3	FP-POS=2 (1529784)	(43) WD-J104019.50 -495129.7	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=2; BUFFER-TIME=26 3; LIFETIME-POS=L P3	USE OFFSET V43S AF	1165 Secs (1165 Secs) [==>]	[2]	
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the quiescent level).</i>										
4	FP-POS=3 (1529784)	(43) WD-J104019.50 -495129.7	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=3; BUFFER-TIME=26 3; LIFETIME-POS=L P3	USE OFFSET V43S AF	1165 Secs (1165 Secs) [==>]	[2]		
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the quiescent level).</i>										
5	FP-POS=4 (1529784)	(43) WD-J104019.50 -495129.7	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=4; BUFFER-TIME=26 1; LIFETIME-POS=L P3	USE OFFSET V43S AF	2464 Secs (2464 Secs) [==>]	[3]		
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the quiescent level).</i>										





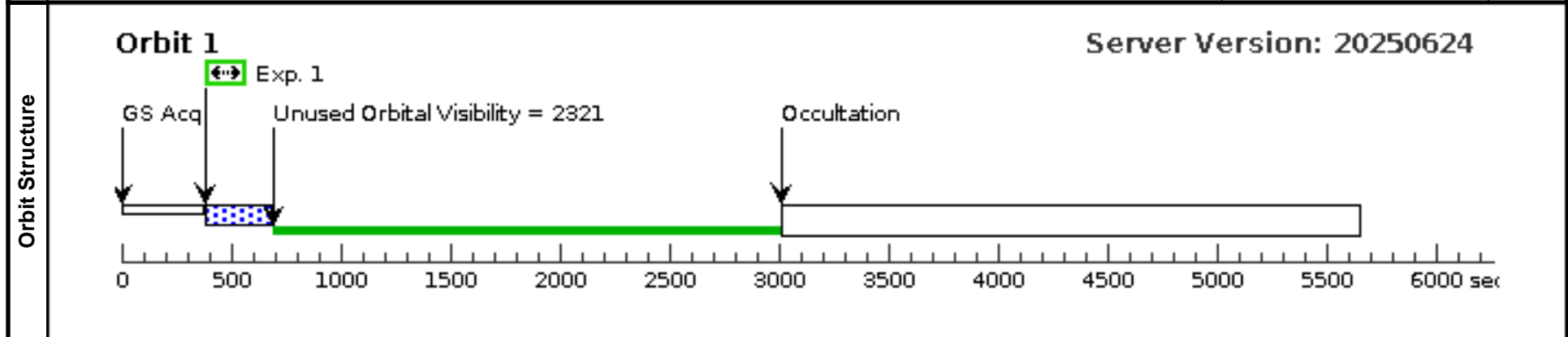
Proposal 17401 - S/C visit for visit 43 (4G) - Accreting white dwarfs as probes of compact binary evolution

Wed Nov 19 17:00:24 GMT 2025

Visit	Proposal 17401, S/C visit for visit 43 (4G), scheduling Diagnostic Status: No Diagnostics Scientific Instruments: S/C Special Requirements: ORIENT 170D TO 45 D <i>Comments: This visit allocates and sets up the ssafe position offset slot for visit 43 which will use that slot. This S/C visit should go 3-5 days before visit 43. The S/C visit will contain only 1 exposure</i>				

Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
	(43)	WD-J104019.50-495129.7	RA: 10 40 19.4785 (160.0811604d) Dec: -49 51 29.43 (-49.85818d) Equinox: J2000	Proper Motion RA: -11.746504535541535 mas/yr Proper Motion Dec: 16.566512133372953 mas/yr Parallax: 0.003783567563747677" Epoch of Position: 2016.0	V=17 Gaia EDR3 5360970516650651136, G=17.1, parallax= 3.78(5) mas.	Reference Frame: ICRS
<i>Comments:</i> Category=STAR Description=[ACCRETION DISK, DWARF NOVA, INTERACTING BINARY, IRREGULAR VARIABLE] Extended=NO						

