



17839 - The Beat Goes On

Cycle: 32, Proposal Category: GO

(Availability Mode: AVAILABLE)

INVESTIGATORS

<i>Name</i>	<i>Institution</i>
Dr. Thomas R. Ayres (PI) (Contact)	University of Colorado at Boulder

VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
10	(1) HD-131156A (2) HD-131156B	COS/FUV COS/NUV	2	16-Aug-2024 08:00:17.0	yes
11	(3) HD-165341A (4) HD-165341B	COS/FUV COS/NUV	2	16-Aug-2024 08:00:18.0	yes

4 Total Orbits Used

ABSTRACT

(This is a joint Chandra/HST program: the Chandra proposal abstract follows.)

Chandra has been following magnetic heartbeats of late-type stars via high-contrast coronal X-rays. Goal is to provide fundamental observational constraints for contemporary and future studies of the underlying magnetic dynamo, whose internal workings remain elusive. The Sun's high-energy modulations play an important space weather role in our heliosphere, as do stellar counterparts for their exoplanets. A 3-year continuation (joint with HST) is proposed for current cycles targets Xi Boo (G8V+K4) and 70 Oph (K0V+K5V). Key questions: (1) origin of diverging branches in rotational period vs. cycle duration, where Sun sits isolated in the middle, possibly in a transitional state; (2) extent of high-energy variability bias, which can affect interpretations of large surveys.

OBSERVING DESCRIPTION

****Impact of Reduced Gyro Mode****

The two targets, Xi Boo and 70 Oph, would be minimally affected by RGM, because there is only a minor orientation constraint on the observations (having to do with protecting the PSA for the NUV/BOA target acquisition). Further, the Visits are only 2 orbits, minimizing the possibility of re-acquisition failures. However, if target acquisition times are increased by RGM, then one or more of the exposures would have to be shortened. This can be done with minimal loss of science.

Phase II Observing Description (for normal 3-gyro operations)

Observational goal is to obtain FUV moderate-resolution spectra of both components of Xi Boo and 70 Oph using Cosmic Origins Spectrograph (COS). The primaries, Xi Boo A and 70 Oph A, are nearly identical in FUV flux levels, about a tenth of well-studied (by STIS) Alpha Cen A. Normally, high-sensitivity COS would be off-limits for these primaries, because their strong Lyman-alpha features would violate FUV detector bright limits. However, the COS G130M-1222 setting places Lyman-alpha in the gap between detector sides A and B, so can be used for cool stars, as long as their other strong emissions (e.g., C II 133 nm) are "safe" (as is the case for Xi Boo A and 70 Oph A). When combined with COS G160M-1533, the spectral range 107-171 nm can be covered, with only two small gaps: 120.7-122.3 nm (Lyman-alpha) and 151.5-153.3 nm (devoid of important lines). COS's 15 km/s resolution is fully adequate for the relatively broad FUV emission features of these cool stars, especially the "hot" lines like Si III 120.6 nm (T~30,000 K), Si IV 139.3 nm (80,000 K), and C IV 154.8 nm (100,000 K). The COS FUV channel is so sensitive that a mere 1000 s exposure obtains S/N~100 (line peak) at Si III, 80 at Si IV, and 65 at C IV, for both primary stars. Each grating setting requires 4 "FP-Splits" to mitigate fixed pattern noise.

The pair of FUV settings plus FP-SPLITS can be done in a single orbit for the primary of each system, including the initial Guide Star and NUV imaging target acquisitions.

At the beginning of the second orbit, the secondary star would be centered by a small offset maneuver from the primary, given the accurate orbital ephemeris. The same dual-setting FUV sequence then would be carried out, extending over the remaining unocculted part of the second orbit.

The COS visits would be once a year for each system (4 HST orbits per year, total): experience with Alpha Cen AB has shown that the FUV variability, while significant, is not as pronounced as that of the X-rays (which are observed twice a year for each system). Helpfully, key coronal forbidden lines Fe XII 134.9 nm (T~2 MK) and Fe XXI 135.4 nm (10 MK) are available in the FUV to assess the stellar coronal states at the times of the COS pointings. For this reason, the HST and Chandra observations need not be simultaneous. In fact, it is better to have them separated, since the FUV coronal forbidden lines would then provide an additional point on the high-energy time line. Given the very deep COS exposures proposed here (especially at 134.9 nm and 135.4 nm, which receive double coverage in the 2 nm overlap between G130M and G160M), the Fe XII coronal forbidden line will be easily detected in the primary stars (based on the previous COS visits), and probably also in the UV-fainter secondaries.

