



## 16208 - Improved Masses for Critical Cepheid Binaries

Cycle: 28, Proposal Category: GO

(UV Initiative)

(Availability Mode: AVAILABLE)

### INVESTIGATORS

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### VISITS

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
1A	(1) V-V1334-CYG WAVE	STIS/CCD STIS/FUV-MAMA	1	24-Nov-2020 13:00:52.0	yes
1B	(1) V-V1334-CYG WAVE	STIS/CCD STIS/FUV-MAMA	1	24-Nov-2020 13:00:53.0	yes
2A	(2) V-S-MUS WAVE	STIS/CCD STIS/FUV-MAMA	1	24-Nov-2020 13:00:54.0	yes

<i>Visit</i>	<i>Targets used in Visit</i>	<i>Configurations used in Visit</i>	<i>Orbits Used</i>	<i>Last Orbit Planner Run</i>	<i>OP Current with Visit?</i>
2B	(2) V-S-MUS WAVE	STIS/CCD STIS/FUV-MAMA	1	24-Nov-2020 13:00:56.0	yes
2C	(2) V-S-MUS WAVE	STIS/CCD STIS/FUV-MAMA	1	24-Nov-2020 13:00:57.0	yes
2D	(2) V-S-MUS WAVE	STIS/CCD STIS/FUV-MAMA	1	24-Nov-2020 13:00:58.0	yes
3A	(3) V-SU-CYG WAVE	STIS/CCD STIS/FUV-MAMA	1	24-Nov-2020 13:00:59.0	yes
3B	(3) V-SU-CYG WAVE	STIS/CCD STIS/FUV-MAMA	1	24-Nov-2020 13:01:00.0	yes

8 Total Orbits Used

## **ABSTRACT**

Cepheids play an important part in the Hubble Constant tension, and in evolutionary calculation benchmarks, which include comparisons with LMC Cepheids. In some cases, they lead to exotic end-stage objects. An important HST UV legacy is the measurement of the Cepheid masses, the fundamental parameter in stellar evolution. The HST high resolution E140H spectra requested in this proposal (a unique capability of HST) will measure the orbital velocity amplitude of hot companions of three Cepheids, V1334 Cyg, S Mus, and SU Cyg. This can be combined with recent infrared interferometry with the Very Large Telescope Interferometer (VLTI) and CHARA which has resolved the systems, providing visual orbits to augment the spectroscopic orbits. From this combination both masses and distances for the Cepheids have been derived, which already challenge current evolutionary tracks, even those incorporating rotation and core convective overshoot on the main sequence. The observations detailed in this proposal will improve the orbital solution through more accurate orbital velocity amplitudes of the hot companions. In addition for SU Cyg the phasing of the short period orbit of the companion will be more precise. The requested observations are predicted to yield mass solutions accurate to a few percent. These will be the most accurate Cepheid mass determinations available to quantitatively test the underlying physics at this stage.

## **OBSERVING DESCRIPTION**

The goal of this program is to obtain STIS FUV E140H spectra of the blue companions of V1334 Cyg, S Mus, and SU Cyg to measure their orbital velocity at several phases and thereby significantly improve our knowledge of the masses and distances of both components. Specifically, for each

target, two observations separated by a few days will be made near both orbital velocity maxima and minima to derive the orbital velocity amplitude of the companion. Each observation takes one orbit resulting in spectra with S/N 20 or better (per 2 pixel resolution element). For SU Cyg, the two spectra at each orbital extremum are requested to be separated by by 1.9 to 2.3 days days which will provide a strong constraint on the short period orbit of the companion (Fig. 5), while for the other targets a 2 to 6 day spacing was requested in the phase 1. For SU Cyg, the hot companion itself is a close binary which requires timing constraints on both the long and the short period orbits; we constrain the phase of the long period orbit for the 1st observation at each epoch using a BETWEEN special requirement, and of the short period by specifying the phase required at the start of the visits. For other systems, with only one orbital period, we simply use the phase constraint.

Updates to the orbital parameters for SU Cyg have resulted in some minor shifts to the scheduling windows relative to the original phase 1 proposal, but these amount to changes of only a few days and do not reduce the overall schedulability.

As was presented in our phase 1, the required phase window for the V1334 Cyg observations is a rather narrow one near the formal start of the cycle. However, our science goals will benefit from executing the V1334 Cyg visits 1A and 1B as early as possible in our requested window, so STScI is encouraged to execute it even before the formal start of Cycl 28 on October 1 should that be feasible.

The purpose of this program is to obtain additional radial velocity measurements of three stars already observed in previous programs, and so all of these targets have previously been observed by STIS using the exact same configurations we will be using here. While the Cepheids themselves are variable, at these FUV wavelengths we are observing not the cepheids themselves, but rather their main-sequence B star companions which are not noticeably variable. Also note that Cepheids have little chromospheric activity so the Lyman alpha brightness and flares are not of any concern.

Cross-correlation of two spectra of the same star taken at different times is the most robust way to measure velocities of the STIS spectra (yielding the sharpest correlation peak and the least sensitivity to data processing such as smoothing). This is an ideal approach for binary stars, since the full orbital velocity amplitude can be measured directly by obtaining spectra at orbital velocity minimum and maximum. Details of the requested timing were justified in the Special Requirements section of the phase 1 proposal.

