



# 17737 - Seasonal dependence of Uranus' upper atmosphere: Tapping 26 years of HST Ly-alpha observations

Cycle: 32, Proposal Category: GO

(UV Initiative)

(Availability Mode: SUPPORTED)

## INVESTIGATORS

<i>Name</i>	<i>Institution</i>
<b>Dr. Lorenz Roth (PI) (ESA Member) (Contact)</b>	<b>Royal Institute of Technology</b>
Dr. Laurent Lamy (CoI) (ESA Member)	Observatoire de Paris - Section de Meudon
Dr. G. Randall Gladstone (CoI) (AdminUSPI)	Southwest Research Institute
Dr. Darrell F. Strobel (CoI)	The Johns Hopkins University
Dr. Nickolay Ivchenko (CoI) (ESA Member)	Royal Institute of Technology
Dr. Henrik Melin (CoI) (ESA Member)	Northumbria University
Mr. Sushen Pravin Joshi (CoI) (ESA Member)	Royal Institute of Technology

## VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) URANUS	STIS/FUV-MAMA	4	04-Oct-2024 09:00:31.0	yes
02	(1) URANUS	STIS/FUV-MAMA	4	04-Oct-2024 09:00:33.0	yes

8 Total Orbits Used

## ABSTRACT

The upper atmospheres of giant planets play key role in various processes such as solar wind interaction, atmospheric escape, magnetosphere-ionosphere coupling, and interaction with ring particles. Voyager2 revealed a largely extended upper atmosphere of Uranus, devoid of hydrocarbons -

unique among the solar system's giant planets. It is not yet well understood how the upper atmosphere varies seasonally and Uranus' orbital period of 84 years makes it generally difficult to obtain observational evidence for seasonal changes. However, IR monitoring of ionospheric H<sub>3</sub><sup>+</sup> emissions suggest a decrease in temperature of the ionized upper atmosphere since 1992. The abundance of H, H<sub>2</sub> and H<sub>3</sub><sup>+</sup> in the upper atmosphere are strongly coupled and indeed HST Ly-alpha observations before (1998) and close to equinox (2011) revealed a similar trend with shrinking and decreasing emissions.

Imaging the H Ly-alpha emissions from Uranus with STIS now in cycle 32 - more than a decade after equinox - offers the unique possibility to test the puzzling cooling trend. The new STIS Ly-alpha images together with the archived STIS images will determine if Uranus' neutral upper atmosphere is indeed undergoing continuous cooling. Additional STIS G140M slit spectra provide precise measurements of Raman shifted features for a characterization of the H<sub>2</sub> component. With only 8 orbits in 2024 this program will complete a unique archive of FUV data on the only ice giant planet with Ly-alpha emission signal observable from Earth and inform a possible large Uranus Orbiter and Probe mission as recommended by the NASA decadal survey.

## **OBSERVING DESCRIPTION**

The idea is to take different kind of imaging and spectral observations combined in one visit to compare the signals and calibrate archival observations of similar type (but never simultaneously taken).

First we want a global image of only Lyman-alpha emissions through slit-less spectral imaging with G140L and the FUV-MAMA

Then we take a long-slit spectrum to get Raman-shifted Lyman-alpha photons longward of the Lyman-alpha line.

And finally, we take an image pair with the clear filter (FUV-MAMA) and the S25SRF2 filter. The difference of these two should also give a Lyman-alpha image but with some uncertainty from the throughput dependency. For our two visits these filter pairs can be calibrated for the first time with the G140L Lyman-alpha image.

