



# 15112 - The low-metallicity starburst NGC346: massive-star population and feedback

Cycle: 25, Proposal Category: GO

(UV Initiative)

(Availability Mode: SUPPORTED)

## INVESTIGATORS

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## VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) NGC-346-W3-OFFSET (2) NGC346-A (3) NGC346-B (4) NGC346-C (7) NGC-346-W3	STIS/CCD STIS/FUV-MAMA	3	07-Aug-2018 13:00:19.0	yes
02	(1) NGC-346-W3-OFFSET (5) NGC346-C2 (6) NGC346-D	STIS/CCD STIS/FUV-MAMA	2	07-Aug-2018 13:00:21.0	yes

5 Total Orbits Used

## ABSTRACT

The Small Magellanic Cloud (SMC) is ideal to study young, massive stars at low metallicity. The compact cluster NGC346 contains about half of all O-type stars in the entire SMC. The massive-star population of this cluster powers N66, the brightest and largest HII region in the SMC.

We propose to use HST-STIS to slice NGC346 with 20 long-slit exposures, in order to obtain the UV spectra of most of the massive early-type stars of this cluster. Archival data of 13 exposures that cover already a minor part of this cluster will be included in our analyses.

Our aim is to quantitatively analyze virtually the whole massive-star population of NGC346. We have already secured the optical spectra of all massive stars in the field with the integral-field spectrograph MUSE at the ESO-VLT. However, for the determination of the stellar-wind parameters, i.e. the mass-loss rates and the wind velocities, ultraviolet spectra are indispensable. Our advanced "Potsdam Wolf-Rayet (PoWR)" code will be used for modeling the stellar and wind spectra in the course of the analysis.

Finally, we will obtain:

- (a) the fundamental stellar and wind parameters of all stars brighter than spectral type B2V in the field, which, e.g., will constrain the initial mass function in this young low-metallicity starburst;
- (b) mass-loss rates of many more OB-type stars at SMC metallicity than hitherto known, allowing to better constrain their metallicity dependence;
- (c) the integrated feedback by ionizing radiation and stellar winds of the whole massive-star population of NGC346, which will be used as input to model the ecology of the giant HII region N66.

These HST UV data will be of high legacy value.

## **OBSERVING DESCRIPTION**

The observations shall map major parts of the young stellar cluster NGC346 in the SMC. We want to take STIS long-slit UV spectra arranged in mosaic of adjacent slit positions. The slit is 52" x 2", and the positions should overlap by 0.2" in dispersion direction. The aim is to extract as many as possible stellar spectra of the blue cluster stars.

The fields are carefully chosen, taking into account that a minor part of the cluster has already been observed with a similar technique with STIS before (GO-8629, PI Bruhweiler), see Figure3 of the Phase I proposal.

For the visual brightness that is requested in the "target" forms we have always given the UV-brightest star in the whole mosaic field (CI\* NGC346 W3, V=12.8mag), although this star is only covered by one of the slit positions.

For the precise positioning of the slit exposures, we use first a reference star for target acquisition and peakup. Then, the slit is positioned with relative offsets using the "STIS PERP-TO-SLIT" pattern.

In addition to the coordinate offsets, we apply an offset of 3" in slit direction in order to account for the spectral offset of the MAMA detector described in Sect. 7.6 of the "STIS Instrument Handbook for Cycle 25" (Version 16.0).

Our mosaic coverage requires a position angle of the slit of about 35deg. We specify here a tolerance of plus/minus 10deg in order to allow for larger windows of visibility. If this would still lead to scheduling problems, we could release this constraint further. In any case, we would ask to contact us when the orbits have been scheduled and the exact position angle of the slit is known, so that we could check and re-adjust the position offsets in order not to miss any interesting star.

Following the special note for Cycle 25, we have structured our 5 orbits into visits of 3 + 2 orbits. If it would finally happen that the orbits can be scheduled back-to-back, the target acquisition and peakup at the beginning of orbit 4 would not be necessary and the time can be distributed for the science exposures.

In each orbit we have left 165s at the end unused, because we noticed that otherwise the visibility is strongly reduced. Of course, finally all available time should be evenly distributed to extend the science exposures to the maximum possible.