Proposal 17839 - Visit 10 - The Beat Goes On

Fri Aug 16 12:00:19 GMT 2024

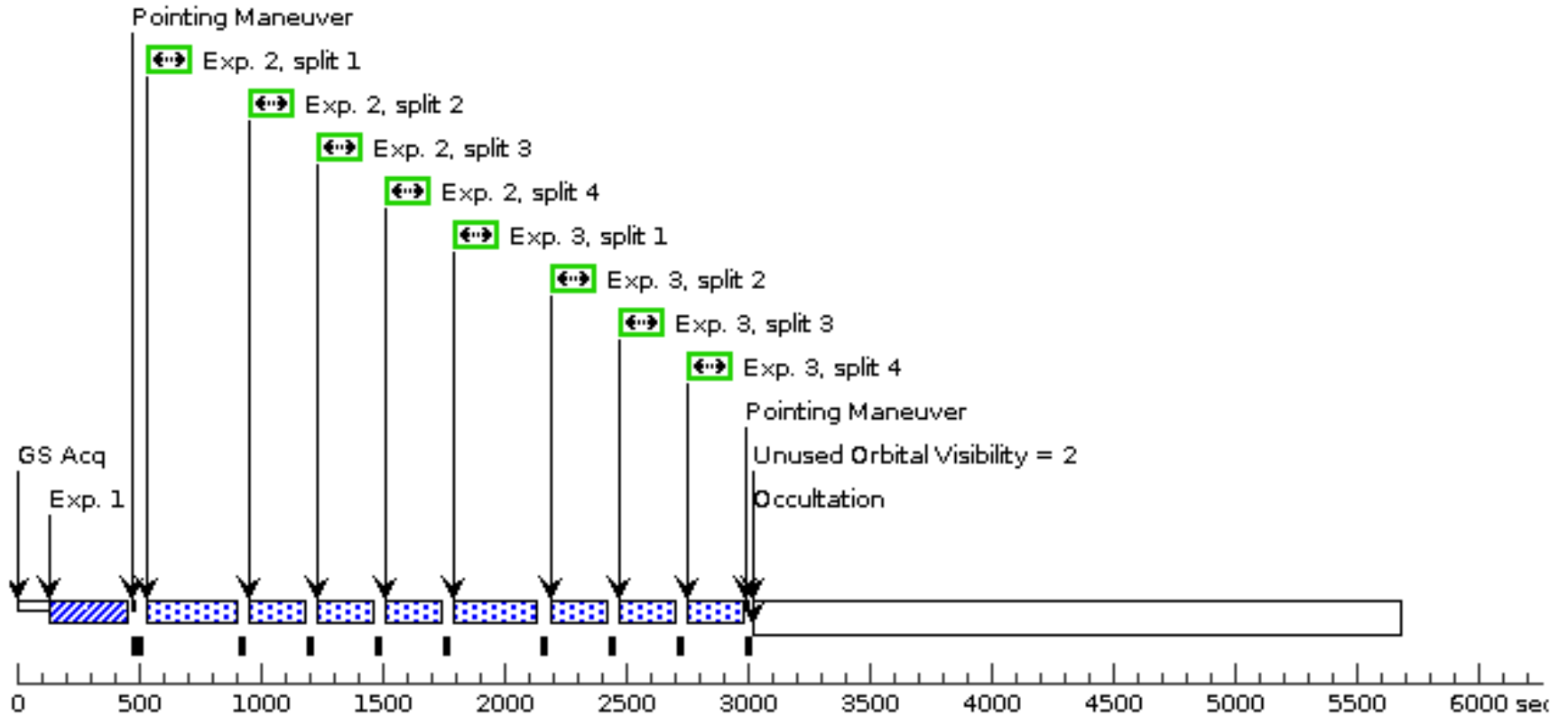
Visit		Proposal 17839, Visit 10, implementation Diagnostic Status: No Diagnostics Scientific Instruments: COS/FUV, COS/NUV Special Requirements: GYRO MODE 1G; ORIENT 130D TO 10 D					
Fixed Targets		#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
		(1)	HD-131156A Alt Name1: XI-BOO-A	RA: 14 51 23.3885 (222.8474521d) Dec: +19 06 1.62 (19.10045d) Equinox: J2000	Proper Motion RA: +0.1275 arcsec/yr Proper Motion Dec: -0.0406 arcsec/yr Parallax: 0.148" Epoch of Position: 2000 Radial Velocity: +1.5 km/sec	V=+4.68+/-0.1	Reference Frame: ICRS
		<i>Comments: Target coords for epoch 2000 from SIMBAD, based on Gaia measurements for a mean epoch of 2015.5 corrected for Gaia proper motions</i> Category=STAR Description=[CORONA, G V-IV] Extended=NO					
		(2)	HD-131156B Alt Name1: XI-BOO-B	Offset from HD-131156A RA Offset: -0.325 Secs Dec Offset: 1.58 Arcsec	Radial Velocity: +2.3 km/sec	V=+6.82+/-0.1	Offset Position (HD-131156B)
		<i>Comments: Target offset based on ephemeris from 6th Orbit Catalog, for 2025.5</i> Category=STAR Description=[CORONA, K V-IV] Extended=NO					

