



## 13465 - Alpha Cen: Climbing out of a Coronal Recession? (year 3 continuation)

Cycle: 21, Proposal Category: GO

(Availability Mode: SUPPORTED)

### INVESTIGATORS

| <i>Name</i>                               | <i>Institution</i>                       | <i>E-Mail</i>                    |
|---|--|----------------------------------|
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### VISITS

| <i>Visit</i> | <i>Targets used in Visit</i>   | <i>Configurations used in Visit</i> | <i>Orbits Used</i> | <i>Last Orbit Planner Run</i> | <i>OP Current with Visit?</i> |
|--------------|--------------------------------|-------------------------------------|--------------------|-------------------------------|-------------------------------|
| 10           | (1) HD128620<br>(2) HD128621-1 | STIS/CCD<br>STIS/FUV-MAMA           | 2                  | 20-Aug-2013 21:27:28.0        | yes                           |
| 11           | (1) HD128620<br>(3) HD128621-2 | STIS/CCD<br>STIS/FUV-MAMA           | 2                  | 20-Aug-2013 21:27:42.0        | yes                           |

4 Total Orbits Used

### ABSTRACT

Nearby Alpha Centauri contains the two best characterized G and K dwarfs, next to the Sun itself, thanks to the accurate orbit, resolved angular diameters, and well understood co-evolutionary state. Alpha Cen A & B also have the best studied stellar X-ray activity cycles, extending back to the 1970's. Present proposal is to continue tracking the evolving multi-decadal high-energy narrative of Alpha Cen with semiannual HRC-I pointings in Cycles 13-15, as solar twin A is expected to be rising to cycle maximum following an extended coronal recession. STIS E140M spectra will support and leverage the broad-band X-ray measurements by probing subcoronal dynamics and providing a low-T boundary condition for DEM modeling, with connection to the X-rays through the FUV Fe XII coronal forbidden line.

## **OBSERVING DESCRIPTION**

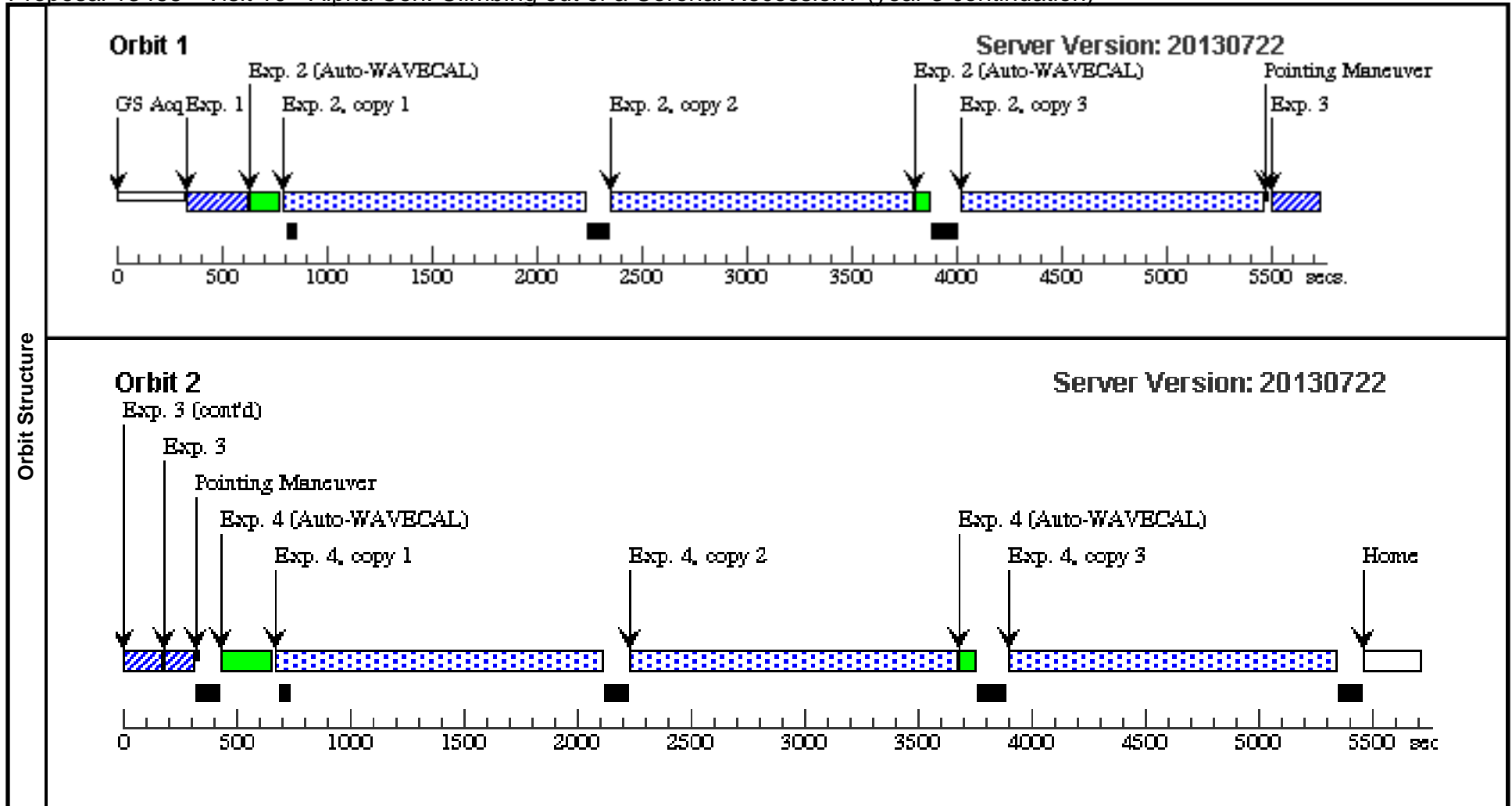
The STIS observations are straightforward. Alpha Cen falls in the HST Continuous Viewing Zone numerous times during the year, allowing the two stars to be captured in a single visit of just two orbits. Two such visits (four orbits total), about six months apart, are contemplated for Cycle 21, compatible with the semiannual pointings by Chandra. There is no need to strictly coordinate the FUV and X-ray visits, because the FUV Fe XII coronal forbidden line can tie the STIS observation into the X-ray framework.

