



## 14332 - A Precision Measurement of the Mass of the Cepheid V350 Sgr

Cycle: 23, Proposal Category: GO

(UV Initiative)

(Availability Mode: AVAILABLE)

### INVESTIGATORS

<i>Name</i>	<i>Institution</i>	<i>E-Mail</i>
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### VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) V350-SGR WAVE	STIS/CCD STIS/FUV-MAMA	3	29-Jul-2016 13:29:49.0	yes
02	(1) V350-SGR WAVE	STIS/CCD STIS/FUV-MAMA	3	29-Jul-2016 13:29:51.0	yes
03	(1) V350-SGR WAVE	STIS/CCD STIS/FUV-MAMA	3	29-Jul-2016 13:29:53.0	yes
04	(1) V350-SGR WAVE	STIS/CCD STIS/FUV-MAMA	3	29-Jul-2016 13:29:55.0	yes

12 Total Orbits Used

## **ABSTRACT**

An important HST UV legacy is the measurement of the masses of Cepheids. HST has provided double-lined spectroscopic binaries since the orbital velocity amplitude of hot companions can be measured on high resolution ultraviolet spectra. STIS UV E140H echelle observations of the Cepheid V350 Sgr will yield a dramatic improvement in the precision of its mass (5% or 0.25 solar masses vs the current 17%). This will allow a unique and critical test of the role of convective overshoot in the evolution of intermediate mass stars, by coupling the measured mass with a luminosity. Furthermore, the very accurate masses (1-2%) recently determined for two Cepheids in eclipsing binaries in the LMC mean the mass--luminosity relation for Cepheids can be compared for two metallicities. This will improve both confidence in the use of Cepheids as primary extragalactic distance indicators and also our understanding of the evolution of intermediate mass stars.

## **OBSERVING DESCRIPTION**

There are four identical visits of three orbits each.

Auto-wavecalcs will be turned off and replaced with deep wavecalcs at the beginning and end of each orbit. All but the first wavecalc hide in the occultations and or guide-star re-acquisitions.

Other than the initial ACQ, the exposures for each orbit are put into a non-interruptable sequence to ensure that the wavecalcs stay close in time to each bracketed science exposure.

In each visit:

1st orbit

Proposal 14332 (STScI Edit Number: 1, Created: Friday, July 29, 2016 12:29:56 PM EST) - Overview

- (1) Acquire target with 3s F28X52OIII ACQ exposure
- (2) 3 s ACQ/PEAK with G430L and 0.2X0.09 aperture
- (2) 50 s WAVE
- (3) 1762 s E140H 1416
- (4) 50 s WAVE

2nd and 3rd orbits

- (1) 50 s WAVE
- (2) 2970 s E140H 1416
- (3) 50 s WAVE

----- Additional Comments -----

Notes to schedulers:

The STIS FUV detector background rate increases with time after detector HV turn-on. Since our target is rather faint, please SCHEDULE EACH VISIT AS THE FIRST FUV MAMA VISIT IN THAT DAY'S SAA FREE BLOCK, and START THE VISIT AS SOON AS POSSIBLE AFTER THAT DAY'S FUV HV RAMPUP. Please don't schedule any two of these visits in the same SAA free block as the second one will see enhanced detector background.

These observations will provide the final epoch of data for this program. They should occur between 20-Jun-2016 and 01-Oct-2016 to cover the required part of the long period orbit. All visits should be grouped within about five days.

Bright object screening:

The "unknown" objects reported by the BOT are just extra catalog entries for the target itself.

At FUV wavelengths, the SED is dominated by the companion, not the Cepheid.

Adopted SED based on the large aperture, low dispersion IUE spectrum SWP44358 + LWR22766; other IUE spectra consistent.

Expect global rate of 157 c/s background + 82 c/s source for E140H, 1416, 0.2X0.09



