



## 17753 - Binarity and Accretion Processes in AGB Stars: A UV Study of pi1 Gru

Cycle: 32, Proposal Category: GO

(UV Initiative)

(Availability Mode: SUPPORTED)

### INVESTIGATORS

<i>Name</i>	<i>Institution</i>
<b>Dr. Raghvendra Sahai (PI) (Contact)</b>	<b>Jet Propulsion Laboratory</b>
Leen Decin (CoI) (ESA Member)	Institute of Astronomy, KU Leuven
Dr. Miguel Montarges (CoI) (ESA Member)	Observatoire de Paris - Section de Meudon
Dr. Alex De Koter (CoI) (ESA Member)	Universiteit van Amsterdam

### VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) PI.01-GRU	COS/FUV COS/NUV	2	10-Sep-2025 13:00:39.0	yes
02	(1) PI.01-GRU	COS/FUV COS/NUV	2	10-Sep-2025 13:00:40.0	yes

4 Total Orbits Used

### ABSTRACT

The dramatic changes in the geometry and dynamics of the mass-ejection during the death throes of most stars that evolve in a Hubble time (i.e., with initial masses  $\sim 1-8 M_{\text{sun}}$ ), are widely believed to be due to binarity-induced formation of bound disks that power high-speed collimated jets. But observational evidence of this process are lacking. We propose to obtain multi-epoch UV spectra of the nearby (160 pc) late-spectral-type AGB star, pi1 Gru, that belongs to a newly-discovered class of AGB stars showing strong (and often variable) FUV fluxes -- suggestive of variable accretion of matter onto an accretion disk in a binary system. A STIS spectroscopic study of the prototype of this class, Y Gem, showing the presence of

Proposal 17753 (STScI Edit Number: 0, Created: Wednesday, September 10, 2025, 12:00:40PM Eastern Standard Time) - Overview  
flickering and high-velocity infall and outflows, supports this hypothesis. ALMA observations of pi1 Gru reveal strong departures from spherical symmetry in its dusty molecular wind (spiral structure, fast polar bubbles) and the presence of a close companion at ~6 AU, confirmed by optical imaging using SPHERE/ZIMPOL.

We will use the UV spectra of pi1 Gru to fit the UV continuum and derive the temperature and size of the accretion hot-spot. Using CLOUDY modelling, we will fit the emission lines to provide constraints on the physical parameters of the accreting flow. The line profiles, if resolved, will reveal the nature of the companion star. Our study will probe the early formation phase of bound disks in binary AGB stars, and provide valuable insights into the binary-induced accretion processes that make such disks, which have long been theorized to drive collimated outflows that produce aspherical morphologies in post-AGB objects.

### **OBSERVING DESCRIPTION**

We will obtain multi-epoch UV spectra of pi1 Gru with COS using the G140L (center=1280 Å) and G230L (center=2950 & 3360 Å) gratings in TIME-TAG mode. The average GALEX

FUV (NUV) flux of pi1 Gru is 277(2680) uJy. For this source, using COS ETC, we find:

FUV: Using the G140L grating, with Cen.Wave. setting  $\lambda=1280$  Å (B[916-1195], A[1282-2148]) (spectral resolution  $R \sim 2000 - 5000$ ), we will get a signal-to-noise ratio, S/N, of  $\sim 10$  for the continuum with an exposure time of 12 min.

