



17421 - Shocking detections: Characterising exocometary shock fronts by tracking star-grazing comets using the UV AIII line

Cycle: 31, Proposal Category: GO

(UV Initiative)

(Availability Mode: SUPPORTED)

INVESTIGATORS

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Dr. Grant Kennedy (CoI) (ESA Member)	The University of Warwick
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Mr. Theo Vrignaud (CoI) (ESA Member)	CNRS, Institut d'Astrophysique de Paris

VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) BETA-PIC WAVE	STIS/CCD STIS/NUV-MAMA	3	20-Jun-2024 15:05:00.0	yes
02	(1) BETA-PIC WAVE	STIS/CCD STIS/NUV-MAMA	3	20-Jun-2024 15:05:04.0	yes

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
03	(1) BETA-PIC WAVE	STIS/CCD STIS/NUV-MAMA	4	20-Jun-2024 15:05:07.0	yes

10 Total Orbits Used

ABSTRACT

Exocomets are small icy bodies, which orbit stars other than the Sun and sublimate when close to their star, producing a tail of dust and gas (like solar system comets). During transit, the gas causes spectroscopic absorption signatures that are routinely detected in several systems. The most prolific case is beta Pic, which continuously displays narrow, variable, non-photospheric absorption features superimposed on the much broader stellar lines. The features vary on short timescales (hours) and are seen at optical (e.g. CaII) and UV (e.g. FeII, SiIV) wavelengths.

Previous HST observations of beta Pic in the UV showed the presence of AlIII lines in the exocomet gas. The mere existence of AlIII is puzzling as beta Pic is unable to twice photoionise Al. Hydrodynamical models suggest that these highly-ionised species are formed in a shock where compression and thus collisions within the shock surface are sufficient to generate AlIII. This is theorised to occur when the exocomets get sufficiently close to the star (a few stellar radii).

We propose to monitor beta Pic to characterise the doppler movement of the AlIII lines and from this directly measure the exocomets' acceleration. This we use to calculate the comet-star distances. The highly ionised species are thought only to be present in the shock fronts of exocomets close to the star. Measuring the distance between the star and the exocomets will allow us to directly test this hypothesis. The observations cover a wavelength range with other species thought to only emerge in shocks. Measuring the transit ingress/egress times of these species will give us the composition of the shock as a function of radial distance.

OBSERVING DESCRIPTION

One line summary: We plan to observe exocometary absorption events in the beta Pictoris planetary system with STIS and to track these features by monitoring how the Al III absorption line varies as a function of time.

The AlIII line is the only line capable of unequivocally detecting an exocometary shocks, given that the second most favourable line (SiIV, which we previously analysed) yields results right at the edge of statistical significance.

In order to measure the acceleration as accurately as possible and to avoid blended absorption features, we will use the higher resolution and sensitivity of STIS (compared to COS) using the E230H setup giving. Our setup with the central wavelength set at 1763 Ang, will provide a wavelength coverage from 1624 to 1901Ang. This wavelength range includes many very interesting lines, in which GHRS data has shown detections of transiting exocomets. The line list in this wavelength range includes Fe II, Al II, Si, SiII}, CI, and NiII. Due to the brightness of beta Pic (our target), STIS is in addition required to avoid saturating the detector. Due to the brightness of the object we will be observing in ACCUM mode.

Justification for visits and orbits:

To ensure the two main science goals (1. Test the hypothesis that AIII are caused by shocks and 2. Measure the species dependent size of the shock front) are met we deem it necessary to obtain observations on three separate visits. Three visits ensures completely different exocometry absorption signatures and would avoid the possible conclusion that a detections from a single visit is an outlier. We would further avoid the blending of exocometary signatures. We initially wrote two visits, but believe it is more important to have three visits (with less orbits) than 2 visits with more orbits due to the unpredictable nature of the exocomet absorption features. This change was also in part motivated by the strong encouragement from STScI to avoid 3+ orbit visits.

There will in total be three visits. The two first visits will consist of 3 orbits and the last visit will consist of 4 orbits.

Proposal 17421 - Visit 01 - Shocking detections: Characterising exocometary shock fronts by tracking star-grazing comets using the U...

Visit	Proposal 17421, Visit 01, completed Thu Jun 20 19:05:08 GMT 2024 Diagnostic Status: Warning Scientific Instruments: STIS/NUV-MAMA, STIS/CCD Special Requirements: BETWEEN 16-APR-2024:23:00:00 AND 17-APR-2024:04:00:00																	
	Diagnosics (Visit 01) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN (Visit 01) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN																	
Fixed Targets	<table border="1"> <thead> <tr> <th>#</th> <th>Name</th> <th>Target Coordinates</th> <th>Targ. Coord. Corrections</th> <th>Fluxes</th> <th>Miscellaneous</th> </tr> </thead> <tbody> <tr> <td>(1)</td> <td>BETA-PIC</td> <td>RA: 05 47 17.0953 (86.8212304d) Dec: -51 03 58.15 (-51.06615d) Equinox: J2000</td> <td>Proper Motion RA: 4.932985550191305E-4 sec of time/yr Proper Motion Dec: 0.0831 arcsec/yr Epoch of Position: 2015.5</td> <td>V=3.86</td> <td>Reference Frame: ICRS</td> </tr> </tbody> </table>						#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous	(1)	BETA-PIC	RA: 05 47 17.0953 (86.8212304d) Dec: -51 03 58.15 (-51.06615d) Equinox: J2000	Proper Motion RA: 4.932985550191305E-4 sec of time/yr Proper Motion Dec: 0.0831 arcsec/yr Epoch of Position: 2015.5	V=3.86	Reference Frame: ICRS
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Comments: This object was generated by the targetselector and retrieved from the SIMBAD database. Category=STAR Description=[A4-A9 V-IV, DISK, EXTRA-SOLAR PLANETARY SYSTEM] Extended=NO																		