#	Label	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
1		(43) WD-J104019.50-495129.7	S/C, DATA, V1			POS TARG 232.723 0,-237.515; SAVE OFFSET V43 SAF; SPEC COM INSTR ECSLOTSET; QESIPARM ANGL E 331.7; QESIPARM DIST 9.576		310 Secs (310 Secs) [==>]	[1]



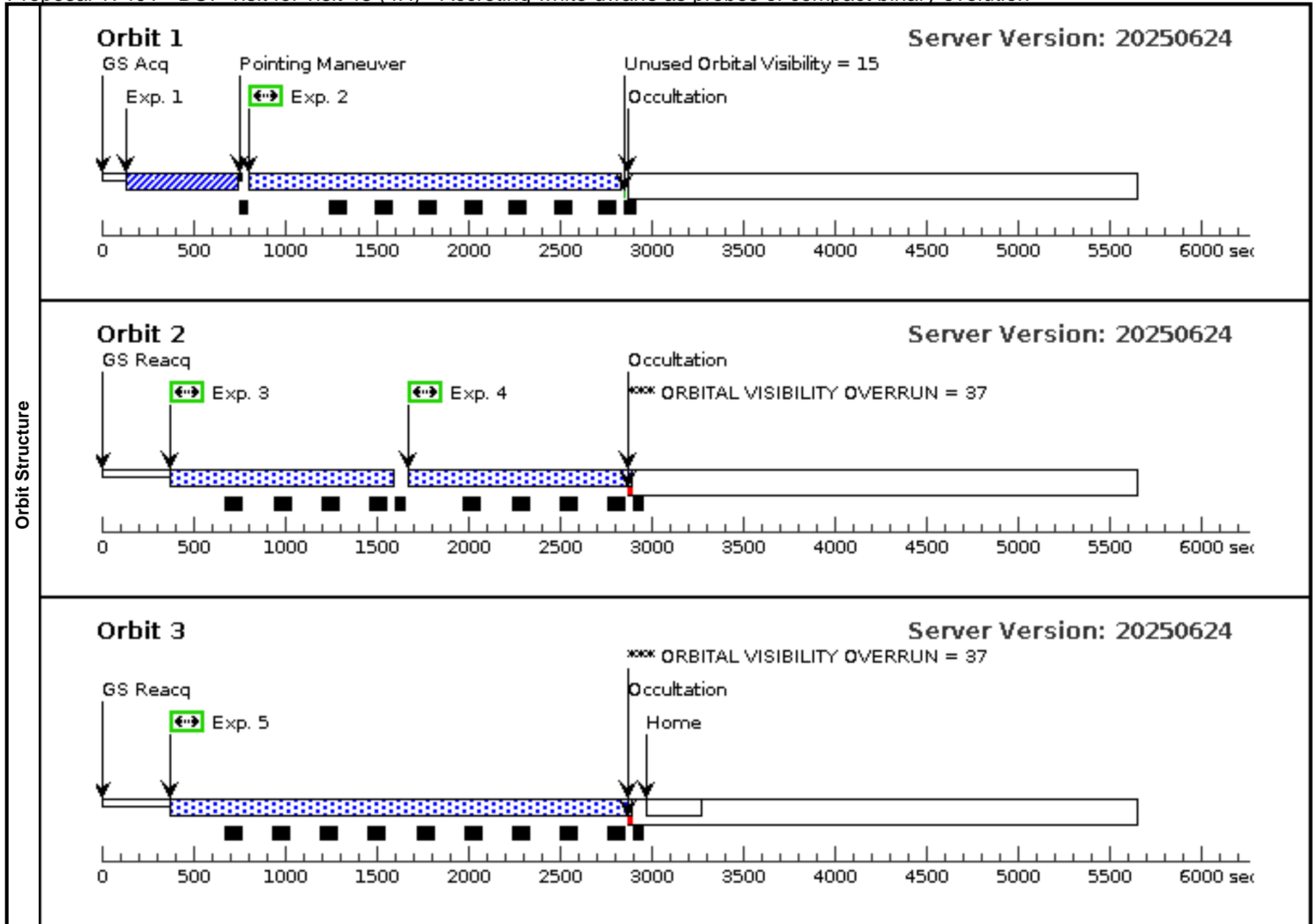
Proposal 17401 - BOP visit for visit 43 (4H) - Accreting white dwarfs as probes of compact binary evolution