In each HST/STIS visit, the binary companions are observed sequentially, beginning with Alp Cen A, the brighter of the two stars. The target is acquired with the CCD and F25ND5, followed by a 4.3 ks exposure with the E140M-1425 medium-res echelle through the photometric slot (0.2x0.2), which delivers  $R=40,000$  and good sensitivity (peak  $S/N=60$  per resol at the tops of the important Si IV and C IV resonance lines). The observation is split into three subexposures to mitigate instrumental drifts. The exposure depth is sufficient to capture the key Fe XII 1242 coronal forbidden line, which as mentioned earlier is used to tie the STIS FUV measurements into the X-ray timeline. Following the E140M exposure, a  $\sim 5''$  offset maneuver is performed to Alp Cen B, and a peak-up is performed in dispersed light (G430L) with the 0.3x0.05ND filtered slit to precisely locate the target. This particular slit was chosen because the default peak-up sweeps out a sufficient area (0.3"x0.3") to capture the B component, given that the time-dependent separation of the binary companions is accurately known (to the 0.1" level; based on five years of measurements by Chandra HRC-I). After the peak-up, a second 4.3 ks E140M exposure through the photometric aperture is obtained on B. Given that the AB orbital separation is changing relatively rapidly in the present epoch, Alp Cen B was specified as two separate targets, with different offsets, to reflect the orbital evolution over the six months between the two STIS visits.

Proposal 13465 - Visit 10 - Alpha Cen: Climbing out of a Coronal Recession? (year 3 continuation)

