



16046 - Understanding the Instability Strip of Accreting Pulsating White Dwarfs: Followup to a Recent COS Result

Cycle: 27, Proposal Category: GO

(UV Initiative)

(Availability Mode: AVAILABLE)

INVESTIGATORS

<i>Name</i>	<i>Institution</i>	<i>E-Mail</i>
Prof. Paula Szkody (PI) (Contact)	University of Washington	szkody@astro.washington.edu
Prof. Boris T. Gaensicke (CoI) (ESA Member)	The University of Warwick	boris.gaensicke@warwick.ac.uk
Dr. Patrick Godon (CoI)	Villanova University	patrick.godon@villanova.edu
Prof. Edward M. Sion (CoI)	Villanova University	edward.sion@villanova.edu
Dr. Stella Kafka (CoI)	American Association Of Variable Star Observers	skafka@aavso.org
Dr. Odette Fabiola Toloza Castillo (CoI) (ESA Member)	The University of Warwick	odette.toloza@warwick.ac.uk
Dr. Keaton Bell (CoI)	University of Washington	keatonb@uw.edu

VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
01	(1) V-V386-SER	COS/FUV COS/NUV	3	11-Feb-2020 17:00:16.0	yes
1B	(2) TARGET01-SAFE-TARGET	COS/FUV COS/NUV	3	11-Feb-2020 17:00:18.0	yes
1A	(1) V-V386-SER	S/C	1	11-Feb-2020 17:00:19.0	yes

7 Total Orbits Used

ABSTRACT

Analysis of August, 2019 COS time-tag ultraviolet spectra shows great promise for confirmation of the physical properties defining the observed wide instability strip for pulsating white dwarfs in close binaries accreting material from their companions. The dwarf nova V386 Ser containing a pulsating white dwarf underwent its first rare large amplitude outburst in January 2019. The August UV spectra 7 months after this outburst implies a large He abundance from a white dwarf $\sim 5000\text{K}$ hotter than at quiescence, with a pulsation period 6 times shorter than its quiescent one. Both the temperature and period fit with a mode driven by a He enriched composition, exactly what is predicted as the cause of the wide instability strip. For the first time, the observations match theoretical predictions! We propose to monitor the cooling during the next 6 months to further constrain the HeII and H/HeI convection zones in pulsating white dwarfs. This information is unique to the few white dwarf pulsators that undergo dwarf nova outbursts, as their heating and cooling only takes months-few years (versus millions of years in single white dwarf pulsators), presenting a rare opportunity to track how a deepening convection zone begins to drive pulsation modes with longer periods. Unfortunately, the January outburst missed the normal Cycle 26 deadline and the August observation (a Cycle 26 mid proposal) took place too late for results to provide a strong case for the Cycle 27 deadline. Due to the decades long recurrence time between outbursts in dwarf novae with pulsating white dwarfs, this is likely the last object HST can follow from outburst to quiescence.

OBSERVING DESCRIPTION

We will use COS with the G140L grating (setting 800) in time-tag mode for 3 consecutive HST orbits.

This same configuration was used for the 2 orbits in our August 2019 observation which provided a useful spectrum with coverage between 900-1900Å and detection of 3% amplitude pulsations. The AAVSO shows the system is currently at 18.5 magnitude, about one-half magnitude above quiescence and a half-magnitude fainter than the previous Aug observation. The 3 orbits will provide an adequate spectrum for determining the temperature which is expected to be somewhat cooler than the 20,000K determined in August. To maximize the determination of pulsations, the orbits should be consecutive. To maximize the cooling rate, the observations should be conducted as soon as possible within the window when the object is observable from the ground (Feb-March). Target acquisition will be done using the ACQ/IMAGE in the NUV.

This is the first outburst since the discovery 19 years ago, meaning the outburst recurrence time is several decades. However, we will determine the precise quiescent magnitude within 24 hrs of the HST observation as done in the past.

While the science is optimized if the observations take place as soon as possible (before the object cools to its quiescent temperature), if reduced gyros impact the observation windows, the observations can be made up to August 2020 when it will no longer be visible from the ground.

