



# 1712 - JWST beholds the multiple-merger assembly of the most luminous quasar

Cycle: 1, Proposal Category: GO

## INVESTIGATORS

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Dr. Hyunsung Jun (CoI)	Seoul National University
Andrew Blain (CoI) (ESA Member)	University of Leicester

## OBSERVATIONS

<i>Folder</i>	<i>Observation</i>	<i>Label</i>	<i>Observing Template</i>	<i>Science Target</i>
Observation Folder				
	1	MIRI-BKG	MIRI Medium Resolution Spectroscopy	(3) W2246-0526-MIRI-MRS-BKG
	2	MIRI	MIRI Medium Resolution Spectroscopy	(1) W2246-0526-MIRI-MRS
	3	NIRSPEC	NIRSpec IFU Spectroscopy	(2) W2246-0526-NIRSPEC-IFU

## ABSTRACT

The most luminous quasars open a window to study the physics that govern the assembly of mass at the nodes of the cosmic web at a time when the Universe was ramping up to its period of peak activity. In particular, “cosmic morning”  $z \sim 3\text{--}5$  obscured quasars may be signposting the locations at which the most super-massive black holes were being built via the accretion of in-falling galaxies, before emerging in a blast of radiation. While long-predicted by simulations, the detection of the tidal structures that could pinpoint the dynamics of these systems is at the limit of current facilities. Deep ALMA observations of the [CII] line in the most luminous galaxy known, W2246–0526, a hyper-luminous obscured quasar at  $z = 4.6$  surrounded by multiple neighbor galaxies, have now revealed the complex kinematics of this merger system with unparalleled detail at this epoch. Companion galaxies, tidal streams, proto-spiral arms, gas clumps and bubbles, are all resolved spatially and spectrally. However, the [CII] emission does not allow for an unambiguous characterization of the ionizing sources, as it arises from virtually every interstellar medium phase. Here we request JWST NIRSpec IFU and MIRI MRS observations of the W2246–0526 system to identify the sources of gas ionization and heating, establish whether the multi-phase [CII] gas is kinematically coupled to the purely ionized phase, and reveal the transition between the starlight-dominated spectrum of W2246–0526's host and the hot dust continuum from the central obscured quasar. These observations will test the paradigm of galaxy-quasar feeding and feedback in this truly unique source, adding an early gem to the JWST legacy.

## **OBSERVING DESCRIPTION**

This proposal aims to spatially and kinematically resolve the ionized gas phase in W2246-0526, a hyper-luminous obscured quasar at  $z=4.6$  that is part of multiple-merger system of galaxies. NIRSpec IFU observations of the H-alpha line will be obtained down to a SFR of  $\sim 2 \text{ Msun/yr/kpc}^2$ , allowing the identification of emission in faint, tidal streamers connecting the members of the merger over a region of  $\sim 9'' \times 6''$ . Other rest-frame optical lines will be targeted (see below), allowing for a characterization of the ionizing sources in W2246-0526 and its companion galaxies via emission line diagnostics. In addition, a single  $3'' \times 3''$  pointing with MIRI MRS will map the transition from starlight-dominated to hot dust dominated emission in W2246-0526.

Two high-resolution spectral gratings will be used with NIRSpec, covering the rest-frame  $\sim 3000\text{--}9000\text{\AA}$  at a resolution of  $\sim 150 \text{ km/s}$ , needed to kinematically resolve the components of the merger. The emission lines to be detected are [OII], H-beta, [OIII], H-alpha and [NII]. Since the quasar at the center of W2246-0526 dominates the emission at observed-frame  $\sim > 5\mu\text{m}$ , all the spectral settings (A, B, C) of at least channels 1 and 2 of the MRS are needed in order to probe the transition between the continua from the host and the AGN.

The available V3PA ranges to observe W2246-0526 are limited to two  $\sim 20\text{deg}$  wide windows during the year, which narrowly sets the orientation the NIRSpec mosaic and allows us to define a more efficient  $3 \times 2$  rectangular tiling instead of a more generic squared  $3 \times 3$  tiling.

Proposal 1712 - Targets - JWST beholds the multiple-merger assembly of the most luminous quasar

#	Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous
(1)	W2246-0526-MIRI-MRS	RA: 22 46 7.5500 (341.5314583d) Dec: -05 26 34.83 (-5.44301d) Equinox: J2000		
<i>Comments:</i> Category=Galaxy Description=[Infrared galaxies, Quasars, Starburst galaxies] Extended=YES				
(2)	W2246-0526-NIRSPEC-IFU	RA: 22 46 7.6600 (341.5319167d) Dec: -05 26 34.87 (-5.44302d) Equinox: J2000		
<i>Comments:</i> Category=Galaxy Description=[Infrared galaxies, Quasars, Starburst galaxies] Extended=YES				
(3)	W2246-0526-MIRI-MRS- BKG	RA: 22 46 6.5000 (341.5270833d) Dec: -05 27 38.63 (-5.46073d) Equinox: J2000		
<i>Comments:</i> Category=Galaxy Description=[Infrared galaxies, Quasars, Starburst galaxies] Extended=YES				

Fixed Targets

Proposal 1712 - Observation 1 - JWST beholds the multiple-merger assembly of the most luminous quasar

Thu Jun 15 23:00:59 GMT 2023

<b>Observation</b>	<b>Proposal 1712, Observation 1: MIRI-BKG</b> <b>Diagnostic Status: Warning</b> Observing Template: MIRI Medium Resolution Spectroscopy Background Observation For: [MIRI (Obs 2)]																																																																																																																																													
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Proposal 1712 - Observation 1 - JWST beholds the multiple-merger assembly of the most luminous quasar

