



6410 - A Giant Planet Candidate Orbiting a Young, Massive White Dwarf

Cycle: 3, Proposal Category: GO

INVESTIGATORS

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OBSERVATIONS

<i>Folder</i>	<i>Observation</i>	<i>Label</i>	<i>Observing Template</i>	<i>Science Target</i>
1				
	1	1 NIRSpec IFU	NIRSpec IFU Spectroscopy	(1) SDSS-J071816.41+373139.1

ABSTRACT

From a sample of nearby, young, massive white dwarfs with archival Spitzer/IRAC photometry, we have identified one source with 4.5 micron flux excesses that are inconsistent with debris disks or background galaxy contamination. This source is best explained by a binary system unresolved in Spitzer's 2" PSF consisting of a white dwarf primary and a hot, young secondary with 5 Jupiter mass. If this planetary interpretation is correct, then we expect the white dwarf primary and planetary secondary in this system to have F444W ~ 18 and angular separation in excess of 0.2". The white dwarf primary would have had $M_{\text{star}} \sim 8 M_{\text{Sun}}$ and been classified as B stars on the main sequence. Observing this or analogous systems from the ground with GPI 2.0 or SPHERE would require currently impossible-to-achieve H-band contrast ratios in excess of 10^{10} at 0.2". We propose to spectrally and likely spatially resolve this system with JWST NIRSpec IFU to confirm or reject the planetary explanation for its photometry. If the proposed observation recover point sources with planet-like spectra, then it will be the first unambiguously planet-mass objects found orbiting objects that are or were main sequence B stars. The exquisite resolution and sensitivity of the JWST applied to this spatially and spectrally resolvable system will produce high signal-to-noise ratio near-infrared emission spectra of planet-mass objects with precise system age.

OBSERVING DESCRIPTION

The program plans to observe a massive white dwarf at around 4.5 microns with the NIRSpec IFU. A giant planet is expected to be orbiting this white dwarf according to Spitzer's photometry. The white dwarf is massive, faint, and located at 85 pc. It has an expected brightness around 18 mag (~10 μJy) and the expected giant planet is around 19 mag (~4 μJy) at 4.5 micron, which is not a high-contrast imaging. The planet--white dwarf separation is expected to be larger than 0.1", likely resolvable by NIRSpec IFU.

- NIRSpec IFU:

G395M/F290LP

4-point nodding

NRSIRS2RAPID readout

4727s exposure

Proposal 6410 - Targets - A Giant Planet Candidate Orbiting a Young, Massive White Dwarf

Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous
	(1)	SDSS-J071816.41+373139.1	RA: 07 18 16.3781 (109.5682421d) Dec: +37 31 38.68 (37.52741d) Equinox: J2000	Proper Motion RA: -0.0029319271152844282 sec of time/yr Proper Motion Dec: -0.035815999945043586 arcsec/yr Parallax: 0.01182" Epoch of Position: 2015.5	Comments: This object was generated by the targetselector and retrieved from the SIMBAD database. $J = 17.70$, Spitzer 3.6 = 18.08, Spitzer 4.5 = 17.66, distance = 85 pc Category=Star Description=[White dwarfs] Extended=NO

Proposal 6410 - Observation 1 - A Giant Planet Candidate Orbiting a Young, Massive White Dwarf

Fri Mar 01 09:01:47 GMT 2024

Observation	<p>Proposal 6410, Observation 1: 1 NIRSpec IFU</p> <p>Diagnostic Status: Warning</p> <p>Observing Template: NIRSpec IFU Spectroscopy</p>											
Diagnostics	(Visit 1:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.											
Fixed Targets	#	Name	Target Coordinates			Targ. Coord. Corrections			Miscellaneous			
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Template	TA Method											
	NONE											
Dithers	#	Dither Type		Size	Starting Point			Number of Points	Points			
	1	4-POINT-NOD										
Spectral Elements	#	Grating/Filter	Readout Pattern	Groups/Int	Integrations/Exp	Leakcal	Dither	Autocal	Total Dithers	Total Integrations	Total Exposure Time	ETC Wkbk.Calc ID
	1	G395M/F290LP	NRSIRS2RAPID	80	1	false	true	NONE	4	4	4726.8	176484.50