Wed Aug 21 01:27:51 GMT 2013

| Visit  | <b>Proposal 13465, Visit 10, implementation</b><br><b>Diagnostic Status: No Diagnostics</b><br>Scientific Instruments: STIS/CCD, STIS/FUV-MAMA<br>Special Requirements: CVZ; BETWEEN 12-FEB-2014:00:00:00 AND 13-FEB-2014:23:59:59<br><i>Comments: This is first CVZ opportunity roughly 6 months after final Visit of program 13060 (of which this is a third year continuation)</i> |  |  |  |  |                       |                             |  |  |       |
|--|---|--|--|--|--|-----------------------|-----------------------------|--|--|-------|
|  | Fixed Targets   | #  | Name   | Target Coordinates   | Targ. Coord. Corrections   | Fluxes                | Miscellaneous               |  |  |       |
| (1)  |   | HD128620<br>Alt Name1: ALP-CEN-A   | RA: 14 39 29.3800 (219.8724167d)<br>Dec: -60 49 58.40 (-60.83289d)<br>Equinox: J2000 | Proper Motion RA: -3.82 arcsec/yr<br>Proper Motion Dec: +0.31 arcsec/yr<br>Parallax: 0.747"<br>Epoch of Position: 2013.480<br>Radial Velocity: 24 km/sec | V=-0.1+/-0.1<br>FUV peak fluxes are 5E-12 per Angstrom (excluding Lyman-alpha, which is ~ 8E-11) | Reference Frame: ICRS |                             |  |  |       |
| <i>Comments: The target coords for epoch 2013.480 were obtained from a global fit to high-precision Chandra HRC images, and includes the parallactic shift at that time. The effective proper motion of Alpha Cen A was determined from the best-fit Center-of-Mass proper motion of the system, based on eight years of Chandra HRC-I measurements, accounting for parallax in each epoch; together with a orbital solution from a recent ephemeris modified slightly to better fit the HRC relative orbit in recent years: Alpha Cen AB are moving toward a close approach (on the sky) in 2016, consequently the orbital motion of A has a significant influence on its apparent proper motion.</i> |   |  |  |  |  |                       |                             |  |  |       |
| (2)  | HD128621-1<br>Alt Name1: ALP-CEN-B  | Offset from HD128620<br>RA Offset: -0.587 Secs<br>Dec Offset: 0.603 Arcsec |  | V=1.33+/-0.1   | Offset Position (HD128621-1)   |                       |                             |  |  |       |
| <i>Comments: Offset of B relative to A determined from the slightly modified ephemeris, which closely matches the relative orbit as recorded by Chandra HRC in recent years.</i>   |   |  |  |  |  |                       |                             |  |  |       |
| Exposures  | #   | Label (ETC Run)  | Target   | Config,Mode,Aperture   | Spectral Els.  | Opt. Params.          | Special Reqs.               | Groups   | Exp. Time (Total)/[Actual Dur.]  | Orbit |
|  | 1   | (STIS.ta.235 462)  | (1) HD128620   | STIS/CCD, ACQ, F25ND5  | MIRROR   |                       | GS ACQ SCENARI<br>O BASE1B3 |  | 0.1 Secs (0.1 Secs)<br>[==>]   | [1]   |
|  | <i>Comments: Castelli and Kurucz Models:G2V;Time to Saturation (for a single exposure) = 2.3 seconds ; Optimum SNR = 179</i>  |  |  |  |  |                       |                             |  |  |       |
|  | 2   | (STIS.sp.23 5468)  | (1) HD128620   | STIS/FUV-MAMA, ACCUM,<br>0.2X0.2   | E140M<br>1425 A  |                       |                             |  | 1425 Secs X 3 (4275 Secs)<br>[==>(Copy 1)]<br>[==>(Copy 2)]<br>[==>(Copy 3)] | [1]   |
|  | <i>Comments: input=StarCAT ETC file for ALP-CEN-A; Exposure time= 4.3 ks at wavelength 1393.80 Å gives: SNR = 58 (per resol); Brightest Pixel (at 1215.40 Å)= 1.8 cps; peak S/N~200 at Lyman-alpha</i>  |  |  |  |  |                       |                             |  |  |       |
| 3  | (STIS.sp.23 5634)   | (2) HD128621-1   | STIS/CCD, ACQ/PEAK,<br>0.3X0.05ND  | G430L<br>4300 A  |  |                       |                             | 0.1 Secs (0.1 Secs)<br>[==>]   | [1]  |       |
| <i>Comments: dispersed light peak-up assuming Castelli-Kurucz Models KOV, V=1.33 mag: at wavelength 5000 Å in 0.1 s, SNR ~ 15 (per resol); Time to Saturation (for a single exposure) ~ 60 seconds. ETC calculation done with 52x0.05 slit, then scaled to 0.3x0.05ND aperture by factor 0.001 (ND=3), as verified by comparing throughputs indicated by separate ETC TARG ACQ runs (without disperser) with ND ap and clear apertures.</i>  |   |  |  |  |  |                       |                             |  |  |       |
| 4  | (STIS.sp.23 5468)   | (2) HD128621-1   | STIS/FUV-MAMA, ACCUM,<br>0.2X0.2   | E140M<br>1425 A  |  |                       |                             | 1425 Secs X 3 (4275 Secs)<br>[==>(Copy 1)]<br>[==>(Copy 2)]<br>[==>(Copy 3)] | [2]  |       |
| <i>Comments: ETC run is for ALP-CEN-A, which has very similar spectrum to ALP-CEN-B in observed flux densities</i>   |   |  |  |  |  |                       |                             |  |  |       |



Proposal 13465 - Visit 11 - Alpha Cen: Climbing out of a Coronal Recession? (year 3 continuation)