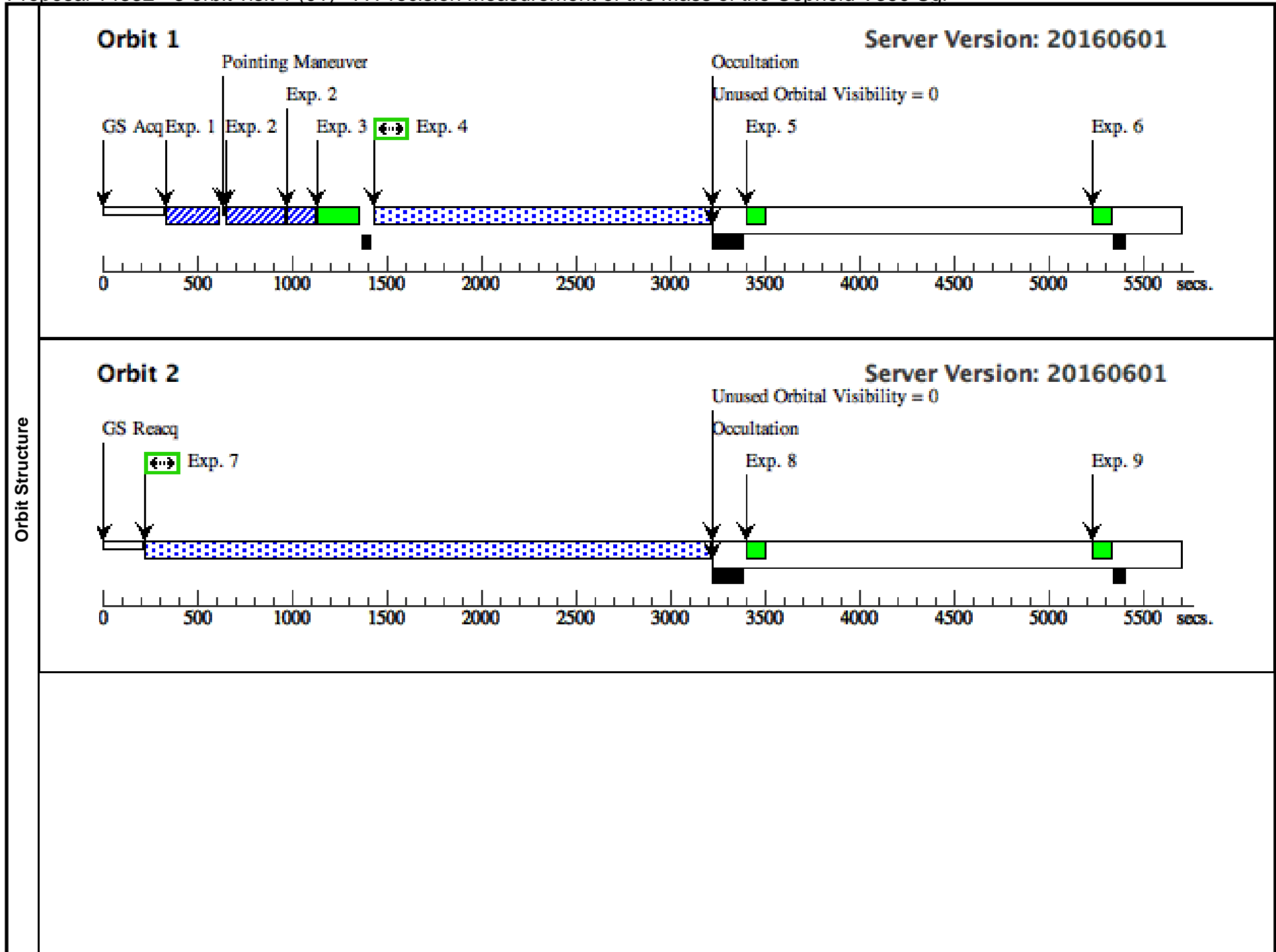
Proposal 14332 - 3 orbit visit 1 (01) - A Precision Measurement of the Mass of the Cepheid V350 Sgr