Because the wavelength/velocity accuracy is the most important objective of this program, in our phase 1 we had requested extra deep wavecals. We request 125s wavecals instead of the default 25s to compensate for the expected fading of the lamp since the last time the default was adjusted. To ensure that the accuracy of the faint FUV wavecal lamp is not degraded by a high dark current, it would be best if these visits could be scheduled as the first STIS FUV MAMA observation in a given SAA free period, as this avoids the dramatic increase in dark current that occurs when the FUV

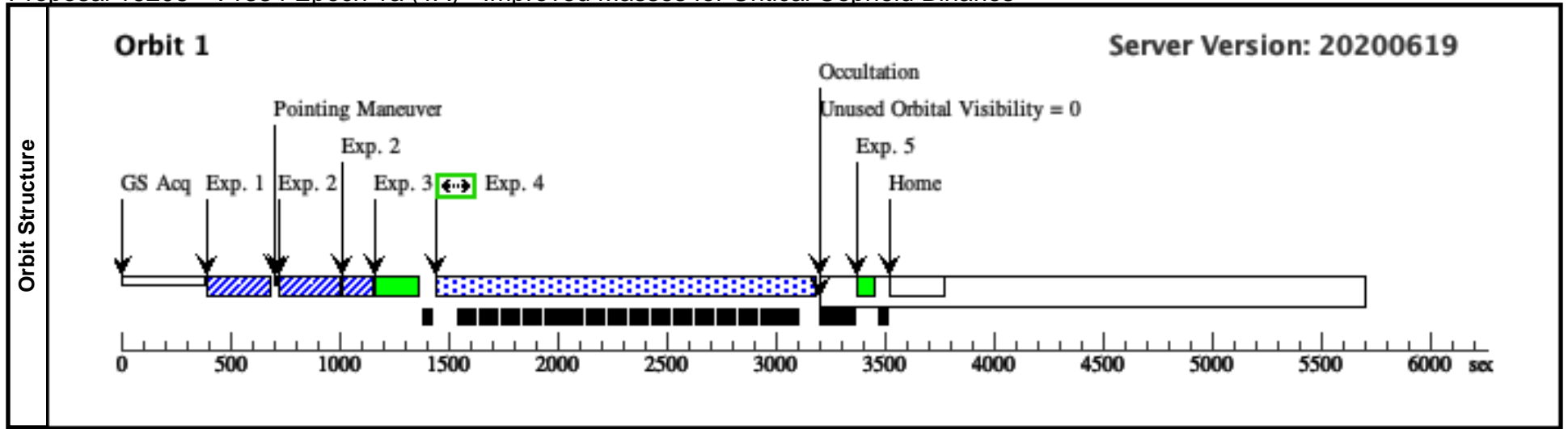
MAMA high voltage has been on for more than an orbit or so. This is especially important for SU Cyg as the ISM lines in that target cannot be used to verify the velocity zero point to the required accuracy.

We checked feasibility under one-gyro by setting the gyro mode to 1G and rerunning the orbit and visit planners. If this program had to be implemented under 1 gyro guiding, we'd need to shorten the exposure times by a few hundred seconds to account for the shortened visibilities. While this would modestly degrade our S/N, we should be able to still meet our core science goals. The scheduling would also become somewhat more difficult due to scheduling restrictions, especially for our first epoch of V1334 Cyg observations, where our visits would need to be executed in the very beginning of October to meet our phase windows, but for all targets some windows are available that meet all of our timing requirements.

Proposal 16208 - V1334 Epoch 1a (1A) - Improved Masses for Critical Cepheid Binaries