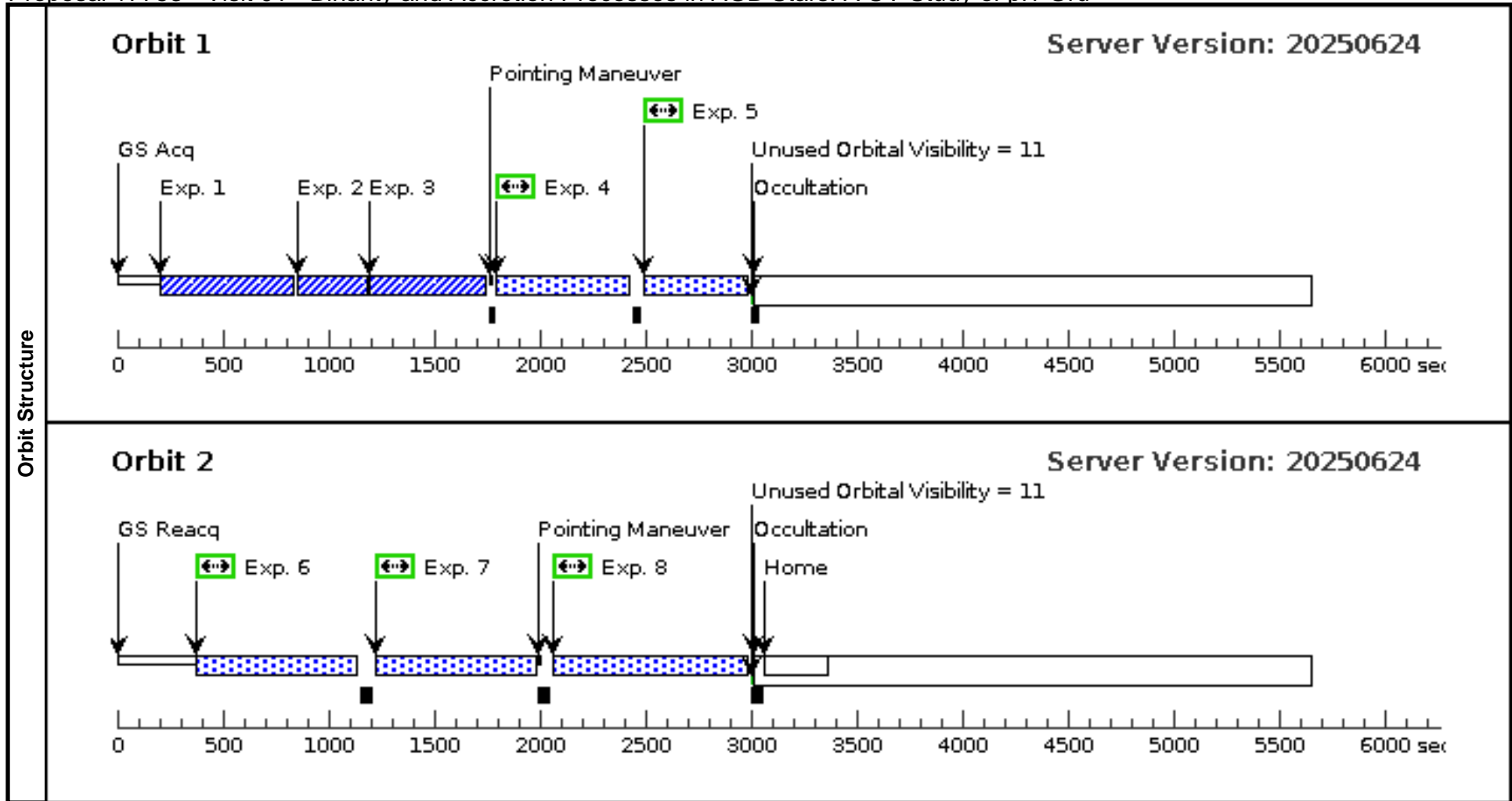
NUV: Using the G230L grating with two Cen.Wave settings: (i) setting  $\lambda=3360$  Å (A[2060-2457], B[3162 -3560]) and (ii) setting  $\lambda=2950$  Å (A[1651-2049], A[2751 3150]) ( $R \sim 1550 - 2900$ ), we get an S/N of  $\sim 10$  for the continuum with an exposure time of  $\sim 1.1$  min. As seen from the study of Y Gem (Setal18), the S/N for the lines will be much higher than for the continuum.

The above selection of gratings and setting has been made in order to provide nearly fullwavelength coverage at  $\lambda > 916$  Å in the FUV and NUV (for  $\lambda < 3200$  Å) bands (since the G230L NUVB spectrum is expected to have negligible second-order contamination shortward of 3200 Å.) There will be a  $\sim 300$  Å gap (between 2460 Å and 2750 Å) in the NUV spectra that we obtain, but it does not impact our ability to reach the science goals of our study. Stripe C of the COS/ MAMA detector may provide potentially useful data over wavelength bands 1900-2100 Å for Cen.Wave.=2950 Å and 2164 - 2361 Å for Cen.Wave.=3360, with about 5% contamination from the first order spectra (ISR COS 2010-01), but these are not critical for our proposed study.

Proposal 17753 - Visit 01 - Binarity and Accretion Processes in AGB Stars: A UV Study of pi1 Gru

Wed Sep 10 17:00:40 GMT 2025

Visit	<b>Proposal 17753, Visit 01, implementation</b> <b>Diagnostic Status: No Diagnostics</b> Scientific Instruments: COS/FUV, COS/NUV Special Requirements: (none)																																																																																										
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Proposal 17753 - Visit 02 - Binarity and Accretion Processes in AGB Stars: A UV Study of pi1 Gru

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<b>Visit</b>	<b>Proposal 17753, Visit 02, implementation</b> <b>Diagnostic Status: Warning</b> Scientific Instruments: COS/FUV, COS/NUV Special Requirements: AFTER 01 BY 300 D TO 370 D																																																																																																		
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