Special Requirements

Background Limited. Background no more than 10th percentile above minimum

2 After 1

Sequence Observations 1, 2, Non-interruptible

Group Observations 1, 2, 3 within 10 Days

Proposal 1712 - Observation 2 - JWST beholds the multiple-merger assembly of the most luminous quasar

Thu Jun 15 23:00:59 GMT 2023

<b>Observation</b>	<b>Proposal 1712, Observation 2: MIRI</b> <b>Diagnostic Status: Warning</b> Observing Template: MIRI Medium Resolution Spectroscopy Background Observations:[MIRI-BKG (Obs 1)]												
	(Visit 2:1) Warning (Form): Overheads are provisional until the Visit Planner has been run. (Visit 2:1) Informational (Form): Visit schedulable, but most scheduling windows are when JWST is pointed in direction of greatest micrometeoroid impact risk. This is likely due to scheduling special requirements.												
<b>Diagnosics</b>													
<b>Fixed Targets</b>	#	Name	Target Coordinates		Targ. Coord. Corrections				Miscellaneous				
	(1)	W2246-0526-MIRI-MRS	RA: 22 46 7.5500 (341.5314583d) Dec: -05 26 34.83 (-5.44301d) Equinox: J2000										
Comments: Category=Galaxy Description=[Infrared galaxies, Quasars, Starburst galaxies] Extended=YES													
<b>Acquisition</b>	#	Target											
	1	NONE											
<b>Template</b>	AcqFilter	Primary Channel			Simultaneous Imaging			Imager Subarray		Grating Wheel Direction			
		ALL			YES			FULL		NEUTRAL			
<b>Dithers</b>	#	Dither Type				Optimized For				Direction			
	1	4-Point				EXTENDED SOURCE				NEGATIVE			
<b>Spectral Elements</b>	#	Wavelength Range	Detector	Filter	Readout Pattern	Groups/Int	Integrations/E xp	Exposures/Dit h	Dither	Total Dithers	Total Integrations	Total Exposure Time	ETC Wkbk.Calc ID
	1		IMAGER	F1000W	SLOWR1	15	1	1	Dither 1	4	4	1433.395	56037
	1	LONG(C)	MRSLONG		SLOWR1	15	1	1	Dither 1	4	4	1433.395	56037
	1	LONG(C)	MRSSHORT		SLOWR1	15	1	1	Dither 1	4	4	1433.395	56037
	2		IMAGER	F770W	SLOWR1	23	1	1	Dither 1	4	4	2197.873	56037
	2	MEDIUM(B)	MRSLONG		SLOWR1	23	1	1	Dither 1	4	4	2197.873	56037
	2	MEDIUM(B)	MRSSHORT		SLOWR1	23	1	1	Dither 1	4	4	2197.873	56037
	3		IMAGER	F560W	SLOWR1	31	1	1	Dither 1	4	4	2962.35	56037
	3	SHORT(A)	MRSLONG		SLOWR1	31	1	1	Dither 1	4	4	2962.35	56037
	3	SHORT(A)	MRSSHORT		SLOWR1	31	1	1	Dither 1	4	4	2962.35	56037

Proposal 1712 - Observation 2 - JWST beholds the multiple-merger assembly of the most luminous quasar

Special Requirements

Background Limited. Background no more than 10th percentile above minimum

2 After 1

Sequence Observations 1, 2, Non-interruptible

Group Observations 1, 2, 3 within 10 Days

Proposal 1712 - Observation 3 - JWST beholds the multiple-merger assembly of the most luminous quasar

Thu Jun 15 23:00:59 GMT 2023

<b>Observation</b>	<b>Proposal 1712, Observation 3: NIRSPEC</b> <b>Diagnostic Status: Warning</b> Observing Template: NIRSPEC IFU Spectroscopy											
	(Visit 3:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.											
<b>Fixed Targets</b>	<b>#</b>	<b>Name</b>	<b>Target Coordinates</b>			<b>Targ. Coord. Corrections</b>			<b>Miscellaneous</b>			
	(2)	W2246-0526-NIRSPEC-IFU	RA: 22 46 7.6600 (341.5319167d) Dec: -05 26 34.87 (-5.44302d) Equinox: J2000									
<b>Template</b>	<b>TA Method</b>											
	NONE											
<b>Mosaic</b>	<b>Rows</b>	<b>Columns</b>	<b>Row Overlap %</b>	<b>Column Overlap %</b>	<b>Row shift (deg)</b>	<b>Column shift (deg)</b>	<b>Tile Order</b>					
	2	3	5.0	5.0	0.0	0.0	DEFAULT					
<b>Dithers</b>	<b>#</b>	<b>Dither Type</b>	<b>Size</b>	<b>Starting Point</b>			<b>Number of Points</b>	<b>Points</b>				
	1	4-POINT-DITHER										
<b>Spectral Elements</b>	<b>#</b>	<b>Grating/Filter</b>	<b>Readout Pattern</b>	<b>Groups/Int</b>	<b>Integrations/Exp</b>	<b>Leakcal</b>	<b>Dither</b>	<b>Autocal</b>	<b>Total Dithers</b>	<b>Total Integrations</b>	<b>Total Exposure Time</b>	<b>ETC Wkbk.Calc ID</b>
	1	G235H/F170LP	NRSIRS2	15	1	false	true	NONE	4	4	4435.023	56097
	2	G235H/F170LP	NRSIRS2	15	1	true	false	NONE	1	1	1108.756	
	3	G395H/F290LP	NRSIRS2	6	1	false	true	NONE	4	4	1809.022	56097
	4	G395H/F290LP	NRSIRS2	6	1	true	false	NONE	1	1	452.256	

Proposal 1712 - Observation 3 - JWST beholds the multiple-merger assembly of the most luminous quasar

Special Requirements

Group Observations 1, 2, 3 within 10 Days