Tue Nov 24 18:01:00 GMT 2020

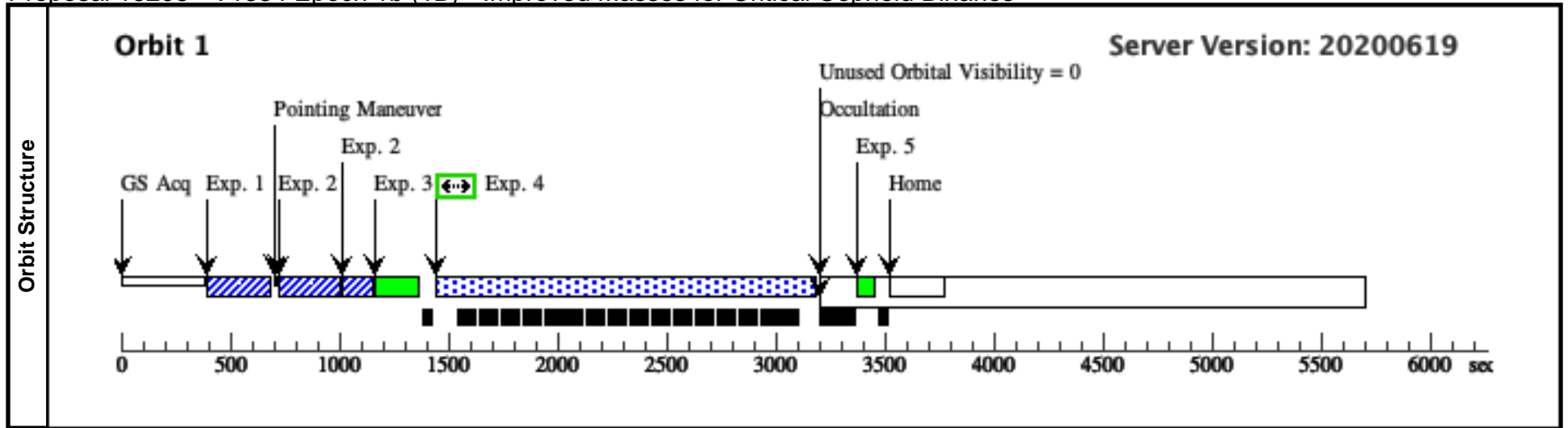
Visit	<b>Proposal 16208, V1334 Epoch 1a (1A), completed</b> <b>Diagnostic Status: No Diagnostics</b> Scientific Instruments: STIS/CCD, STIS/FUV-MAMA Special Requirements: Period 1932.8 D AND ZERO-PHASE HJD2453316.75										
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous				
		(1)	V-V1334-CYG	RA: 21 19 22.1822 (319.8424258d) Dec: +38 14 14.87 (38.23746d) Equinox: J2000	Proper Motion RA: 3.1743915473177853E-4 sec of time/yr Proper Motion Dec: 3.26E-4 arcsec/yr Parallax: 0.0011506" Epoch of Position: 2015.5	V=5.882 SpT=F1II+B7.0V	Reference Frame: ICRS				
	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i> Category=STAR Description=[B6-B9.5 V-IV, CEPHEID, COMPOSITE SPECTRAL TYPE, F0-F2, PULSATING VARIABLE] Extended=NO										
Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
	1	F25ND3 AC Q (STIS.ta.466 076)	(1) V-V1334-CYG	STIS/CCD, ACQ, F25ND3	MIRROR		PHASE 0.87 TO 0.01	Sequence 1-5 Non-Int in V1334 Epoch 1a (1A)	0.2 Secs (0.2 Secs) [==>]	[1]	
	2	G230LB AC Q/PEAK (STIS.sp.14 44996)	(1) V-V1334-CYG	STIS/CCD, ACQ/PEAK, 0.2X0.09	G230LB 2375 A			Sequence 1-5 Non-Int in V1334 Epoch 1a (1A)	1 Secs (1 Secs) [==>]	[1]	
	<i>Comments: Need to allow for variability of the Cepheid in the NUV.</i> Is gives 559,000 e- from source as calculated from IUE spectra, would need 46s to saturate.										
	3	wavecal	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1271 A			Sequence 1-5 Non-Int in V1334 Epoch 1a (1A)	34 Secs (34 Secs) [==>]	[1]	
	4	E140H 1271 V1334 Cyg (STIS.sp.14 44987)	(1) V-V1334-CYG	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1271 A	WAVECAL=NO; BUFFER-TIME=99		Sequence 1-5 Non-Int in V1334 Epoch 1a (1A)	1728 Secs (1728 Secs) [==>]	[1]	
	<i>Comments: ETC predicts global rate of 18958 cnts/s using IUE LGAP spectra</i> Observed GLOBRATE from previous 4 STIS/E140H/1271/0.2x0.09 observations ranges from ~ 16400 to 17900 cnts/s; Count rate variations are believed to be entirely due to focus driven throughput variations of 0.2x0.09 aperture. So while the minimum buffer time doesn't provide a lot of margin for time-tag mode, the history of previous observations gives us confidence we won't overfill the buffer.										
	5	wavecal	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1271 A			Sequence 1-5 Non-Int in V1334 Epoch 1a (1A)	34 Secs (34 Secs) [==>]	[1]	



Proposal 16208 - V1334 Epoch 1b (1B) - Improved Masses for Critical Cepheid Binaries

Tue Nov 24 18:01:00 GMT 2020

Visit	<b>Proposal 16208, V1334 Epoch 1b (1B), completed</b> <b>Diagnostic Status: No Diagnostics</b> Scientific Instruments: STIS/CCD, STIS/FUV-MAMA Special Requirements: AFTER 1A BY 2 D TO 6 D										
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous				
		(1)	V-V1334-CYG	RA: 21 19 22.1822 (319.8424258d) Dec: +38 14 14.87 (38.23746d) Equinox: J2000	Proper Motion RA: 3.1743915473177853E-4 sec of time/yr Proper Motion Dec: 3.26E-4 arcsec/yr Parallax: 0.0011506" Epoch of Position: 2015.5	V=5.882 SpT=F1II+B7.0V	Reference Frame: ICRS				
	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i> Category=STAR Description=[B6-B9.5 V-IV, CEPHEID, COMPOSITE SPECTRAL TYPE, F0-F2, PULSATING VARIABLE] Extended=NO										
Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
	1	F25ND3 AC Q (STIS.ta.466 076)	(1) V-V1334-CYG	STIS/CCD, ACQ, F25ND3	MIRROR			Sequence 1-5 Non-Int in V1334 Epoch 1b (1B)	0.2 Secs (0.2 Secs) [==>]	[1]	
	2	G230LB AC Q/PEAK (STIS.sp.14 44996)	(1) V-V1334-CYG	STIS/CCD, ACQ/PEAK, 0.2X0.09	G230LB 2375 A			Sequence 1-5 Non-Int in V1334 Epoch 1b (1B)	1 Secs (1 Secs) [==>]	[1]	
	<i>Comments: Need to allow for variability of the Cepheid in the NUV.</i> Is gives 559,000 e- from source as calculated from IUE spectra, would need 46s to saturate.										
	3	wavecal	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1271 A			Sequence 1-5 Non-Int in V1334 Epoch 1b (1B)	34 Secs (34 Secs) [==>]	[1]	
	4	E140H 1271 V1334 Cyg (STIS.sp.14 44987)	(1) V-V1334-CYG	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1271 A	WAVECAL=NO; BUFFER-TIME=99		Sequence 1-5 Non-Int in V1334 Epoch 1b (1B)	1728 Secs (1728 Secs) [==>]	[1]	
	<i>Comments: ETC predicts global rate of 18958 cnts/s using IUE LGAP spectra</i> Observed GLOBRATE from previous 4 STIS/E140H/1271/0.2x0.09 observations ranges from ~ 16400 to 17900 cnts/s; variations are believed to be entirely due to focus driven throughput variations of 0.2x0.09 aperture. So while the minimum buffer time doesn't provide a lot of margin for time-tag mode, the history of previous observations gives us confidence we won't overfill the buffer.										
	5	wavecal	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1271 A			Sequence 1-5 Non-Int in V1334 Epoch 1b (1B)	34 Secs (34 Secs) [==>]	[1]	

