



1178 - Ram Pressure Stripping in ESO 137-001

Cycle: 1, Proposal Category: GTO

INVESTIGATORS

<i>Name</i>	<i>Institution</i>
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OBSERVATIONS

<i>Folder</i>	<i>Observation</i>	<i>Label</i>	<i>Observing Template</i>	<i>Science Target</i>
Observation Folder				
	1	Inner Tail [MRS+Imaging]	MIRI Medium Resolution Spectroscopy	(18) Group INNER-TAIL
	2	Inner Tail 2 [MRS+Imaging]	MIRI Medium Resolution Spectroscopy	(22) Group INNER-TAIL-2
	3	Inner Tail 3 [MRS+Imaging]	MIRI Medium Resolution Spectroscopy	(23) Group INNER-TAIL-3
	4	Outer Tail [MRS+Imaging]	MIRI Medium Resolution Spectroscopy	(21) Group OUTER-TAIL
	5	Background	MIRI Medium Resolution Spectroscopy	(10) BACKGROUND

ABSTRACT

Once thought to be rare, the number of known ram pressure stripping (RPS) events has been steadily rising, observed as truncated or disturbed gaseous disks or one-sided tails in the X-ray through the radio. These events hold key information regarding the relation between galaxy transformation and environment. We propose MIRI MRS observations of ESO 137-001, a well studied local galaxy ($z=0.01625$) with a spectacular double ram pressure stripped tail. At both high spectral ($R\sim 2700$) and spatial (~ 0.1 arcsec) resolution over 5-28.8m, we will detect multiple

transitions of rotational H₂ lines as well as a suite of fine structure lines at high significance. From these observables, we will deduce the kinematics and the temperature/density structure of warm and hot gas components in the tail on sub-kpc scales as well as the excitation mechanism(s) responsible. This information will reveal how the (star forming) interstellar medium of the host galaxy responds to strong RPS and how the stripped gas subsequently interacts with the intra-cluster medium. Notably, the detailed state of H₂ will identify the spatial extent of shocked gas and constrain the mechanisms and timescales for the cooling of molecular gas, revealing whether star-forming regions in the tail were formed in situ or from molecular gas stripped directly from the galactic disk. Additionally, high resolution MIRI 7.7m imaging obtained simultaneously with MRS pointings in the far-tail (~40 kpc from the main galaxy) will fall back on the main galaxy and near- to mid-tail regions, providing a measure of the aromatic features and a SFR indicator.

FRIEDMAN_0001

OBSERVING DESCRIPTION

This program targets multiple regions at the galaxy-tail interface and along the stripped tail of ESO 137-001 with the MIRI MRS. The full MRS wavelength range with all channels will be observed. The 16 science targets have been grouped into three inner tail region target groups (14 targets in target group INNER-TAIL, INNER-TAIL-2, and INNER-TAIL-3) and an outer tail region target group (2 targets plus a background). Our science goal is to detect and characterize warm molecular hydrogen emission and fine structure lines in key regions, selected HII regions with evidence for star formation and/or cold molecular gas, at the galaxy-tail interface and along the RPS tail. Our primary objective is the H₂ S(1) emission line at 17.035 μ m (rest). As this line is in the MRS channel 3 at $z=0.01625$, our pointings and dither strategy are optimized for this channel. Secondary objectives include the H₂ S(2)-S(7) lines, fine structure lines such as [NeII]-[NeVI], [SIII], [OIV], [FeII], etc. and PAH features. The H₂ S(0) emission line, which falls at the low sensitivity edge of channel 4 long, will be recovered if possible using spatial and spectral binning.

One dedicated background pointing, with the same exposure setup as our science pointings, is included in the target group OUTER-TAIL for calibration purposes in order to ensure proper subtraction of the sky/telescope background. This background pointing must be taken during the same epoch as our observations and therefore we request that the program be executed within a 2 day period in a restricted PA range.