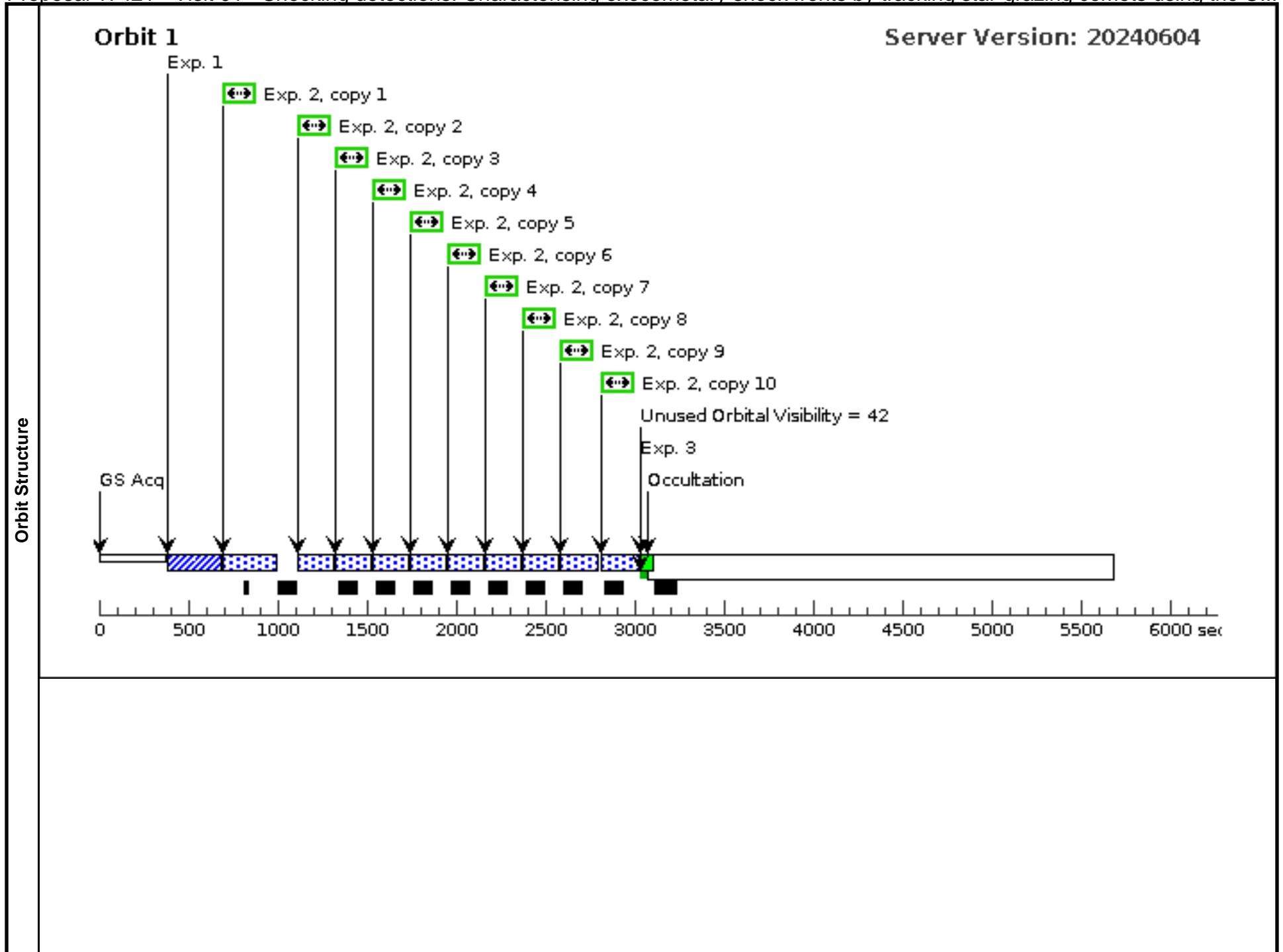
Proposal 17421 - Visit 01 - Shocking detections: Characterising exocometary shock fronts by tracking star-grazing comets using the U...

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3	WAVECAL WAVE		STIS/NUV-MAMA, ACCUM, 0.2X0.2	E230H 1763 A			Sequence 1-3 Non-Int in Visit 01	[==>]	[1]
4	(1890162)	(1) BETA-PIC	STIS/NUV-MAMA, ACCUM, 0.2X0.2	E230H 1763 A	WAVECAL=NO		Sequence 4-5 Non-Int in Visit 01	200 Secs X 13 (2392 Secs) [==>184.0 Secs (Copy 1)] [==>184.0 Secs (Copy 2)] [==>184.0 Secs (Copy 3)] [==>184.0 Secs (Copy 4)] [==>184.0 Secs (Copy 5)] [==>184.0 Secs (Copy 6)] [==>184.0 Secs (Copy 7)] [==>184.0 Secs (Copy 8)] [==>184.0 Secs (Copy 9)] [==>184.0 Secs (Copy 10)] [==>184.0 Secs (Copy 11)] [==>184.0 Secs (Copy 12)] [==>184.0 Secs (Copy 13)]	[2]
5	WAVECAL WAVE		STIS/NUV-MAMA, ACCUM, 0.2X0.2	E230H 1763 A			Sequence 4-5 Non-Int in Visit 01	[==>]	[2]

Exposures

Proposal 17421 - Visit 01 - Shocking detections: Characterising exocometary shock fronts by tracking star-grazing comets using the U...