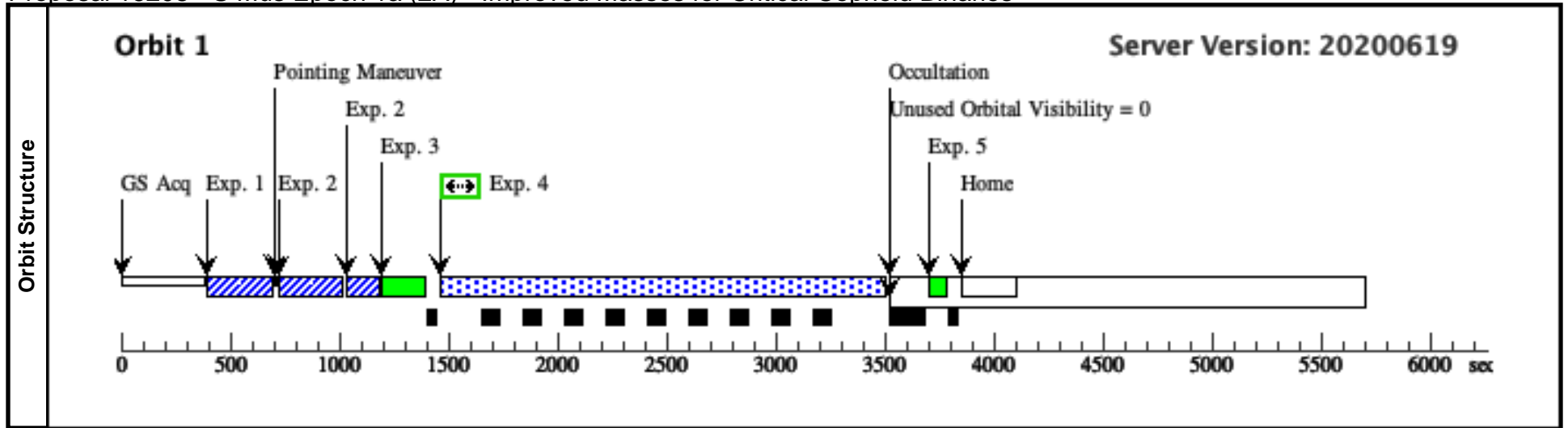




Proposal 16208 - S Mus Epoch 1a (2A) - Improved Masses for Critical Cepheid Binaries

Tue Nov 24 18:01:00 GMT 2020

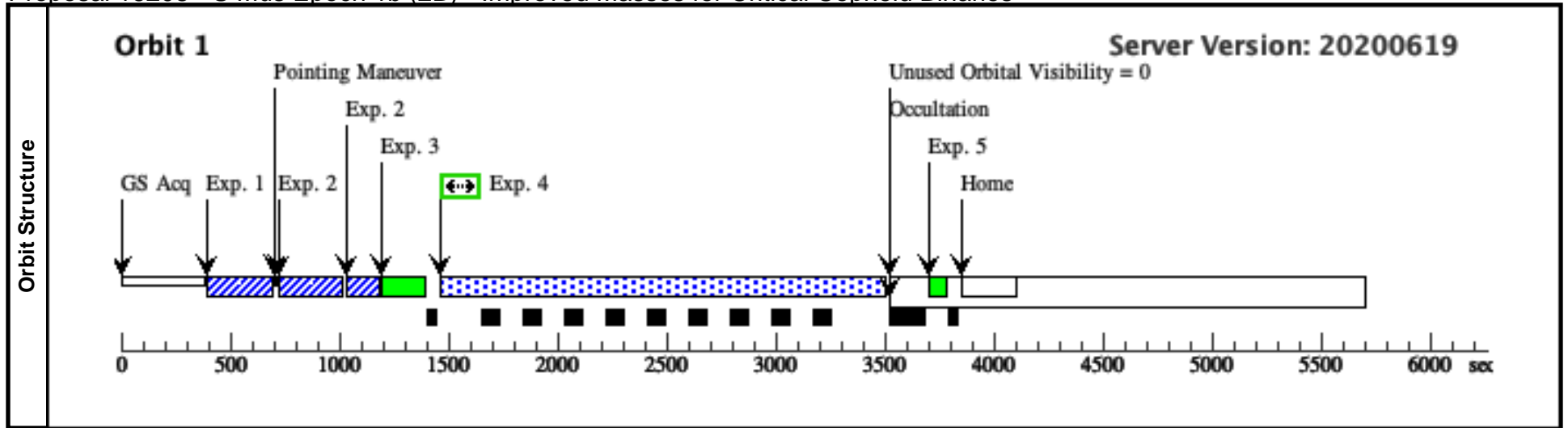
Visit	<b>Proposal 16208, S Mus Epoch 1a (2A), completed</b> <b>Diagnostic Status: No Diagnostics</b> Scientific Instruments: STIS/CCD, STIS/FUV-MAMA Special Requirements: Period 506.3 D AND ZERO-PHASE HJD2457165.9									
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous			
		(2)	V-S-MUS	RA: 12 12 46.9951 (183.1958129d) Dec: -70 09 6.46 (-70.15179d) Equinox: J2000	Proper Motion RA: -0.001573943250845688 sec of time/yr Proper Motion Dec: -0.0016740001001380733 arcsec/yr Parallax: 0.0011314" Epoch of Position: 2015.5	V=8.33 SpT=F6Ib+B5V; R=7.39	Reference Frame: ICRS			
	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i> Category=STAR Description=[B3-B5 V-IV, CEPHEID, COMPOSITE SPECTRAL TYPE, F3-F9, PULSATING VARIABLE] Extended=NO									
Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	ACQ (STIS.ta.615 207)	(2) V-S-MUS	STIS/CCD, ACQ, F25ND3	MIRROR		PHASE 0.92 TO 0.0 1	Sequence 1-5 Non-Int in S Mus Epoch 1a (2A)	1.0 Secs (1 Secs) [==>]	[1]
	2	ACQ/PEAK (STIS.sp.14 44998)	(2) V-S-MUS	STIS/CCD, ACQ/PEAK, 0.2X0.09	G230LB 2375 A			Sequence 1-5 Non-Int in S Mus Epoch 1a (2A)	2 Secs (2 Secs) [==>]	[1]
	<i>Comments: expect 360,000 e- from source in 2 sec exposure</i>									
	3	E140H 1234 WAVE	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1234 A			Sequence 1-5 Non-Int in S Mus Epoch 1a (2A)	34 Secs (34 Secs) [==>]	[1]
	4	E140H 1234 S Mus (STIS.sp.14 44985)	(2) V-S-MUS	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1234 A	BUFFER-TIME=19 0; WAVECAL=NO		Sequence 1-5 Non-Int in S Mus Epoch 1a (2A)	2029 Secs (2029 Secs) [==>]	[1]
<i>Comments: ETC predicts global rate of 8209 cnts/s using IUE LGAP spectra                      Observed GLOBRATE from previous STIS/E140H/1234/0.2x0.09 observations ranges from ~ 6991 to 8224 cnts/s;                      variations are believed to be entirely due to focus driven throughput variations of 0.2x0.09 aperture.</i>										
5	E140H 1234 WAVE	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1234 A			Sequence 1-5 Non-Int in S Mus Epoch 1a (2A)	34 Secs (34 Secs) [==>]	[1]	



