



17533 - Identifying a New Source of r-Process Nucleosynthesis with HST

Cycle: 31, Proposal Category: GO

(Availability Mode: SUPPORTED)

INVESTIGATORS

<i>Name</i>	<i>Institution</i>
Jillian Rastinejad (PI) (Contact)	University of Maryland
Dr. Wen-fai Fong (CoI)	Northwestern University
Prof. Andrew James Levan (CoI) (ESA Member)	Radboud University Nijmegen
Prof. Nial Rahil Tanvir (CoI) (ESA Member)	University of Leicester
Dr. Peter Blanchard (CoI)	Harvard University
Prof. Ashley Villar (CoI)	Harvard University

VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
05	(3) 250108A	WFC3/UVIS	2	08-Sep-2025 10:00:14.0	yes
06	(3) 250108A	WFC3/IR	2	08-Sep-2025 10:00:15.0	yes
07	(3) 250108A	WFC3/UVIS	2	08-Sep-2025 10:00:15.0	yes
08	(3) 250108A	WFC3/IR	2	08-Sep-2025 10:00:16.0	yes

8 Total Orbits Used

ABSTRACT

The origin of the Universe's heavy r-process elements is a fundamental issue which affects chemical enrichment across cosmic time and the development of complex life. It is now confirmed that neutron star mergers are responsible for at least some of the r-process budget, but there is gaining observational and theoretical support for an additional and faster production channel: the core-collapse of rapidly-rotating, massive stars. In

particular, energetic supernovae with associated long-duration gamma-ray bursts (GRB-SNe) are the strongest candidates for a fast channel. Here, their massive neutronized accretion disks mix with the outer, Ni-dominated ejecta producing dramatic reddening at late times (~50-200 days). Even for the most nearby events, HST is necessary to detect the SN at these epochs when the color of an r-process enriched SN will be distinguishable from an event with no r-process, especially against the background light of a star-forming host galaxy. Yet, as these GRB-SNe r-process models gain traction, no tailored observations with the relevant sensitivity or wavelength coverage exist. We propose a first-of-its-kind study to monitor the late-time evolution of one nearby ($z < 0.3$) GRB-SN as a crucial test of current models. We will obtain 12 HST orbits over 5 epochs in F606W and F160W. The detection of r-process nucleosynthesis in a GRB-SN would be transformative for our understanding of when and where heavy elements enter galaxies like our Milky Way. Given both the long time baseline and rarity of nearby LGRBs, we request long-term and carry-over status.