Wed Nov 19 17:00:24 GMT 2025

Visit	<p>Proposal 17401, BOP visit for visit 43 (4H), withdrawn</p> <p>Diagnostic Status: Warning</p> <p>Scientific Instruments: COS/FUV, COS/NUV</p> <p>Special Requirements: SCHED 100%; ORIENT 170D TO 45 D; BETWEEN 01-NOV-2025:00:00:00 AND 01-DEC-2025:00:00:00</p> <p><i>Comments: This visit is for BOP checkign the safe target only and should not execute onboard HST.</i></p>					
Diagnostics	<p>(BOP visit for visit 43 (4H)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p> <p>(BOP visit for visit 43 (4H)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p>					
Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
<p>(43)</p> <p><i>Comments:</i> <i>Category=STAR</i> <i>Description=[ACCRETION DISK, DWARF NOVA, INTERACTING BINARY, IRREGULAR VARIABLE]</i> <i>Extended=NO</i></p>	WD-J104019.50-495129.7	RA: 10 40 19.4785 (160.0811604d) Dec: -49 51 29.43 (-49.85818d) Equinox: J2000	Proper Motion RA: -11.746504535541535 mas/yr Proper Motion Dec: 16.566512133372953 mas/yr Parallax: 0.003783567563747677" Epoch of Position: 2016.0	V=17 Gaia EDR3 5360970516650651136, G=17.1, parallax= 3.78(5) mas.	Reference Frame: ICRS	
	WD-J104019.50-4951-SAFE-TARGET	Offset from WD-J104019.50-495129.7 RA Offset: -0.46 Secs Dec Offset: 8.432 Arcsec	V=17	Offset Position (WD-J104019.50-4951-SAFE-TARGET)		
<p><i>Comments: This target is a blank piece of sky which is the bright object safe pointing and is 9.576 arcseconds away at PA 331.7 degrees from WD-J104019</i></p> <p><i>Category=UNIDENTIFIED</i> <i>Description=[BLANK FIELD]</i> <i>Extended=NO</i></p>						

Proposal 17401 - BOP visit for visit 43 (4H) - Accreting white dwarfs as probes of compact binary evolution

#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
Exposures	1	ACQ/IMAG E (1895722)	(44) WD-J104019.50 -4951-SAFE-TARG ET	COS/NUV, ACQ/IMAGE, PSA	MIRRORB			150 Secs (150 Secs) [==>]	[1]	
	<i>Comments: Using mirror B to avoid exceeding count rate per pixel in case the system is brighter than expected (but still safe for the detector). The exposure time computed assuming a black body with $T_{\text{eff}}=17000\text{K}$, scaled to $FUV=17$, and has been increase by an amount equal to the duration of the full eclipse (90 seconds). This will allow to reach the required S/N even if an eclipse occurs during the acquisition. 50 seconds more have been added to be conservative.</i>									
	2	FP-POS=1 (1529784)	(44) WD-J104019.50 -4951-SAFE-TARG ET	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=1; BUFFER-TIME=24 5; LIFETIME-POS=L P3			1827 Secs (1827 Secs) [==>]	[1]
	<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the quiescent level).</i>									
	3	FP-POS=2 (1529784)	(44) WD-J104019.50 -4951-SAFE-TARG ET	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=2; BUFFER-TIME=26 3; LIFETIME-POS=L P3			1165 Secs (1165 Secs) [==>]	[2]
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the quiescent level).</i>										
4	FP-POS=3 (1529784)	(44) WD-J104019.50 -4951-SAFE-TARG ET	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=3; BUFFER-TIME=26 3; LIFETIME-POS=L P3			1165 Secs (1165 Secs) [==>]	[2]	
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the quiescent level).</i>										
5	FP-POS=4 (1529784)	(44) WD-J104019.50 -4951-SAFE-TARG ET	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=4; BUFFER-TIME=26 1; LIFETIME-POS=L P3			2464 Secs (2464 Secs) [==>]	[3]	
<i>Comments: Buffer-time set conservatively short in case the system is brighter (which will still be safe, but the buffer would overflow if set to the appropriate value for the quiescent level).</i>										



