



# 6751 - Resolving the early phase of co-evolution of galaxies and supermassive black holes within cosmic web filaments

Cycle: 3, Proposal Category: GO

## INVESTIGATORS

<i>Name</i>	<i>Institution</i>
<b>Dr. Hideki Umehata (PI)</b>	<b>Nagoya University</b>
Prof. Kotaro Kohno (CoI)	University of Tokyo, Institute of Astronomy
Dr. Yoichi Tamura (CoI)	Nagoya University
Dr. Yuichi Matsuda (CoI)	National Astronomical Observatory of Japan (NAOJ)
Dr. Kouichiro Nakanishi (CoI)	National Astronomical Observatory of Japan (NAOJ)
Prof. Rob J. Ivison (CoI) (ESA Member)	European Southern Observatory - Germany
Prof. Ian Smail (CoI) (ESA Member)	Durham Univ.
Dr. Mariko Kubo (CoI)	Tohoku University, Astronomical Institute
Dr. Mark Swinbank (CoI) (ESA Member)	Durham Univ.
Mr. Masato Hagimoto (CoI)	Nagoya University
Mr. Chihiro Imamura (CoI)	Nagoya University
HUANG Shuo (CoI)	National Astronomical Observatory of Japan (NAOJ)
Norika Okauchi (CoI)	Nagoya University

## OBSERVATIONS

<i>Folder</i>	<i>Observation</i>	<i>Label</i>	<i>Observing Template</i>	<i>Science Target</i>
Observation Folder				
	1	MIRI F770W (a)	MIRI Imaging	(1) ADF22a
	2	MIRI F770W (b)	MIRI Imaging	(2) ADF22b

## **ABSTRACT**

The cosmic web filaments are thought to fuel galaxies and SMBHs and thus growing galaxies within a remarkable gas filament provides a unique laboratory. In this regard, invaluable targets have been discovered at a  $z=3.1$  proto-cluster core. Numerous bright dusty star-forming galaxies (DSFGs), often hosting X-ray AGNs, are situated within Mpc-scale Lyman-alpha filaments. These findings reinforce the idea that the cosmic web plays a key role in fueling the high level activity. To delve deeper into the mechanisms driving the rapid concurrent growth of galaxies and SMBHs, two critical aspects remain unexplored: (a) detailed views on the gas kinematics within the host galaxy and (b) precise evaluations of the (obscured) AGN activity. Building on a successful pilot survey that revealed spiral arms, bar flows, and a bright dusty core in the brightest DSFG, we now propose (i) [CII] mapping at 0.15" resolution and (ii) JWST/MIRI imaging for the six brightest DSFGs. Leveraging extensive existing JWST/ALMA data, we aim to uncover the drivers behind intense starburst activity and elucidate how the co-evolution of galaxies and SMBHs unfolds during these early, obscured stages. (ALMA Cy11 ID: 2024.1.00335.S)

## **OBSERVING DESCRIPTION**

We will observe 6 DSFGs with MIRI in the imaging mode. Our aim is to detect hot dust emission associated with AGN activity. For this, two MIRI filter bands are proposed - F770W and F2100W. For each, two fields are requested to cover the 6 targets in total. The highest spatial resolution F770W filter (0.25 arcsec or 2kpc) is key to measure the spatial transition between the rest-frame near-IR emission from stars and that from heated dust at high redshift. The 2100W filter is vital to detect or put a constraint on the hot dust associated with a various levels of AGN luminosity. The fiducial 4-point dither pattern is adopted to ensure optimal PSF reconstruction. On the FULL array, the FAST readout mode is used for the detector, with 43 groups and 4 integrations per dither point for the F770W exposures, and 25 groups and 8 integrations for the F2100W exposures, yielding 1900s and 2300s per pointing for F770W and F2100W, respectively. Here the reduced count rate is considered for F2100W observation. According to the JWST ETC, we will achieve a 5sigma point source limit of 25.0, and 22.4 for F770W and F2100W (AB mag). These are  $>1$ dex improvement compared to Spitzer, and allows us to detect AGN component at mid-IR wavelengths above 5 sigma for the bright DSFGs (for a various range of AGN activity. cf: Scientific Justification).

Proposal 6751 - Targets - Resolving the early phase of co-evolution of galaxies and supermassive black holes within cosmic web filam...

Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous
	(1)	ADF22a	RA: 22 17 33.7000 (334.3904167d) Dec: +00 17 48.80 (.29689d) Equinox: J2000		
	<i>Comments:</i> Category=Clusters of Galaxies Description=[High-redshift clusters]				
(2)	ADF22b	RA: 22 17 33.7267 (334.3905279d) Dec: +00 15 58.20 (.26617d) Equinox: J2000			
<i>Comments:</i> Category=Galaxy Description=[Active galactic nuclei, Starburst galaxies]					

Proposal 6751 - Observation 1 - Resolving the early phase of co-evolution of galaxies and supermassive black holes within cosmic we...