Wed Aug 21 01:27:53 GMT 2013

| Visit  | <b>Proposal 13465, Visit 11, implementation</b><br><b>Diagnostic Status: No Diagnostics</b><br>Scientific Instruments: STIS/CCD, STIS/FUV-MAMA<br>Special Requirements: CVZ; BETWEEN 24-JUL-2014:00:00:00 AND 27-JUL-2014:23:59:59<br><i>Comments: This is next available CVZ opportunity roughly six months after first Visit.</i> |  |  |  |  |                       |                             |  |  |       |
|--|---|--|--|--|--|-----------------------|-----------------------------|--|--|-------|
|  | Fixed Targets   | #  | Name   | Target Coordinates   | Targ. Coord. Corrections   | Fluxes                | Miscellaneous               |  |  |       |
| (1)  |   | HD128620<br>Alt Name1: ALP-CEN-A   | RA: 14 39 29.3800 (219.8724167d)<br>Dec: -60 49 58.40 (-60.83289d)<br>Equinox: J2000 | Proper Motion RA: -3.82 arcsec/yr<br>Proper Motion Dec: +0.31 arcsec/yr<br>Parallax: 0.747"<br>Epoch of Position: 2013.480<br>Radial Velocity: 24 km/sec | V=-0.1+/-0.1<br>FUV peak fluxes are 5E-12 per Angstrom (excluding Lyman-alpha, which is ~ 8E-11) | Reference Frame: ICRS |                             |  |  |       |
| <i>Comments: The target coords for epoch 2013.480 were obtained from a global fit to high-precision Chandra HRC images, and includes the parallactic shift at that time. The effective proper motion of Alpha Cen A was determined from the best-fit Center-of-Mass proper motion of the system, based on eight years of Chandra HRC-I measurements, accounting for parallax in each epoch; together with a orbital solution from a recent ephemeris modified slightly to better fit the HRC relative orbit in recent years: Alpha Cen AB are moving toward a close approach (on the sky) in 2016, consequently the orbital motion of A has a significant influence on its apparent proper motion.</i> |   |  |  |  |  |                       |                             |  |  |       |
| (3)  | HD128621-2<br>Alt Name1: ALP-CEN-B  | Offset from HD128620<br>RA Offset: -0.558 Secs<br>Dec Offset: 0.974 Arcsec |  | V=1.33+/-0.1   | Offset Position (HD128621-2)   |                       |                             |  |  |       |
| <i>Comments: Offset of B relative to A determined from the slightly modified ephemeris, which closely matches the relative orbit as recorded by Chandra HRC in recent years.</i>   |   |  |  |  |  |                       |                             |  |  |       |
| Exposures  | #   | Label (ETC Run)  | Target   | Config,Mode,Aperture   | Spectral Els.  | Opt. Params.          | Special Reqs.               | Groups   | Exp. Time (Total)/[Actual Dur.]  | Orbit |
|  | 1   | (STIS.ta.235 462)  | (1) HD128620   | STIS/CCD, ACQ, F25ND5  | MIRROR   |                       | GS ACQ SCENARI<br>O BASE1B3 |  | 0.1 Secs (0.1 Secs)<br>[==>]   | [1]   |
|  | <i>Comments: Castelli and Kurucz Models:G2V;Time to Saturation (for a single exposure) = 2.3 seconds ; Optimum SNR = 179</i>  |  |  |  |  |                       |                             |  |  |       |
|  | 2   | (STIS.sp.23 5468)  | (1) HD128620   | STIS/FUV-MAMA, ACCUM,<br>0.2X0.2   | E140M<br>1425 A  |                       |                             |  | 1425 Secs X 3 (4275 Secs)<br>[==>(Copy 1)]<br>[==>(Copy 2)]<br>[==>(Copy 3)] | [1]   |
|  | <i>Comments: input=StarCAT ETC file for ALP-CEN-A; Exposure time= 4.3 ks at wavelength 1393.80 Å gives: SNR = 58 (per resol); Brightest Pixel (at 1215.40 Å)= 1.8 cps; peak S/N~200 at Lyman-alpha</i>  |  |  |  |  |                       |                             |  |  |       |
| 3  | (STIS.sp.23 5634)   | (3) HD128621-2   | STIS/CCD, ACQ/PEAK,<br>0.3X0.05ND  | G430L<br>4300 A  |  |                       |                             | 0.1 Secs (0.1 Secs)<br>[==>]   | [1]  |       |
| <i>Comments: dispersed light peak-up assuming Castelli-Kurucz Models KOV, V=1.33 mag: at wavelength 5000 Å in 0.1 s, SNR ~ 15 (per resol); Time to Saturation (for a single exposure) ~ 60 seconds. ETC calculation done with 52x0.05 slit, then scaled to 0.3x0.05ND aperture by factor 0.001 (ND=3), as verified by comparing throughputs indicated by separate ETC TARG ACQ runs (without disperser) with ND ap and clear apertures.</i>  |   |  |  |  |  |                       |                             |  |  |       |
| 4  | (STIS.sp.23 5468)   | (3) HD128621-2   | STIS/FUV-MAMA, ACCUM,<br>0.2X0.2   | E140M<br>1425 A  |  |                       |                             | 1425 Secs X 3 (4275 Secs)<br>[==>(Copy 1)]<br>[==>(Copy 2)]<br>[==>(Copy 3)] | [2]  |       |
| <i>Comments: ETC run is for ALP-CEN-A, which has very similar spectrum to ALP-CEN-B in observed flux densities</i>   |   |  |  |  |  |                       |                             |  |  |       |