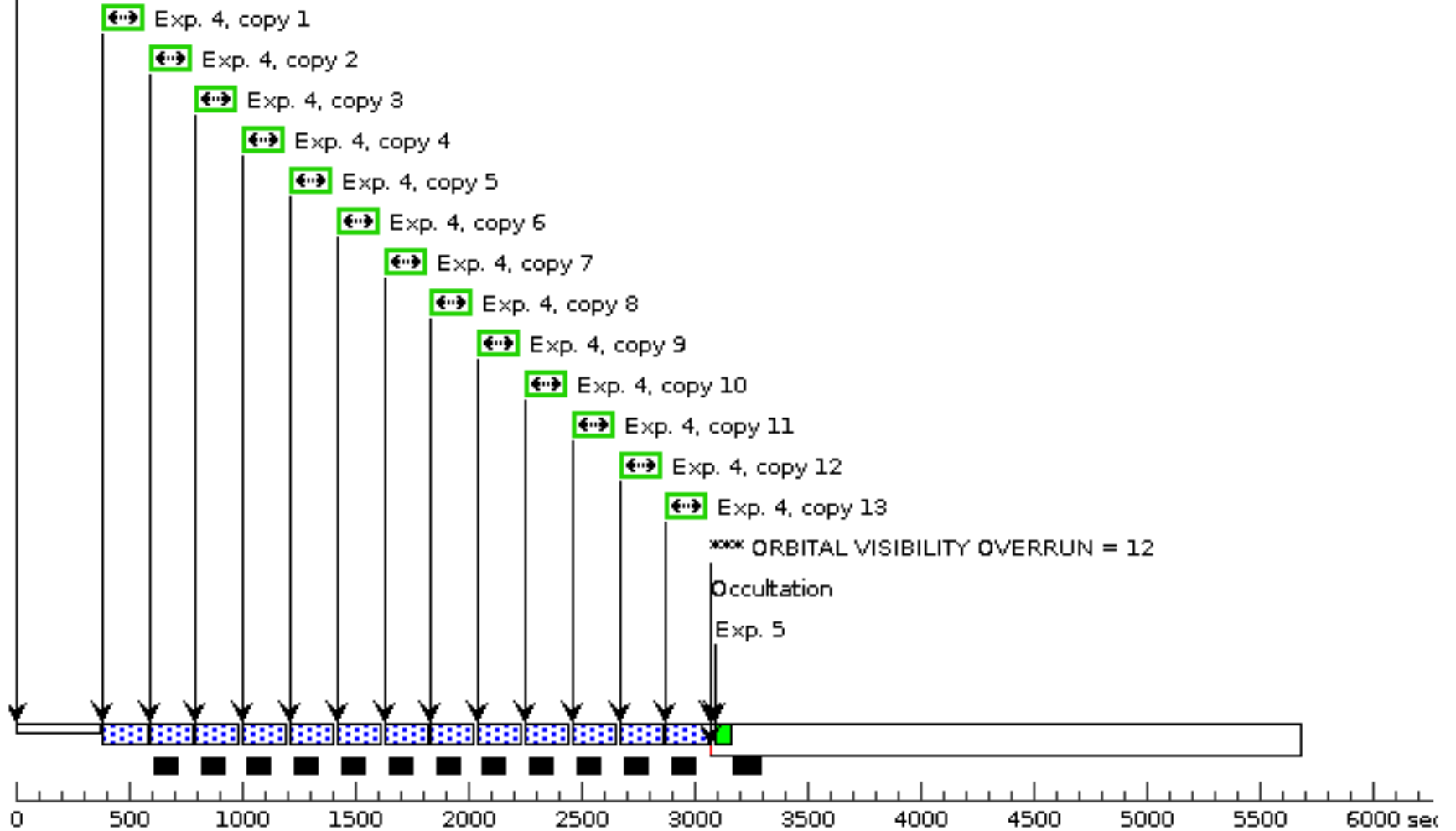
6	(1890162)	(1) BETA-PIC	STIS/NUV-MAMA, ACCUM, 0.2X0.2	E230H 1763 A	WAVECAL=NO	Sequence 6-7 Non-In t in Visit 01	200 Secs X 13 (2392 Secs) [==>184.0 Secs (Copy 1)] [==>184.0 Secs (Copy 2)] [==>184.0 Secs (Copy 3)] [==>184.0 Secs (Copy 4)] [==>184.0 Secs (Copy 5)] [==>184.0 Secs (Copy 6)] [==>184.0 Secs (Copy 7)] [==>184.0 Secs (Copy 8)] [==>184.0 Secs (Copy 9)] [==>184.0 Secs (Copy 10)] [==>184.0 Secs (Copy 11)] [==>184.0 Secs (Copy 12)] [==>184.0 Secs (Copy 13)]	[3]
7	WAVECAL WAVE	STIS/NUV-MAMA, ACCUM, 0.2X0.2	E230H 1763 A	WAVECAL=NO	Sequence 6-7 Non-In t in Visit 01	[==>]	[3]	