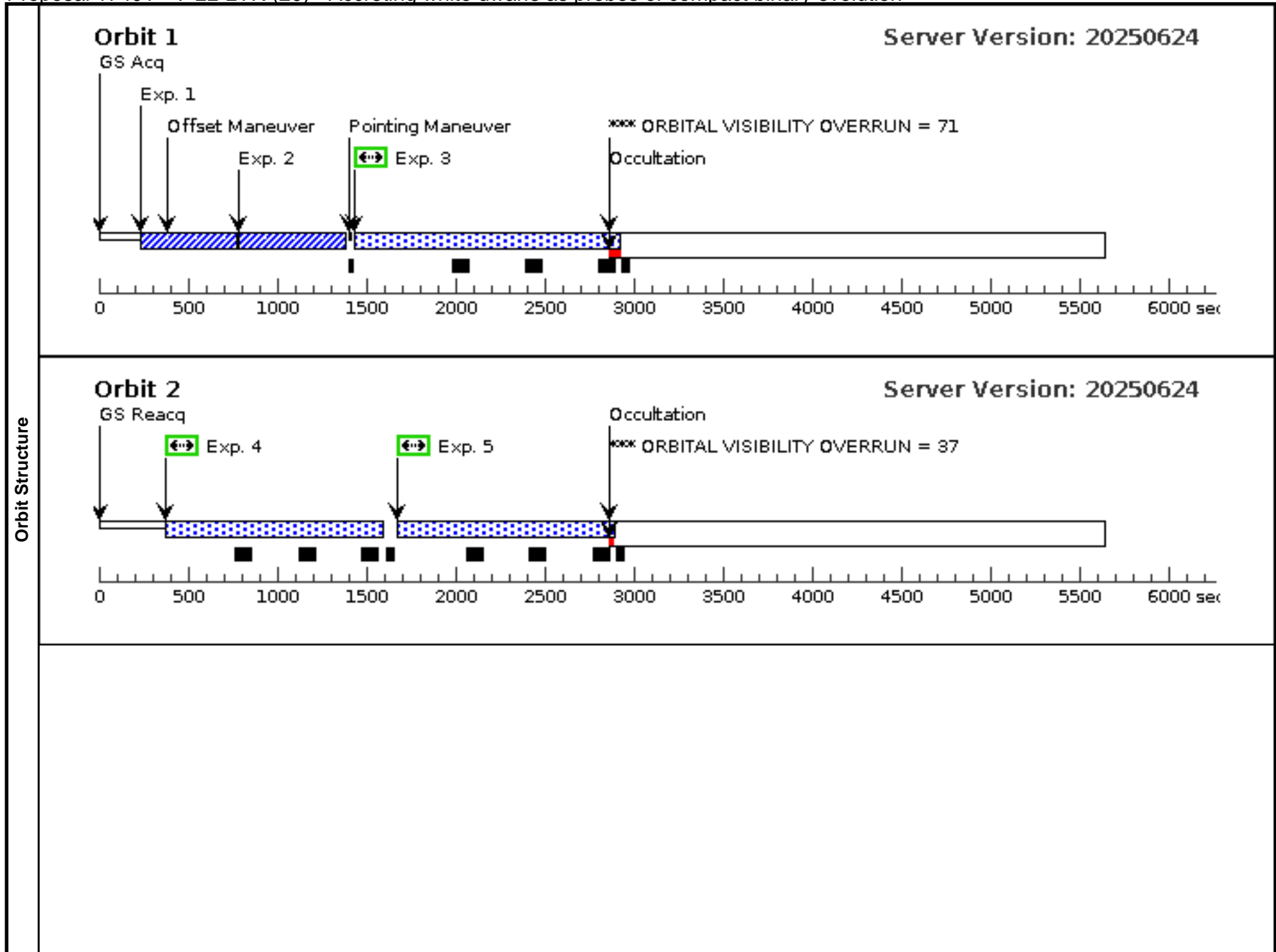
Proposal 17401 - V-LL-LYR (20) - Accreting white dwarfs as probes of compact binary evolution