MIRI imaging, taken simultaneously with MRS mosaic in the outer tail, will provide a F770W image of the galaxy ESO 137-001 and its RPS tail out to a distance of ~25 kpc. This imaging requires a position angle of ~328-342 deg (V3). Since this imaging can improve spatial registration of the MRS exposures, we include it in the inner tail region as well at a minimum increase to program duration.

Given that our targets are extended, the absolute pointing accuracy of JWST is adequate for our needs and target acquisition is not required for this program.

Proposal 1178 - Targets - Ram Pressure Stripping in ESO 137-001

#	Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous
(2)	KNOT-2	RA: 16 13 24.9680 (243.3540333d) Dec: -60 45 43.51 (-60.76209d) Equinox: J2000		
<p><i>Comments: Coordinates from Sivanandam+10, Table 4.</i></p> <p><i>We are observing one background (fixed target 10, observation 5) for all MRS pointings.</i></p> <p>Category=ISM Description=[H II regions, Molecular gas] Extended=YES</p>				
(3)	KNOT-3	RA: 16 13 26.4800 (243.3603333d) Dec: -60 45 32.85 (-60.75912d) Equinox: J2000		
<p><i>Comments: Coordinates from Sivanandam+10, Table 4.</i></p> <p><i>We are observing one background (fixed target 10, observation 5) for all MRS pointings.</i></p> <p>Category=ISM Description=[H II regions] Extended=YES</p>				
(4)	KNOT-4	RA: 16 13 25.4500 (243.3560417d) Dec: -60 45 35.19 (-60.75977d) Equinox: J2000		
<p><i>Comments: Coordinates from Sivanandam+10, Table 4.</i></p> <p><i>We are observing one background (fixed target 10, observation 5) for all MRS pointings.</i></p> <p>Category=ISM Description=[H II regions, Molecular gas] Extended=YES</p>				
(5)	KNOT-5	RA: 16 13 23.9300 (243.3497083d) Dec: -60 45 53.60 (-60.76489d) Equinox: J2000		
<p><i>Comments: Coordinates from Sivanandam+10, Table 4.</i></p> <p><i>We are observing one background (fixed target 10, observation 5) for all MRS pointings.</i></p> <p>Category=ISM Description=[H II regions] Extended=YES</p>				
(6)	KNOT-6	RA: 16 13 24.5400 (243.3522500d) Dec: -60 46 5.92 (-60.76831d) Equinox: J2000		
<p><i>Comments: Coordinates from Sivanandam+10, Table 4.</i></p> <p><i>We are observing one background (fixed target 10, observation 5) for all MRS pointings.</i></p> <p>Category=ISM Description=[H II regions, Molecular gas] Extended=YES</p>				
(7)	KNOT-7	RA: 16 13 24.1300 (243.3505417d) Dec: -60 45 33.46 (-60.75929d) Equinox: J2000		
<p><i>Comments: Coordinates from Sivanandam+10, Table 4.</i></p> <p><i>We are observing one background (fixed target 10, observation 5) for all MRS pointings.</i></p> <p>Category=ISM Description=[H II regions, Molecular gas] Extended=YES</p>				