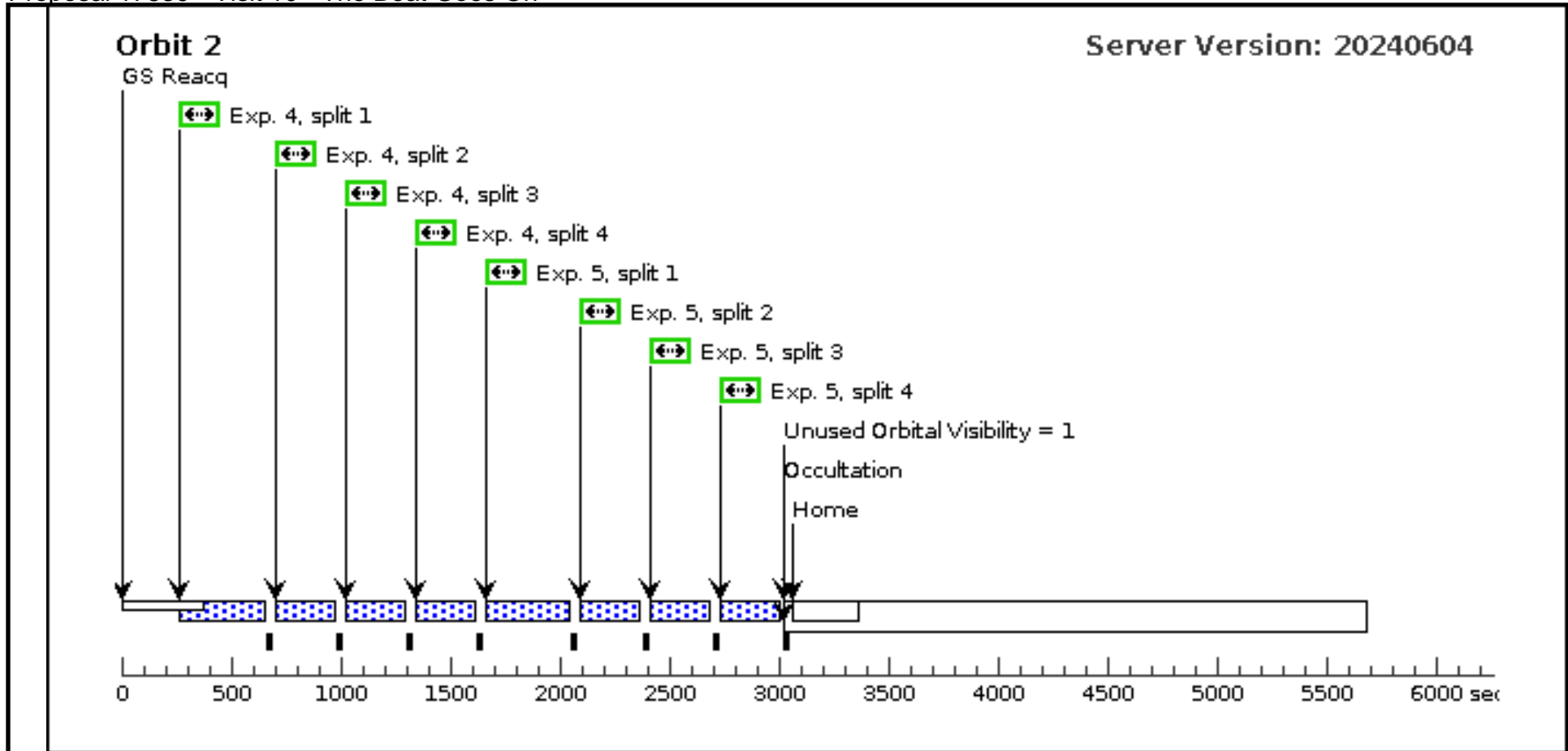
Proposal 17839 - Visit 10 - The Beat Goes On

Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
	1	(COS.ta.136 7079)	(1) HD-131156A	COS/NUV, ACQ/IMAGE, BOA	MIRRORB					11 Secs (11 Secs)	
										[==>]	[1]
	<i>Comments: Castelli-Kurucz Models G8V, renormalized to vegamag = 4.68 in filter Johnson/V</i>										
	2	(COS.sp.136 7281)	(1) HD-131156A	COS/FUV, TIME-TAG, PSA	G130M 1222 A	BUFFER-TIME=24 00; FP-POS=ALL; LIFETIME-POS=L P4				177 Secs (708 Secs)	
										[==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[1]
<i>Comments: h_hd131156_uvsum_1m_51163_etc.dat from StarCAT (STIS E140M-1425)</i>											
3	(COS.sp.136 7282)	(1) HD-131156A	COS/FUV, TIME-TAG, PSA	G160M 1533 A	BUFFER-TIME=40 00; FP-POS=ALL; LIFETIME-POS=L P4				176 Secs (704 Secs)		
									[==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[1]	
<i>Comments: h_hd131156_uvsum_1m_51163_etc.dat from StarCAT (STIS E140M-1425)</i>											
4	(COS.sp.136 7319)	(2) HD-131156B	COS/FUV, TIME-TAG, PSA	G130M 1222 A	BUFFER-TIME=93 00; FP-POS=ALL; LIFETIME-POS=L P4				216 Secs (864 Secs)		
									[==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[2]	
<i>Comments: B fluxes scaled by 0.14 from A [h_hd131156_uvsum_1m_51163_etc.dat from StarCAT (STIS E140M-1425)]</i>											
5	(COS.sp.136 7320)	(2) HD-131156B	COS/FUV, TIME-TAG, PSA	G160M 1533 A	BUFFER-TIME=13 400; FP-POS=ALL; LIFETIME-POS=L P4				216 Secs (864 Secs)		
									[==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[2]	
<i>Comments: B fluxes scaled by 0.14 from A [h_hd131156_uvsum_1m_51163_etc.dat from StarCAT (STIS E140M-1425)]</i>											