Proposal 16208 - S Mus Epoch 1b (2B) - Improved Masses for Critical Cepheid Binaries

Tue Nov 24 18:01:01 GMT 2020

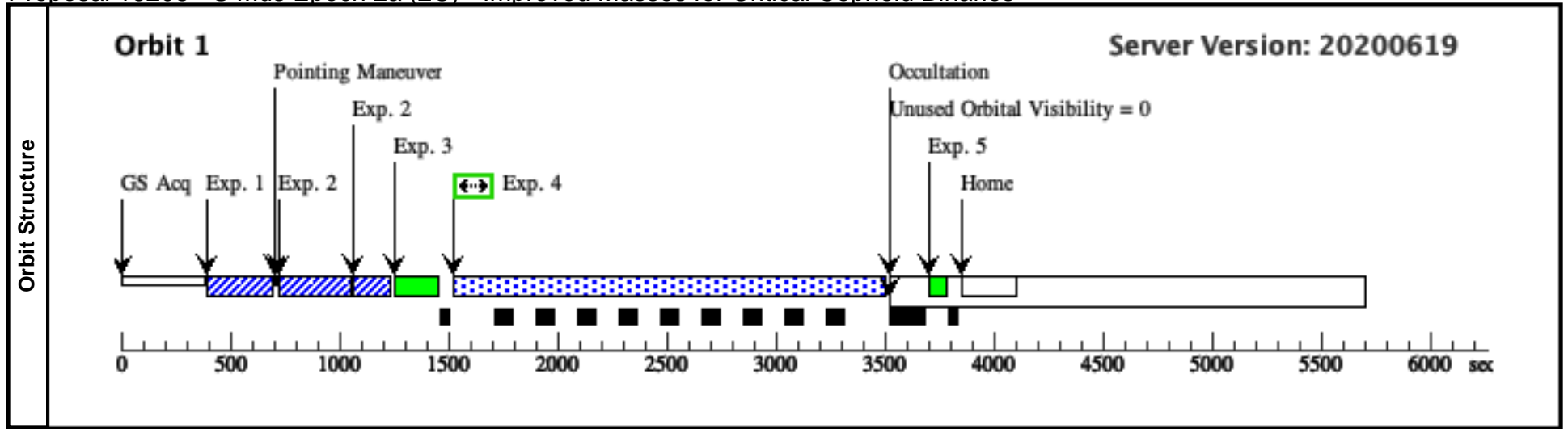
Visit	<b>Proposal 16208, S Mus Epoch 1b (2B), completed</b> <b>Diagnostic Status: No Diagnostics</b> Scientific Instruments: STIS/CCD, STIS/FUV-MAMA Special Requirements: AFTER 2A BY 2 D TO 6 D									
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous			
		(2)	V-S-MUS	RA: 12 12 46.9951 (183.1958129d) Dec: -70 09 6.46 (-70.15179d) Equinox: J2000	Proper Motion RA: -0.001573943250845688 sec of time/yr Proper Motion Dec: -0.0016740001001380733 arcsec/yr Parallax: 0.0011314" Epoch of Position: 2015.5	V=8.33 SpT=F6Ib+B5V; R=7.39	Reference Frame: ICRS			
	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i> Category=STAR Description=[B3-B5 V-IV, CEPHEID, COMPOSITE SPECTRAL TYPE, F3-F9, PULSATING VARIABLE] Extended=NO									
Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	ACQ (STIS.ta.615 207)	(2) V-S-MUS	STIS/CCD, ACQ, F25ND3	MIRROR			Sequence 1-5 Non-Int in S Mus Epoch 1b (2B)	1.0 Secs (1 Secs) [==>]	[1]
	2	ACQ/PEAK (STIS.sp.14 44998)	(2) V-S-MUS	STIS/CCD, ACQ/PEAK, 0.2X0.09	G230LB 2375 A			Sequence 1-5 Non-Int in S Mus Epoch 1b (2B)	2 Secs (2 Secs) [==>]	[1]
	<i>Comments: expect 360,000 e- from source in 2 sec exposure</i>									
	3	E140H 1234 WAVE	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1234 A			Sequence 1-5 Non-Int in S Mus Epoch 1b (2B)	34 Secs (34 Secs) [==>]	[1]
	4	E140H 1234 S Mus (STIS.sp.14 44985)	(2) V-S-MUS	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1234 A	BUFFER-TIME=19 0; WAVECAL=NO		Sequence 1-5 Non-Int in S Mus Epoch 1b (2B)	2029 Secs (2029 Secs) [==>]	[1]
<i>Comments: ETC predicts global rate of 8209 cnts/s using IUE LGAP spectra                      Observed GLOBRATE from previous STIS/E140H/1234/0.2x0.09 observations ranges from ~ 6991 to 8224 cnts/s;                      variations are believed to be entirely due to focus driven throughput variations of 0.2x0.09 aperture.</i>										
5	E140H 1234 WAVE	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1234 A			Sequence 1-5 Non-Int in S Mus Epoch 1b (2B)	34 Secs (34 Secs) [==>]	[1]	