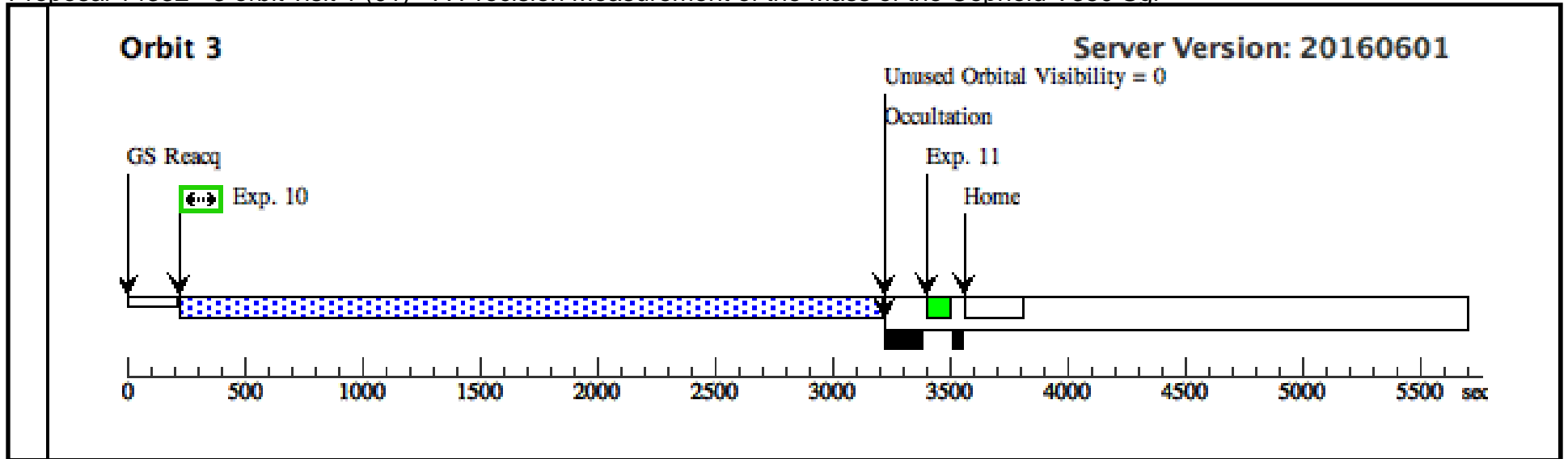
Fri Jul 29 17:29:56 GMT 2016

<b>Visit</b>	<p><b>Proposal 14332, 3 orbit visit 1 (01), implementation</b></p> <p><b>Diagnostic Status: No Diagnostics</b></p> <p>Scientific Instruments: STIS/CCD, STIS/FUV-MAMA</p> <p>Special Requirements: BETWEEN 20-JUN-2016 AND 01-OCT-2016; GROUP 01,02,03,04 WITHIN 5D</p> <p><i>Comments: Note to schedulers: The STIS FUV detector background rate increases with time after detector HV turn-on. Since our target is rather faint (about 3-4X average background), please schedule each visit as the first FUV MAMA visit in that day's SAA free block of orbits, and start the visits as soon as possible after the HV is turned on for the day.</i></p>					
	<b>Fixed Targets</b>	<b>#</b>	<b>Name</b>	<b>Target Coordinates</b>	<b>Targ. Coord. Corrections</b>	<b>Fluxes</b>
(1)		V350-SGR	RA: 18 45 17.4990 (281.3229125d)	Proper Motion RA: 1.40 mas/yr	V=7.47	Reference Frame: ICRS
		Alt Name1: HD173297	Dec: -20 38 50.57 (-20.64738d)	Proper Motion Dec: -4.40 mas/yr	FLAM(1417)=5e-14	
		Alt Name2: HIP92013	Equinox: J2000	Epoch of Position: 2000		
	<p><i>Comments: Coordinates are from Hipparcos, adjusted for proper motion to epoch 2000.</i></p> <p><i>At FUV wavelengths, the SED is dominated by the ~ B9 or A0 companion, not the Cepheid.</i></p> <p><i>Our adopted SED is based on the large aperture, low dispersion IUE spectrum SWP44358 + LWR22766. A few other IUE spectra taken at different epochs are consistent with these, ruling out any large scale FUV variability.</i></p> <p><i>Expect global rate of 157 c/s background + 82 c/s source for E140H, 1416, 0.2X0.09; ETC ID STIS.sp.507182.</i></p>					

Proposal 14332 - 3 orbit visit 1 (01) - A Precision Measurement of the Mass of the Cepheid V350 Sgr

Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
	1	ACQ (STIS.ta.512 994)	(1) V350-SGR	STIS/CCD, ACQ, F28X50OIII	MIRROR				3 Secs (3 Secs) [==>]	[1]	
	2	acq/peak (STIS.sp.51 2995)	(1) V350-SGR	STIS/CCD, ACQ/PEAK, 0.2X0.09	G430L 4300 A			Sequence 2-5 Non-Int in 3 orbit visit 1 (01)	3 Secs (3 Secs) [==>]	[1]	
	<i>Comments: Expect 1e7 e- per dwell point from the source</i>										
	3	wave before	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A			Sequence 2-5 Non-Int in 3 orbit visit 1 (01)	50 Secs (50 Secs) [==>]	[1]	
	4	E140H (STIS.sp.50 7182)	(1) V350-SGR	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1416 A	WAVECAL=NO; BUFFER-TIME=88 1		Sequence 2-5 Non-Int in 3 orbit visit 1 (01)	1762 Secs (1762 Secs) [==>]	[1]	
	<i>Comments: At FUV wavelengths, the SED is dominated by the companion, not the Cepheid. Adopted SED based on the large aperture, low dispersion IUE spectrum SWP44358 + LWR22766; other IUE spectra consistent. Expect global rate of 157 c/s background + 82 c/s source for E140H, 1416, 0.2X0.09</i>										
	5	wave after	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A			Sequence 2-5 Non-Int in 3 orbit visit 1 (01)	50 Secs (50 Secs) [==>]	[1]	
	6	wave before	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A		NEW ALIGNMENT	Sequence 6-8 Non-Int in 3 orbit visit 1 (01)	50 Secs (50 Secs) [==>]	[1]	
	7	E140H (STIS.sp.50 7182)	(1) V350-SGR	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1416 A	WAVECAL=NO; BUFFER-TIME=14 85		Sequence 6-8 Non-Int in 3 orbit visit 1 (01)	2970 Secs (2970 Secs) [==>]	[2]	
	8	wave after	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A			Sequence 6-8 Non-Int in 3 orbit visit 1 (01)	50 Secs (50 Secs) [==>]	[2]	
9	wave before	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A		NEW ALIGNMENT	Sequence 9-11 Non-Int in 3 orbit visit 1 (01)	50 Secs (50 Secs) [==>]	[2]		
10	E140H (STIS.sp.50 7182)	(1) V350-SGR	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1416 A	WAVECAL=NO; BUFFER-TIME=14 85		Sequence 9-11 Non-Int in 3 orbit visit 1 (01)	2970 Secs (2970 Secs) [==>]	[3]		
11	wave after	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A			Sequence 9-11 Non-Int in 3 orbit visit 1 (01)	50 Secs (50 Secs) [==>]	[3]		







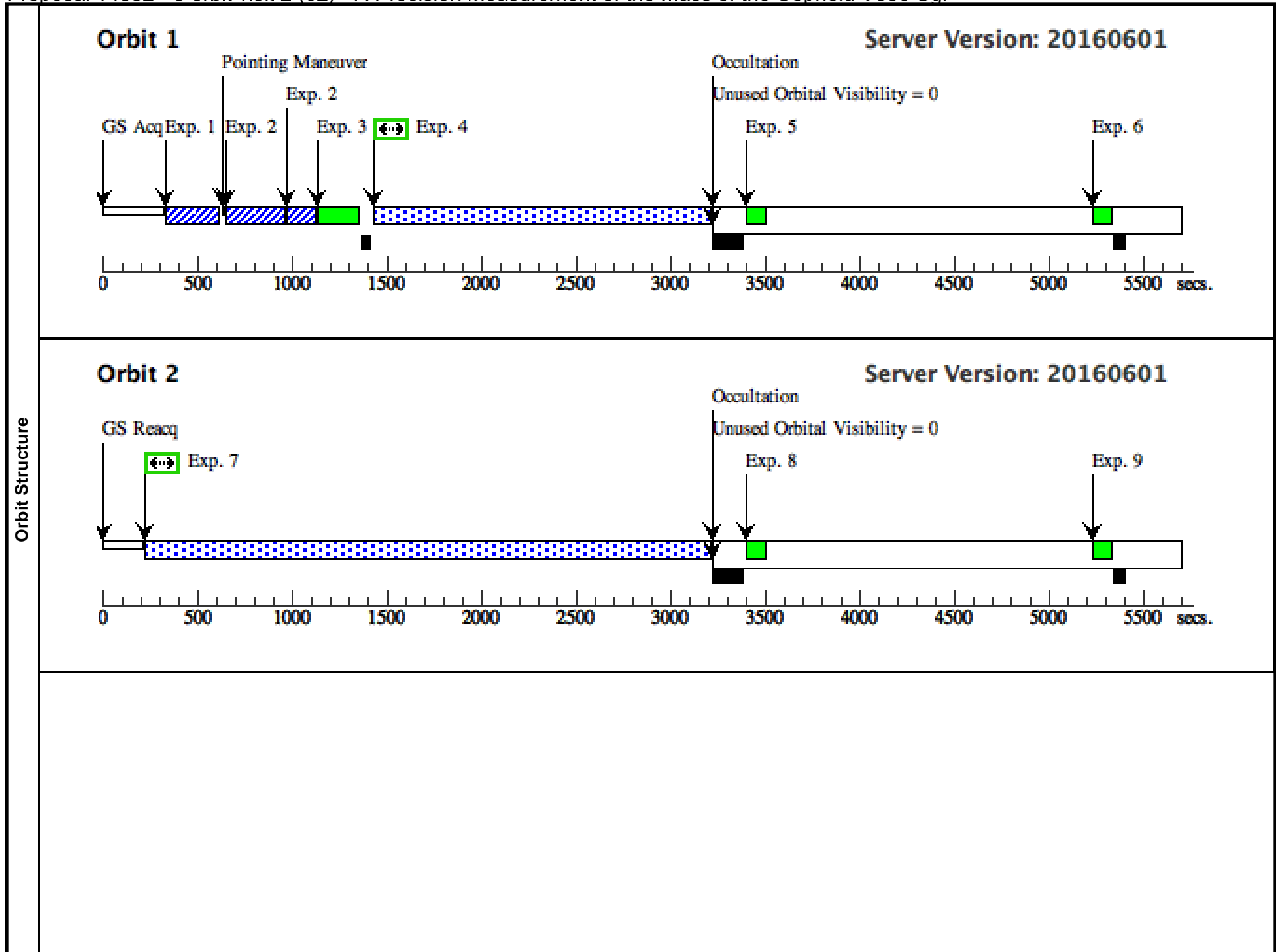
Proposal 14332 - 3 orbit visit 2 (02) - A Precision Measurement of the Mass of the Cepheid V350 Sgr

