



5394 - Untangling the Nature of the Kennicutt Relation at the Low End: the Iconic Case of the Giant Low Surface Brightness Galaxy Malin 1

Cycle: 3, Proposal Category: GO

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Dr. Rongmon Bordoloi (CoI) (US Admin CoI)	North Carolina State University

OBSERVATIONS

<i>Folder</i>	<i>Observation</i>	<i>Label</i>	<i>Observing Template</i>	<i>Science Target</i>
Malin 1				
	1	Malin 1	MIRI Imaging	(1) Malin_1

<i>Folder</i>	<i>Observation</i>	<i>Label</i>	<i>Observing Template</i>	<i>Science Target</i>
	2	Background	MIRI Imaging	(2) Sky_background

ABSTRACT

Star formation is the basis for galaxy evolution, but how this process occurs in extremely low gas density environments is not understood. The Kennicutt relation correlates the star formation rate surface density to the gas surface density in galaxies. On the low end of the relation, there is a break in the profile and a steepening of the slope where HI efficiently converts to H₂. What is unclear is the physical process or ISM conditions that cause the steepness in the relation. Either galaxies have an HI-dominated disk with low H₂ filling factor or there is low star formation efficiency within the H₂ in this regime. At the extreme low end of the Kennicutt relation sits low surface brightness galaxies (LSBGs), which can be used to distinguish between these two scenarios. Recent work has placed a stringent upper limit on the H₂ gas surface density within the giant LSBG Malin 1, seeming to point towards a low H₂ filling factor. However, the true answer cannot be determined without quantifying the amount of H₂ gas within the disk. Here we propose to observe Malin 1 with MIRI in the F770W dust emission filter and F1280W dust continuum filter to map out the gaseous disk (HI and H₂) using dust as a tracer. These observations will enable us to determine the spatially-resolved gas surface density of Malin 1 to the disk edge. By comparing the total gas surface density to the previously measured HI gas surface density, we will identify the cause of the steepening in the Kennicutt relation through the difference of the two values. This work will deepen our understanding of the star formation process in the most extreme gas conditions.

OBSERVING DESCRIPTION

We will obtain MIRI/F770W and F1280W imaging of the giant low surface brightness galaxy Malin 1 to identify where the gas (HI and H₂) resides using PAH emission as a tracer. These observations will allow us to map the gas surface density to extremely low values in the disk and compare the dust morphology with high surface brightness galaxies. We will utilize a 4-point extended source dither pattern. The values of groups and integrations have been chosen to achieve a minimum SNR~5 at the disk edge for the F770W dust emission filter and the F1280W dust continuum filter. We use the FULL subarray and the FASTR1 readout. To cover the full galaxy, we use a 1x2 mosaic. Background observations are taken at the same exposure time with a 4-point dither. We constrain the PA to be the same for each filter observation of science and background, respectively. The background and science observations are grouped and non-interruptable.

Proposal 5394 - Targets - Untangling the Nature of the Kennicutt Relation at the Low End: the Iconic Case of the Giant Low Surface B...

#	Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous
(1)	Malin_1	RA: 12 36 59.3500 (189.2472917d) Dec: +14 19 49.16 (14.33032d) Equinox: J2000		
<i>Comments:</i> <i>Category=Galaxy</i> <i>Description=[Low Surface Brightness galaxies, Spiral arms]</i> <i>Extended=YES</i>				
(2)	Sky_background	RA: 12 37 7.9650 (189.2831875d) Dec: +14 19 26.24 (14.32396d) Equinox: J2000		
<i>Comments: background region to the east of Malin 1</i> <i>Category=Calibration</i> <i>Description=[Telescope/sky background]</i> <i>Extended=NO</i>				

Proposal 5394 - Observation 1 - Untangling the Nature of the Kennicutt Relation at the Low End: the Iconic Case of the Giant Low Sur...

Mon Mar 31 13:00:08 GMT 2025

Observation	Proposal 5394, Observation 1: Malin 1 Diagnostic Status: Warning Observing Template: MIRI Imaging Background Observations:[Background (Obs 2)]																																										
Diagnostics	(Malin 1 (Obs 1)) Warning (Form): Target requiring background exposure selected for template that doesn't require background exposure (Visit 1:1) Warning (Form): Overheads are provisional until the Visit Planner has been run. (Visit 1:2) Warning (Form): Overheads are provisional until the Visit Planner has been run. (Malin 1 (Obs 1)) Informational (Form): The Visit Planner and Spike may produce different schedulability results.																																										
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Dithers	<table border="1"> <thead> <tr> <th>#</th> <th>Dither Type</th> <th>Starting Point</th> <th>Number of Points</th> <th>Points</th> <th>Starting Set</th> <th>Number of Sets</th> <th>Optimized For</th> <th>Direction</th> <th>Pattern Size</th> </tr> </thead> <tbody> <tr> <td>1</td> <td>CYCLING</td> <td>6</td> <td>4</td> <td></td> <td></td> <td></td> <td></td> <td></td> <td>DEFAULT</td> </tr> </tbody> </table>										#	Dither Type	Starting Point	Number of Points	Points	Starting Set	Number of Sets	Optimized For	Direction	Pattern Size	1	CYCLING	6	4						DEFAULT													
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Proposal 5394 - Observation 1 - Untangling the Nature of the Kennicutt Relation at the Low End: the Iconic Case of the Giant Low Sur...

Special Requirements

Group Visits within 53.0 Days
Visits Same PA

Group Observations 1, 2, Non-interruptible
Same Aperture PA 1

Proposal 5394 - Observation 2 - Untangling the Nature of the Kennicutt Relation at the Low End: the Iconic Case of the Giant Low Sur...

Mon Mar 31 13:00:08 GMT 2025

Observation	Proposal 5394, Observation 2: Background Diagnostic Status: Warning Observing Template: MIRI Imaging Background Observation For: [Malin 1 (Obs 1)]										
	(Visit 2:1) Warning (Form): Overheads are provisional until the Visit Planner has been run. (Background (Obs 2)) Informational (Form): The Visit Planner and Spike may produce different schedulability results.										
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	FULL										
Dithers	#	Dither Type	Starting Point	Number of Points	Points	Starting Set	Number of Sets	Optimized For	Direction	Pattern Size	
	1	CYCLING	6	4						DEFAULT	
Spectral Elements	#	Filter	Readout Pattern	Groups/Int	Integrations/Exp	Exposures/Dith	Dither	Total Dithers	Total Integrations	Total Exposure Time	ETC Wkbk.Calc ID
	1	F770W	FASTR1	17	10	1	Dither 1	4	40	1986.929	172028.23
	2	F1280W	FASTR1	32	10	1	Dither 1	4	40	3651.953	172028.26
Special Requirements	Group Observations 1, 2, Non-interruptible Same Aperture PA 2										