Proposal 16208 - S Mus Epoch 2a (2C) - Improved Masses for Critical Cepheid Binaries

Tue Nov 24 18:01:01 GMT 2020

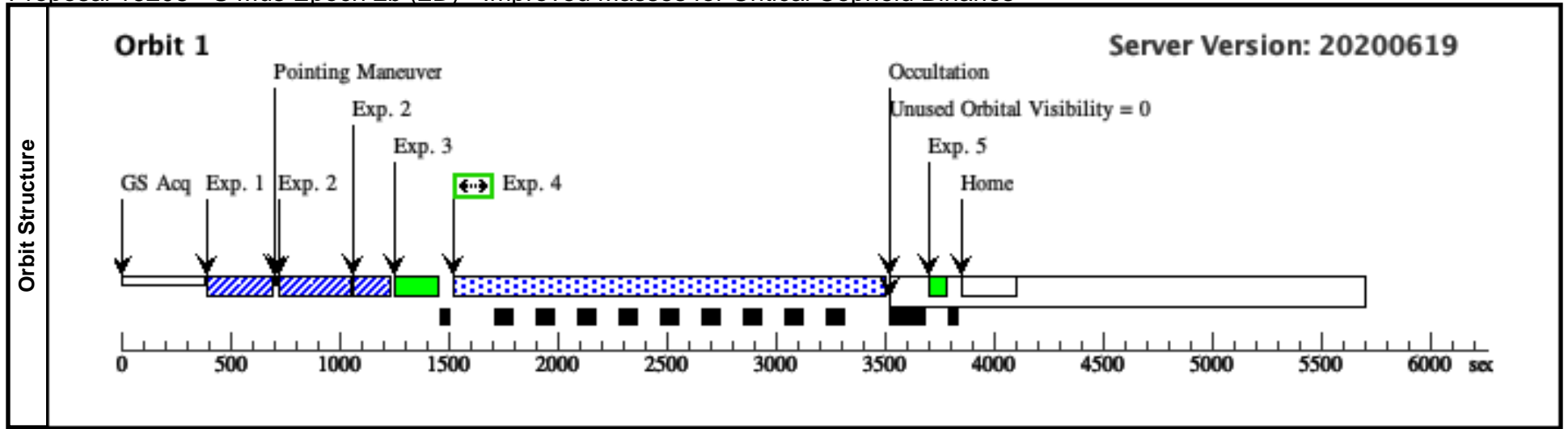
Visit	<b>Proposal 16208, S Mus Epoch 2a (2C), scheduling</b> <b>Diagnostic Status: No Diagnostics</b> Scientific Instruments: STIS/CCD, STIS/FUV-MAMA Special Requirements: Period 506.3 D AND ZERO-PHASE HJD2457165.9										
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous				
		(2)	V-S-MUS	RA: 12 12 46.9951 (183.1958129d) Dec: -70 09 6.46 (-70.15179d) Equinox: J2000	Proper Motion RA: -0.001573943250845688 sec of time/yr Proper Motion Dec: -0.0016740001001380733 arcsec/yr Parallax: 0.0011314" Epoch of Position: 2015.5	V=8.33 SpT=F6Ib+B5V; R=7.39	Reference Frame: ICRS				
	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i> Category=STAR Description=[B3-B5 V-IV, CEPHEID, COMPOSITE SPECTRAL TYPE, F3-F9, PULSATING VARIABLE] Extended=NO										
Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit	
	1	ACQ (STIS.ta.615 207)	(2) V-S-MUS	STIS/CCD, ACQ, F25ND3	MIRROR		PHASE 0.38 TO 0.5 2	Sequence 1-5 Non-Int in S Mus Epoch 2a (2C)	1.0 Secs (1 Secs) [==>]	[1]	
	2	ACQ/PEAK (STIS.sp.14 72321)	(2) V-S-MUS	STIS/CCD, ACQ/PEAK, 0.2X0.09	G230LB 2375 A			Sequence 1-5 Non-Int in S Mus Epoch 2a (2C)	5 Secs (5 Secs) [==>]	[1]	
		<i>Comments: expect 360,000 e- from source in 2 sec exposure expect 900,405e- from source in 5 sec exposure</i>									
	3	E140H 1234 WAVE	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1234 A			Sequence 1-5 Non-Int in S Mus Epoch 2a (2C)	34 Secs (34 Secs) [==>]	[1]	
	4	E140H 1234 S Mus (STIS.sp.14 44985)	(2) V-S-MUS	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1234 A	BUFFER-TIME=190; WAVECAL=NO		Sequence 1-5 Non-Int in S Mus Epoch 2a (2C)	1969 Secs (1969 Secs) [==>]	[1]	
	<i>Comments: ETC predicts global rate of 8209 cnts/s using IUE LGAP spectra Observed GLOBRATE from previous STIS/E140H/1234/0.2x0.09 observations ranges from ~ 6991 to 8224 cnts/s; variations are believed to be entirely due to focus driven throughput variations of 0.2x0.09 aperture.</i>										
5	E140H 1234 WAVE	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1234 A			Sequence 1-5 Non-Int in S Mus Epoch 2a (2C)	34 Secs (34 Secs) [==>]	[1]		