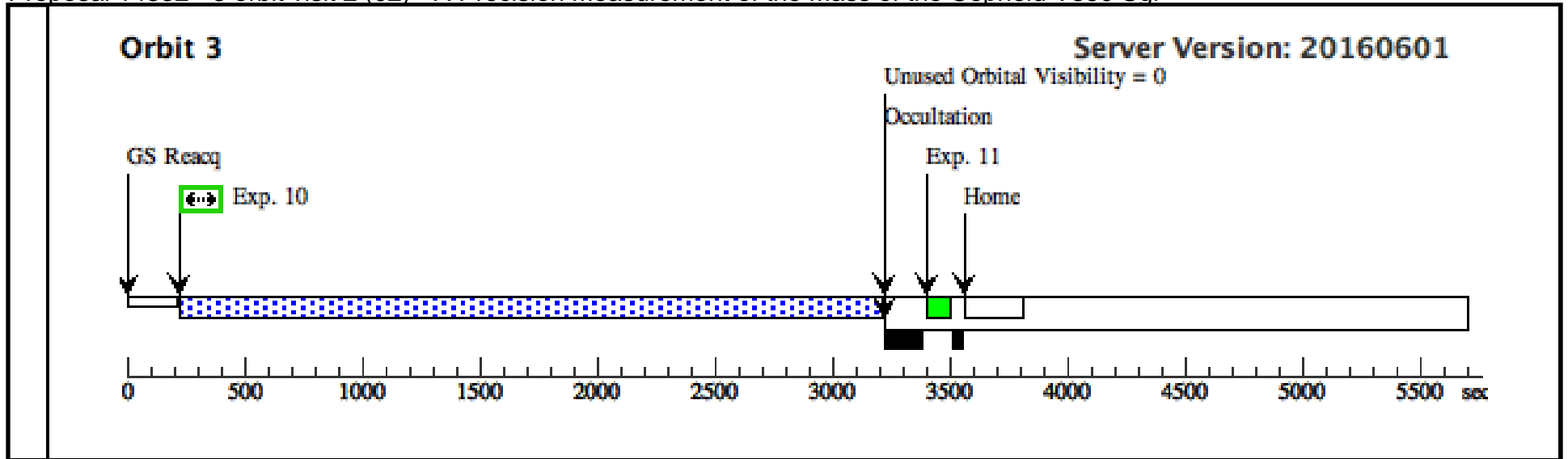
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<b>Visit</b>	<p><b>Proposal 14332, 3 orbit visit 2 (02), implementation</b></p> <p><b>Diagnostic Status: No Diagnostics</b></p> <p>Scientific Instruments: STIS/CCD, STIS/FUV-MAMA</p> <p>Special Requirements: BETWEEN 20-JUN-2016 AND 01-OCT-2016</p> <p><i>Comments: Note to schedulers: The STIS FUV detector background rate increases with time after detector HV turn-on. Since our target is rather faint (about 3-4X average background), please schedule each visit as the first FUV MAMA visit in that day's SAA free block of orbits, and start the visits as soon as possible after the HV is turned on for the day.</i></p>					
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		Alt Name2: HIP92013	Equinox: J2000	Epoch of Position: 2000		
	<p><i>Comments: Coordinates are from Hipparcos, adjusted for proper motion to epoch 2000.</i></p> <p><i>At FUV wavelengths, the SED is dominated by the ~ B9 or A0 companion, not the Cepheid.</i></p> <p><i>Our adopted SED is based on the large aperture, low dispersion IUE spectrum SWP44358 + LWR22766. A few other IUE spectra taken at different epochs are consistent with these, ruling out any large scale FUV variability.</i></p> <p><i>Expect global rate of 157 c/s background + 82 c/s source for E140H, 1416, 0.2X0.09; ETC ID STIS.sp.507182.</i></p>					

