



7335 - Forever Blowing Bubbles: What Powers a 24-kpc Ionized Gas Nebula Around a Normal Galaxy at $z=6$?

Cycle: 4, Proposal Category: GO

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OBSERVATIONS

<i>Folder</i>	<i>Observation</i>	<i>Label</i>	<i>Observing Template</i>	<i>Science Target</i>
Observation Folder				
	1		NIRSpec IFU Spectroscopy	(1) GS-53.00804-27.85477
	2		NIRSpec IFU Spectroscopy	(1) GS-53.00804-27.85477

ABSTRACT

It is thought that powerful galactic outflows may be responsible for disrupting and quenching star formation in massive galaxies at high redshift. This proposal will explore outflows in ID518794, a quenching galaxy hosting a 24 kpc extended [OIII] emission nebula at $z = 6$. This shell-like nebula, the most extended observed across the GOODS extragalactic fields, is extraordinary in size and luminosity ($2 \times 10^{10} L_{\text{Sun}}$), but is even more notable since it is centered not around a massive quasar, but rather a more typical ($10^{10} M_{\text{Sun}}$) galaxy with evidence of a recent drop in star-formation rate. This unique system is at least 5-10 times more rare than quiescent galaxies at $z = 4-5$, and we may be observing a short-lived phase of a powerful gas outflow just before the onset of long-term quenching. To test this hypothesis, we seek to understand whether the giant nebula is an outflow, an AGN ionization echo (a "voorwerp"), or tidally stripped material, and then to confirm the quenching nature of the central galaxy. To achieve these goals, we propose NIRSpec/IFS observations with

the prism and g395h dispersers to

- 1) investigate the star-formation history and stellar-population properties of the central galaxy
- 2) search for evidence of an active galactic nucleus in ID518794 and
- 3) measure the kinematics, chemistry and physical conditions of the gas, in both the host galaxy and the giant nebula.

If confirmed to host a massive, 24-kpc outflow, this system may help to explain how AGNs are able to quench massive galaxies in the first 1-2 billion years after the Big Bang.

OBSERVING DESCRIPTION

This proposal aims to obtain NIRSpec/IFS spatially resolved spectroscopy in clear/prism (3 hours) and f290lp/g395h (7.5 hours).

The targets are a massive $z=6$ galaxy (10.2 dex M_{sun}) and its extended, surrounding nebula (4 arcsec).

To ensure observing the nebula in a single (dithered) NIRSpec/IFS pointing, we further require target acquisition.

To optimize telescope usage, we designed the two observations to be consecutive, i.e. sharing the same target acquisition. This further constrains the two observing setups to have the same aperture angle, which greatly simplifies the analysis, due to the NIRSpec/IFS PSF being elliptical and aligned with the aperture angle.

Proposal 7335 - Targets - Forever Blowing Bubbles: What Powers a 24-kpc Ionized Gas Nebula Around a Normal Galaxy at z=6?

#	Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous
(1)	GS-53.00804-27.85477	RA: 03 32 1.9405 (53.0080854d) Dec: -27 51 17.31 (-27.85481d) Equinox: J2000	Epoch of Position: 2024	
<p><i>Comments:</i> <i>Category=Galaxy</i> <i>Description=[Active galactic nuclei, Galaxy halos, High-redshift galaxies]</i> <i>Extended=YES</i></p>				
Fixed Targets	(2)	GAIA5057502833478542336 RA: 03 32 3.4165 (53.0142354d) Dec: -27 50 58.83 (-27.84967d) Equinox: J2000	Proper Motion RA: -4.3214 mas/yr Proper Motion Dec: 13.76 mas/yr Epoch of Position: 2023	
	<p><i>Comments: F115W = 18.9 AB mag</i> <i>F150W = 18.8 AB mag</i> <i>F410M and F335M Centroid [2023] from own measurement in ICRSd:</i> <i>53.0142356 -27.8496760</i> <i>By comparing with F814W [2018?] I get a pm of 25 mas/year; but not sure the epoch is really 2018?</i> <i>GAIA DRI 03 32 3.4185 -27 50 58.74</i> <i>GAIA Entry (Epoch 2015--2016). Note: no PM available in GAIA!</i> <i># ID RA[J2000 deg] eRA[mas] Dec[J2000 deg] eDec[mas] Gmag</i> <i>5057502833478542336 053.0142439078 0.681 -27.8496492533 1.159 19.332</i> <i>Category=Calibration</i> <i>Description=[Target acquisition test]</i> <i>Extended=NO</i></p>			

Proposal 7335 - Observation 1 - Forever Blowing Bubbles: What Powers a 24-kpc Ionized Gas Nebula Around a Normal Galaxy at z=6?

Mon Jul 28 19:00:11 GMT 2025

Observation	Proposal 7335, Observation 1 Diagnostic Status: Warning Observing Template: NIRSpec IFU Spectroscopy <i>Comments: PA Restrictions arise from avoiding spoiler stars in the NIRSpec/MSA footprint. The range 194.4-208.4 was excluded because of a bright spoiler in the NIRCcam parallel. We deemed this PA range fairly small, so it does not greatly reduce the schedulability of the proposed programme. At the same time, a deep pure-parallel observation may benefit from avoiding bright galaxies. The range V3 PA 282.263 to 89.163 is required by WATA target acquisition (TA), to achieve a slew distance <40 arcsec between the the slit S1600A1 at the location of the TA target and the IFS aperture at the location of the science target; this range is more permissive than our spoiler-star constraints, so it has not been added to the Special Requirements.</i>																																			
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Proposal 7335 - Observation 2 - Forever Blowing Bubbles: What Powers a 24-kpc Ionized Gas Nebula Around a Normal Galaxy at z=6?

Mon Jul 28 19:00:11 GMT 2025

Observation	Proposal 7335, Observation 2 Diagnostic Status: Warning Observing Template: NIRSpec IFU Spectroscopy <i>Comments: PA Restrictions arise from avoiding spoiler stars in the NIRSpec/MSA footprint. The range 194.4-208.4 was excluded because of a bright spoiler in the NIRCcam parallel. We deemed this PA range fairly small, so it does not greatly reduce the schedulability of the proposed programme. At the same time, a deep pure-parallel observation may benefit from avoiding bright galaxies. The range V3 PA 282.263 to 89.163 is required by WATA target acquisition (TA), to achieve a slew distance <40 arcsec between the the slit S1600A1 at the location of the TA target and the IFS aperture at the location of the science target; this range is more permissive than our spoiler-star constraints, so it has not been added to the Special Requirements.</i>																																			
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