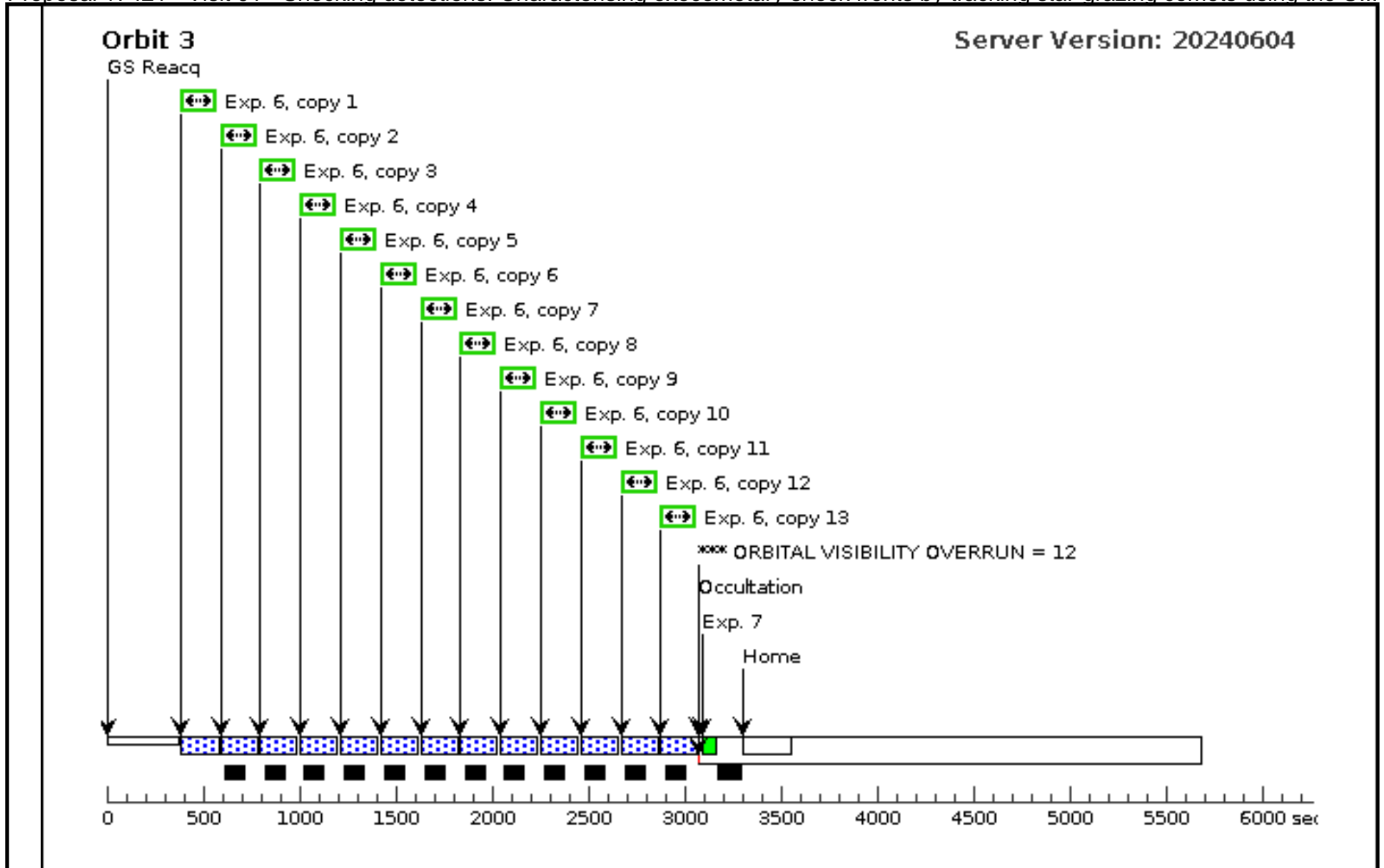


Orbit 2

Server Version: 20240604

GS Reacq





Proposal 17421 - Visit 02 - Shocking detections: Characterising exocometary shock fronts by tracking star-grazing comets using the U...