Proposal 14332 - 3 orbit visit 2 (02) - A Precision Measurement of the Mass of the Cepheid V350 Sgr

Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
	1	ACQ (STIS.ta.512 994)	(1) V350-SGR	STIS/CCD, ACQ, F28X50OIII	MIRROR				3 Secs (3 Secs) [==>]	[1]	
	2	acq/peak (STIS.sp.51 2995)	(1) V350-SGR	STIS/CCD, ACQ/PEAK, 0.2X0.09	G430L 4300 A			Sequence 2-5 Non-Int in 3 orbit visit 2 (02)	3 Secs (3 Secs) [==>]	[1]	
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	3	wave before	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A			Sequence 2-5 Non-Int in 3 orbit visit 2 (02)	50 Secs (50 Secs) [==>]	[1]	
	4	E140H (STIS.sp.50 7182)	(1) V350-SGR	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1416 A	WAVECAL=NO; BUFFER-TIME=88 1		Sequence 2-5 Non-Int in 3 orbit visit 2 (02)	1762 Secs (1762 Secs) [==>]	[1]	
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	7	E140H (STIS.sp.50 7182)	(1) V350-SGR	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1416 A	WAVECAL=NO; BUFFER-TIME=14 85		Sequence 6-8 Non-Int in 3 orbit visit 2 (02)	2970 Secs (2970 Secs) [==>]	[2]	
	8	wave after	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A			Sequence 6-8 Non-Int in 3 orbit visit 2 (02)	50 Secs (50 Secs) [==>]	[2]	
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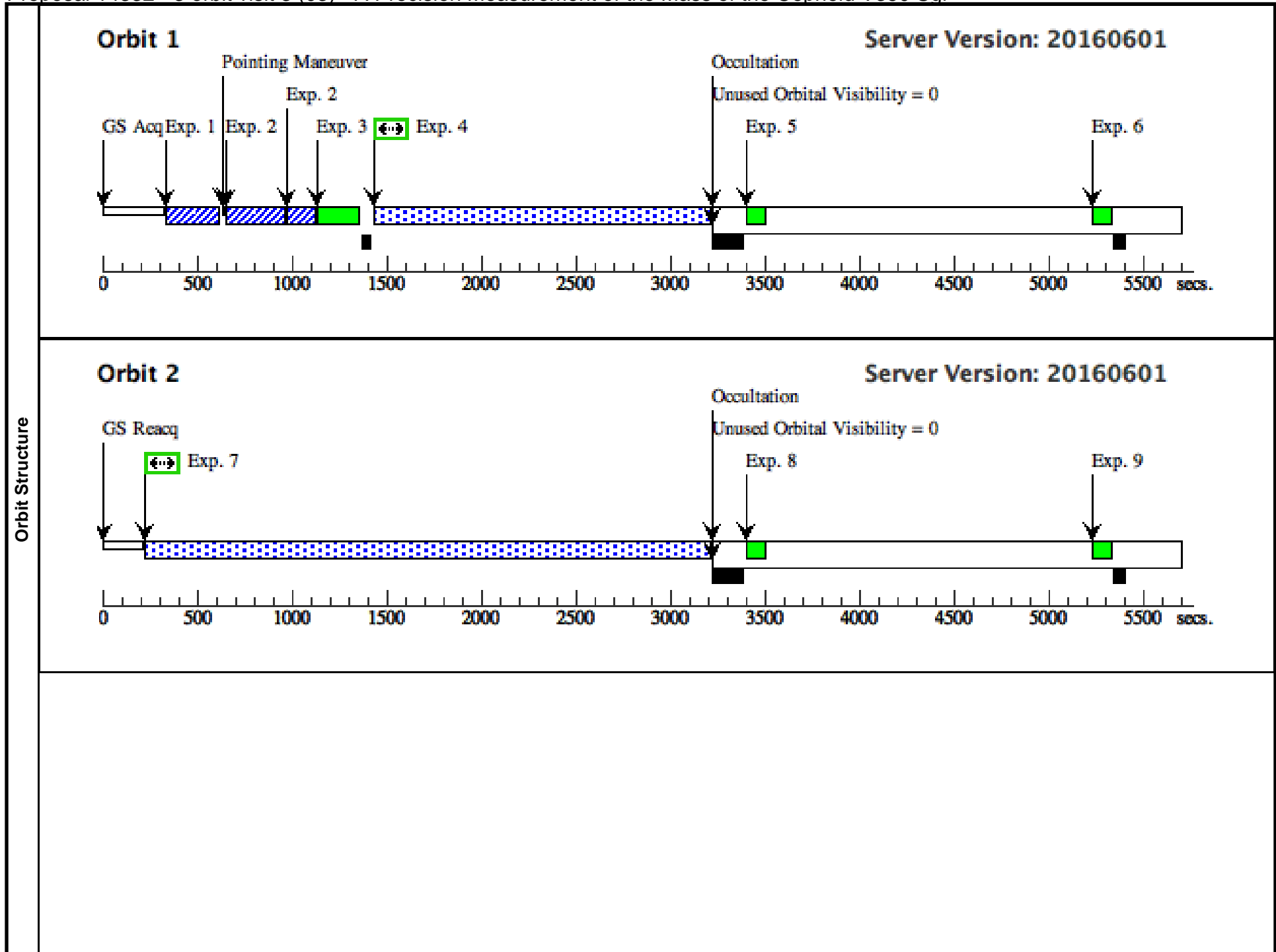
Proposal 14332 - 3 orbit visit 3 (03) - A Precision Measurement of the Mass of the Cepheid V350 Sgr