Proposal 15112 - Visit 01 - The low-metallicity starburst NGC346: massive-star population and feedback

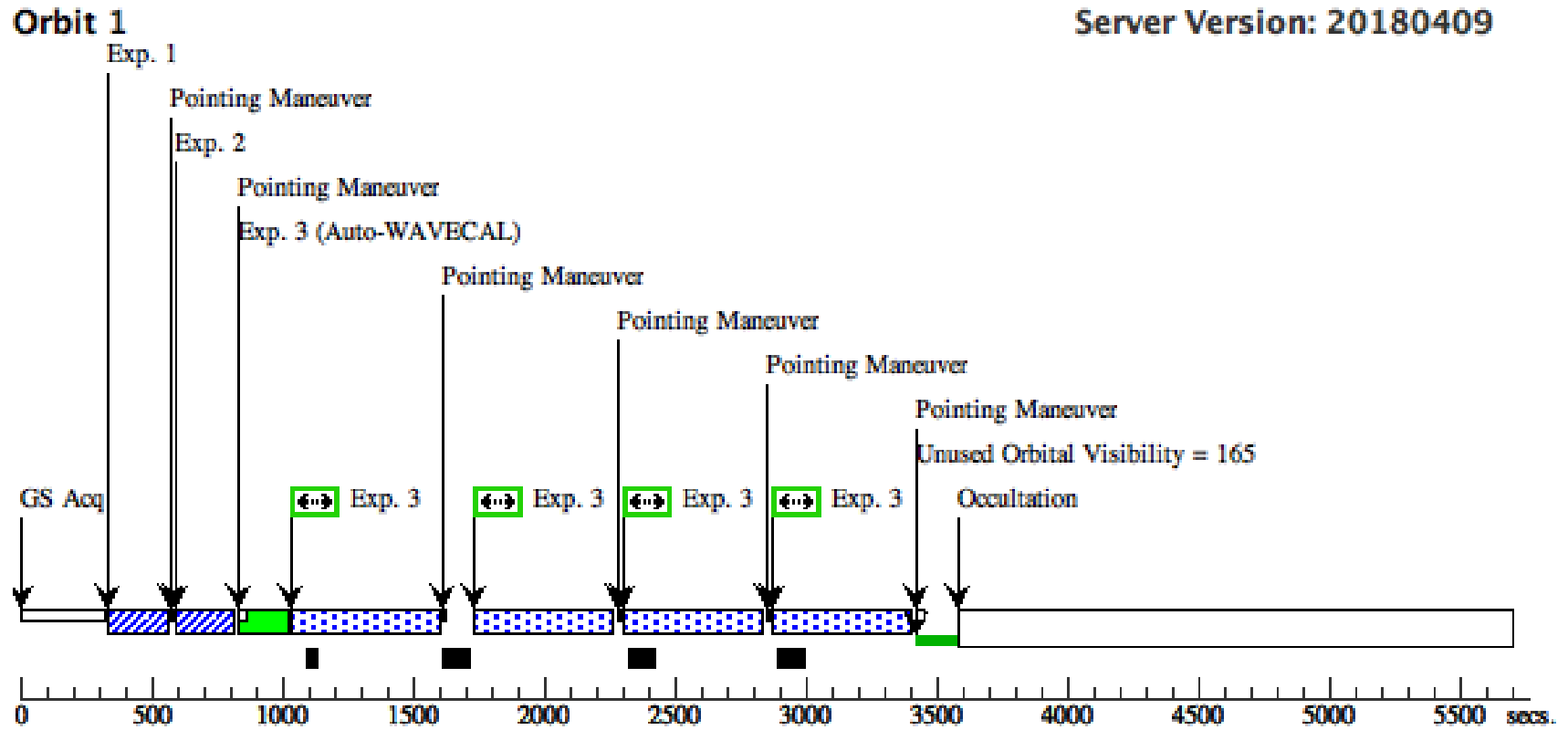
Tue Aug 07 17:00:22 GMT 2018

Visit	<b>Proposal 15112, Visit 01, implementation</b> <b>Diagnostic Status: No Diagnostics</b> Scientific Instruments: STIS/CCD, STIS/FUV-MAMA Special Requirements: ORIENT 70D TO 90 D						
	#	Primary Pattern	Secondary Pattern	Exposures			
Patterns	(1)	Pattern Type=STIS-PERP-TO-SLIT    Coordinate Frame=POS-TARG Purpose=MOSAIC    Pattern Orientation=0.0 Number Of Points=4    Angle Between Sides= Point Spacing=1.8    Center Pattern=false Line Spacing=		(3), (4)			
	(2)	Pattern Type=STIS-PERP-TO-SLIT    Coordinate Frame=POS-TARG Purpose=MOSAIC    Pattern Orientation=0.0 Number Of Points=3    Angle Between Sides= Point Spacing=1.8    Center Pattern=false Line Spacing=		(5)			
Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous	
	(1)	NGC-346-W3-OFFSET	RA: 00 59 0.7500 (14.7531250d) Dec: -72 10 28.20 (-72.17450d) Equinox: J2000		V=12.8	Reference Frame: ICRS	
	<i>Comments: Category=STAR Description=[OF] Extended=NO</i>						