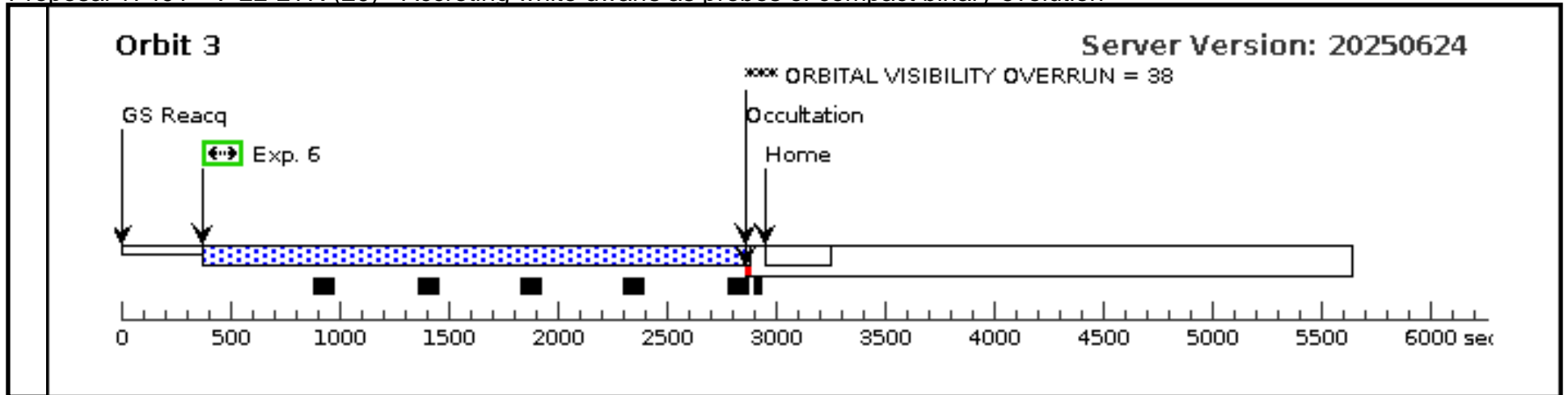
Wed Nov 19 17:00:24 GMT 2025

Visit	<p>Proposal 17401, V-LL-LYR (20), completed</p> <p>Diagnostic Status: Warning</p> <p>Scientific Instruments: COS/FUV</p> <p>Special Requirements: SCHED 100%; ORIENT 220D TO 90 D; BETWEEN 18-MAY-2025:00:00:00 AND 02-JUN-2025:00:00:00</p> <p><i>Comments: This should be scheduled in the evening local-time on a Thursday or Friday. Flags need to be cleared during the work day. Weekends & holidays are to be avoided since the CS must clear the target within 24 hours of HST execution.</i></p>																	
	<p>Diagnosics</p> <p>(V-LL-LYR (20)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p> <p>(V-LL-LYR (20)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p> <p>(V-LL-LYR (20)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p>																	
Fixed Targets	<table border="1"> <thead> <tr> <th>#</th> <th>Name</th> <th>Target Coordinates</th> <th>Targ. Coord. Corrections</th> <th>Fluxes</th> <th>Miscellaneous</th> </tr> </thead> <tbody> <tr> <td>(20)</td> <td>V-LL-LYR</td> <td>RA: 18 35 12.8110 (278.8033792d) Dec: +38 20 4.22 (38.33451d) Equinox: J2000</td> <td>Proper Motion RA: -5.892977038105109 mas/yr Proper Motion Dec: -7.856846564309298 mas/yr Parallax: 0.001185522968873011" Epoch of Position: 2016.0</td> <td>V=18.5 Gaia EDR3 2097903507608061 440, G=17.5, parallax=1.18(7) mas.</td> <td>Reference Frame: ICRS</td> </tr> </tbody> </table>						#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous	(20)	V-LL-LYR	RA: 18 35 12.8110 (278.8033792d) Dec: +38 20 4.22 (38.33451d) Equinox: J2000	Proper Motion RA: -5.892977038105109 mas/yr Proper Motion Dec: -7.856846564309298 mas/yr Parallax: 0.001185522968873011" Epoch of Position: 2016.0	V=18.5 Gaia EDR3 2097903507608061 440, G=17.5, parallax=1.18(7) mas.	Reference Frame: ICRS
	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous												