Fixed Targets

Proposal 1178 - Targets - Ram Pressure Stripping in ESO 137-001

(8)	KNOT-11	RA: 16 13 22.8300 (243.3451250d) Dec: -60 45 20.55 (-60.75571d) Equinox: J2000
<p><i>Comments: Coordinates from Sivanandam+10, Table 4.</i></p> <p><i>We are observing one background (fixed target 10, observation 5) for all MRS pointings.</i></p> <p><i>Category=ISM</i> <i>Description=[H II regions, Molecular gas]</i> <i>Extended=YES</i></p>		
(10)	BACKGROUND	RA: 16 13 12.2116 (243.3008817d) Dec: -60 45 5.69 (-60.75158d) Equinox: J2000
<p><i>Comments: Chosen to have no 8um emission.</i></p> <p><i>Category=Calibration</i> <i>Description=[Telescope/sky background]</i> <i>Extended=YES</i></p>		
(11)	ESO137-001-TILE-2	RA: 16 13 26.1575 (243.3589896d) Dec: -60 45 42.12 (-60.76170d) Equinox: J2000
<p><i>Comments: Tile in mosaic of galaxy-tail interface covering the base of the ram pressure stripped tail of ESO137-001. Mosaic was designed using the APT mosaic tool with 15% overlap, but then the tiles were split into individual targets in order to include them in Target Group INNER-TAIL. This allows us to observe the mosaic and Knots 2, 3, 4, 5, 6, 7, 11 in the same visit as they are within the visit splitting distance.</i></p> <p><i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i></p> <p><i>Category=Galaxy</i> <i>Description=[Spiral galaxies, Tidal tails]</i> <i>Extended=YES</i></p>		
(13)	ESO137-001-TILE-5	RA: 16 13 26.4339 (243.3601412d) Dec: -60 45 47.68 (-60.76324d) Equinox: J2000
<p><i>Comments: Tile in mosaic of galaxy-tail interface covering the base of the ram pressure stripped tail of ESO137-001. Mosaic was designed using the APT mosaic tool with 15% overlap, but then the tiles were split into individual targets in order to include them in Target Group INNER-TAIL. This allows us to observe the mosaic and Knots 2, 3, 4, 5, 6, 7, 11 in the same visit as they are within the visit splitting distance.</i></p> <p><i>We are observing one background (fixed target 10, observation 5) for all MRS pointings.</i></p> <p><i>Category=Galaxy</i> <i>Description=[Spiral galaxies, Tidal tails]</i> <i>Extended=YES</i></p>		
(14)	ESO137-001-TILE-6	RA: 16 13 27.1105 (243.3629604d) Dec: -60 45 45.88 (-60.76274d) Equinox: J2000
<p><i>Comments: Tile in mosaic of galaxy-tail interface covering the base of the ram pressure stripped tail of ESO137-001. Mosaic was designed using the APT mosaic tool with 15% overlap, but then the tiles were split into individual targets in order to include them in Target Group INNER-TAIL. This allows us to observe the mosaic and Knots 2, 3, 4, 5, 6, 7, 11 in the same visit as they are within the visit splitting distance.</i></p> <p><i>We are observing one background (fixed target 10, observation 5) for all MRS pointings.</i></p> <p><i>Category=Galaxy</i> <i>Description=[Spiral galaxies, Tidal tails]</i> <i>Extended=YES</i></p>		