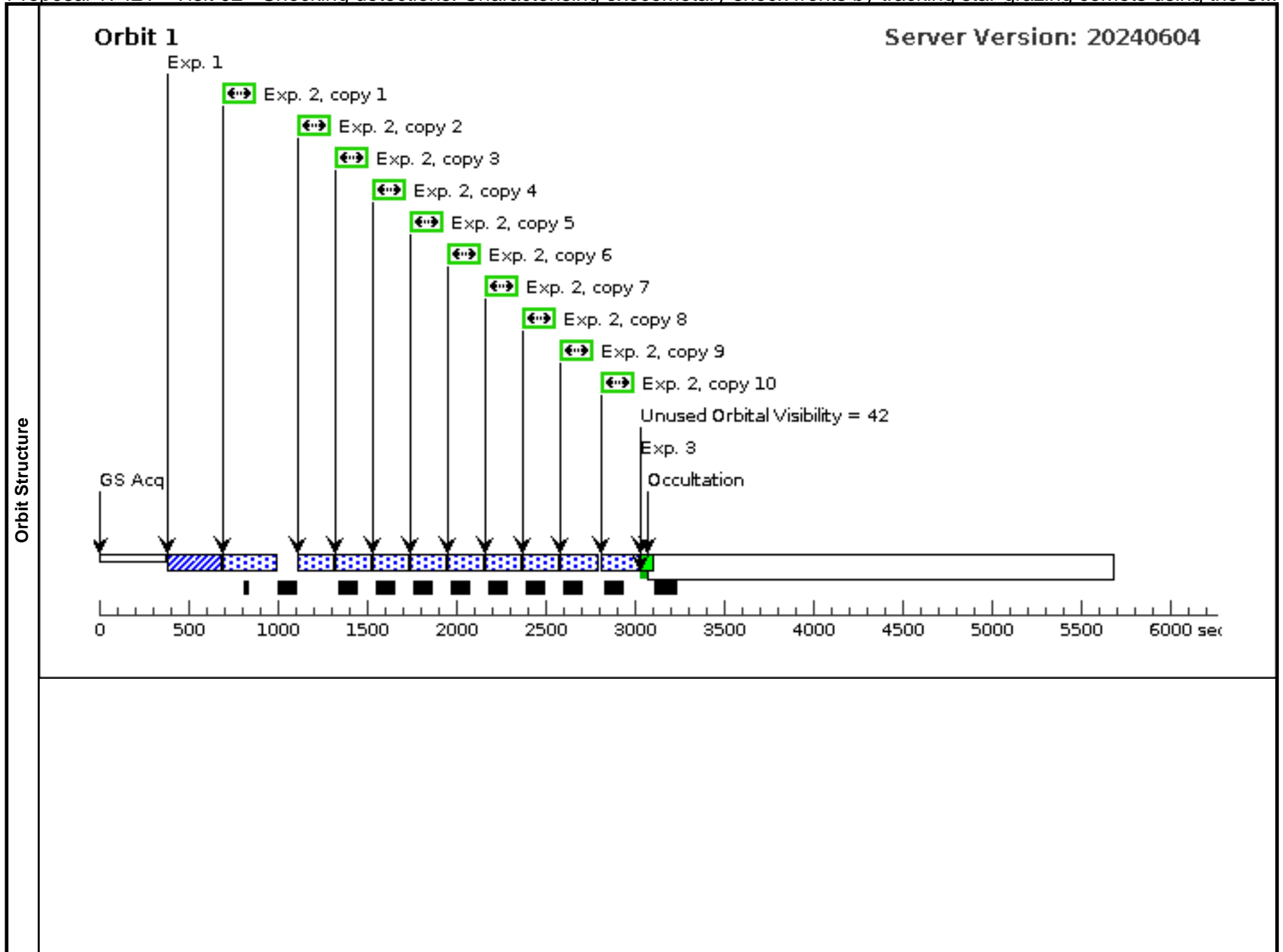
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	Diagnosics (Visit 02) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN (Visit 02) Warning (Orbit Planner): ORBITAL VISIBILITY OVERRUN																	
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Exposures	1	ACQ (1890147)	(1) BETA-PIC	STIS/CCD, ACQ, F25ND5	MIRROR		Sequence 1-3 Non-Int in Visit 02	1 Secs (1 Secs) [==>]	[1]
	2	(1890162)	(1) BETA-PIC	STIS/NUV-MAMA, ACCUM, 0.2X0.2	E230H 1763 A	WAVECAL=NO	Sequence 1-3 Non-Int in Visit 02	200 Secs X 10 (1886 Secs) [==>186.0 Secs (Copy 1)] [==>186.0 Secs (Copy 2)] [==>186.0 Secs (Copy 3)] [==>186.0 Secs (Copy 4)] [==>186.0 Secs (Copy 5)] [==>186.0 Secs (Copy 6)] [==>186.0 Secs (Copy 7)] [==>186.0 Secs (Copy 8)] [==>199.0 Secs (Copy 9)] [==>199.0 Secs (Copy 10)]	[1]
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	4	(1890162)	(1) BETA-PIC	STIS/NUV-MAMA, ACCUM, 0.2X0.2	E230H 1763 A	WAVECAL=NO	Sequence 4-5 Non-Int in Visit 02	200 Secs X 13 (2392 Secs) [==>184.0 Secs (Copy 1)] [==>184.0 Secs (Copy 2)] [==>184.0 Secs (Copy 3)] [==>184.0 Secs (Copy 4)] [==>184.0 Secs (Copy 5)] [==>184.0 Secs (Copy 6)] [==>184.0 Secs (Copy 7)] [==>184.0 Secs (Copy 8)] [==>184.0 Secs (Copy 9)] [==>184.0 Secs (Copy 10)] [==>184.0 Secs (Copy 11)] [==>184.0 Secs (Copy 12)] [==>184.0 Secs (Copy 13)]	[2]
	5	WAVECAL WAVE		STIS/NUV-MAMA, ACCUM, 0.2X0.2	E230H 1763 A		Sequence 4-5 Non-Int in Visit 02	[==>]	[2]