OBSERVING DESCRIPTION

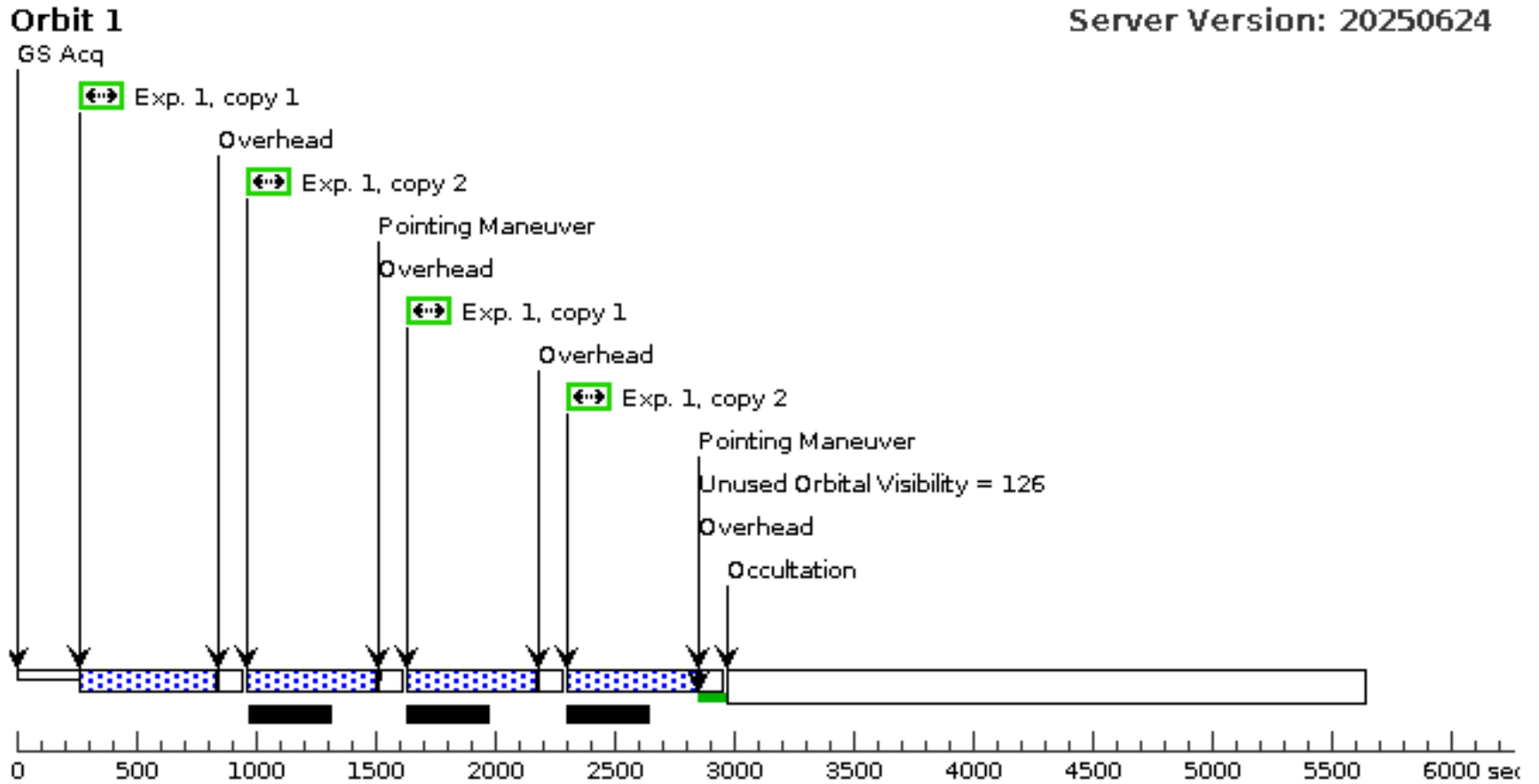
This is a non-disruptive target-of-opportunity proposal with long-term and carry-over status. This proposal is for 12 orbits (8 orbits awarded in Cycle 31, 4 in Cycle 32) of HST imaging of a nearby GRB supernova (SN), a promising but untested site to observe the formation of elements via rapid neutron capture (r-process) nucleosynthesis. These observations will observe the SN's color evolution by observing contemporaneously in one optical (F606W) filter and one near-IR (F160W) filter over ~50 - 600 days following the GRB, when ground-based telescopes are no longer sensitive to the SN's expected brightness. Observations will be obtained across 5 visits, with the first 4 visits consisting of one orbit each of WFC3/UVIS F606W and WFC3/IR F160W imaging. The final visit will be 2 orbits in each filter in order to provide a deep template for image subtraction with earlier epochs, intended to remove host galaxy light contamination.