Proposal 1178 - Targets - Ram Pressure Stripping in ESO 137-001

(16)	ESO137-001-TILE-8	RA: 16 13 26.7103 (243.3612929d) Dec: -60 45 53.24 (-60.76479d) Equinox: J2000
<p><i>Comments: Tile in mosaic of galaxy-tail interface covering the base of the ram pressure stripped tail of ESO137-001. Mosaic was designed using the APT mosaic tool with 15% overlap, but then the tiles were split into individual targets in order to include them in Target Group INNER-TAIL. This allows us to observe the mosaic and Knots 2, 3, 4, 5, 6, 7, 11 in the same visit as they are within the visit splitting distance.</i></p> <p><i>We are observing one background (fixed target 10, observation 5) for all MRS pointings.</i></p> <p>Category=Galaxy Description=[Spiral galaxies, Tidal tails] Extended=YES</p>		
(17)	ESO137-001-TILE-9	RA: 16 13 27.3869 (243.3641121d) Dec: -60 45 51.44 (-60.76429d) Equinox: J2000
<p><i>Comments: Tile in mosaic of galaxy-tail interface covering the base of the ram pressure stripped tail of ESO137-001. Mosaic was designed using the APT mosaic tool with 15% overlap, but then the tiles were split into individual targets in order to include them in Target Group INNER-TAIL. This allows us to observe the mosaic and Knots 2, 3, 4, 5, 6, 7, 11 in the same visit as they are within the visit splitting distance.</i></p> <p><i>We are observing one background (fixed target 10, observation 5) for all MRS pointings.</i></p> <p>Category=Galaxy Description=[Spiral galaxies, Tidal tails] Extended=YES</p>		
(18)	Group INNER-TAIL	
<p><i>Comments:</i> <i>Target Selection=[5 KNOT-5, 6 KNOT-6, 16 ESO137-001-TILE-8, 17 ESO137-001-TILE-9]</i></p>		
(19)	OUTER-TAIL-1	RA: 16 13 14.6400 (243.3110000d) Dec: -60 44 45.30 (-60.74592d) Equinox: J2000
<p><i>Comments: We are observing one background (fixed target 10, observation 5) for all MRS pointings.</i></p> <p>Category=ISM Description=[H II regions, Molecular gas]</p>		
(20)	OUTER-TAIL-2	RA: 16 13 14.3800 (243.3099167d) Dec: -60 44 39.40 (-60.74428d) Equinox: J2000
<p><i>Comments: We are observing one background (fixed target 10, observation 5) for all MRS pointings.</i></p> <p>Category=ISM Description=[H II regions, Molecular gas]</p>		
(21)	Group OUTER-TAIL	
<p><i>Comments:</i> <i>Target Selection=[19 OUTER-TAIL-1, 20 OUTER-TAIL-2]</i></p>		
(22)	Group INNER-TAIL-2	
<p><i>Comments:</i> <i>Target Selection=[2 KNOT-2, 11 ESO137-001-TILE-2, 13 ESO137-001-TILE-5, 14 ESO137-001-TILE-6]</i></p>		
(23)	Group INNER-TAIL-3	
<p><i>Comments:</i> <i>Target Selection=[3 KNOT-3, 4 KNOT-4, 7 KNOT-7, 8 KNOT-11]</i></p>		

Proposal 1178 - Observation 1 - Ram Pressure Stripping in ESO 137-001

Mon Apr 03 19:00:27 GMT 2023

Observation	Proposal 1178, Observation 1: Inner Tail [MRS+Imaging] Diagnostic Status: Warning Observing Template: MIRI Medium Resolution Spectroscopy <i>Comments: We are observing one background (fixed target 10, observation 4) for all MRS pointings. We request that all observations be taken within two days to minimize variation in the background.</i>																																																																																																						
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Proposal 1178 - Observation 1 - Ram Pressure Stripping in ESO 137-001

Special Requirements

Group Observations 1, 2, 3, 4, 5 within 2 Days

Proposal 1178 - Observation 2 - Ram Pressure Stripping in ESO 137-001

Mon Apr 03 19:00:27 GMT 2023

Observation	Proposal 1178, Observation 2: Inner Tail 2 [MRS+Imaging] Diagnostic Status: Warning Observing Template: MIRI Medium Resolution Spectroscopy <i>Comments: We are observing one background (fixed target 10, observation 4) for all MRS pointings. We request that all observations be taken within two days to minimize variation in the background.</i>																																																																																																						
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Proposal 1178 - Observation 2 - Ram Pressure Stripping in ESO 137-001