Orbit Structure

Orbit 1





Proposal 17839 - Visit 11 - The Beat Goes On

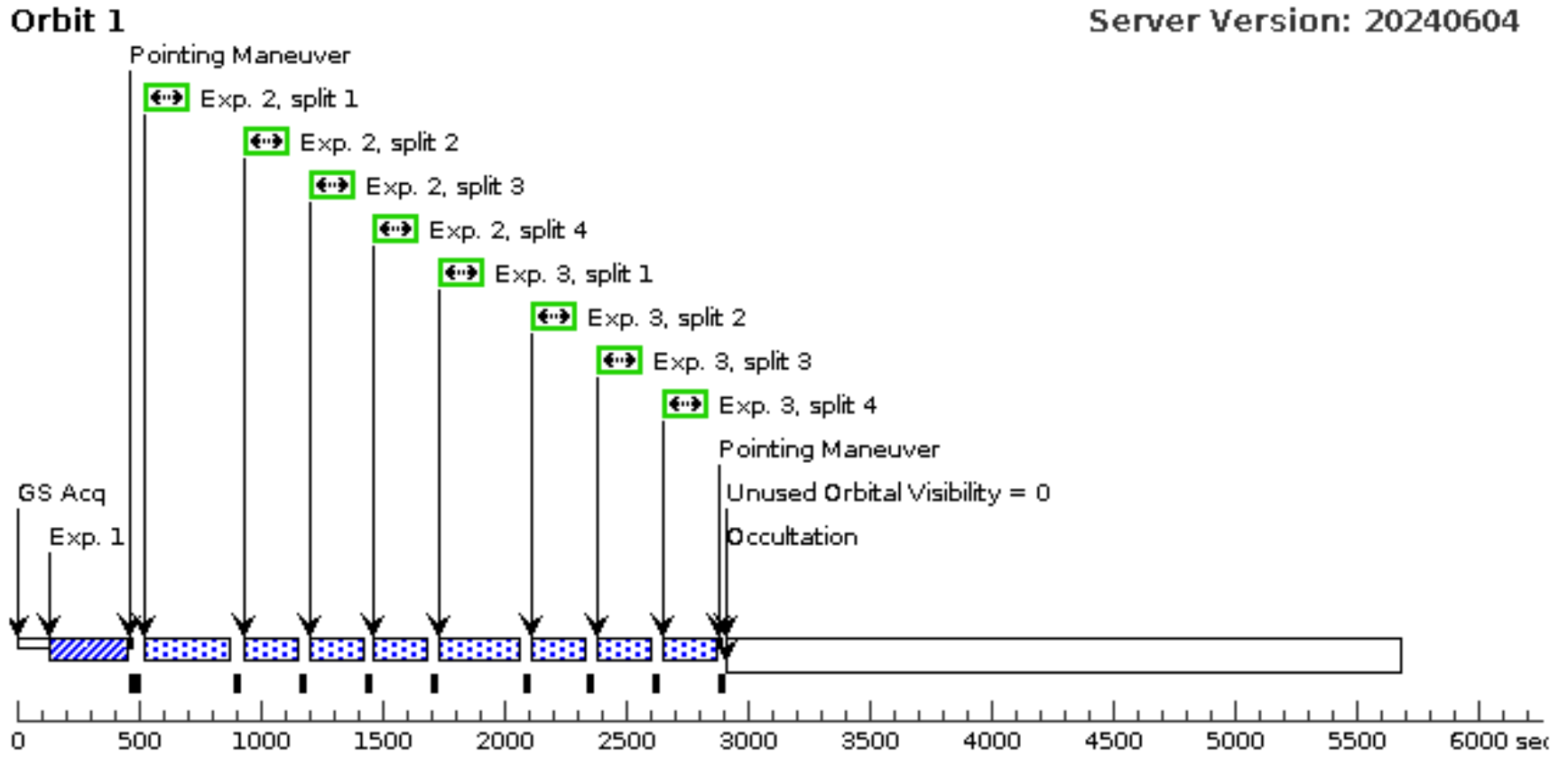
Fri Aug 16 12:00:19 GMT 2024

Visit	Proposal 17839, Visit 11, implementation				
	Diagnostic Status: No Diagnostics				
	Scientific Instruments: COS/FUV, COS/NUV				
	Special Requirements: GYRO MODE 1G; SCHED 100%; ORIENT 315D TO 195 D				

Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
	(3)	HD-165341A Alt Name1: 70-OPH-A	RA: 18 05 27.2484 (271.3635350d) Dec: +02 30 0.53 (2.50015d) Equinox: J2000	Proper Motion RA: +0.2065 arcsec/yr Proper Motion Dec: -1.1075 arcsec/yr Parallax: 0.1956" Epoch of Position: 2000 Radial Velocity: -9.7 km/sec	V=+4.03+/-0.1	Reference Frame: ICRS
<i>Comments: Target coords for epoch 2000 from SIMBAD, based on Gaia measurements for a mean epoch of 2015.5 corrected for Gaia proper motions</i> Category=STAR Description=[CORONA, K V-IV] Extended=NO						
(4)	HD-165341B Alt Name1: 70-OPH-B	Offset from HD-165341A RA Offset: 0.4 Secs Dec Offset: -3.06 Arcsec	Radial Velocity: -10.0 km/sec	V=+6.07+/-0.1	Offset Position (HD-165341B)	
<i>Comments: Target offset based on ephemeris from 6th Orbit Catalog, for 2025.5</i> Category=STAR Description=[CORONA, K V-IV] Extended=NO						

Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
		1	(COS.ta.136 7302)	(3) HD-165341A	COS/NUV, ACQ/IMAGE, BOA	MIRRORB		GS ACQ SCENARI O BASE103		8 Secs (8 Secs) [==>]
<i>Comments: Castelli-Kurucz Models K0V, renormalized to vegamag = 4.22 in filter Johnson/V (from BSC and WDS)</i>										
	2	(COS.sp.136 7312)	(3) HD-165341A	COS/FUV, TIME-TAG, PSA	G130M 1222 A	BUFFER-TIME=28 00; FP-POS=ALL; LIFETIME-POS=L P4			164 Secs (656 Secs) [==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[1]
	3	(COS.sp.136 7313)	(3) HD-165341A	COS/FUV, TIME-TAG, PSA	G160M 1533 A	BUFFER-TIME=43 00; FP-POS=ALL; LIFETIME-POS=L P4			163 Secs (652 Secs) [==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[1]
	4	(COS.sp.136 7325)	(4) HD-165341B	COS/FUV, TIME-TAG, PSA	G130M 1222 A	BUFFER-TIME=92 00; FP-POS=ALL; LIFETIME-POS=L P4			202 Secs (808 Secs) [==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[2]
	5	(COS.sp.136 7326)	(4) HD-165341B	COS/FUV, TIME-TAG, PSA	G160M 1533 A	BUFFER-TIME=12 700; FP-POS=ALL; LIFETIME-POS=L P4			202 Secs (808 Secs) [==>(Split 1)] [==>(Split 2)] [==>(Split 3)] [==>(Split 4)]	[2]

Orbit Structure



Orbit 2

GS Reacq