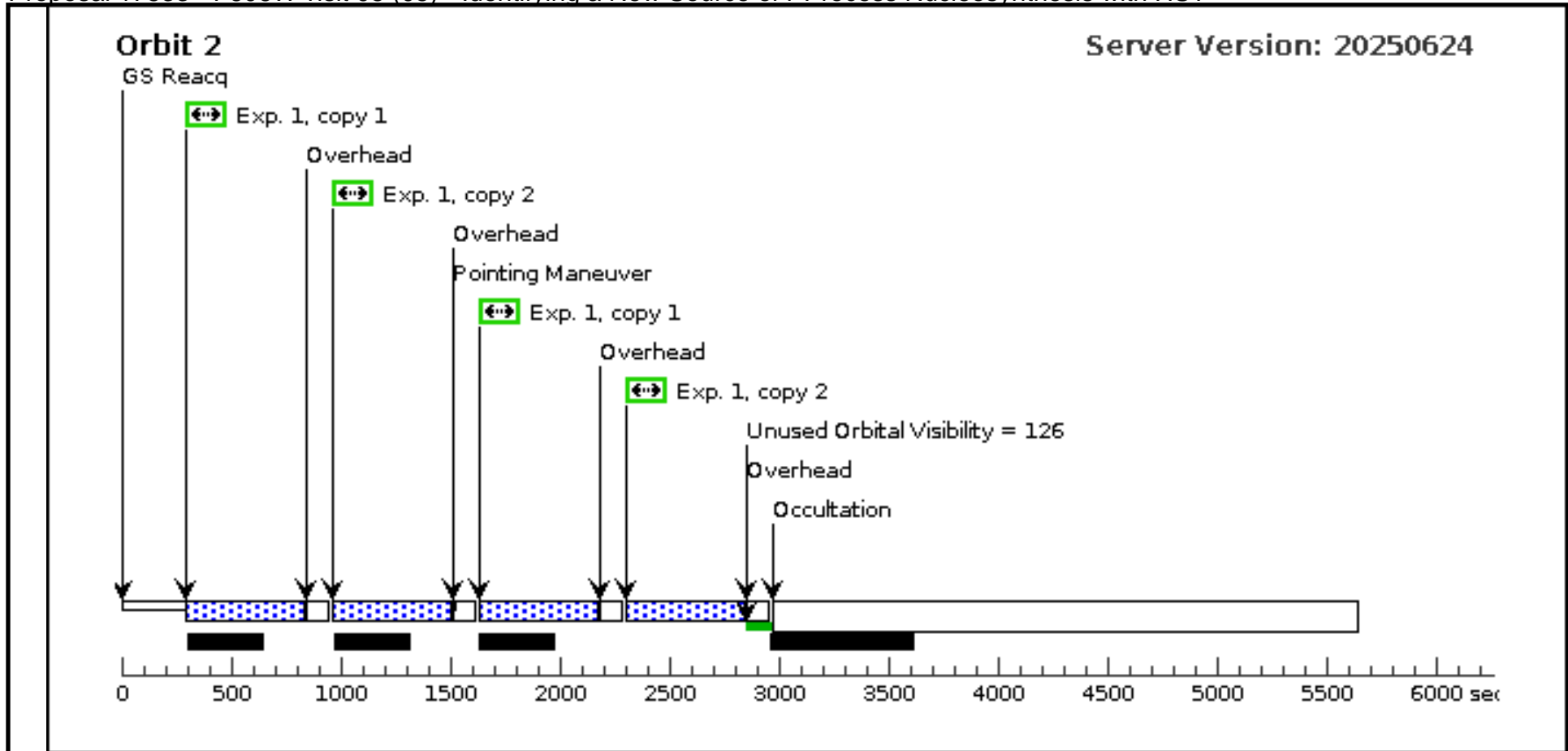
Proposal 17533 - F606W visit 05 (05) - Identifying a New Source of r-Process Nucleosynthesis with HST

Mon Sep 08 14:00:16 GMT 2025

Visit	Proposal 17533, F606W visit 05 (05), implementation Diagnostic Status: No Diagnostics Scientific Instruments: WFC3/UVIS Special Requirements: BETWEEN 01-NOV-2025:00:00:00 AND 25-NOV-2025:00:00:00; TOO RESPONSE TIME 15.0D										
	Patterns	#	Primary Pattern				Secondary Pattern			Exposures	
		(3)	Pattern Type=WFC3-UVIS-DITHER-BOX Purpose=DITHER Number Of Points=4 Point Spacing=0.173 Line Spacing=0.112				Coordinate Frame=POS-TARG Pattern Orientation=23.884 Angle Between Sides=81.785 Center Pattern=false			(1)	
Fixed Targets	#	Name	Target Coordinates		Targ. Coord. Corrections		Fluxes		Miscellaneous		
	(3)	250108A	RA: 03 42 28.3900 (55.6182917d) Dec: -22 30 21.20 (-22.50589d) Equinox: J2000				V=27+/-0.5		Reference Frame: ICRS		
Comments: Expected magnitude based on supernova modeling. We have not been able to get ground-based imaging since March due to Sun constraint. Category=STAR Description=[GAMMA RAY BURSTER, SUPERNOVA]											
Exposures	#	Label	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]		Orbit
	1		(3) 250108A	WFC3/UVIS, ACCUM, UVIS-CENTER	F606W		POS TARG -65,-65; GS ACQ SCENARI O ONEB103	Pattern 3, Exps 1-1 i n F606W visit 05 (05)) (3)	540 Secs X 2 (4320 Secs) [==>(Pattern 1, Copy 1)] [==>(Pattern 1, Copy 2)] [==>(Pattern 2, Copy 1)] [==>(Pattern 2, Copy 2)] [==>(Pattern 3, Copy 1)] [==>(Pattern 3, Copy 2)] [==>(Pattern 4, Copy 1)] [==>(Pattern 4, Copy 2)]		[1] [2]