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<p><i>Comments:</i> <i>Category=STAR</i> <i>Description=[ACCRETION DISK, DWARF NOVA, EMISSION LINE STAR, INTERACTING BINARY, IRREGULAR VARIABLE]</i> <i>Extended=NO</i></p>																		

Proposal 17401 - V-LL-LYR (20) - Accreting white dwarfs as probes of compact binary evolution

Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	ACQ/PEAK XD (1528116)	(20) V-LL-LYR	COS/FUV, ACQ/PEAKXD, PSA	G140L 1105 A	LIFETIME-POS=LP 4	USE OFFSET V20S AF		90 Secs (90 Secs) [==>]	[1]
	<i>Comments: Computed assuming a WD models with $T_{\text{eff}}=35000\text{K}$, scaled to $FUV=17.7$ (as expected for the given distance to the system). The exposure time allows the acquisition in case the system is up to one magnitude fainter than expected.</i>									
	2	ACQ/PEAK D (1528116)	(20) V-LL-LYR	COS/FUV, ACQ/PEAKD, PSA	G140L 1105 A	CENTER=FLUX-W T-FLR; NUM-POS=5; STEP-SIZE=0.9; LIFETIME-POS=L P4	USE OFFSET V20S AF		90 Secs (90 Secs) [==>]	[1]
	<i>Comments: Computed assuming a WD models with $T_{\text{eff}}=35000\text{K}$, scaled to $FUV=17.7$ (as expected for the given distance to the system). The exposure time allows the acquisition in case the system is up to one magnitude fainter than expected.</i>									
	3	FP-POS=1 (1528122)	(20) V-LL-LYR	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=1; BUFFER-TIME=40 9; LIFETIME-POS=L P3	USE OFFSET V20S AF		1339 Secs (1331 Secs) [==>1331.0 Secs]	[1]
	4	FP-POS=2 (1528122)	(20) V-LL-LYR	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=2; BUFFER-TIME=35 4; LIFETIME-POS=L P3	USE OFFSET V20S AF		1174 Secs (1169 Secs) [==>1169.0 Secs]	[2]
5	FP-POS=3 (1528122)	(20) V-LL-LYR	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=3; BUFFER-TIME=35 4; LIFETIME-POS=L P3	USE OFFSET V20S AF		1173 Secs (1168 Secs) [==>1168.0 Secs]	[2]	
6	FP-POS=4 (1528122)	(20) V-LL-LYR	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=4; BUFFER-TIME=47 3; LIFETIME-POS=L P3	USE OFFSET V20S AF		2471 Secs (2462 Secs) [==>2462.0 Secs]	[3]	