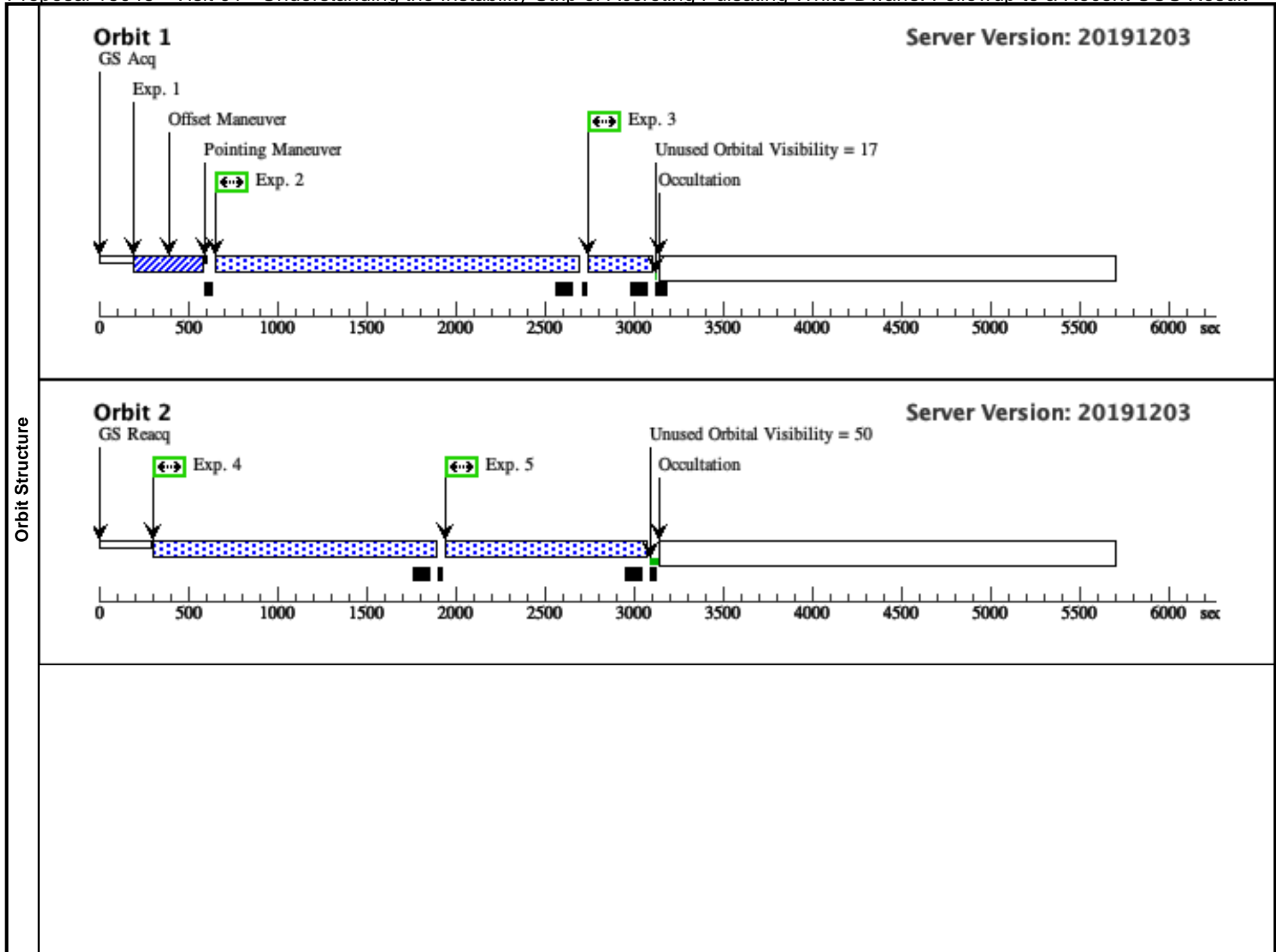
Proposal 16046 - Visit 01 - Understanding the Instability Strip of Accreting Pulsating White Dwarfs: Followup to a Recent COS Result