Orbit Structure



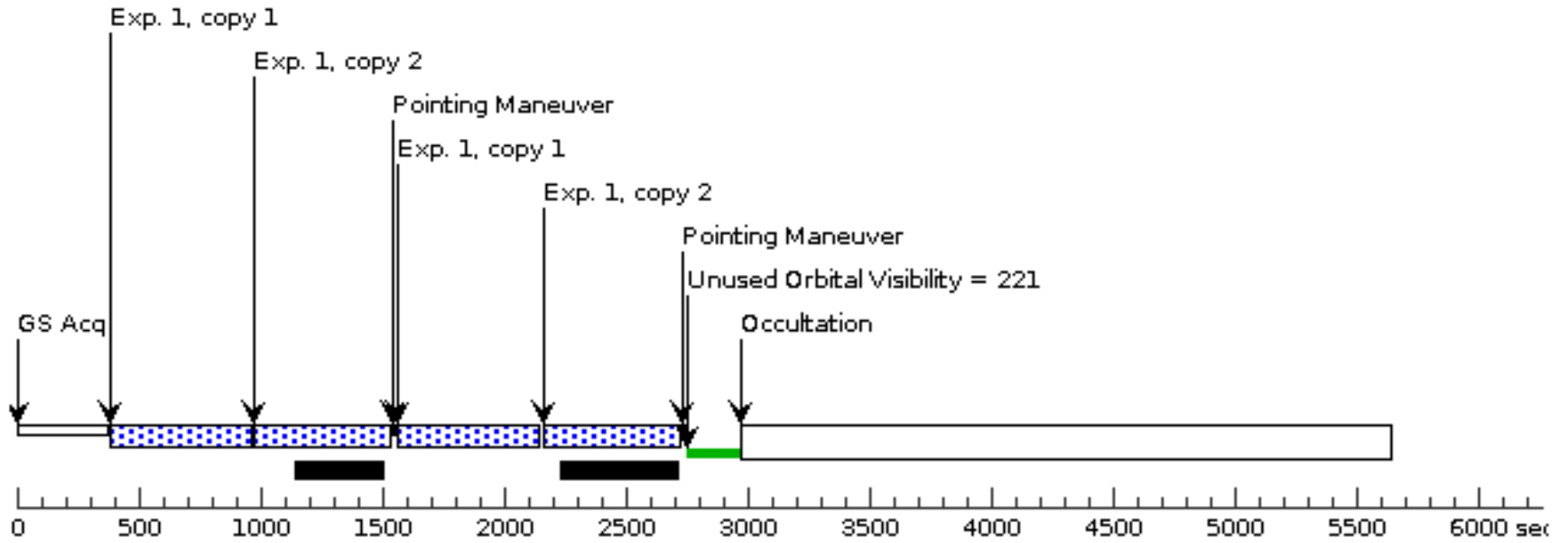


Proposal 17533 - F160W visit 06 (06) - Identifying a New Source of r-Process Nucleosynthesis with HST