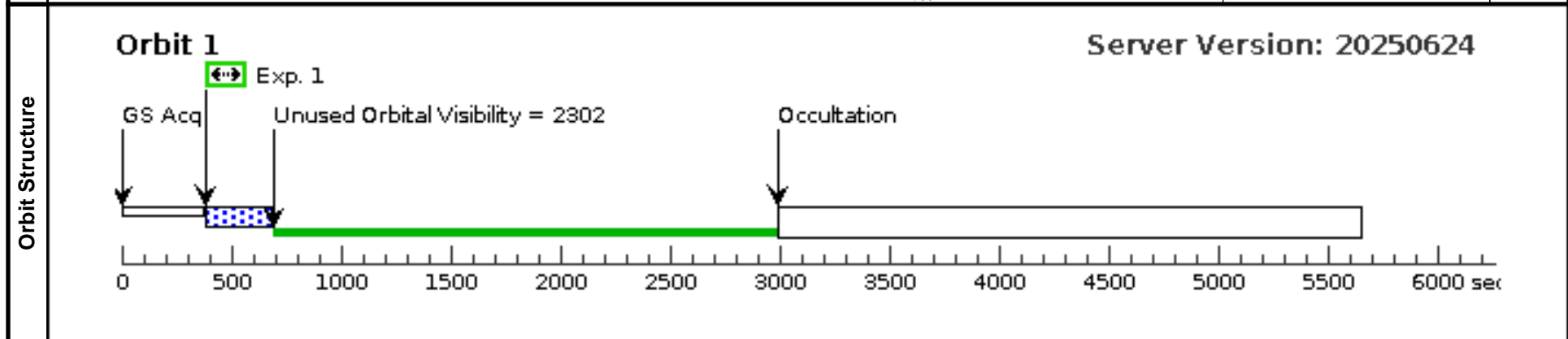
Proposal 17401 - S/C visit for visit 20 (2A) - Accreting white dwarfs as probes of compact binary evolution

Wed Nov 19 17:00:24 GMT 2025

Visit	Proposal 17401, S/C visit for visit 20 (2A), completed Diagnostic Status: No Diagnostics Scientific Instruments: S/C Special Requirements: ORIENT 220D TO 90 D <i>Comments: This visit allocates and sets up the safe position offset slot for visit 20 which will use that slot. This S/C visit should go 3-5 days before visit 20. The S/C visit will contain only 1 exposure</i>				

Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
	(20)	V-LL-LYR	RA: 18 35 12.8110 (278.8033792d) Dec: +38 20 4.22 (38.33451d) Equinox: J2000	Proper Motion RA: -5.892977038105109 mas/yr Proper Motion Dec: -7.856846564309298 mas/yr Parallax: 0.001185522968873011" Epoch of Position: 2016.0	V=18.5 Gaia EDR3 2097903507608061440, G=17.5, parallax=1.18(7) mas.	Reference Frame: ICRS
<i>Comments:</i> Category=STAR Description=[ACCRETION DISK, DWARF NOVA, EMISSION LINE STAR, INTERACTING BINARY, IRREGULAR VARIABLE] Extended=NO						

#	Label	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
1		(20) V-LL-LYR	S/C, DATA, V1			POS TARG 229.132 8,-241.0575; SAVE OFFSET V20 SAF; SPEC COM INSTR ECSLOTSET; QESIPARM ANGL E 19.6; QESIPARM DIST 9. 739		310 Secs (310 Secs) [==>]	[1]



Proposal 17401 - BOP visit for visit 20 (2B) - Accreting white dwarfs as probes of compact binary evolution

