



# 5086 - Confirming Population III or a Direct Collapse Black Hole in the halo of GN-z11 at $z=10.6$

Cycle: 3, Proposal Category: GO

## INVESTIGATORS

| <i>Name</i>                                     | <i>Institution</i>                                     |
|---|--|
| <b>Prof. Roberto Maiolino (PI) (ESA Member)</b> | <b>University of Cambridge</b>                         |
| Dr. Jan Scholtz (CoI) (ESA Member)              | University of Cambridge, Kavli Institute for Cosmology |
| Dr. Emma Curtis-Lake (CoI) (ESA Member)         | University of Hertfordshire                            |
| Dr. Stefano Carniani (CoI) (ESA Member)         | Scuola Normale Superiore, Pisa                         |
| Prof. Andrew Bunker (CoI) (ESA Member)          | University of Oxford                                   |
| Dr. Joris Witstok (CoI) (ESA Member)            | University of Cambridge                                |
| Dr. Hannah Uebler (CoI) (ESA Member)            | University of Cambridge                                |
| Dr. Francesco D'Eugenio (CoI) (ESA Member)      | University of Cambridge, Kavli Institute for Cosmology |
| Dr. Giacomo Venturi (CoI) (ESA Member)          | Scuola Normale Superiore, Pisa                         |
| Dr. Xihan Ji (CoI) (ESA Member)                 | University of Cambridge                                |
| Dr. Giovanni Cresci (CoI) (ESA Member)          | INAF - Osservatorio Astrofisico di Arcetri             |
| Dr. Santiago Arribas (CoI) (ESA Member)         | Consejo Superior de Investigaciones Cientificas        |
| Mr. Giovanni Mazzolari (CoI) (ESA Member)       | Universita di Bologna                                  |
| Dr. Aayush Saxena (CoI) (ESA Member)            | University of Oxford                                   |
| Dr. Sandro Tacchella (CoI) (ESA Member)         | University of Cambridge                                |
| Dr. Michele Perna (CoI) (ESA Member)            | Centro de Astrobiología (CAB)                          |
| Dr. Eiichi Egami (CoI) (US Admin CoI)           | University of Arizona                                  |
| Mr. Callum Witten (CoI) (ESA Member)            | University of Cambridge                                |
| Dr. Nimisha Kumari (CoI) (ESA Member)           | Space Telescope Science Institute - ESA - JWST         |
| Dr. Daniel J. Eisenstein (CoI)                  | Harvard University                                     |
| Prof. Brant Robertson (CoI)                     | University of California - Santa Cruz                  |

| <i>Name</i>                             | <i>Institution</i>                      |
|---|---|
| Dr. Stephane Charlot (CoI) (ESA Member) | CNRS, Institut d'Astrophysique de Paris |
| Dr. Zhiyuan Ji (CoI)                    | University of Arizona                   |
| Dr. Mirko Curti (CoI) (ESA Member)      | European Southern Observatory - Germany |

## OBSERVATIONS

| <i>Folder</i>      | <i>Observation</i> | <i>Label</i> | <i>Observing Template</i> | <i>Science Target</i> |
|--------------------|--------------------|--------------|---------------------------|-----------------------|
| Observation Folder |                    |              |                           |                       |
|                    | 1                  | IFS          | NIRSpec IFU Spectroscopy  | (1) GNz11             |

## ABSTRACT

Previous observations have revealed HeII emission in the halo of the luminous galaxy GN-z11 at  $z=10.6$ , which could be tracing either Population III stars (formed in a pocket of pristine gas, as expected by some models) or an accreting Direct Collapse Black Hole (DCBH), in the earliest (pristine) phases of its formation. However, the detection is still tentative (5 sigma). We propose to obtain a deep NIRSpec-IFS observation (40 hours on-source), with the high resolution grating, with the primary goal of confirming the HeII detection in the halo of GN-z11 and discriminate between the PopIII and DCBH scenarios. The confirmation of the detection of either of these long-sought objects would be a groundbreaking discovery. In addition, the observation will tackle a wealth of additional science cases, such as: trace the ionization cone and outflow in GN-z11, which is key for understanding the onset of AGN feedback in early galaxies; determine the C/O abundance in the host galaxy of GN-z11 and in its circumgalactic medium, possibly confirming signatures of PopIII enrichment; accurately measure the black hole mass of GN-z11; disentangle the nitrogen emission from the Broad Line Region in GN-z11 from the emission coming from of the host galaxy, hence constraining the N/O abundance in the nuclear region of GN-z11 and its host galaxy; confirm the presence of WR stars and their spatial distribution, hence providing tight constraints on the age of the stellar population in GN-z11.

## OBSERVING DESCRIPTION

We propose observations for GN-z11 using the NIRSpec/IFS mode to explore the 2D spatially resolved properties of its ISM and the PopIII cluster 0.5 arcseconds away.. To this goal we plan to use the medium resolution ( $R\sim 2700$ ) G235H/F170LP spectral configurations to target the following lines: CIII]1907,09, NIII], NIV] and HeII1640 in GN-z11 and its surrounding.

These lines have measured fluxes in the range of  $5-22 \times 10^{-19}$  erg/s/cm<sup>2</sup> from NIRSPEC/MSA observations using PRISM and R1000 gratings in previous IFS and MSA observations (Bunker et al 2023, Maiolino et al 2023, Scholtz et al 2023). Considering these values and assuming a linewidth

## JWST Proposal 5086 (Created: Wednesday, March 6, 2024 at 1:01:25 PM Eastern Standard Time) - Overview

of  $\sim 120$  km/s, with 40 hours of exposure time for G235H/F170LP the ETC gives an integrated S/N  $> 10$  (depending on the line) at the central  $\sim 0.6$  arcsec region, then allowing us to explore its 2D properties beyond that zone.