Tue Sep 17 16:01:32 GMT 2024

<b>Observation</b>	<p><b>Proposal 6751, Observation 1: MIRI F770W (a)</b></p> <p><b>Diagnostic Status: Warning</b></p> <p>Observing Template: MIRI Imaging</p>										
<b>Diagnostics</b>	(Visit 1:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.										
<b>Fixed Targets</b>	<b>#</b>	<b>Name</b>	<b>Target Coordinates</b>			<b>Targ. Coord. Corrections</b>		<b>Miscellaneous</b>			
	(1)	ADF22a	RA: 22 17 33.7000 (334.3904167d) Dec: +00 17 48.80 (.29689d) Equinox: J2000								
	<p><i>Comments:</i>                  Category=Clusters of Galaxies                  Description=High-redshift clusters]</p>										
<b>Template</b>	<p><b>Subarray</b> FULL</p>										
<b>Dithers</b>	<b>#</b>	<b>Dither Type</b>	<b>Starting Point</b>	<b>Number of Points</b>	<b>Points</b>	<b>Starting Set</b>	<b>Number of Sets</b>	<b>Optimized For</b>	<b>Direction</b>	<b>Pattern Size</b>	
	1	4-Point-Sets				1	1	POINT SOURCE	POSITIVE	DEFAULT	
<b>Spectral Elements</b>	<b>#</b>	<b>Filter</b>	<b>Readout Pattern</b>	<b>Groups/Int</b>	<b>Integrations/Exp</b>	<b>Exposures/Dith</b>	<b>Dither</b>	<b>Total Dithers</b>	<b>Total Integrations</b>	<b>Total Exposure Time</b>	<b>ETC Wkbk.Calc ID</b>
	1	F770W	FASTR1	43	4	1	Dither 1	4	16	1942.528	
	2	F2100W	FASTR1	25	8	1	Dither 1	4	32	2297.733	
<b>Special Requirements</b>	<p>Aperture PA Range 70 to 85 Degrees (V3 65.16455103 to 80.16455103)                  Aperture PA Range 240 to 250 Degrees (V3 235.16455103 to 245.16455103)</p>										

Proposal 6751 - Observation 2 - Resolving the early phase of co-evolution of galaxies and supermassive black holes within cosmic we...

Tue Sep 17 16:01:32 GMT 2024

<b>Observation</b>	<p><b>Proposal 6751, Observation 2: MIRI F770W (b)</b></p> <p><b>Diagnostic Status: Warning</b></p> <p>Observing Template: MIRI Imaging</p>										
<b>Diagnostics</b>	(Visit 2:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.										
<b>Fixed Targets</b>	<b>#</b>	<b>Name</b>	<b>Target Coordinates</b>			<b>Targ. Coord. Corrections</b>		<b>Miscellaneous</b>			
	(2)	ADF22b	RA: 22 17 33.7267 (334.3905279d) Dec: +00 15 58.20 (.26617d) Equinox: J2000								
	<p><i>Comments:</i>  <i>Category=Galaxy</i>  <i>Description=[Active galactic nuclei, Starburst galaxies]</i></p>										
<b>Template</b>	<p><b>Subarray</b> FULL</p>										
<b>Dithers</b>	<b>#</b>	<b>Dither Type</b>	<b>Starting Point</b>	<b>Number of Points</b>	<b>Points</b>	<b>Starting Set</b>	<b>Number of Sets</b>	<b>Optimized For</b>	<b>Direction</b>	<b>Pattern Size</b>	
	1	4-Point-Sets				1	1	POINT SOURCE	POSITIVE	DEFAULT	
<b>Spectral Elements</b>	<b>#</b>	<b>Filter</b>	<b>Readout Pattern</b>	<b>Groups/Int</b>	<b>Integrations/Exp</b>	<b>Exposures/Dith</b>	<b>Dither</b>	<b>Total Dithers</b>	<b>Total Integrations</b>	<b>Total Exposure Time</b>	<b>ETC Wkbk.Calc ID</b>
	1	F770W	FASTR1	43	4	1	Dither 1	4	16	1942.528	
	2	F2100W	FASTR1	25	8	1	Dither 1	4	32	2297.733	
<b>Special Requirements</b>	<p>Aperture PA Range 70 to 85 Degrees (V3 65.16455103 to 80.16455103)                  Aperture PA Range 240 to 250 Degrees (V3 235.16455103 to 245.16455103)</p>										