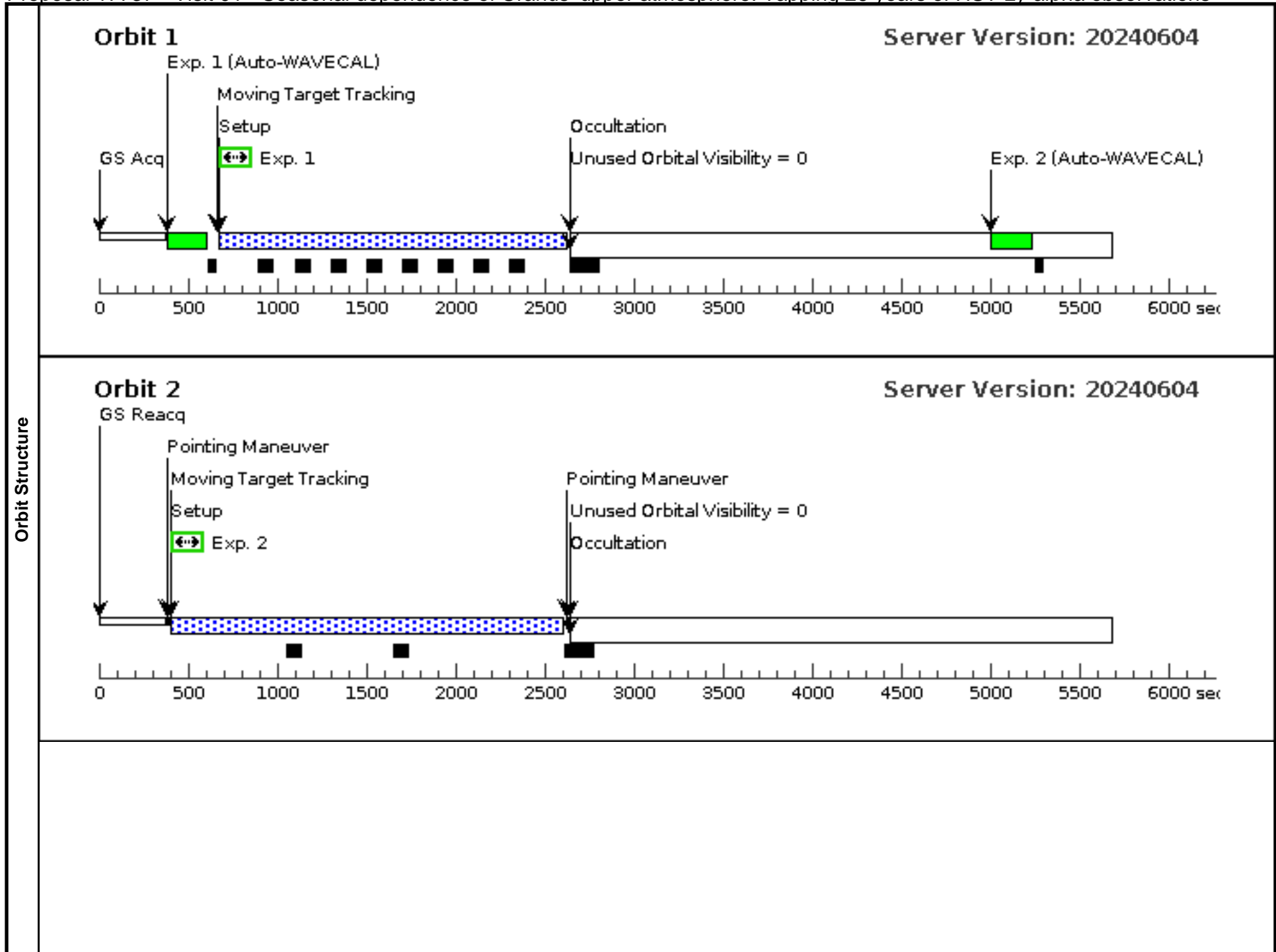
We do not do acquisitions for the spectral observations. For the slitless exposure exact pointing is not required at all. For the long-slit exposures with the 0.5 arcsec slit, even with an offset of up to 1 arcsec along x the slit will still intersect with a good fraction of the Uranus' 3.6 arcsec wide disk.