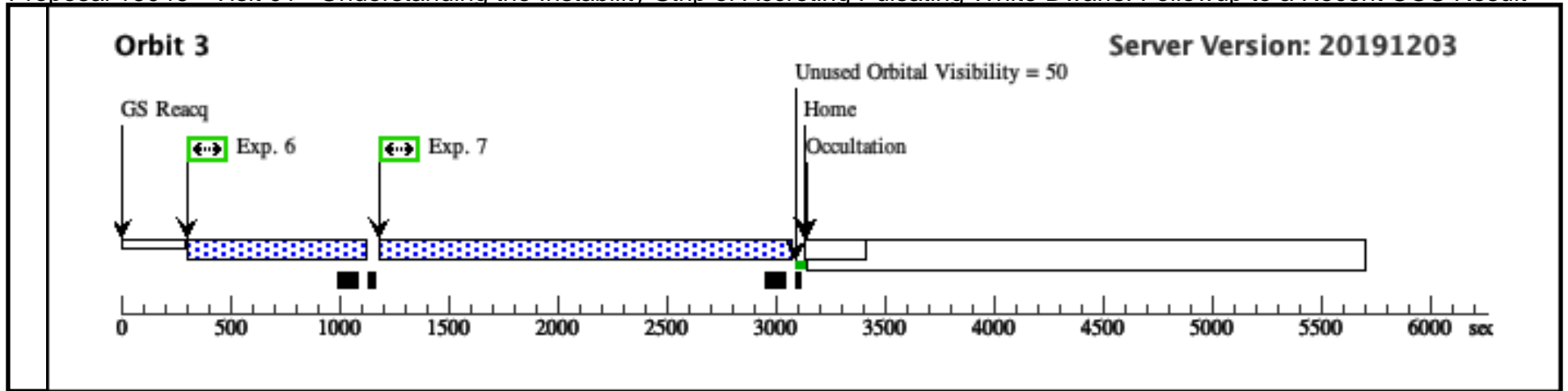
Tue Feb 11 22:00:19 GMT 2020

Visit	Proposal 16046, Visit 01, implementation				
	Diagnostic Status: No Diagnostics				
	Scientific Instruments: COS/FUV, COS/NUV				
	Special Requirements: ORIENT 115D TO 355 D; AFTER 1A BY 3 D TO 5 D				
<i>Comments: This should be scheduled in evening-local time. Flags need to be cleared during the work day.</i>					

Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
	(1)	V-V386-SER	RA: 16 10 33.6206 (242.6400858d) Dec: -01 02 23.51 (-1.03986d) Equinox: J2000	Proper Motion RA: -5.34554702006616E-4 sec of time/yr Proper Motion Dec: -0.01910200003294449 arcsec/yr Parallax: 0.00415" Epoch of Position: 2015.5	V=18.5+/-0.5	Reference Frame: ICRS
	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i>					
	Category=STAR Description=[DWARF NOVA, INTERACTING BINARY, PULSATING VARIABLE] Extended=NO					

Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
	1	V386-Ser A CQ/IMAGE (COS.ta.134 8079)	(1) V-V386-SER	COS/NUV, ACQ/IMAGE, PSA	MIRRORB			USE OFFSET V1AS AF		50 Secs (50 Secs) [==>]	[1]
	2	V386-Ser P1 (COS.sp.139 2253)	(1) V-V386-SER	COS/FUV, TIME-TAG, PSA	G140L 800 A		BUFFER-TIME=17 30; FP-POS=1; FLASH=YES	USE OFFSET V1AS AF		1841 Secs (1841 Secs) [==>]	[1]
	3	V386-Ser P2 (COS.sp.139 2253)	(1) V-V386-SER	COS/FUV, TIME-TAG, PSA	G140L 800 A		BUFFER-TIME=20 0; FP-POS=2; FLASH=YES	USE OFFSET V1AS AF		310 Secs (310 Secs) [==>]	[1]
	4	V386-Ser P2 (COS.sp.139 2253)	(1) V-V386-SER	COS/FUV, TIME-TAG, PSA	G140L 800 A		BUFFER-TIME=14 30; FP-POS=2; FLASH=YES	USE OFFSET V1AS AF		1541 Secs (1541 Secs) [==>]	[2]
	5	V386-Ser P3 (COS.sp.139 2253)	(1) V-V386-SER	COS/FUV, TIME-TAG, PSA	G140L 800 A		BUFFER-TIME=97 0; FP-POS=3; FLASH=YES	USE OFFSET V1AS AF		1080 Secs (1080 Secs) [==>]	[2]
	6	V386-Ser P3 (COS.sp.139 2253)	(1) V-V386-SER	COS/FUV, TIME-TAG, PSA	G140L 800 A		BUFFER-TIME=66 0; FP-POS=3; FLASH=YES	USE OFFSET V1AS AF		771 Secs (771 Secs) [==>]	[3]
	7	V386-Ser-P 4 (COS.sp.139 2253)	(1) V-V386-SER	COS/FUV, TIME-TAG, PSA	G140L 800 A		BUFFER-TIME=17 30; FP-POS=4; FLASH=YES	USE OFFSET V1AS AF		1840 Secs (1840 Secs) [==>]	[3]