	(2)	NGC346-A	Offset from NGC-346-W3-OFFSET RA Offset: 0.359 Secs Dec Offset: -0.91 Arcsec		V=12.8+/-0.04 abmag=12.8 in filter WFC3/UVI S/F225W	Offset Position (NGC346-A)	
	<i>Comments: Category=STELLAR CLUSTER Description=[OPEN CLUSTER]</i>						
	(3)	NGC346-B	Offset from NGC-346-W3-OFFSET RA Offset: 5.636 Secs Dec Offset: 5.12 Arcsec		V=12.8 abmag>13.06 in filter WFC3/U VIS/F225W	Offset Position (NGC346-B)	
	<i>Comments: Category=STELLAR CLUSTER Description=[OPEN CLUSTER]</i>						
(4)	NGC346-C	Offset from NGC-346-W3-OFFSET RA Offset: 3.236 Secs Dec Offset: -17.88 Arcsec		V=12.8 abmag>13.06 in filter WFC3/U VIS/F225W	Offset Position (NGC346-C)		
<i>Comments: Category=STELLAR CLUSTER Description=[OPEN CLUSTER]</i>							
(7)	NGC-346-W3	RA: 00 59 0.7500 (14.7531250d) Dec: -72 10 28.20 (-72.17450d) Equinox: J2000		V=12.8	Reference Frame: ICRS		
<i>Comments: Category=STAR Description=[OF] Extended=NO</i>							

Proposal 15112 - Visit 01 - The low-metallicity starburst NGC346: massive-star population and feedback

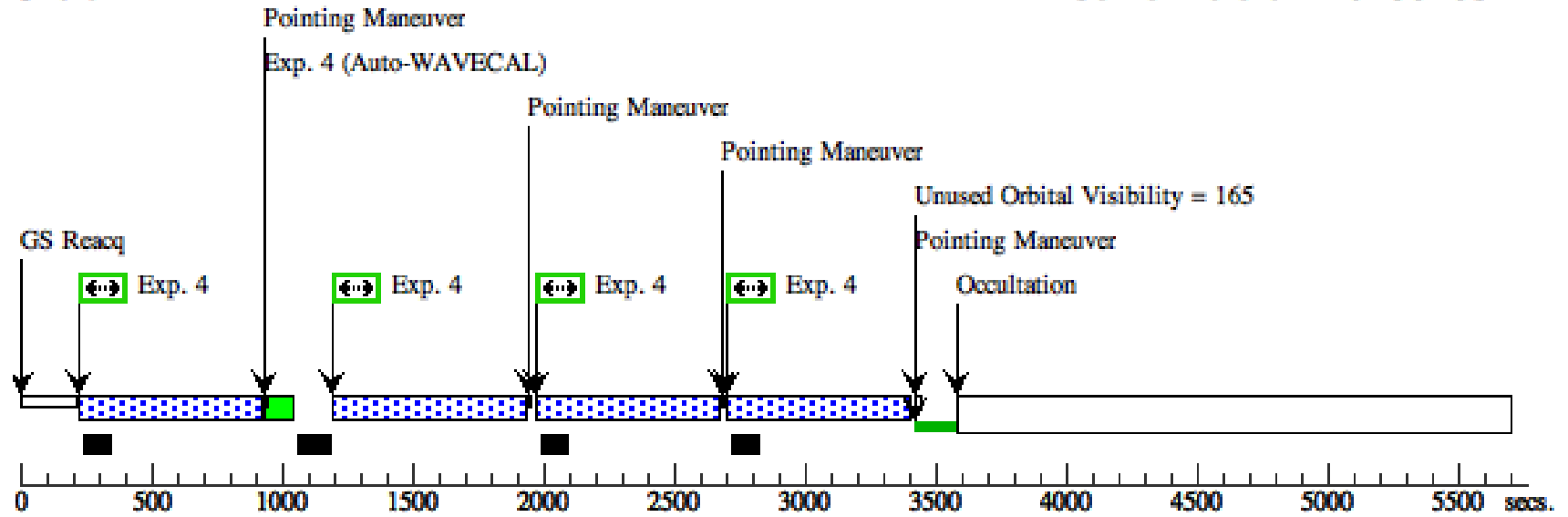
Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	(STIS.ta.100 5884)	(1) NGC-346-W3-O FFSET	STIS/CCD, ACQ, F28X50LP	MIRROR				1.0 Secs (1 Secs) [==>]	[1]
	2	(STIS.ta.100 5893)	(1) NGC-346-W3-O FFSET	STIS/CCD, ACQ/PEAK, 52X0.1	MIRROR				0.5 Secs (0.5 Secs) [==>]	[1]
	3	(STIS.sp.93 9708)	(3) NGC346-B	STIS/FUV-MAMA, ACCUM, 52X2	G140L 1425 A		POS TARG 0.,3.	Pattern 1, Exps 3-3 i n Visit 01 (1)	520 Secs (2080 Secs) [==>(Pattern 1)] [==>(Pattern 2)] [==>(Pattern 3)] [==>(Pattern 4)]	[1]
	4	(STIS.sp.93 9708)	(4) NGC346-C	STIS/FUV-MAMA, ACCUM, 52X2	G140L 1425 A		POS TARG 0.,3.	Pattern 1, Exps 4-4 i n Visit 01 (1)	686 Secs (2744 Secs) [==>(Pattern 1)] [==>(Pattern 2)] [==>(Pattern 3)] [==>(Pattern 4)]	[2]
	5	(STIS.sp.93 9708)	(2) NGC346-A	STIS/FUV-MAMA, ACCUM, 52X2	G140L 1425 A		POS TARG 0.,3.	Pattern 2, Exps 5-5 i n Visit 01 (2)	625 Secs (1875 Secs) [==>(Pattern 1)] [==>(Pattern 2)] [==>(Pattern 3)]	[3]
	6	(STIS.ta.100 5884)	(7) NGC-346-W3	STIS/CCD, ACQ, F28X50LP	MIRROR				1.0 Secs (1 Secs) [==>]	[3]
	7	(STIS.sp.11 77436)	(7) NGC-346-W3	STIS/FUV-MAMA, ACCUM, 52X0.2	G140L 1425 A				225 Secs (225 Secs) [==>]	[3]