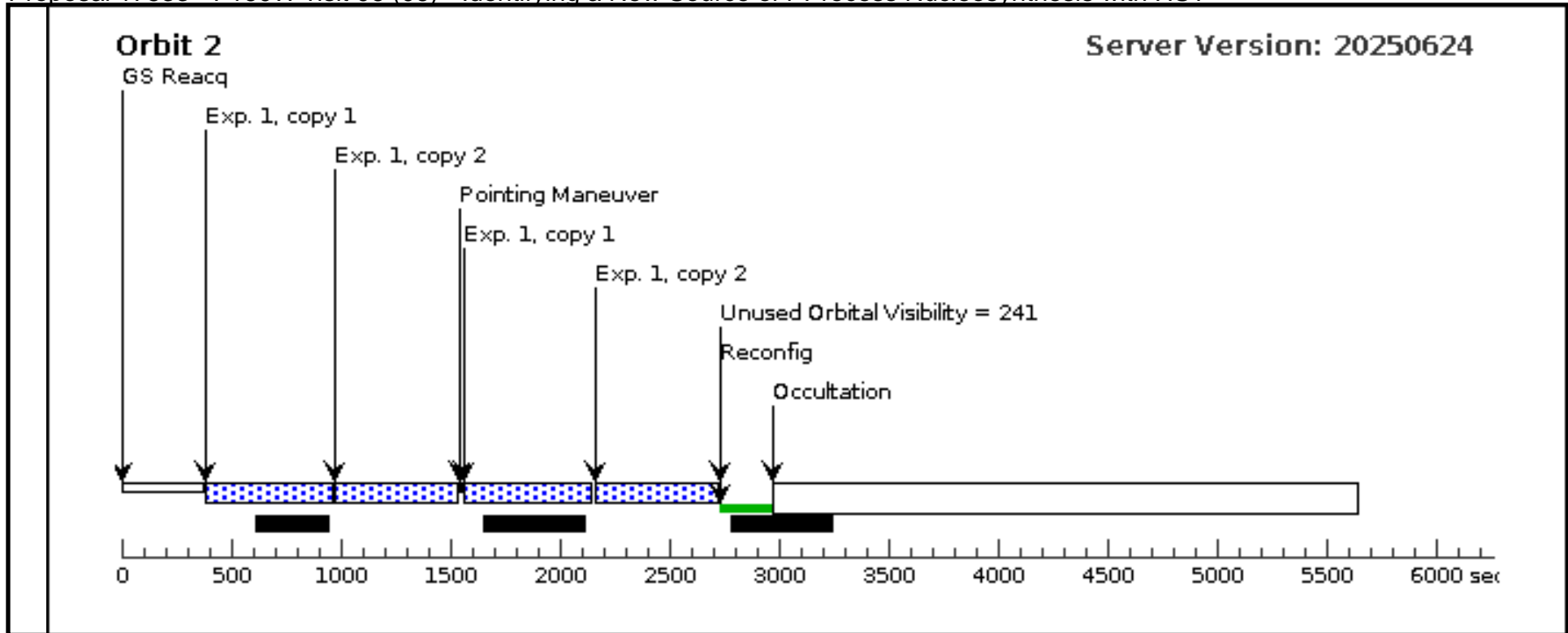
Mon Sep 08 14:00:16 GMT 2025

Visit	Proposal 17533, F160W visit 06 (06), implementation Diagnostic Status: No Diagnostics Scientific Instruments: WFC3/IR Special Requirements: AFTER 05 BY 0 D TO 3 D; TOO RESPONSE TIME 15.0D										
	Patterns	#	Primary Pattern				Secondary Pattern			Exposures	
		(1)	Pattern Type=WFC3-IR-DITHER-BOX-MIN Purpose=DITHER Number Of Points=4 Point Spacing=0.572 Line Spacing=0.365	Coordinate Frame=POS-TARG Pattern Orientation=18.528 Angle Between Sides=74.653 Center Pattern=false							(1)
Fixed Targets	#	Name	Target Coordinates		Targ. Coord. Corrections		Fluxes		Miscellaneous		
	(3)	250108A	RA: 03 42 28.3900 (55.6182917d) Dec: -22 30 21.20 (-22.50589d) Equinox: J2000				V=27+/-0.5		Reference Frame: ICRS		
	<i>Comments: Expected magnitude based on supernova modeling. We have not been able to get ground-based imaging since March due to Sun constraint.</i> Category=STAR Description=[GAMMA RAY BURSTER, SUPERNOVA]										
Exposures	#	Label	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]		Orbit
	1		(3) 250108A	WFC3/IR, MULTIACCUM, IR-FIX	F160W	NSAMP=12; SAMP-SEQ=SPAR S50	POS TARG 40,35	Pattern 1, Exps 1-1 in F160W visit 06 (06) (1)	552.937252 Secs X 2 (4423.498 Secs)		
									[=>(Pattern 1, Copy 1)] [=>(Pattern 1, Copy 2)] [=>(Pattern 2, Copy 1)] [=>(Pattern 2, Copy 2)]	[1]	
									[=>(Pattern 3, Copy 1)] [=>(Pattern 3, Copy 2)] [=>(Pattern 4, Copy 1)] [=>(Pattern 4, Copy 2)]	[2]	