Special Requirements

Group Observations 1, 2, 3, 4, 5 within 2 Days

Proposal 1178 - Observation 3 - Ram Pressure Stripping in ESO 137-001

Mon Apr 03 19:00:27 GMT 2023

Observation	<p>Proposal 1178, Observation 3: Inner Tail 3 [MRS+Imaging]</p> <p>Diagnostic Status: Warning</p> <p>Observing Template: MIRI Medium Resolution Spectroscopy</p> <p><i>Comments: We are observing one background (fixed target 10, observation 4) for all MRS pointings. We request that all observations be taken within two days to minimize variation in the background.</i></p>												
Diagnostics	(Visit 3:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.												
Fixed Targets	#	Name	Target Coordinates				Targ. Coord. Corrections			Miscellaneous			
Acquisition	#											Target	
Template	AcqFilter		Primary Channel				Simultaneous Imaging			Imager Subarray			
Dithers	#	Dither Type				Optimized For			Direction				
Spectral Elements	#	Wavelength Range	Detector	Filter	Readout Pattern	Groups/Int	Integrations/Exp	Exposures/Dith	Dither	Total Dithers	Total Integrations	Total Exposure Time	ETC Wkbk.Calc ID
	1	LONG(C)	MRSLONG		FASTR1	44	1	1	Dither 1	4	4	488.407	25204
	1	LONG(C)	MRSSHORT		FASTR1	44	1	1	Dither 1	4	4	488.407	25204
	2	MEDIUM(B)	MRSLONG		FASTR1	44	1	1	Dither 1	4	4	488.407	25204
	2	MEDIUM(B)	MRSSHORT		FASTR1	44	1	1	Dither 1	4	4	488.407	25204
	3	SHORT(A)	MRSLONG		FASTR1	44	1	1	Dither 1	4	4	488.407	25204
	3	SHORT(A)	MRSSHORT		FASTR1	44	1	1	Dither 1	4	4	488.407	25204

Proposal 1178 - Observation 3 - Ram Pressure Stripping in ESO 137-001

Special Requirements

Group Observations 1, 2, 3, 4, 5 within 2 Days

Proposal 1178 - Observation 4 - Ram Pressure Stripping in ESO 137-001

Mon Apr 03 19:00:27 GMT 2023

Observation	Proposal 1178, Observation 4: Outer Tail [MRS+Imaging] Diagnostic Status: Warning Observing Template: MIRI Medium Resolution Spectroscopy <i>Comments: PA is set to 328-342 in order for the simultaneous imaging from this target group to fall back on the main science target.</i> <i>We are observing one background (fixed target 10, observation 4) for all MRS pointings. We request that all observations be taken within two days to minimize variation in the background.</i>																																																																																																																																													
Diagnostics	(Visit 4:1) Warning (Form): Overheads are provisional until the Visit Planner has been run. (Visit 4:1) Informational (Form): Visit schedulable, but most scheduling windows are when JWST is pointed in direction of greatest micrometeoroid impact risk. This is likely due to scheduling special requirements.																																																																																																																																													
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Proposal 1178 - Observation 4 - Ram Pressure Stripping in ESO 137-001

Special Requirements

Aperture PA Range 327 to 342 Degrees (V3 327.0 to 342.0)