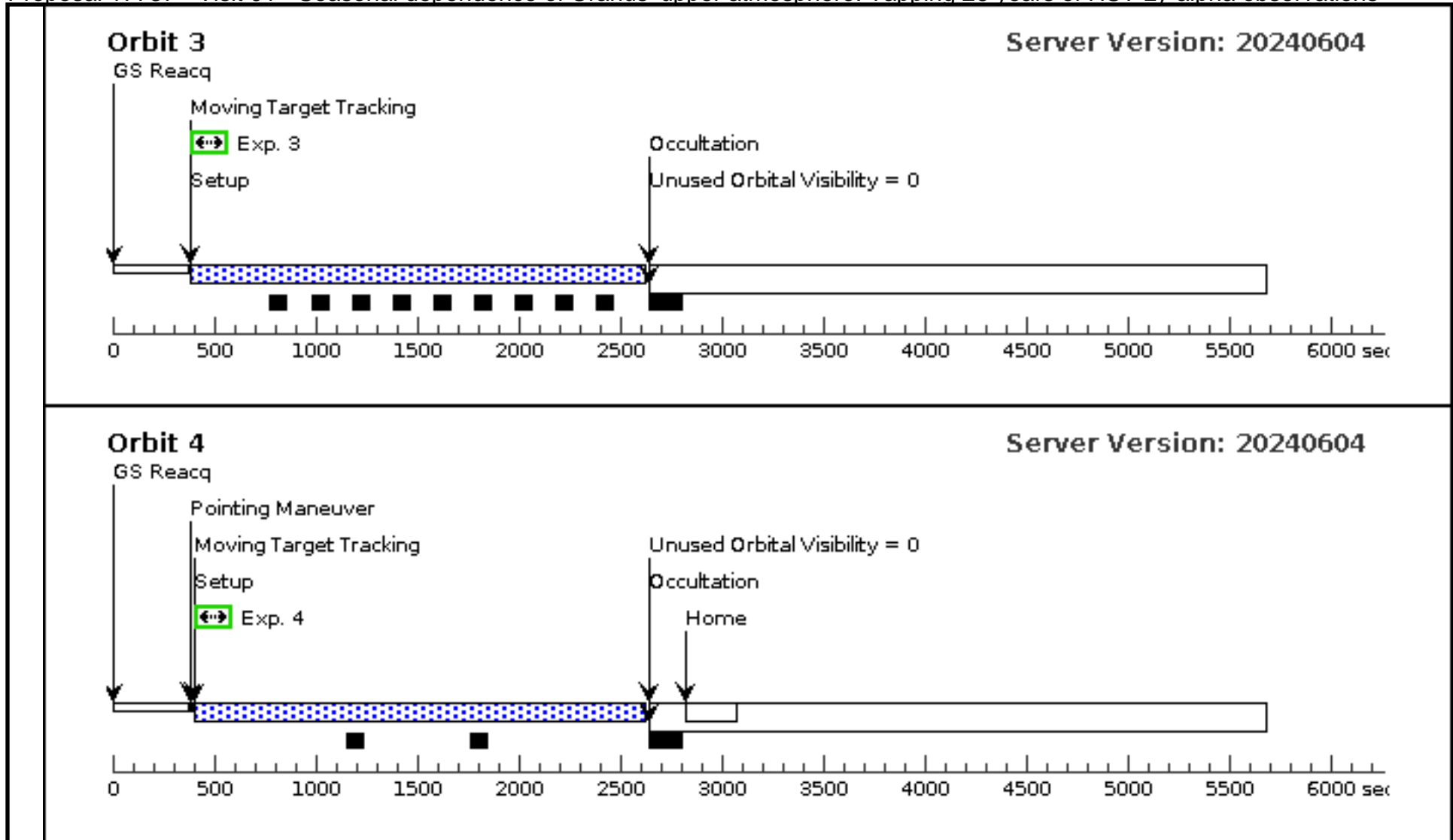
One of the visits should optimally happen simultaneously to the observations by JWST from program 5073.

Proposal 17737 - Visit 01 - Seasonal dependence of Uranus' upper atmosphere: Tapping 26 years of HST Ly-alpha observations