Proposal 16208 - S Mus Epoch 2b (2D) - Improved Masses for Critical Cepheid Binaries

Tue Nov 24 18:01:01 GMT 2020

Visit	<b>Proposal 16208, S Mus Epoch 2b (2D), scheduling</b> <b>Diagnostic Status: No Diagnostics</b> Scientific Instruments: STIS/CCD, STIS/FUV-MAMA Special Requirements: AFTER 2C BY 2 D TO 6 D									
	Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous			
		(2)	V-S-MUS	RA: 12 12 46.9951 (183.1958129d) Dec: -70 09 6.46 (-70.15179d) Equinox: J2000	Proper Motion RA: -0.001573943250845688 sec of time/yr Proper Motion Dec: -0.0016740001001380733 arcsec/yr Parallax: 0.0011314" Epoch of Position: 2015.5	V=8.33 SpT=F6Ib+B5V; R=7.39	Reference Frame: ICRS			
	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i> Category=STAR Description=[B3-B5 V-IV, CEPHEID, COMPOSITE SPECTRAL TYPE, F3-F9, PULSATING VARIABLE] Extended=NO									
Exposures	#	Label (ETC Run)	Target	Config,Mode,Aperture	Spectral Els.	Opt. Params.	Special Reqs.	Groups	Exp. Time (Total)/[Actual Dur.]	Orbit
	1	ACQ (STIS.ta.615 207)	(2) V-S-MUS	STIS/CCD, ACQ, F25ND3	MIRROR			Sequence 1-5 Non-Int in S Mus Epoch 2b (2D)	1.0 Secs (1 Secs) [==>]	[1]
	2	ACQ/PEAK (STIS.sp.14 72321)	(2) V-S-MUS	STIS/CCD, ACQ/PEAK, 0.2X0.09	G230LB 2375 A			Sequence 1-5 Non-Int in S Mus Epoch 2b (2D)	5 Secs (5 Secs) [==>]	[1]
	<i>Comments: expect 360,000 e- from source in 2 sec exposure expect 900,405e- from source in 5 sec exposure</i>									
	3	E140H 1234 WAVE	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1234 A			Sequence 1-5 Non-Int in S Mus Epoch 2b (2D)	34 Secs (34 Secs) [==>]	[1]
	4	E140H 1234 S Mus (STIS.sp.14 44985)	(2) V-S-MUS	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1234 A	BUFFER-TIME=190; WAVECAL=NO		Sequence 1-5 Non-Int in S Mus Epoch 2b (2D)	1969 Secs (1969 Secs) [==>]	[1]
<i>Comments: ETC predicts global rate of 8209 cnts/s using IUE LGAP spectra Observed GLOBRATE from previous STIS/E140H/1234/0.2x0.09 observations ranges from ~ 6991 to 8224 cnts/s; variations are believed to be entirely due to focus driven throughput variations of 0.2x0.09 aperture.</i>										
5	E140H 1234 WAVE	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1234 A			Sequence 1-5 Non-Int in S Mus Epoch 2b (2D)	34 Secs (34 Secs) [==>]	[1]	

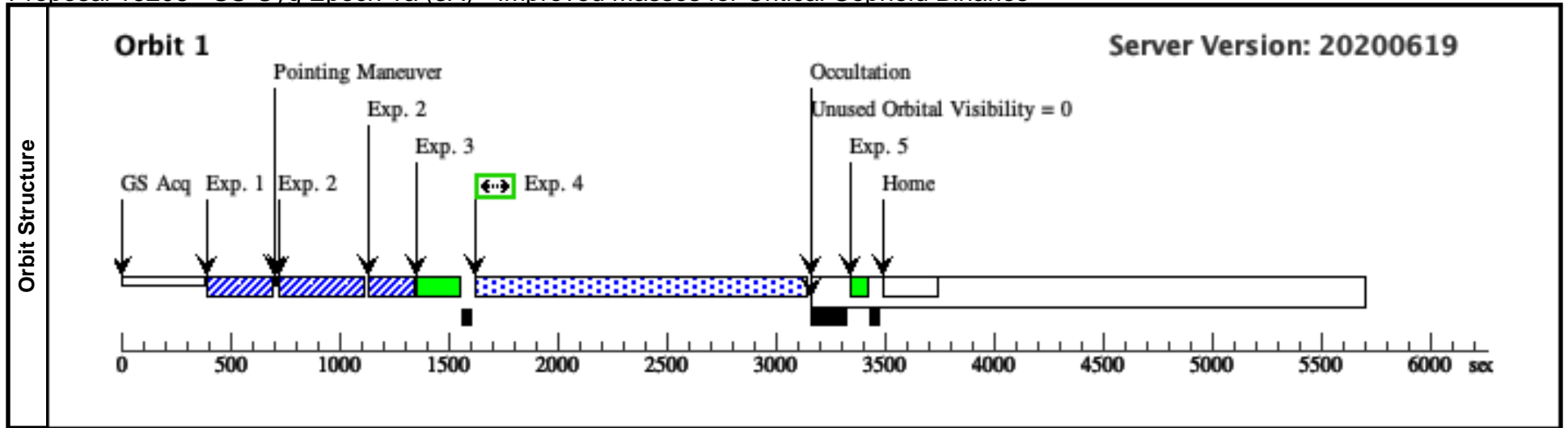




Proposal 16208 - SU Cyg Epoch 1a (3A) - Improved Masses for Critical Cepheid Binaries

