



# 7035 - Brown Dwarfs in NGC 602 in the SMC - An opportunity to characterize a substellar IMF at low metallicity

Cycle: 4, Proposal Category: GO

## INVESTIGATORS

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Dr. Mario Giuseppe Guarcello (CoI) (ESA Member)	
Dr. Anna Faye McLeod (CoI) (ESA Member)	Durham University

## OBSERVATIONS

<i>Folder</i>	<i>Observation</i>	<i>Label</i>	<i>Observing Template</i>	<i>Science Target</i>
NIRCam				
	6		NIRCam Imaging	(1) NGC-602
NIRSpec				

<i>Folder</i>	<i>Observation</i>	<i>Label</i>	<i>Observing Template</i>	<i>Science Target</i>
	3	APA80	NIRSpec MultiObject Spectroscopy	(2) APT_cat

## ABSTRACT

We propose to exploit the synergy of NIRCcam and NIRSpec to characterize, for the first time, the initial mass function (IMF) down to 15 Jupiter masses in NGC 602, a low-metallicity (1/5 Solar) young star cluster in the Small Magellanic Cloud (SMC). NGC 602 formed in an isolated location in the “wing” of the SMC, still it hosts numerous OB stars. It is the only low-metallicity young (about 2 Myr) star cluster where, with JWST, planetary-mass Brown Dwarfs have been detected, thus offering an opportunity to study the impact of stellar feedback from OB stars across the cluster region. We will collect tens of NIRSpec spectra of young, low-metallicity Brown Dwarfs. By targeting a completely unexplored parameter space (young and low metallicity), we will assemble a unique dataset that will have long-lasting legacy value for all the studies of Brown Dwarf formation, especially at sub-Solar metallicity. Only combining deep NIRCcam photometry and NIRSpec MOS it is possible to derive accurate spectral-type-based masses and measure ages better than 0.5 Myr using evolutionary models and spectral templates. This dataset will add an invaluable piece to the puzzle of the formation of sub stellar objects at conditions that are different than Solar, and will address the environmental condition impact on the young star cluster IMF.

## OBSERVING DESCRIPTION

The goals of this proposal are deep NIRCcam photometry and NIRSpec MOS to measure the sub-stellar initial mass function (IMF) down to 15 Jupiter masses and obtain the first spectra of around 45 young Brown Dwarfs (BDs) outside the Milky Way.

NIRCcam photometry:

To fulfill our science goals of characterizing the sub-stellar IMF down to 15 Jupiter masses we request NIRCcam photometry in 6 different filters: F140M, F162M, F182M, F277W, F360M, F444W. This combination of medium and wide-band filters is necessary to separate BDs from main sequence stars using CMDs and two-color diagrams by targeting H<sub>2</sub>O absorption bands (F140M, F182M) + continuum (F162M) and CH<sub>4</sub> (F360M), which are much deeper in BDs compared to low-mass stars. Two continuum filters (F277W, F444W) will show infrared excess in PMS stars and reveal the very different SED shape of background galaxies. To reach a S/N >10 in all filters a total integration time per filter of 41,744s is needed. To optimally cover the cluster region and simultaneously observe a background field West of NGC 602 possible due to the large field of view of NIRCcam, we will use the FULLBOX 6 dither pattern. Additionally, we apply 9 subpixel dithers allowing us to optimally sample the PSF, which is particularly important at short wavelengths, where the NIRCcam detector is under sampled. We chose the readout pattern MEDIUM2 with 8 groups and 1 integration each, which is a good compromise between data volume, to avoid data excess warnings allowing maximum flexibility for

scheduling, and dynamic range. We will utilize readout groups 0 and 1 as well as the archival data of program GO-2662 to recover most of the bright, saturated point sources.

NIRSpec MOS:

To obtain around 45 BD spectra and 30 low-mass PMS spectra in PRISM mode 3 individual MOS configurations are required. With a total exposure time of 27,835s per configuration a continuum  $S/N > 15$  can be reached for all objects above 50 Jupiter masses. We use the NRSIRS2 readout pattern, which suppresses the  $1/f$  noise thus improving the  $S/N$ . We apply 3-point nods to subtract the local background. Each observation requires 6 integrations with 21 groups each leaving the exposure time per integration at  $\sim 1550$ s, which is sufficiently short to avoid severe impacts from cosmic rays, as recommended by the NIRSpec team. The many isolated point sources in the region are sufficient for MSATA.