Fri Oct 04 13:00:33 GMT 2024

<b>Visit</b>	<b>Proposal 17737, Visit 01, completed</b> <b>Diagnostic Status: Warning</b> Scientific Instruments: STIS/FUV-MAMA Special Requirements: BETWEEN 02-OCT-2024:02:00:00 AND 02-OCT-2024:10:00:00									
	(Visit 01) Warning (Form): A target acquisition should probably be performed before doing spectroscopy or coronagraphy with STIS or COS. (Visit 01) Informational (Form): The Visit Planner and Spike may produce different schedulability results.									
<b>Diagnosics</b>										
<b>Solar System Targets</b>	<b>#</b>	<b>Name</b>	<b>Level 1</b>	<b>Level 2</b>	<b>Level 3</b>	<b>Window</b>	<b>Ephem Center</b>			
	(1)	URANUS	STD=URANUS				EARTH			
<i>Comments: Description=Planet Uranus Extended=YES</i>										
<b>Exposures</b>	<b>#</b>	<b>Label (ETC Run)</b>	<b>Target</b>	<b>Config,Mode,Aperture</b>	<b>Spectral Els.</b>	<b>Opt. Params.</b>	<b>Special Reqs.</b>	<b>Groups</b>	<b>Exp. Time (Total)/[Actual Dur.]</b>	<b>Orbit</b>
	1	Spectral Ly man-alpha (189449)	(1) URANUS	STIS/FUV-MAMA, TIME-TAG, 25MAMAD1	G140L 1425 A	BUFFER-TIME=20 0			2100 Secs (1917 Secs) [=>1917.0 Secs ]	[1]
	2	Spectral Raman (189449)	(1) URANUS	STIS/FUV-MAMA, TIME-TAG, 52X0.5D1	G140M 1272 A	BUFFER-TIME=60 0			2100 Secs (2134 Secs) [=>2134.0 Secs ]	[2]
	3	Clear image (189449)	(1) URANUS	STIS/FUV-MAMA, TIME-TAG, 25MAMA	MIRROR	BUFFER-TIME=20 0			2100 Secs (2037 Secs) [=>2037.0 Secs ]	[3]
	4	Filter image (189449)	(1) URANUS	STIS/FUV-MAMA, TIME-TAG, F25SRF2	MIRROR	BUFFER-TIME=60 0			2100 Secs (2056 Secs) [=>2056.0 Secs ]	[4]





Proposal 17737 - Visit 02 - Seasonal dependence of Uranus' upper atmosphere: Tapping 26 years of HST Ly-alpha observations

Fri Oct 04 13:00:34 GMT 2024

<b>Visit</b>	<b>Proposal 17737, Visit 02, implementation</b> <b>Diagnostic Status: Warning</b> Scientific Instruments: STIS/FUV-MAMA Special Requirements: BETWEEN 24-OCT-2024:13:30:00 AND 24-OCT-2024:14:30:00									
	(Visit 02) Warning (Form): A target acquisition should probably be performed before doing spectroscopy or coronagraphy with STIS or COS. (Visit 02) Warning (Orbit Planner): SAME POS MAY NOT BE APPROPRIATE (Visit 02) Informational (Form): The Visit Planner and Spike may produce different schedulability results.									
<b>Diagnosics</b>										
<b>Solar System Targets</b>	#	Name	Level 1	Level 2	Level 3	Window	Ephem Center			
	(1)	URANUS	STD=URANUS				EARTH			
Comments: Description=Planet Uranus Extended=YES										
<b>Exposures</b>	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	Spectral Ly man-alpha (189449)	(1) URANUS	STIS/FUV-MAMA, TIME-TAG, 25MAMA	G140L 1425 A	BUFFER-TIME=20 0	POS TARG 3,-2.5		2100 Secs (1917 Secs) [==>1917.0 Secs ]	[1]
	2	Spectral Ra man (189449)	(1) URANUS	STIS/FUV-MAMA, TIME-TAG, 52X2	G140M 1272 A	BUFFER-TIME=60 0	POS TARG 0,-6		2100 Secs (2134 Secs) [==>2134.0 Secs ]	[2]
	3	Clear image (189449)	(1) URANUS	STIS/FUV-MAMA, TIME-TAG, 25MAMA	MIRROR	BUFFER-TIME=20 0	POS TARG 3,-4		2100 Secs (2037 Secs) [==>2037.0 Secs ]	[3]
	4	Filter image (189449)	(1) URANUS	STIS/FUV-MAMA, TIME-TAG, F25SRF2	MIRROR	BUFFER-TIME=60 0	SAME POS AS 3		2100 Secs (2056 Secs) [==>2056.0 Secs ]	[4]

