



4201 - The Butterfly Effect: Determining the distribution of ices across a young disk to constrain planet formation

Cycle: 2, Proposal Category: GO

INVESTIGATORS

<i>Name</i>	<i>Institution</i>
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Dr. Melissa McClure (CoI) (ESA Member)	Universiteit Leiden
Mr. Ardjan Sturm (CoI) (ESA Member)	Universiteit Leiden

OBSERVATIONS

<i>Folder</i>	<i>Observation</i>	<i>Label</i>	<i>Observing Template</i>	<i>Science Target</i>
Observation Folder				
	1	MIRI science	MIRI Medium Resolution Spectroscopy	(1) IRAS04302-MIRI
	2	MIRI background	MIRI Medium Resolution Spectroscopy	(2) MIRI-BACKGROUND
	3	NIRSpec	NIRSpec IFU Spectroscopy	(3) IRAS04302-NIRSPEC

ABSTRACT

Young disks (Class I) orbiting nascent stars are the birthplaces of planets. In order to link the elemental composition of exoplanets, that JWST will provide in plenty, to their formation history, one needs to determine the chemical composition of both the gas and solids across the disk. With ALMA a stunning progress has been made to map the gas content of disks, but the distribution and compositions of ices remains poorly known. We aim to target IRAS 04302 (the Butterfly Star) with MIRI MRS and NIRSpec IFU, to map the major carbon and oxygen carriers in the ice phase: CO, CO₂, CH₄ and H₂O. The exact edge-on geometry and large size makes this young disk uniquely suited to unveil for the first time the vertical and radial

structure of the ices. We will obtain the radial and vertical location of the CO, CO₂, H₂O snow surfaces, which are setting the C/O ratio of the gas and dust. These observations will demonstrate the impact that JWST will have on the studies of planet formation.

OBSERVING DESCRIPTION

We target a Class I protostar IRAS-04302 with MIRI MRS in full spectral coverage of 5-28 micron, and with NIRSpec IFU in high spectral resolution grating/filter combination: G395H/F290LP and G235H/F170LP. This spectral setup covers H₂O, CO₂, CO, CH₄ features, essential for determining the C/O ratio of the ice mantle.

17 groups per sub-band, per dither, per mosaic tile is spent with MIRI (in slow readout mode) resulting in total science time of 4.1 hrs, and 10 groups per grating, per dither, per mosaic tile on NIRSpec resulting in 2.92 hrs of science observation. For MIRI a mosaic of 3x1 is selected to cover the entire extent of the disk and the inner envelope. Since Spitzer images show more extended emission at 3.6 micron compared to 5.8 micron, we use a 3x3 mosaic for NIRSpec to maximize spatial information. A 4-point dither is used for both instruments to maximize the spatial and spectral sampling. A dedicated sky background is targeted with MIRI.

Proposal 4201 - Targets - The Butterfly Effect: Determining the distribution of ices across a young disk to constrain planet formation

#	Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous
(1)	IRAS04302-MIRI	RA: 04 33 16.5000 (68.3187500d) Dec: +22 53 20.40 (22.88900d) Equinox: J2000		
<i>Comments:</i> <i>Category=ISM</i> <i>Description=[Interstellar dust, Protostars]</i> <i>Extended=YES</i>				
(2)	MIRI-BACKGROUND	RA: 04 33 16.0100 (68.3167083d) Dec: +23 03 50.40 (23.06400d) Equinox: J2000		
<i>Comments:</i> <i>Category=Calibration</i> <i>Description=[Telescope/sky background]</i> <i>Extended=YES</i>				
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Proposal 4201 - Observation 1 - The Butterfly Effect: Determining the distribution of ices across a young disk to constrain planet forma...

Mon Jul 10 23:01:58 GMT 2023

Observation	Proposal 4201, Observation 1: MIRI science Diagnostic Status: Warning Observing Template: MIRI Medium Resolution Spectroscopy Background Observations:[MIRI background (Obs 2)]																																																																																																																																													
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Special Requirements

Sequence Observations 1, 2, Non-interruptible

Proposal 4201 - Observation 2 - The Butterfly Effect: Determining the distribution of ices across a young disk to constrain planet forma...

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Proposal 4201 - Observation 3 - The Butterfly Effect: Determining the distribution of ices across a young disk to constrain planet forma...

Mon Jul 10 23:01:58 GMT 2023

Observation	<p>Proposal 4201, Observation 3: NIRSpec</p> <p>Diagnostic Status: Warning</p> <p>Observing Template: NIRSpec IFU Spectroscopy</p>											
Diagnostics	(Visit 3:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.											
Fixed Targets	#	Name	Target Coordinates			Targ. Coord. Corrections			Miscellaneous			
	(3)	IRAS04302-NIRSPEC	RA: 04 33 16.5000 (68.3187500d) Dec: +22 53 20.40 (22.88900d) Equinox: J2000									
	<p><i>Comments:</i> <i>Category=ISM</i> <i>Description=[Interstellar dust, Protostars]</i> <i>Extended=YES</i></p>											
Template	TA Method											
	NONE											
Mosaic	Rows	Columns	Row Overlap %	Column Overlap %	Row shift (deg)	Column shift (deg)	Tile Order					
	3	3	10.0	10.0	0.0	-25.0	DEFAULT					
Dithers	#	Dither Type		Size	Starting Point		Number of Points	Points				
	1	4-POINT-DITHER										
Spectral Elements	#	Grating/Filter	Readout Pattern	Groups/Int	Integrations/Exp	Leakcal	Dither	Autocal	Total Dithers	Total Integrations	Total Exposure Time	ETC Wkbk.Calc ID
	1	G395H/F290LP	NRSIRS2RAPID	10	1	false	true	NONE	4	4	641.911	63038.4
	2	G235H/F170LP	NRSIRS2RAPID	10	1	false	true	NONE	4	4	641.911	