Tue Nov 24 18:01:01 GMT 2020

Visit	<b>Proposal 16208, SU Cyg Epoch 1a (3A), scheduling</b> <b>Diagnostic Status: No Diagnostics</b> Scientific Instruments: STIS/CCD, STIS/FUV-MAMA Special Requirements: BETWEEN 03-FEB-2021:00:00:00 AND 23-MAY-2021:00:00:00; Period 4.675083036 D AND ZERO-PHASE HJD2444832.1224 <i>Comments: To ensure that the accuracy of the faint FUV wavecal lamp is not degraded by a high dark current, it would be best if these visits could be scheduled as the first STIS FUV MAMA observation in a given SAA free period, as this avoids the dramatic increase in dark current that occurs when the FUV MAMA high voltage has been on for more than an orbit or so. This is especially important for SU Cyg as the ISM lines in that target cannot be used to verify the velocity zero point to the required accuracy.</i>																																																																																									
	Fixed Targets	<table border="1"> <thead> <tr> <th>#</th> <th>Name</th> <th>Target Coordinates</th> <th>Targ. Coord. Corrections</th> <th>Fluxes</th> <th>Miscellaneous</th> </tr> </thead> <tbody> <tr> <td>(3)</td> <td>V-SU-CYG</td> <td>RA: 19 44 48.7343 (296.2030596d) Dec: +29 15 52.85 (29.26468d) Equinox: J2000</td> <td>Proper Motion RA: -1.4160645255034498E-4 sec of time/yr Proper Motion Dec: -0.0031549999221169855 arcsec/yr Parallax: 0.0011695" Epoch of Position: 2015.5</td> <td>V=6.44 SpT=F2Iab:+B8.0V; I=5.84</td> <td>Reference Frame: ICRS</td> </tr> </tbody> </table> <p><i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i>                      Category=STAR                      Description=[B6-B9.5 V-IV, CEPHEID, COMPOSITE SPECTRAL TYPE, F0-F2, PULSATING VARIABLE]                      Extended=NO</p>										#	Name	Target Coordinates	Targ. Coord. Corrections	Fluxes	Miscellaneous	(3)	V-SU-CYG	RA: 19 44 48.7343 (296.2030596d) Dec: +29 15 52.85 (29.26468d) Equinox: J2000	Proper Motion RA: -1.4160645255034498E-4 sec of time/yr Proper Motion Dec: -0.0031549999221169855 arcsec/yr Parallax: 0.0011695" Epoch of Position: 2015.5	V=6.44 SpT=F2Iab:+B8.0V; I=5.84	Reference Frame: ICRS																																																																			
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2	ACQ/PEAK (STIS.sp.14 72320)	(3) V-SU-CYG	STIS/CCD, ACQ/PEAK, 0.2X0.09	G230LB 2375 A			Sequence 1-5 Non-Int in SU Cyg Epoch 1a (3A)	10 Secs (10 Secs) [==>]	[1]																																																																																	
<i>Comments: 358,477 e- from source in 3 s exposure                              1,194,923 e- from source in 10s exposure</i>																																																																																										
3	E140H 1234 WAVE	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1234 A			Sequence 1-5 Non-Int in SU Cyg Epoch 1a (3A)	34 Secs (34 Secs) [==>]	[1]																																																																																	
4	E140H 1234 SU Cyg (STIS.sp.14 44984)	(3) V-SU-CYG	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1234 A	BUFFER-TIME=75 4; WAVECAL=NO		Sequence 1-5 Non-Int in SU Cyg Epoch 1a (3A)	1507 Secs (1507 Secs) [==>]	[1]																																																																																	
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5	E140H 1234 WAVE	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1234 A			Sequence 1-5 Non-Int in SU Cyg Epoch 1a (3A)	34 Secs (34 Secs) [==>]	[1]																																																																																	



# Proposal 16208 - SU Cyg Epoch 1b (3B) - Improved Masses for Critical Cepheid Binaries

Tue Nov 24 18:01:01 GMT 2020

Visit	<b>Proposal 16208, SU Cyg Epoch 1b (3B), scheduling</b> <b>Diagnostic Status: No Diagnostics</b> Scientific Instruments: STIS/CCD, STIS/FUV-MAMA Special Requirements: AFTER 3A BY 1.9 D TO 2.3 D <i>Comments: To ensure that the accuracy of the faint FUV wavecal lamp is not degraded by a high dark current, it would be best if these visits could be scheduled as the first STIS FUV MAMA observation in a given SAA free period, as this avoids the dramatic increase in dark current that occurs when the FUV MAMA high voltage has been on for more than an orbit or so. This is especially important for SU Cyg as the ISM lines in that target cannot be used to verify the velocity zero point to the required accuracy.</i>																																																																																									
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3	E140H 1234 WAVE	WAVE	STIS/FUV-MAMA, ACCUM, 0.2X0.09	E140H 1234 A			Sequence 1-5 Non-Int in SU Cyg Epoch 1b (3B)	34 Secs (34 Secs) [==>]	[1]																																																																																	
4	E140H 1234 SU Cyg (STIS.sp.14 44984)	(3) V-SU-CYG	STIS/FUV-MAMA, TIME-TAG, 0.2X0.09	E140H 1234 A	BUFFER-TIME=75 4; WAVECAL=NO		Sequence 1-5 Non-Int in SU Cyg Epoch 1b (3B)	1507 Secs (1507 Secs) [==>]	[1]																																																																																	
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