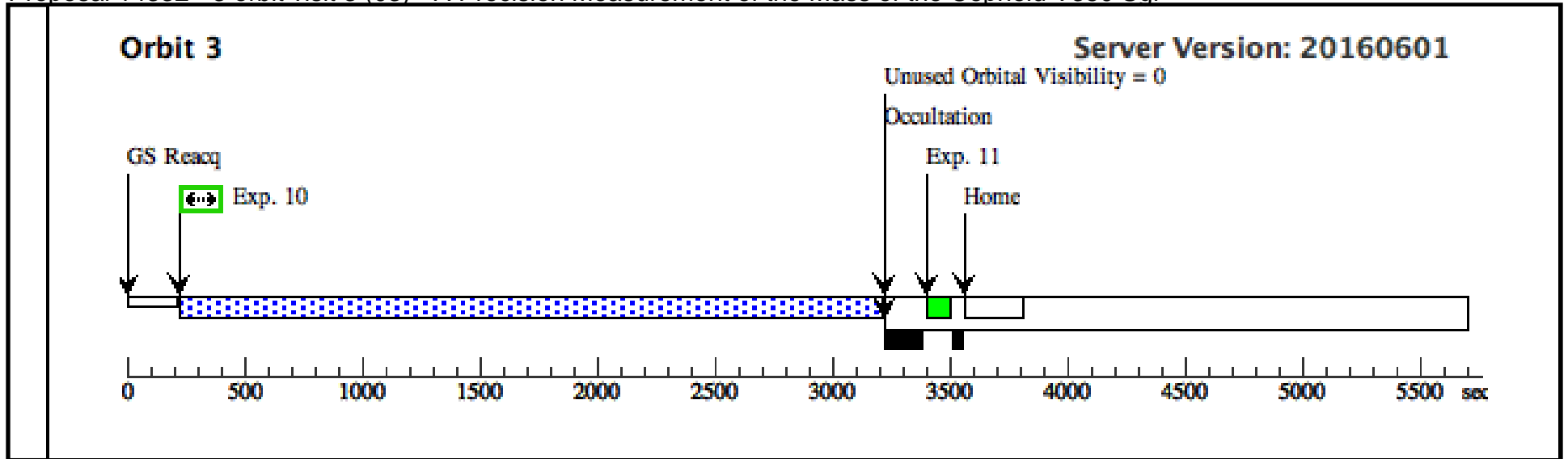
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<b>Visit</b>	<p><b>Proposal 14332, 3 orbit visit 3 (03), implementation</b></p> <p><b>Diagnostic Status: No Diagnostics</b></p> <p>Scientific Instruments: STIS/CCD, STIS/FUV-MAMA</p> <p>Special Requirements: BETWEEN 20-JUN-2016 AND 01-OCT-2016</p> <p><i>Comments: Note to schedulers: The STIS FUV detector background rate increases with time after detector HV turn-on. Since our target is rather faint (about 3-4X average background), please schedule each visit as the first FUV MAMA visit in that day's SAA free block of orbits, and start the visits as soon as possible after the HV is turned on for the day.</i></p>					
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Proposal 14332 - 3 orbit visit 3 (03) - A Precision Measurement of the Mass of the Cepheid V350 Sgr