No target acquisition is included, as the accuracy of the direct JWST pointing is enough to guarantee that the target is well centered in the NIRSpec IFU FoV.

To minimize the redout noise we use NRSIRS2 readout pattern with 25 groups per integration. To each a total exposure time of  $\sim 40$  hr, we use 38 dither positions with 2 integrations per exposure. We select a cycling medium dither pattern, which provides a good compromise between an amplitude ( $\sim 0.5''$ ) large enough to "jump" the failed open microshutters, and to deal with other sources of background. This dither pattern also provides a good sub-pixel sampling, and allows to reconstruct the datacube with spaxels of 50 mas.

No extra background exposures are included. At high resolution we expect the background to be low and, in addition, there will plenty of spaxels free from galaxy emission suitable to derive the background.

We do not ask for MSA leakage exposures, as we have verified that there are not very bright targets on the MSA FoV that could contaminate the IFU spectra. In addition, the presence of possible leakage would mainly affect the continuum, but our main goal is the analysis of the emission lines. We restricted PA to avoid bright stars in the FoV of the MSA masks.

Proposal 5086 - Targets - Confirming Population III or a Direct Collapse Black Hole in the halo of GN-z11 at z=10.6

| Fixed Targets  | #   | Name  | Target Coordinates  | Targ. Coord. Corrections | Miscellaneous |
|--|-----|-------|---|--------------------------|---------------|
|  | (1) | GNz11 | RA: 12 36 25.4501 (189.1060421d)<br>Dec: +62 14 31.35 (62.24204d)<br>Equinox: J2000 |                          |               |
| <i>Comments:</i><br>Category= <i>Galaxy</i><br>Description= <i>[Active galactic nuclei, High-redshift galaxies, Lyman-alpha galaxies, Protogalaxies]</i> |     |       |   |                          |               |

Proposal 5086 - Observation 1 - Confirming Population III or a Direct Collapse Black Hole in the halo of GN-z11 at z=10.6

Wed Mar 06 18:01:25 GMT 2024

|                             |  |                       |   |                   |                         |                                 |               |                         |                      |                           |                            |                         |
|-----------------------------|--|-----------------------|---|-------------------|-------------------------|---------------------------------|---------------|-------------------------|----------------------|---------------------------|----------------------------|-------------------------|
| <b>Observation</b>          | <p><b>Proposal 5086, Observation 1: IFS</b></p> <p><b>Diagnostic Status: Warning</b></p> <p>Observing Template: NIRSpec IFU Spectroscopy</p>   |                       |   |                   |                         |                                 |               |                         |                      |                           |                            |                         |
| <b>Diagnostics</b>          | <p>(Visit 1:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.</p> <p>(Visit 1:2) Warning (Form): Overheads are provisional until the Visit Planner has been run.</p>  |                       |   |                   |                         |                                 |               |                         |                      |                           |                            |                         |
| <b>Fixed Targets</b>        | <b>#</b>   | <b>Name</b>           | <b>Target Coordinates</b>   |                   |                         | <b>Targ. Coord. Corrections</b> |               |                         | <b>Miscellaneous</b> |                           |                            |                         |
|                             | (1)  | GNz11                 | RA: 12 36 25.4501 (189.1060421d)<br>Dec: +62 14 31.35 (62.24204d)<br>Equinox: J2000 |                   |                         |                                 |               |                         |                      |                           |                            |                         |
|                             | <p><i>Comments:</i><br/> <i>Category=Galaxy</i><br/> <i>Description=[Active galactic nuclei, High-redshift galaxies, Lyman-alpha galaxies, Protogalaxies]</i></p>  |                       |   |                   |                         |                                 |               |                         |                      |                           |                            |                         |
| <b>Template</b>             | <b>TA Method</b>   |                       |   |                   |                         |                                 |               |                         |                      |                           |                            |                         |
|                             | NONE   |                       |   |                   |                         |                                 |               |                         |                      |                           |                            |                         |
| <b>Dithers</b>              | <b>#</b>   | <b>Dither Type</b>    |   | <b>Size</b>       | <b>Starting Point</b>   |                                 |               | <b>Number of Points</b> | <b>Points</b>        |                           |                            |                         |
|                             | 1  | CYCLING               |   | MEDIUM            | 1                       |                                 |               | 39                      |                      |                           |                            |                         |
| <b>Spectral Elements</b>    | <b>#</b>   | <b>Grating/Filter</b> | <b>Readout Pattern</b>  | <b>Groups/Int</b> | <b>Integrations/Exp</b> | <b>Leakcal</b>                  | <b>Dither</b> | <b>Autocal</b>          | <b>Total Dithers</b> | <b>Total Integrations</b> | <b>Total Exposure Time</b> | <b>ETC Wkbk.Calc ID</b> |
|                             | 1  | G235H/F170LP          | NRSIRS2   | 25                | 2                       | false                           | true          | NONE                    | 39                   | 78                        | 143379.611                 | 169782                  |
| <b>Special Requirements</b> | <p>Group Visits within 53.0 Days<br/>                 Aperture PA Range 48 to 58 Degrees (V3 269.02835083 to 279.02835083)<br/>                 Aperture PA Range 144 to 159 Degrees (V3 5.02835083 to 20.02835083)<br/>                 Aperture PA Range 235 to 283 Degrees (V3 96.02835083 to 144.02835083)<br/>                 Aperture PA Range 296 to 305 Degrees (V3 157.02835083 to 166.02835083)<br/>                 Visits Same PA</p> |                       |   |                   |                         |                                 |               |                         |                      |                           |                            |                         |