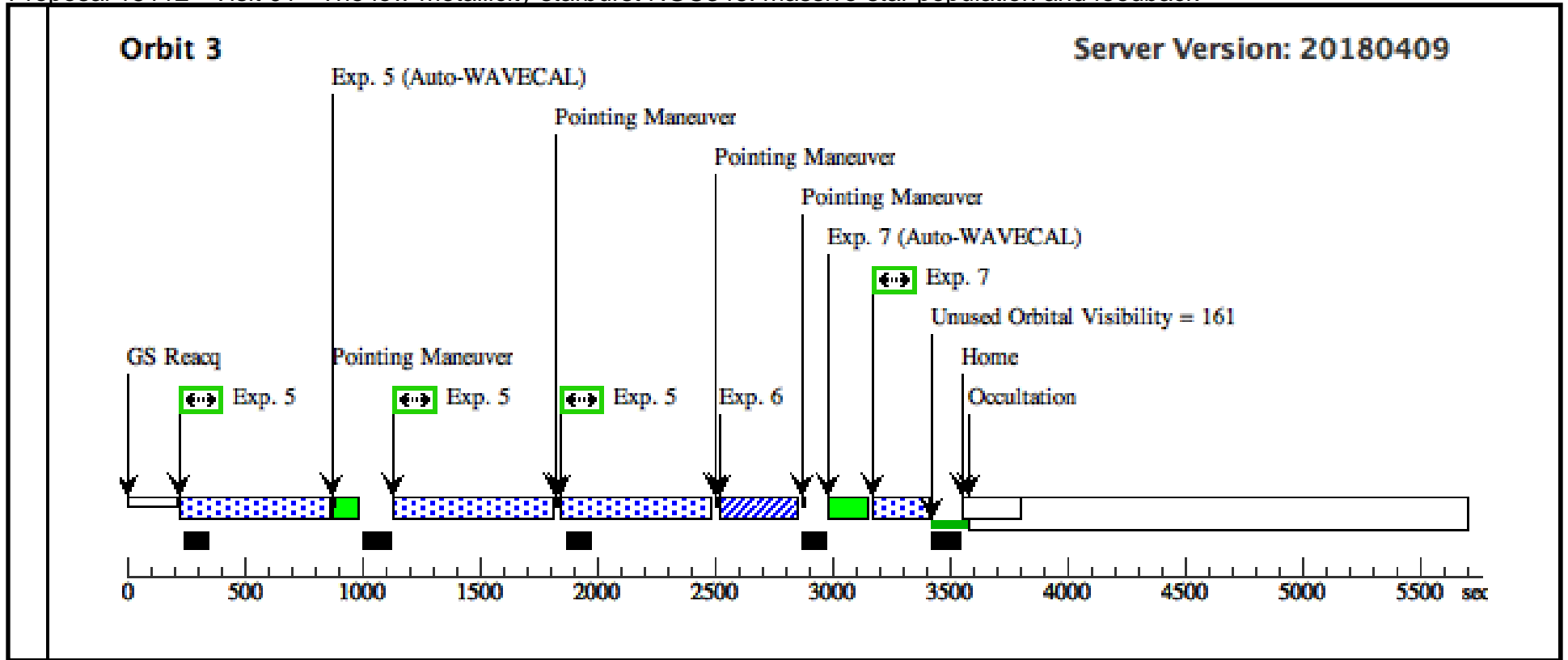
Orbit Structure



**Orbit 2**

**Server Version: 20180409**







Proposal 15112 - Visit 02 - The low-metallicity starburst NGC346: massive-star population and feedback

Tue Aug 07 17:00:23 GMT 2018

Visit	<b>Proposal 15112, Visit 02, implementation</b> <b>Diagnostic Status: No Diagnostics</b> Scientific Instruments: STIS/CCD, STIS/FUV-MAMA Special Requirements: ORIENT 70D TO 90 D										
	Patterns	#	Primary Pattern			Secondary Pattern			Exposures		
		(1)	Pattern Type=STIS-PERP-TO-SLIT Purpose=MOSAIC Number Of Points=4 Point Spacing=1.8 Line Spacing=	Coordinate Frame=POS-TARG Pattern Orientation=0.0 Angle Between Sides= Center Pattern=false							(3), (4)
Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous					
	(1)	NGC-346-W3-OFFSET	RA: 00 59 0.7500 (14.7531250d) Dec: -72 10 28.20 (-72.17450d) Equinox: J2000		V=12.8	Reference Frame: ICRS					
		<i>Comments:</i> Category=STAR Description=[OF] Extended=NO									
	(5)	NGC346-C2	Offset from NGC-346-W3-OFFSET RA Offset: 4.525 Secs Dec Offset: -21.98 Arcsec		V=12.8 abmag>13.06 in filter WFC3/U VIS/F225W	Offset Position (NGC346-C2)					
	<i>Comments:</i> Category=STELLAR CLUSTER Description=[OPEN CLUSTER]										
(6)	NGC346-D	Offset from NGC-346-W3-OFFSET RA Offset: 6.736 Secs Dec Offset: -19.88 Arcsec		V=12.8 abmag>13.06 in filter WFC3/U VIS/F225W	Offset Position (NGC346-D)						
	<i>Comments:</i> Category=STELLAR CLUSTER Description=[OPEN CLUSTER]										
Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
	1	(STIS.ta.100 5884)	(1) NGC-346-W3-OFFSET	STIS/CCD, ACQ, F28X50LP	MIRROR				1.0 Secs (1 Secs) [==>]	[1]	
	2	(STIS.ta.100 5893)	(1) NGC-346-W3-OFFSET	STIS/CCD, ACQ/PEAK, 52X0.1	MIRROR				0.5 Secs (0.5 Secs) [==>]	[1]	
	3	(STIS.sp.93 9708)	(5) NGC346-C2	STIS/FUV-MAMA, ACCUM, 52X2	G140L 1425 A		POS TARG 0.,3.	Pattern 1, Exps 3-3 in Visit 02 (1)	520 Secs (2080 Secs) [==>(Pattern 1)] [==>(Pattern 2)] [==>(Pattern 3)] [==>(Pattern 4)]	[1]	
	4	(STIS.sp.93 9708)	(6) NGC346-D	STIS/FUV-MAMA, ACCUM, 52X2	G140L 1425 A		POS TARG 0.,3.	Pattern 1, Exps 4-4 in Visit 02 (1)	686 Secs (2744 Secs) [==>(Pattern 1)] [==>(Pattern 2)] [==>(Pattern 3)] [==>(Pattern 4)]	[2]	

Orbit Structure