Orbit 1



Orbit Structure

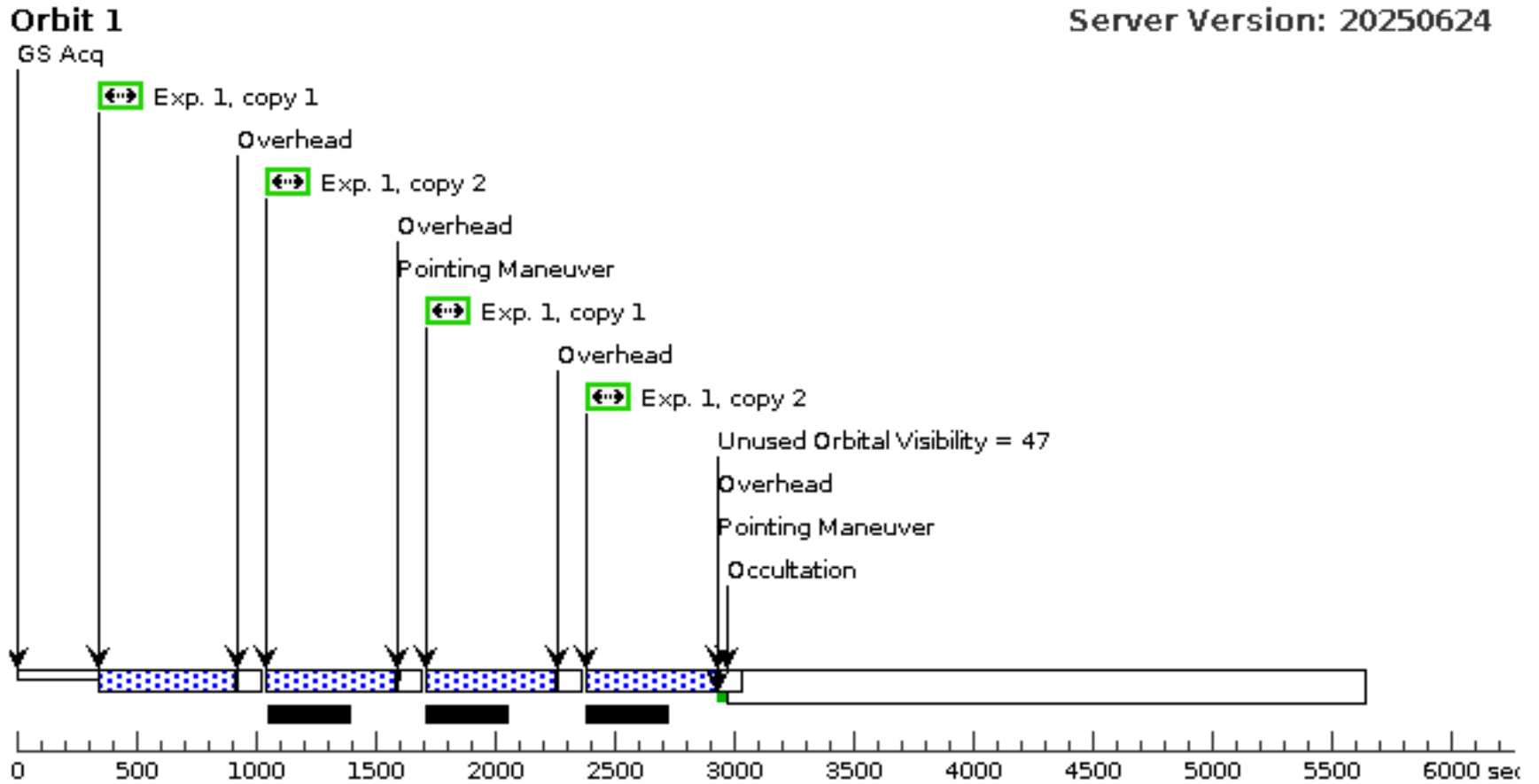


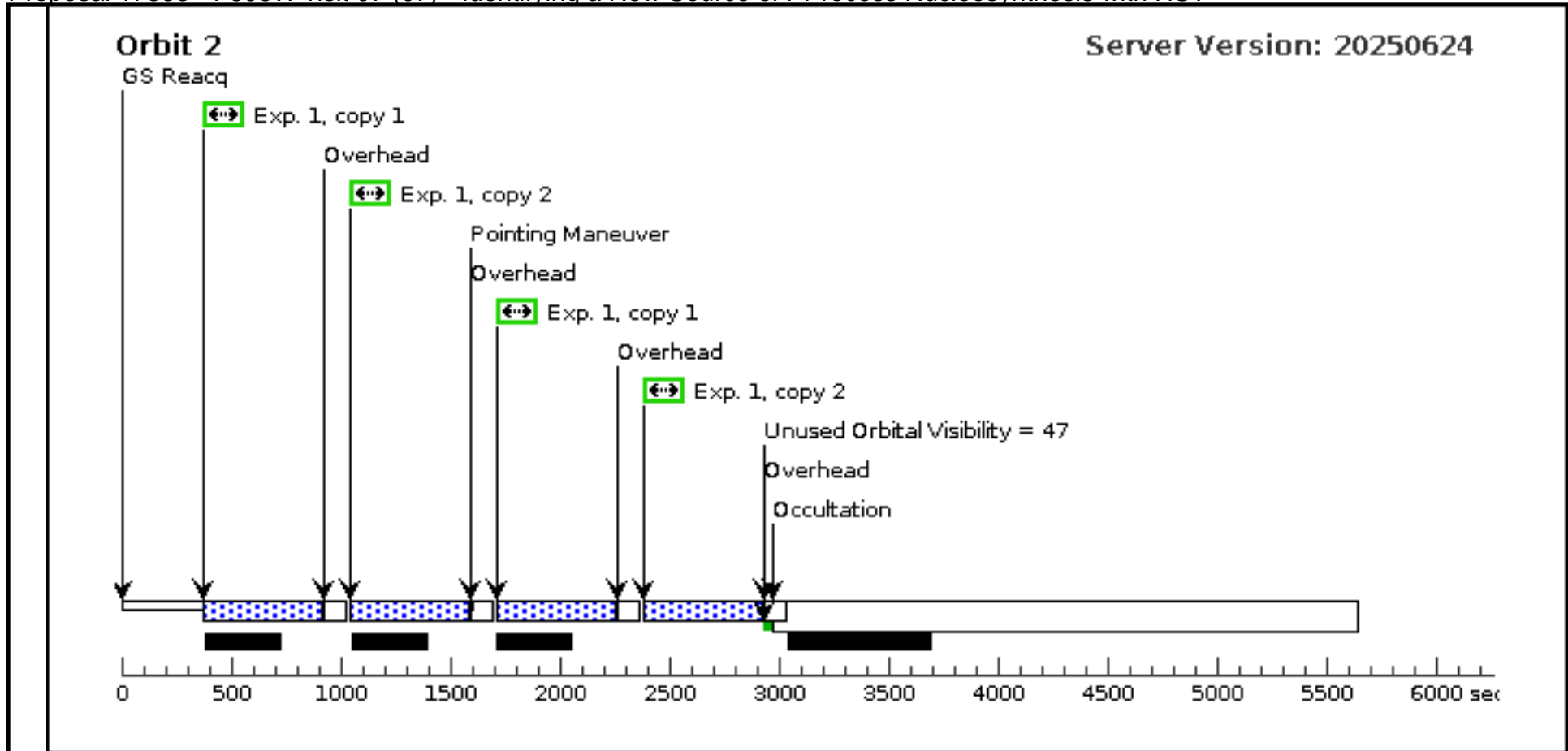
Proposal 17533 - F606W visit 07 (07) - Identifying a New Source of r-Process Nucleosynthesis with HST

