



1617 - The Boundary of Galaxy Formation: Constraints from the Ancient Star Formation of the Isolated, Extremely Low-Mass Galaxy Leo P

Cycle: 1, Proposal Category: GO

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OBSERVATIONS

| <i>Folder</i> | <i>Observation</i> | <i>Label</i> | <i>Observing Template</i> | <i>Science Target</i> |
|---------------|--------------------|--------------|---------------------------|-----------------------|
| Leo P | | | | |
| | 1 | LeoP | NIRCam Imaging | (1) LEO-P |

ABSTRACT

In the early universe, at the lowest masses where gravity is in a tug-of-war with the energy injected by internal events (stellar feedback) and outside events (reionization, environmental effects), the dominant factors that govern the growth of the smallest structures are still speculative. Archeological studies of low-mass galaxies in the Local Group (LG) have striven to answer this question through detailed analyses of the oldest stars. Yet, the LG is, by definition, a complicated system and the history of its satellite galaxies are intertwined with their evolution in the shadow of their massive host. The only laboratory to truly measure how galaxies at the limit of structure formation grow is a very low-mass system that is isolated.

Leo P is an isolated, extremely low-mass ($M^* \sim 10^5 M_{\text{sun}}$), metal poor ($< 3\%$ Solar), gas-rich galaxy just outside the LG at 1.6 Mpc. It is the quintessential system to test theories of how the smallest structures in our universe survive and grow. We will image the resolved stars in Leo P to below the old main sequence turn-off, a depth unreachable with HST but achievable with JWST, to derive the ancient star formation history of the galaxy. The star formation history will test if (i) early star formation in Leo P was "quelled", as predicted nearly uniformly by reionization models, and then re-ignited, (ii) Leo P has a delayed onset in star formation suggesting the accretion history of baryons may be dependent on environment at low masses, or (iii) Leo P has constant star formation across all epochs which is not predicted by cosmological models in this mass range and which would also be in tension with reionization models.

OBSERVING DESCRIPTION

We will image Leo P with NIRCcam with the short-wavelength filters F090W and F150W, with simultaneous imaging in the long-wavelength filter F277W. The target galaxy has an angular size that fits within one NIRCcam module; we elect to use only module B for the observations to reduce data transfer rates and total data volume. Our science goals require the observations to reach below the old main sequence turn-off with a $\text{SNR}=5$ for a metal-poor, 13 Gyr old star ($m_{\text{F090W}}=29.55$ mag based on the MIST stellar evolution models). We will use a DEEP8 readout with 5 groups, 6 integrations for the F150W filter and 10 integrations for the F090W filter. We use a 7-point subpixel dither pattern to mitigate cosmic rays and bad pixels, improve PSF sampling, and reduce flat-field errors thereby maximizing the SNR. The JWST ETC v1.5.2 suggests a total science time of 29.4 hrs to achieve our photometric depth requirement. Total requested time with overhead from the APT is 35.7 hours. The background variability is $> 5\%$. Given the faint stars we are imaging and the requirement of precision photometry, we request the background limited constraint of $< 10\%$ above minimum. We place the bright foreground in a chip away from the main stellar disk to avoid persistence image effects and Dragon's Breath with orient constraints to ensure the majority of the stellar disk falls on the Module B chips. With the scheduling constraints, Leo P is observable for 19 days per year.

Proposal 1617 - Targets - The Boundary of Galaxy Formation: Constraints from the Ancient Star Formation of the Isolated, Extremely ...

| Fixed Targets | # | Name | Target Coordinates | Targ. Coord. Corrections | Miscellaneous |
|-----------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|-----|-------|-------------------------------------------------------------------------------------|--------------------------|---------------|
| | (1) | LEO-P | RA: 10 21 42.5094 (155.4271225d) Dec: +18 05 16.09 (18.08780d) Equinox: J2000 | | |
| <i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i> <i>Category=Galaxy</i> <i>Description=[Dwarf irregular galaxies, Field galaxies]</i> | | | | | |

Proposal 1617 - Observation 1 - The Boundary of Galaxy Formation: Constraints from the Ancient Star Formation of the Isolated, Extr...

Fri May 27 05:02:13 GMT 2022

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|-----------------------------|-----------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|----------------------------|-------------------------------------------------------------------------------------|------------------------|---------------------------------|-----------------------------|---------------------------|----------------------|----------------------------|---------------------------|
| Observation | <p>Proposal 1617, Observation 1: LeoP</p> <p>Diagnostic Status: Warning</p> <p>Observing Template: NIRCcam Imaging</p> | | | | | | | | | |
| Diagnostics | <p>(Visit 1:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.</p> <p>(Visit 1:2) Warning (Form): Overheads are provisional until the Visit Planner has been run.</p> | | | | | | | | | |
| Fixed Targets | # | Name | Target Coordinates | | Targ. Coord. Corrections | | | Miscellaneous | | |
| | (1) | LEO-P | RA: 10 21 42.5094 (155.4271225d) Dec: +18 05 16.09 (18.08780d) Equinox: J2000 | | | | | | | |
| | <p><i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i></p> <p><i>Category=Galaxy</i></p> <p><i>Description=[Dwarf irregular galaxies, Field galaxies]</i></p> | | | | | | | | | |
| Template | Module | | | | Subarray | | | | | |
| | B | | | | FULL | | | | | |
| Dithers | # | Primary Dither Type | | Primary Dithers | | Subpixel Dither Type | | Dither Size | | Subpixel Positions |
| | 1 | NONE | | | | STANDARD | | | | 7 |
| Spectral Elements | # | Short Filter | Long Filter | Readout Pattern | Groups/Int | Integrations/Exp | Total Integrations | Total Dithers | Total Exposure Time | ETC Wkbk.Calc ID |
| | 1 | F090W | F277W | DEEP8 | 5 | 10 | 70 | 7 | 66814.92 | |
| | 2 | F150W | F277W | DEEP8 | 5 | 6 | 42 | 7 | 40058.889 | |
| Special Requirements | <p>Group Visits within 53.0 Days</p> <p>Aperture PA Range 119.3 to 120 Degrees (V3 119.27547802 to 119.97547802)</p> <p>Visits Same PA</p> <p>Background Limited. Background no more than 10th percentile above minimum</p> | | | | | | | | | |