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	8	wave after	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A			Sequence 6-8 Non-Int in 3 orbit visit 3 (03)	50 Secs (50 Secs) [==>]	[2]	
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Proposal 14332 - 3 orbit visit 4 (04) - A Precision Measurement of the Mass of the Cepheid V350 Sgr

Fri Jul 29 17:29:57 GMT 2016

<b>Visit</b>	<p><b>Proposal 14332, 3 orbit visit 4 (04), implementation</b></p> <p><b>Diagnostic Status: No Diagnostics</b></p> <p>Scientific Instruments: STIS/CCD, STIS/FUV-MAMA</p> <p>Special Requirements: BETWEEN 20-JUN-2016 AND 01-OCT-2016</p> <p><i>Comments: Note to schedulers: The STIS FUV detector background rate increases with time after detector HV turn-on. Since our target is rather faint (about 3-4X average background), please schedule each visit as the first FUV MAMA visit in that day's SAA free block of orbits, and start the visits as soon as possible after the HV is turned on for the day.</i></p>																													
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	Alt Name2: HIP92013	Equinox: J2000	Epoch of Position: 2000																											

Proposal 14332 - 3 orbit visit 4 (04) - A Precision Measurement of the Mass of the Cepheid V350 Sgr

Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
	1	ACQ (STIS.ta.512 994)	(1) V350-SGR	STIS/CCD, ACQ, F28X50OIII	MIRROR				3 Secs (3 Secs) [==>]	[1]	
	2	acq/peak (STIS.sp.51 2995)	(1) V350-SGR	STIS/CCD, ACQ/PEAK, 0.2X0.09	G430L 4300 A			Sequence 2-5 Non-Int in 3 orbit visit 4 (04)	3 Secs (3 Secs) [==>]	[1]	
	<i>Comments: Expect 1e7 e- per dwell point from the source</i>										
	3	wave before	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A			Sequence 2-5 Non-Int in 3 orbit visit 4 (04)	50 Secs (50 Secs) [==>]	[1]	
	4	E140H (STIS.sp.50 7182)	(1) V350-SGR	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1416 A	WAVECAL=NO; BUFFER-TIME=88 1		Sequence 2-5 Non-Int in 3 orbit visit 4 (04)	1762 Secs (1762 Secs) [==>]	[1]	
	<i>Comments: At FUV wavelengths, the SED is dominated by the companion, not the Cepheid. Adopted SED based on the large aperture, low dispersion IUE spectrum SWP44358 + LWR22766; other IUE spectra consistent. Expect global rate of 157 c/s background + 82 c/s source for E140H, 1416, 0.2X0.09</i>										
	5	wave after	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A			Sequence 2-5 Non-Int in 3 orbit visit 4 (04)	50 Secs (50 Secs) [==>]	[1]	
	6	wave before	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A		NEW ALIGNMENT	Sequence 6-8 Non-Int in 3 orbit visit 4 (04)	50 Secs (50 Secs) [==>]	[1]	
	7	E140H (STIS.sp.50 7182)	(1) V350-SGR	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1416 A	WAVECAL=NO; BUFFER-TIME=14 85		Sequence 6-8 Non-Int in 3 orbit visit 4 (04)	2970 Secs (2970 Secs) [==>]	[2]	
	8	wave after	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A			Sequence 6-8 Non-Int in 3 orbit visit 4 (04)	50 Secs (50 Secs) [==>]	[2]	
9	wave before	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A		NEW ALIGNMENT	Sequence 9-11 Non-Int in 3 orbit visit 4 (04)	50 Secs (50 Secs) [==>]	[2]		
10	E140H (STIS.sp.50 7182)	(1) V350-SGR	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1416 A	WAVECAL=NO; BUFFER-TIME=14 85		Sequence 9-11 Non-Int in 3 orbit visit 4 (04)	2970 Secs (2970 Secs) [==>]	[3]		
11	wave after	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1416 A			Sequence 9-11 Non-Int in 3 orbit visit 4 (04)	50 Secs (50 Secs) [==>]	[3]		