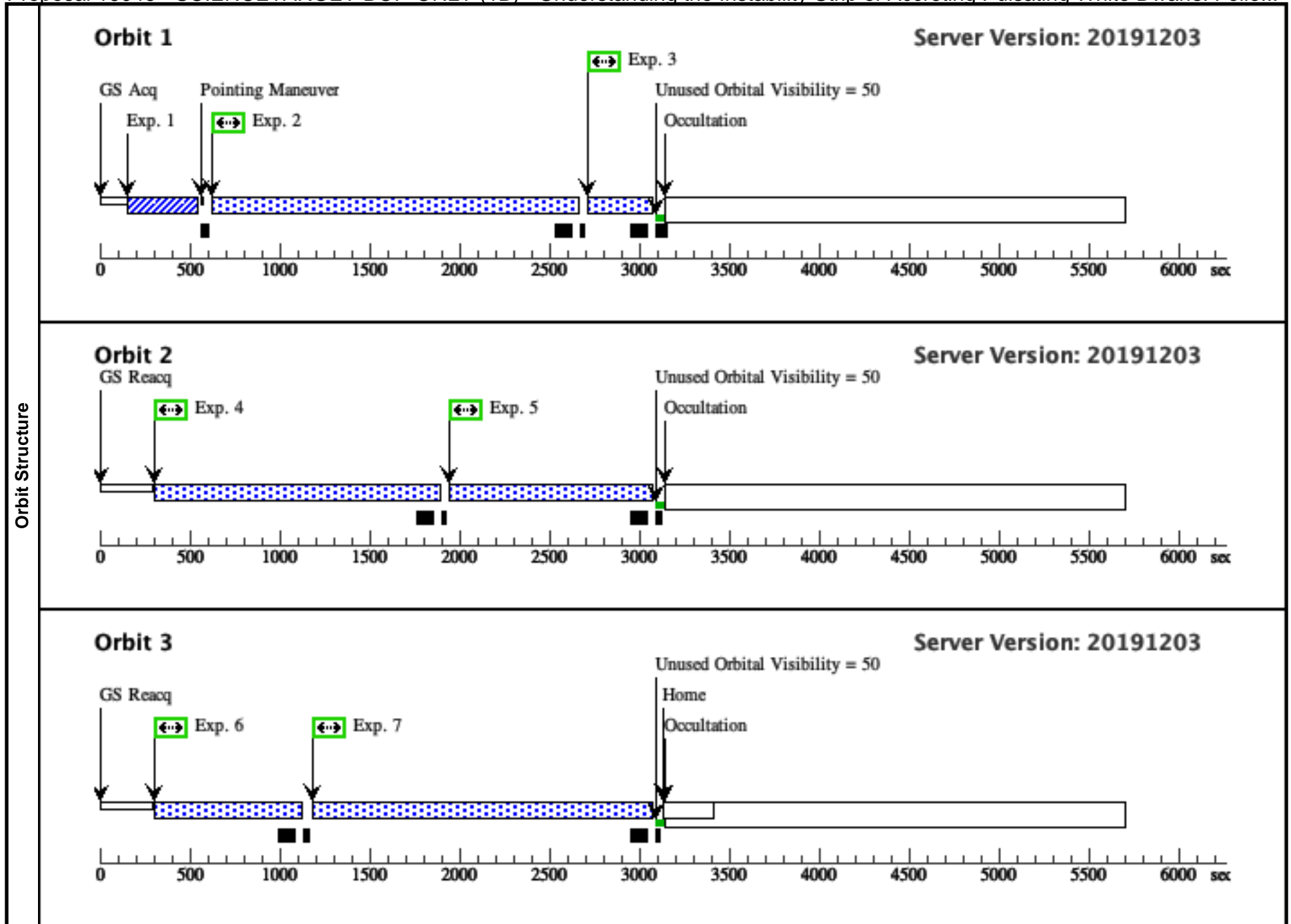
Proposal 16046 - SCIENCETARGET-BOP-ONLY (1B) - Understanding the Instability Strip of Accreting Pulsating White Dwarfs: Follo...