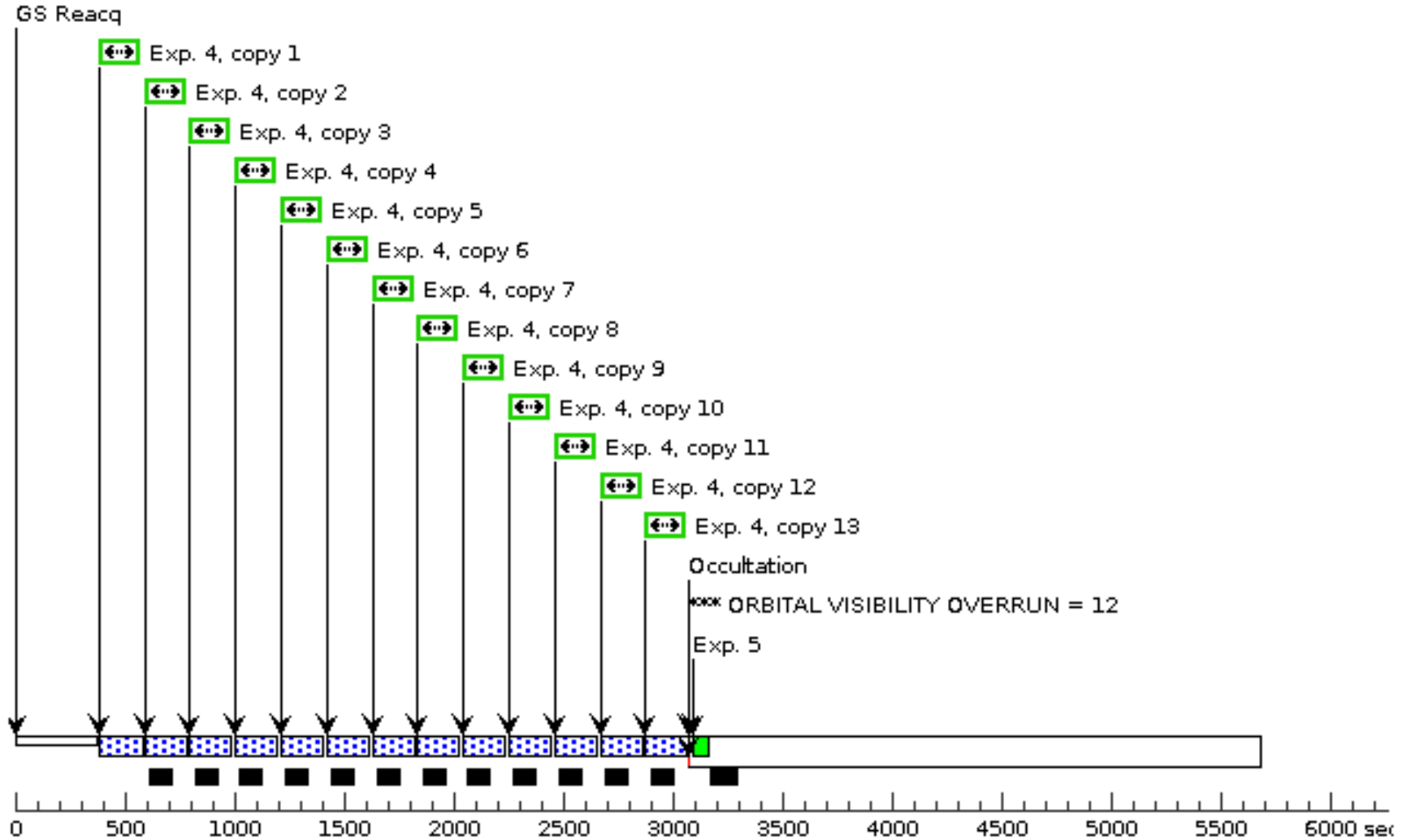
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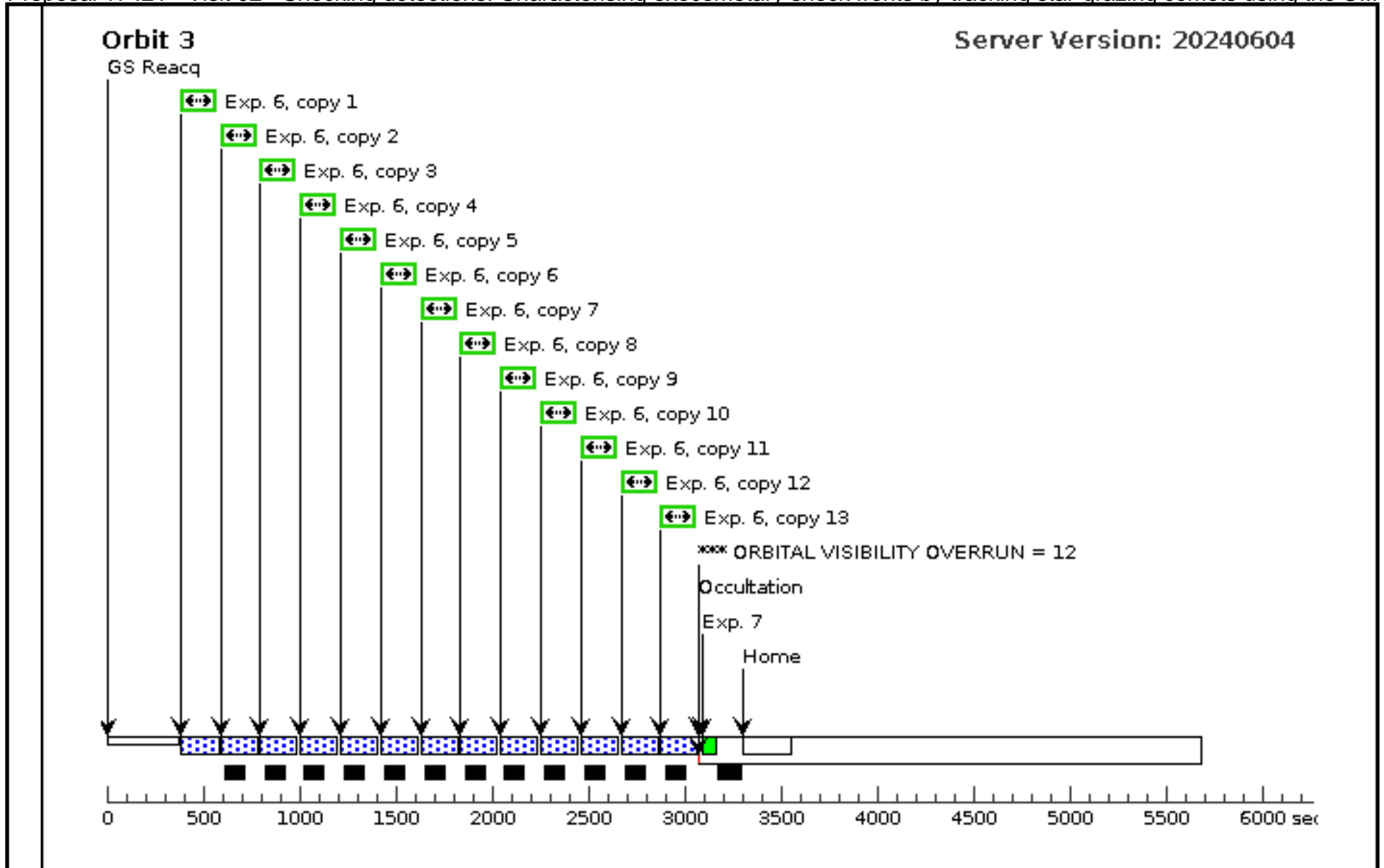
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7	WAVECAL WAVE	STIS/NUV-MAMA, ACCUM, 0.2X0.2	E230H 1763 A		Sequence 6-7 Non-Int in Visit 02	[==>]	[3]