Wed Nov 19 17:00:24 GMT 2025

Visit	<p>Proposal 17401, BOP visit for visit 20 (2B), withdrawn</p> <p>Diagnostic Status: Warning</p> <p>Scientific Instruments: COS/FUV</p> <p>Special Requirements: SCHED 100%; ORIENT 220D TO 90 D; BETWEEN 18-MAY-2025:00:00:00 AND 02-JUN-2025:00:00:00</p> <p><i>Comments: This visit is for BOP checking the safe target only and should not execute onboard HST</i></p>																																		
	<p>Diagnosics</p> <p>(BOP visit for visit 20 (2B)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p> <p>(BOP visit for visit 20 (2B)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p> <p>(BOP visit for visit 20 (2B)) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN</p>																																		
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	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous																													
(20)	V-LL-LYR	RA: 18 35 12.8110 (278.8033792d) Dec: +38 20 4.22 (38.33451d) Equinox: J2000	Proper Motion RA: -5.892977038105109 mas/yr Proper Motion Dec: -7.856846564309298 mas/yr Parallax: 0.001185522968873011" Epoch of Position: 2016.0	V=18.5 Gaia EDR3 2097903507608061440, G=17.5, parallax=1.18(7) mas.	Reference Frame: ICRS																														
<p><i>Comments: Category=STAR Description=[ACCRETION DISK, DWARF NOVA, EMISSION LINE STAR, INTERACTING BINARY, IRREGULAR VARIABLE] Extended=NO</i></p>																																			
(21)	V-LL-LYR-SAFE-TGT	Offset from V-LL-LYR RA Offset: 0.27 Secs Dec Offset: 9.172 Arcsec		V=18.5	Offset Position (V-LL-LYR-SAFE-TGT)																														
<p><i>Comments: This target is a blank piece of sky which is the bright object safe pointing and is 9.739 arcseconds away at PA 19.6 degrees from V-LL-LYR Category=UNIDENTIFIED Description=[BLANK FIELD] Extended=NO</i></p>																																			
<p>Fixed Targets</p>																																			

Proposal 17401 - BOP visit for visit 20 (2B) - Accreting white dwarfs as probes of compact binary evolution

Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	ACQ/PEAK XD (1528116)	(21) V-LL-LYR-SA FE-TGT	COS/FUV, ACQ/PEAKXD, PSA	G140L 1105 A	LIFETIME-POS=LP 4			90 Secs (90 Secs) [==>]	[1]
	<i>Comments: Computed assuming a WD models with Teff=35000K, scaled to FUV=17.7 (as expected for the given distance to the system). The exposure time allows the acquisition in case the system is up to one magnitude fainter than expected.</i>									
	2	ACQ/PEAK D (1528116)	(21) V-LL-LYR-SA FE-TGT	COS/FUV, ACQ/PEAKD, PSA	G140L 1105 A	CENTER=FLUX-W T-FLR; NUM-POS=5; STEP-SIZE=0.9; LIFETIME-POS=L P4			90 Secs (90 Secs) [==>]	[1]
	<i>Comments: Computed assuming a WD models with Teff=35000K, scaled to FUV=17.7 (as expected for the given distance to the system). The exposure time allows the acquisition in case the system is up to one magnitude fainter than expected.</i>									
	3	FP-POS=1 (1528122)	(21) V-LL-LYR-SA FE-TGT	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=1; BUFFER-TIME=40 9; LIFETIME-POS=L P3			1339 Secs (1331 Secs) [==>1331.0 Secs]	[1]
	4	FP-POS=2 (1528122)	(21) V-LL-LYR-SA FE-TGT	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=2; BUFFER-TIME=35 4; LIFETIME-POS=L P3			1174 Secs (1169 Secs) [==>1169.0 Secs]	[2]
5	FP-POS=3 (1528122)	(21) V-LL-LYR-SA FE-TGT	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=3; BUFFER-TIME=35 4; LIFETIME-POS=L P3			1173 Secs (1168 Secs) [==>1168.0 Secs]	[2]	
6	FP-POS=4 (1528122)	(21) V-LL-LYR-SA FE-TGT	COS/FUV, TIME-TAG, PSA	G140L 800 A	FLASH=YES; FP-POS=4; BUFFER-TIME=47 3; LIFETIME-POS=L P3			2471 Secs (2462 Secs) [==>2462.0 Secs]	[3]	