Tue Feb 11 22:00:19 GMT 2020

Visit	<p>Proposal 16046, SCIENCETARGET-BOP-ONLY (1B), implementation</p> <p>Diagnostic Status: No Diagnostics</p> <p>Scientific Instruments: COS/FUV, COS/NUV</p> <p>Special Requirements: ORIENT 115D TO 355 D</p> <p><i>Comments: This visit is for BOP checking the safetarget only and should not execute onboard HST.</i></p>					
Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous
(1)	V-V386-SER	RA: 16 10 33.6206 (242.6400858d) Dec: -01 02 23.51 (-1.03986d) Equinox: J2000	Proper Motion RA: -5.34554702006616E-4 sec of time/yr Proper Motion Dec: -0.01910200003294449 arcsec/yr Parallax: 0.00415" Epoch of Position: 2015.5	V=18.5+/-0.5	Reference Frame: ICRS	
<p><i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i></p> <p><i>Category=STAR</i></p> <p><i>Description=[DWARF NOVA, INTERACTING BINARY, PULSATING VARIABLE]</i></p> <p><i>Extended=NO</i></p>						
(2)	TARGET01-SAFE-TARGET	Offset from V-V386-SER RA Offset: -0.64 Secs Dec Offset: 1.678 Arcsec		V=18.5+/-0.5	Offset Position (TARGET01-SAFE-TARGET)	
<p><i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database. This target is a blank piece of sky which is the bright object safe pointing and is 9.794 arcseconds away at a PA 279.9 degrees from V-V386-SER.</i></p> <p><i>Category=UNIDENTIFIED</i></p> <p><i>Description=[BLANK FIELD]</i></p> <p><i>Extended=NO</i></p>						

Proposal 16046 - SCIENCETARGET-BOP-ONLY (1B) - Understanding the Instability Strip of Accreting Pulsating White Dwarfs: Follo...

Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
	1	V386-Ser A CQ/IMAGE (COS.ta.134 8079)	(2) TARGET01-SAF E-TARGET	COS/NUV, ACQ/IMAGE, PSA	MIRRORB					50 Secs (50 Secs) [==>]	[1]
	2	V386-Ser P1 (COS.sp.139 2253)	(2) TARGET01-SAF E-TARGET	COS/FUV, TIME-TAG, PSA	G140L 800 A	BUFFER-TIME=17 30; FP-POS=1; FLASH=YES				1841 Secs (1841 Secs) [==>]	[1]
	3	V386-Ser P2 (COS.sp.139 2253)	(2) TARGET01-SAF E-TARGET	COS/FUV, TIME-TAG, PSA	G140L 800 A	BUFFER-TIME=20 0; FP-POS=2; FLASH=YES				310 Secs (310 Secs) [==>]	[1]
	4	V386-Ser P2 (COS.sp.139 2253)	(2) TARGET01-SAF E-TARGET	COS/FUV, TIME-TAG, PSA	G140L 800 A	BUFFER-TIME=14 30; FP-POS=2; FLASH=YES				1541 Secs (1541 Secs) [==>]	[2]
	5	V386-Ser P3 (COS.sp.139 2253)	(2) TARGET01-SAF E-TARGET	COS/FUV, TIME-TAG, PSA	G140L 800 A	BUFFER-TIME=97 0; FP-POS=3; FLASH=YES				1080 Secs (1080 Secs) [==>]	[2]
	6	V386-Ser P3 (COS.sp.139 2253)	(2) TARGET01-SAF E-TARGET	COS/FUV, TIME-TAG, PSA	G140L 800 A	BUFFER-TIME=66 0; FP-POS=3; FLASH=YES				771 Secs (771 Secs) [==>]	[3]
	7	V386-Ser-P 4 (COS.sp.139 2253)	(2) TARGET01-SAF E-TARGET	COS/FUV, TIME-TAG, PSA	G140L 800 A	BUFFER-TIME=17 30; FP-POS=4; FLASH=YES				1840 Secs (1840 Secs) [==>]	[3]



Proposal 16046 - Visit 1A - Understanding the Instability Strip of Accreting Pulsating White Dwarfs: Followup to a Recent COS Result