Group Observations 1, 2, 3, 4, 5 within 2 Days

Proposal 1178 - Observation 5 - Ram Pressure Stripping in ESO 137-001

Mon Apr 03 19:00:27 GMT 2023

Observation	Proposal 1178, Observation 5: Background Diagnostic Status: Warning Observing Template: MIRI Medium Resolution Spectroscopy <i>Comments: PA is set to 328-342 in order for the simultaneous imaging from this target group to fall back on the main science target.</i> <i>We are observing one background (fixed target 10, observation 4) for all MRS pointings. We request that all observations be taken within two days to minimize variation in the background.</i>																																																																																																																																													
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<i>Comments: Chosen to have no 8um emission.</i> <i>Category=Calibration</i> <i>Description=[Telescope/sky background]</i> <i>Extended=YES</i>																																																																																																																																														
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Spectral Elements	<table border="1"> <thead> <tr> <th>#</th> <th>Wavelength Range</th> <th>Detector</th> <th>Filter</th> <th>Readout Pattern</th> <th>Groups/Int</th> <th>Integrations/Exp</th> <th>Exposures/Dith</th> <th>Dither</th> <th>Total Dithers</th> <th>Total Integrations</th> <th>Total Exposure Time</th> <th>ETC Wkbk.Calc ID</th> </tr> </thead> <tbody> <tr> <td>1</td> <td></td> <td>IMAGER</td> <td>F770W</td> <td>FASTR1</td> <td>44</td> <td>1</td> <td>1</td> <td>Dither 1</td> <td>2</td> <td>2</td> <td>244.204</td> <td>25204</td> </tr> <tr> <td>1</td> <td>LONG(C)</td> <td>MRSLONG</td> <td></td> <td>FASTR1</td> <td>44</td> <td>1</td> <td>1</td> <td>Dither 1</td> <td>2</td> <td>2</td> <td>244.204</td> <td>25204</td> </tr> <tr> <td>1</td> <td>LONG(C)</td> <td>MRSSHORT</td> <td></td> <td>FASTR1</td> <td>44</td> <td>1</td> <td>1</td> <td>Dither 1</td> <td>2</td> <td>2</td> <td>244.204</td> <td>25204</td> </tr> <tr> <td>2</td> <td></td> <td>IMAGER</td> <td>F770W</td> <td>FASTR1</td> <td>44</td> <td>1</td> <td>1</td> <td>Dither 1</td> <td>2</td> <td>2</td> <td>244.204</td> <td>25204</td> </tr> <tr> <td>2</td> <td>MEDIUM(B)</td> <td>MRSLONG</td> <td></td> <td>FASTR1</td> <td>44</td> <td>1</td> <td>1</td> <td>Dither 1</td> <td>2</td> <td>2</td> <td>244.204</td> <td>25204</td> </tr> <tr> <td>2</td> <td>MEDIUM(B)</td> <td>MRSSHORT</td> <td></td> <td>FASTR1</td> <td>44</td> <td>1</td> <td>1</td> <td>Dither 1</td> <td>2</td> <td>2</td> <td>244.204</td> <td>25204</td> </tr> <tr> <td>3</td> <td></td> <td>IMAGER</td> <td>F770W</td> <td>FASTR1</td> <td>44</td> <td>1</td> <td>1</td> <td>Dither 1</td> <td>2</td> <td>2</td> <td>244.204</td> <td>25204</td> </tr> <tr> <td>3</td> <td>SHORT(A)</td> <td>MRSLONG</td> <td></td> <td>FASTR1</td> <td>44</td> <td>1</td> <td>1</td> <td>Dither 1</td> <td>2</td> <td>2</td> <td>244.204</td> <td>25204</td> </tr> <tr> <td>3</td> <td>SHORT(A)</td> <td>MRSSHORT</td> <td></td> <td>FASTR1</td> <td>44</td> <td>1</td> <td>1</td> <td>Dither 1</td> <td>2</td> <td>2</td> <td>244.204</td> <td>25204</td> </tr> </tbody> </table>												#	Wavelength Range	Detector	Filter	Readout Pattern	Groups/Int	Integrations/Exp	Exposures/Dith	Dither	Total Dithers	Total Integrations	Total Exposure Time	ETC Wkbk.Calc ID	1		IMAGER	F770W	FASTR1	44	1	1	Dither 1	2	2	244.204	25204	1	LONG(C)	MRSLONG		FASTR1	44	1	1	Dither 1	2	2	244.204	25204	1	LONG(C)	MRSSHORT		FASTR1	44	1	1	Dither 1	2	2	244.204	25204	2		IMAGER	F770W	FASTR1	44	1	1	Dither 1	2	2	244.204	25204	2	MEDIUM(B)	MRSLONG		FASTR1	44	1	1	Dither 1	2	2	244.204	25204	2	MEDIUM(B)	MRSSHORT		FASTR1	44	1	1	Dither 1	2	2	244.204	25204	3		IMAGER	F770W	FASTR1	44	1	1	Dither 1	2	2	244.204	25204	3	SHORT(A)	MRSLONG		FASTR1	44	1	1	Dither 1	2	2	244.204	25204	3	SHORT(A)	MRSSHORT		FASTR1	44	1	1	Dither 1	2	2	244.204	25204
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Proposal 1178 - Observation 5 - Ram Pressure Stripping in ESO 137-001

Special Requirements

Aperture PA Range 327 to 342 Degrees (V3 327.0 to 342.0)

Group Observations 1, 2, 3, 4, 5 within 2 Days