Orbit 2

Server Version: 20240604





Proposal 17421 - Visit 03 - Shocking detections: Characterising exocometary shock fronts by tracking star-grazing comets using the U...

Visit	Proposal 17421, Visit 03, implementation Thu Jun 20 19:05:09 GMT 2024 Diagnostic Status: Warning Scientific Instruments: STIS/NUV-MAMA, STIS/CCD Special Requirements: AFTER 02 BY 5 D TO 1000 D																	
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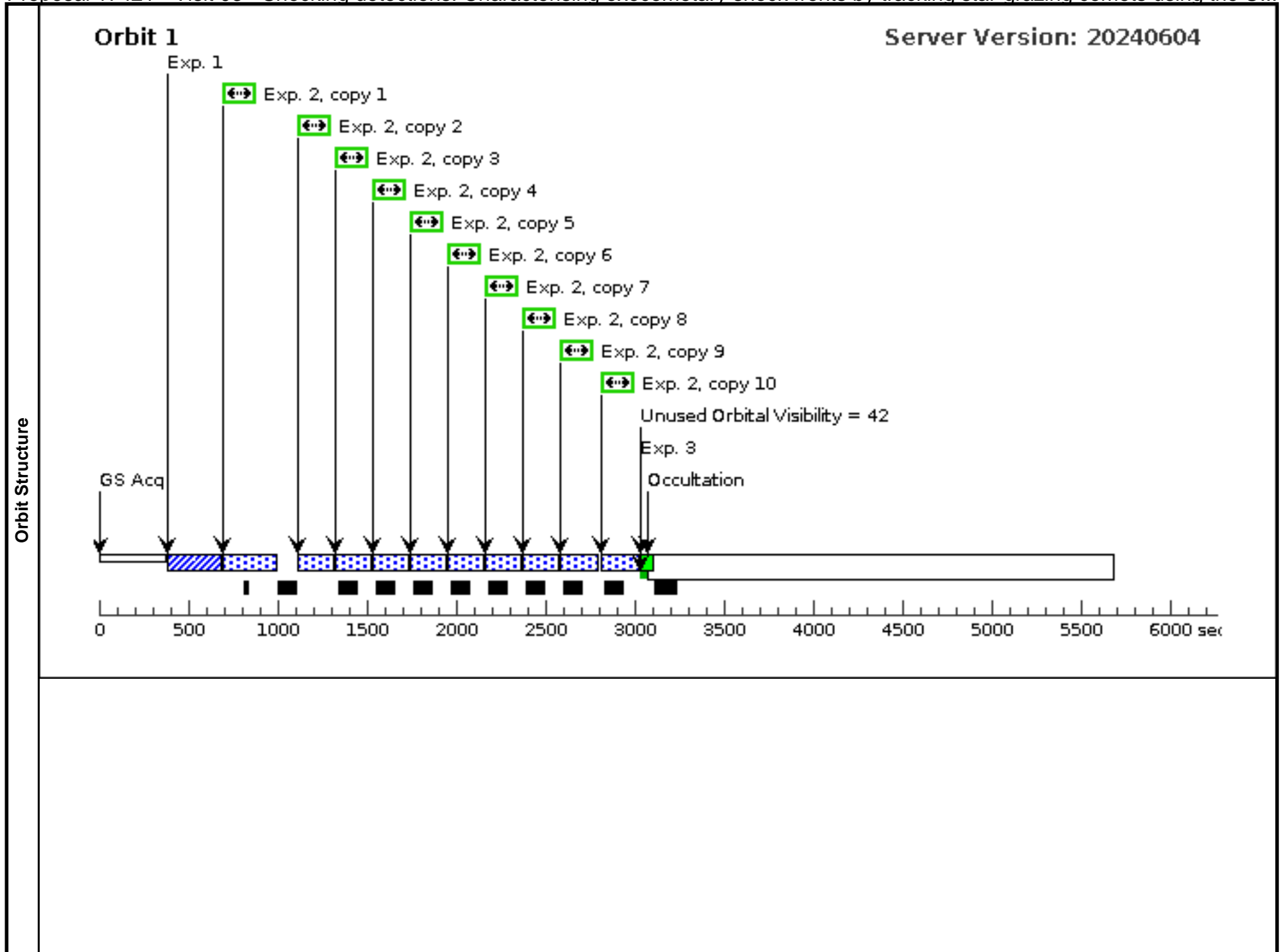
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Exposures