Tue Feb 11 22:00:20 GMT 2020

Visit	Proposal 16046, Visit 1A, implementation Diagnostic Status: No Diagnostics Scientific Instruments: S/C Special Requirements: ORIENT 115D TO 355 D <i>Comments: This visit allocates and set up the safe position offset slot for visit 01 which will use that slot. This S/C visit should go earlier in the week while Visit 01 will be atleast 3 days later. The S/C visit will contain only 1 exposure. Note: weekends are to be avoided since the CS must clear the target within 24 hours of HST execution.</i>																									
	Fixed Targets	<table border="1"> <thead> <tr> <th>#</th> <th>Name</th> <th>Target Coordinates</th> <th>Targ. Coord. Corrections</th> <th>Fluxes</th> <th>Miscellaneous</th> </tr> </thead> <tbody> <tr> <td>(1)</td> <td>V-V386-SER</td> <td> RA: 16 10 33.6206 (242.6400858d) Dec: -01 02 23.51 (-1.03986d) Equinox: J2000 </td> <td> Proper Motion RA: -5.34554702006616E-4 sec of time/yr Proper Motion Dec: -0.01910200003294449 arcsec/yr Parallax: 0.00415" Epoch of Position: 2015.5 </td> <td>V=18.5+/-0.5</td> <td>Reference Frame: ICRS</td> </tr> </tbody> </table>	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous	(1)	V-V386-SER	RA: 16 10 33.6206 (242.6400858d) Dec: -01 02 23.51 (-1.03986d) Equinox: J2000	Proper Motion RA: -5.34554702006616E-4 sec of time/yr Proper Motion Dec: -0.01910200003294449 arcsec/yr Parallax: 0.00415" Epoch of Position: 2015.5	V=18.5+/-0.5	Reference Frame: ICRS	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i> Category=STAR Description=[DWARF NOVA, INTERACTING BINARY, PULSATING VARIABLE] Extended=NO											
#		Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous																				
(1)	V-V386-SER	RA: 16 10 33.6206 (242.6400858d) Dec: -01 02 23.51 (-1.03986d) Equinox: J2000	Proper Motion RA: -5.34554702006616E-4 sec of time/yr Proper Motion Dec: -0.01910200003294449 arcsec/yr Parallax: 0.00415" Epoch of Position: 2015.5	V=18.5+/-0.5	Reference Frame: ICRS																					
Exposures	<table border="1"> <thead> <tr> <th>#</th> <th>Label</th> <th>Target</th> <th>Config,Mode,Aperture</th> <th>Spectral Els.</th> <th>Opt. Params.</th> <th>Special Reqs.</th> <th>Groups</th> <th>Exp. Time (Total)/[Actual Dur.]</th> <th>Orbit</th> </tr> </thead> <tbody> <tr> <td>1</td> <td></td> <td>(1) V-V386-SER</td> <td>S/C, DATA, V1</td> <td></td> <td></td> <td> POS TARG 232.723, -237.515; SAVE OFFSET V1 ASAF; SPEC COM INSTR ECSLOTSET; QESIPARM ANGL E 279.9; QESIPARM DIST 9. 794 </td> <td></td> <td>5 Secs (5 Secs) [==>]</td> <td>[1]</td> </tr> </tbody> </table>	#	Label	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	1		(1) V-V386-SER	S/C, DATA, V1			POS TARG 232.723, -237.515; SAVE OFFSET V1 ASAF; SPEC COM INSTR ECSLOTSET; QESIPARM ANGL E 279.9; QESIPARM DIST 9. 794		5 Secs (5 Secs) [==>]	[1]					
	#	Label	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit																
1		(1) V-V386-SER	S/C, DATA, V1			POS TARG 232.723, -237.515; SAVE OFFSET V1 ASAF; SPEC COM INSTR ECSLOTSET; QESIPARM ANGL E 279.9; QESIPARM DIST 9. 794		5 Secs (5 Secs) [==>]	[1]																	
Orbit Structure																										
	<p>Orbit 1</p> <p>GS Acq Unused Orbital Visibility = 2804</p> <p>Exp. 1</p> <p>Occultation</p> <p>Server Version: 20191203</p> <p>0 500 1000 1500 2000 2500 3000 3500 4000 4500 5000 5500 6000 sec</p>																									