Proposal 7035 - Targets - Brown Dwarfs in NGC 602 in the SMC - An opportunity to characterize a substellar IMF at low metallicity

Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous
	(1)	NGC-602	RA: 01 29 13.1798 (22.3049158d) Dec: -73 33 40.72 (-73.56131d) Equinox: J2000  <i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i> <i>Category=Stellar Cluster</i> <i>Description=[Young star clusters]</i>	Epoch of Position: 2000	
(2)	APT_cat	RA: 01 29 27.8070 (22.3658625d) Dec: -73 33 30.48 (-73.55847d) Equinox: J2000  <i>Comments:</i> <i>Description=[]</i>			

Proposal 7035 - Observation 6 - Brown Dwarfs in NGC 602 in the SMC - An opportunity to characterize a substellar IMF at low metalli...

Thu May 01 13:00:28 GMT 2025

<b>Observation</b>	<p><b>Proposal 7035, Observation 6</b></p> <p><b>Diagnostic Status: Warning</b></p> <p>Observing Template: NIRCam Imaging</p>																																																	
<b>Diagnostics</b>	<p>(Visit 6:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.</p> <p>(Visit 6:2) Warning (Form): Overheads are provisional until the Visit Planner has been run.</p> <p>(Visit 6:3) Warning (Form): Overheads are provisional until the Visit Planner has been run.</p> <p>(Visit 6:4) Warning (Form): Overheads are provisional until the Visit Planner has been run.</p> <p>(Visit 6:5) Warning (Form): Overheads are provisional until the Visit Planner has been run.</p> <p>(Visit 6:6) Warning (Form): Overheads are provisional until the Visit Planner has been run.</p>																																																	
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<b>Special Requirements</b>	<p>Sequence Visits within 53.0 Days</p> <p>Aperture PA Range 169.92542306 to 209.92542306 Degrees (V3 170.0 to 210.0)</p> <p>Aperture PA Range 349.92542306 to 29.92542306 Degrees (V3 350.0 to 30.0)</p> <p>Visits Same PA</p>																																																	

Proposal 7035 - Observation 3 - Brown Dwarfs in NGC 602 in the SMC - An opportunity to characterize a substellar IMF at low metalli...

Thu May 01 13:00:29 GMT 2025

<b>Observation</b>	<p><b>Proposal 7035, Observation 3: APA80</b></p> <p><b>Diagnostic Status: Warning</b></p> <p>Observing Template: NIRSpec MultiObject Spectroscopy</p>																																																						
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Reference Stars	Visit	ID	RA	Dec	Magnitude	Visit	ID	RA	Dec	Magnitude	
	1	2197	22.387087	-73.588213	22.58599175091226 4	1	10865	22.441277	-73.572110	22.68535511487366 6	
	1	2201	22.363707	-73.587955	21.94797291138347 4	1	11167	22.434344	-73.546578	22.86372994657824	
	1	2249	22.383426	-73.586209	22.86422866559553 3	1	11325	22.482622	-73.540826	21.71657115687884 2	
	1	2261	22.399920	-73.585585	22.51451330162965 5	1	11344	22.438898	-73.539801	22.03595482352456 3	
	Visit	ID	RA	Dec	Magnitude	Visit	ID	RA	Dec	Magnitude	
	2	6038	22.314020	-73.579673	23.28213275232708 2	2	7981	22.264138	-73.578863	22.03769433077623 3	
	2	6062	22.310839	-73.578637	23.12323865821752 4	2	7995	22.250214	-73.578227	22.78203538162314	
	2	6297	22.306762	-73.569517	22.91394125960717 2	2	8009	22.255824	-73.577322	22.88530199628390 7	
	2	6478	22.306332	-73.564022	23.29704964380384 5	2	9995	22.234153	-73.533466	22.79704359775866	
Spectral Elements	#	Exposure Specification	MSA Configuration	Nod Pattern	Pointing	Aperture PA	Dispersion Offset (Shutters)	Cross-Dispersion Offset (Shutters)	Total Dithers	Total Integrations	Total Exposure Time
	1	1 (PRISM/CLEAR)	c1	3 Shutter Slitlet	22.417134208333 337 Degrees - 73.568983611111 1 Degrees	79.950820932445 47			3	18	27835.602
	2	1 (PRISM/CLEAR)	c2	3 Shutter Slitlet	22.326219208333 335 Degrees - 73.549541944444 46 Degrees	80.038025386673 67			3	18	27835.602
	3	1 (PRISM/CLEAR)	c3	3 Shutter Slitlet	22.433084083333 334 Degrees - 73.572138055555 57 Degrees	79.935521746056 92			3	18	27835.602
Special Requirements	Group Visits within 53.0 Days Aperture PA Range 50 to 120 Degrees (V3 271.4254303 to 341.4254303) Aperture PA Range 230 to 300 Degrees (V3 91.4254303 to 161.4254303) Visits Same PA MSA Planned Aperture PA 80.0000 to 80.0000 Degrees (V3 301.4254303 to 301.4254303)										