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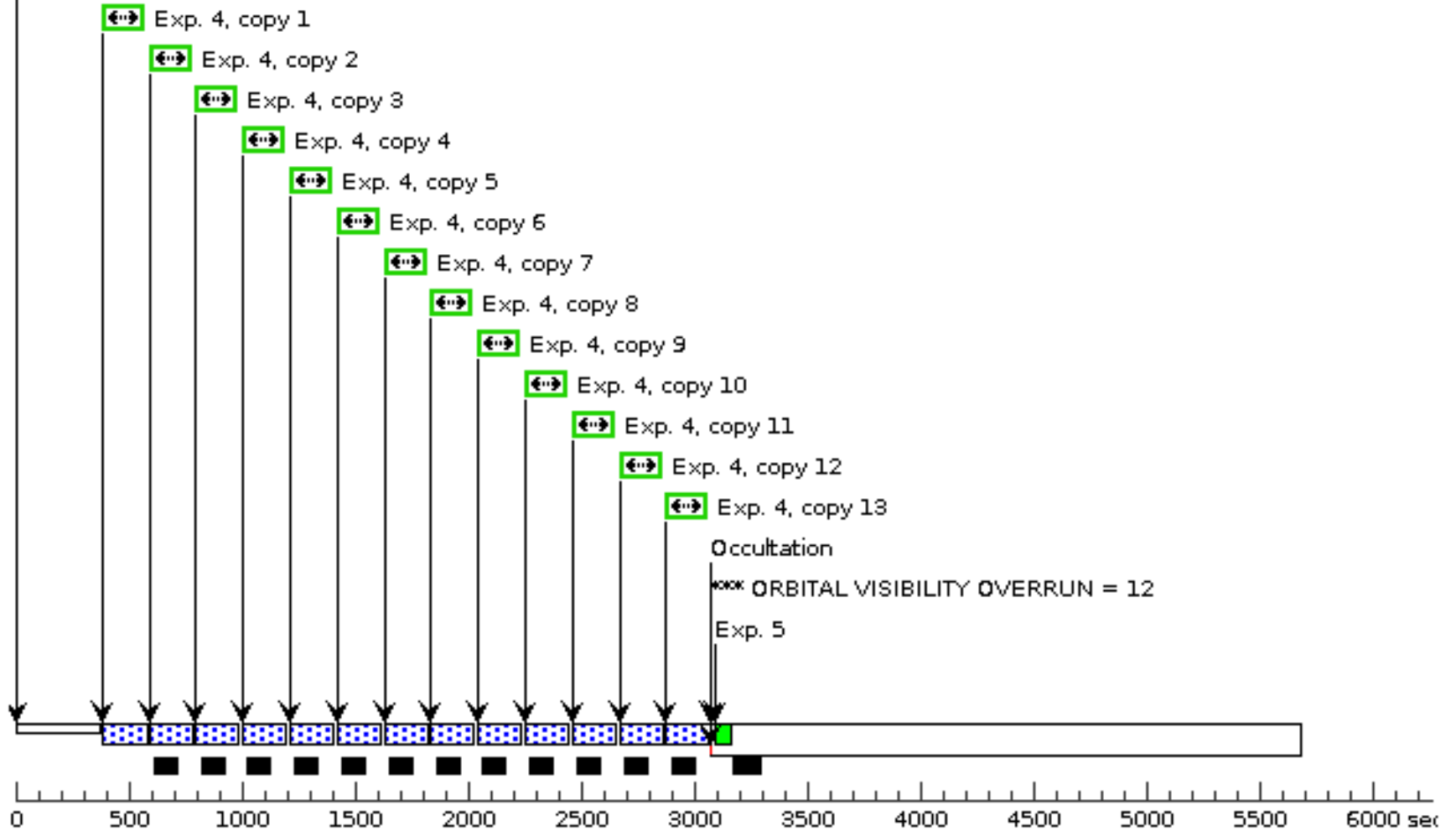
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7		WAVECAL WAVE	STIS/NUV-MAMA, ACCUM, 0.2X0.2	E230H 1763 A		Sequence 6-7 Non-Int in Visit 03	[==>]	[3]
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9		WAVECAL WAVE	STIS/NUV-MAMA, ACCUM, 0.2X0.2	E230H 1763 A		Sequence 8-9 Non-Int in Visit 03	[==>]	[4]



Orbit 2

Server Version: 20240604

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Orbit 3

Server Version: 20240604