Mon Sep 08 14:00:16 GMT 2025

Visit	Proposal 17533, F606W visit 07 (07), implementation Diagnostic Status: No Diagnostics Scientific Instruments: WFC3/UVIS Special Requirements: BETWEEN 01-APR-2026:00:00:00 AND 31-AUG-2026:00:00:00; TOO RESPONSE TIME 15.0D <i>Comments: Deep template observations for image subtraction</i>									
	Patterns	#	Primary Pattern	Secondary Pattern	Exposures					
	(3)	Pattern Type=WFC3-UVIS-DITHER-BOX Purpose=DITHER Number Of Points=4 Point Spacing=0.173 Line Spacing=0.112	Coordinate Frame=POS-TARG Pattern Orientation=23.884 Angle Between Sides=81.785 Center Pattern=false		(1)					
Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous				
	(3)	250108A	RA: 03 42 28.3900 (55.6182917d) Dec: -22 30 21.20 (-22.50589d) Equinox: J2000		V=27+/-0.5	Reference Frame: ICRS				
	<i>Comments: Expected magnitude based on supernova modeling. We have not been able to get ground-based imaging since March due to Sun constraint.</i> Category=STAR Description=[GAMMA RAY BURSTER, SUPERNOVA]									
Exposures	#	Label	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	(3) 250108A		WFC3/UVIS, ACCUM, UVIS-CENTER	F606W		POS TARG -65,-65	Pattern 3, Exps 1-1 in F606W visit 07 (07) (3)	540 Secs X 2 (4320 Secs) [==>(Pattern 1, Copy 1)] [==>(Pattern 1, Copy 2)] [==>(Pattern 2, Copy 1)] [==>(Pattern 2, Copy 2)] [==>(Pattern 3, Copy 1)] [==>(Pattern 3, Copy 2)] [==>(Pattern 4, Copy 1)] [==>(Pattern 4, Copy 2)]	[1] [2]

Orbit Structure





Proposal 17533 - F160W visit 08 (08) - Identifying a New Source of r-Process Nucleosynthesis with HST

Mon Sep 08 14:00:16 GMT 2025

Visit	Proposal 17533, F160W visit 08 (08), implementation Diagnostic Status: No Diagnostics Scientific Instruments: WFC3/IR Special Requirements: AFTER 07 BY 0 D TO 31 D; TOO RESPONSE TIME 15.0D										
	Patterns	#	Primary Pattern				Secondary Pattern			Exposures	
		(1)	Pattern Type=WFC3-IR-DITHER-BOX-MIN Purpose=DITHER Number Of Points=4 Point Spacing=0.572 Line Spacing=0.365	Coordinate Frame=POS-TARG Pattern Orientation=18.528 Angle Between Sides=74.653 Center Pattern=false						(1)	
Fixed Targets	#	Name	Target Coordinates		Targ. Coord. Corrections		Fluxes		Miscellaneous		
	(3)	250108A	RA: 03 42 28.3900 (55.6182917d) Dec: -22 30 21.20 (-22.50589d) Equinox: J2000				V=27+/-0.5		Reference Frame: ICRS		
<i>Comments: Expected magnitude based on supernova modeling. We have not been able to get ground-based imaging since March due to Sun constraint.</i> Category=STAR Description=[GAMMA RAY BURSTER, SUPERNOVA]											
Exposures	#	Label	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]		Orbit
	1		(3) 250108A	WFC3/IR, MULTIACCUM, IR-FIX	F160W	NSAMP=12; SAMP-SEQ=SPARS50	POS TARG 40,35; GS ACQ SCENARIO ONEB103	Pattern 1, Exps 1-1 in F160W visit 08 (08) (1)	552.937252 Secs X 2 (4423.498 Secs)		
									[=>(Pattern 1, Copy 1)]	[1]	
									[=>(Pattern 1, Copy 2)]		
									[=>(Pattern 2, Copy 1)]		
									[=>(Pattern 2, Copy 2)]		
									[=>(Pattern 3, Copy 1)]	[2]	
									[=>(Pattern 3, Copy 2)]		
									[=>(Pattern 4, Copy 1)]		
									[=>(Pattern 4, Copy 2)]		

Orbit Structure

Orbit 1

