



7760 - Testing the gravitational lens origin of SAV events -- MIRI/MFS IFU observations of PKS 1413+135

Cycle: 4, Proposal Category: GO

INVESTIGATORS

<i>Name</i>	<i>Institution</i>
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Dr. Brian Keeney (CoI)	Southwest Research Institute

OBSERVATIONS

<i>Folder</i>	<i>Observation</i>	<i>Label</i>	<i>Observing Template</i>	<i>Science Target</i>
Observation Folder				
	7	MIR Spectroscopy	MIRI Medium Resolution Spectroscopy	(1) PKS1413+135
	8		MIRI Medium Resolution Spectroscopy	(2) PKS1413+135_BKG

ABSTRACT

Time-domain surveys are an exciting new frontier, having discovered tidal disruption events, gamma-ray bursts and fast radio transients. Symmetric achromatic variability (SAV) events are a new class of variability events, identified in 7 AGN out of 1830 monitored regularly by the Owens Valley Radio Observatory. SAV events have lightcurves that are time-symmetric and achromatic. They have been modeled by gravitational milli-lensing, which can occur when relativistically moving jet components in a background AGN, move through caustics created by 10^3 - 10^6 solar mass features in a foreground galaxy. Confirmation of the lensing nature of SAV could give us a powerful new tool, allowing future multiwaveband observations to reveal <0.01 pc scale structure in distant galaxies.

We propose to test this idea with MIRI/MRS IFU observations of PKS 1413+135, which has had four SAV events between 1992-2015. PKS 1413+135 is one of the most puzzling blazars known, due to uncertainties about its host galaxy, redshift and nature. The apparent host galaxy, a

$z=0.247$ edge-on spiral, shows no evidence of activity in its optical spectrum. This is incongruent with the blazar projected on the sky 13 ± 4 mas (52 ± 16 pc) from its isophotal centroid in HST/NICMOS observations. The host also has a GMC complex projected within 25 mas of the AGN, with time-variable absorption. Confirming the lens model would require identifying AGN lines at $z > 0.247$. This can only be done in the infrared, as the AGN suffers 30 mag of extinction in the optical. Existing Spitzer data are too low S/N and resolution to do this. We will also re-examine whether the AGN is centered in the $z=0.247$ spiral.

OBSERVING DESCRIPTION

We propose to obtain mid-infrared spectra and imaging of PKS 1413+135. The most important goal of the mid-IR spectroscopy is to obtain complete coverage between 5-28 μm so that we can obtain both emission lines from the AGN, as well as PAH and other molecular features due to the cloud(s) that may be the lensing screen in the $z=0.247$ spiral. Discovery of AGN lines at $z > 0.247$ will test the gravitational lensing model of Vedantham et al. (2017a) and Peirson et al. (2022). The most important goal of the imaging is to re-examine whether the AGN is de-centered in the spiral, as shown at 3 sigma in HST/NICMOS images.

We have therefore chosen to use MIRI's MRS mode, using all four spectral channels to obtain complete spectral coverage. We have used the ETC to calculate the exposure times so that a $S/N \sim 15-20$ per pixel is achieved (using the Spitzer spectrum as a seed), and we have followed the recommendations of the JWST documentation as far as choosing the detector configuration. The goal of the simultaneous imaging is to both place in context the MRS spectra and also test the conclusion of Perlman et al. (2002) that the blazar is not centered within the $z=0.247$ spiral host galaxy. We have chosen to obtain images in F560W, F770W and F1000W. The F560W and F1800W imaging cover the continuum, while the F1280W imaging will be in the middle of the silicate dust feature at $z=0.247$. We have thus used the ETC to estimate the exposure time required to obtain a quick image that should show both the galaxy and the AGN. In the mid-IR, dust has very little effect.

Proposal 7760 - Targets - Testing the gravitational lens origin of SAV events -- MIRI/MFS IFU observations of PKS 1413+135

Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous
	(1)	PKS1413+135	RA: 14 15 58.8168 (213.9950700d) Dec: +13 20 23.71 (13.33992d) Equinox: J2000 <i>Comments: This object was generated by the targetselector and retrieved from the NED database.</i> <i>Category=Galaxy</i> <i>Description=[Active galactic nuclei, BL Lacertae objects, Einstein rings, Relativistic jets]</i>		
(2)	PKS1413+135_BKG	RA: 14 15 58.8168 (213.9950700d) Dec: +13 20 23.71 (13.33992d) Equinox: J2000 <i>Comments:</i> <i>Category=Calibration</i> <i>Description=[Telescope/sky background]</i> <i>Extended=YES</i>	Epoch of Position: 2000		

Proposal 7760 - Observation 7 - Testing the gravitational lens origin of SAV events -- MIRI/MFS IFU observations of PKS 1413+135

Wed Jun 11 00:00:08 GMT 2025

Observation	Proposal 7760, Observation 7: MIR Spectroscopy Diagnostic Status: Error Observing Template: MIRI Medium Resolution Spectroscopy Background Observations:[Observation 8]																																																																																																						
	(MIR Spectroscopy (Obs 7)) Error (Form): This target requires similar background exposures that are linked in a non-interruptible group/sequence. (Visit 7:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.																																																																																																						
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Special Requirements

Sequence Observations 7, 8, Non-interruptible

Proposal 7760 - Observation 8 - Testing the gravitational lens origin of SAV events -- MIRI/MFS IFU observations of PKS 1413+135

Wed Jun 11 00:00:08 GMT 2025

Observation	Proposal 7760, Observation 8 Diagnostic Status: Warning Observing Template: MIRI Medium Resolution Spectroscopy Background Observation For: [MIR Spectroscopy (Obs 7)]												
	(Visit 8:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.												
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Acquisition	<i>Comments:</i> Category=Calibration Description=[Telescope/sky background] Extended=YES												
	#											Target	
	1											NONE	
Template	AcqFilter	Primary Channel				Simultaneous Imaging			Imager Subarray		Grating Wheel Direction		
	F560W	Imager				YES			FULL		Allow Auto Reorder		
Dithers	#	Dither Type				Optimized For				Direction			
	1	2-Point				BACKGROUND				NEGATIVE			
Spectral Elements	#	Wavelength Range	Detector	Filter	Readout Pattern	Groups/Int	Integrations/E xp	Exposures/Dit h	Dither	Total Dithers	Total Integrations	Total Exposure Time	ETC Wkbk.Calc ID
	1		IMAGER	F1800W	FASTR1	30	2	1	Dither 1	2	4	338.555	
	1	SHORT(A)	MRSLONG		FASTR1	30	2	1	Dither 1	2	4	338.555	
	1	SHORT(A)	MRSSHORT		FASTR1	30	2	1	Dither 1	2	4	338.555	
	2		IMAGER	F1280W	FASTR1	30	2	1	Dither 1	2	4	338.555	
	2	MEDIUM(B)	MRSLONG		FASTR1	30	2	1	Dither 1	2	4	338.555	
	2	MEDIUM(B)	MRSSHORT		FASTR1	30	2	1	Dither 1	2	4	338.555	
	3		IMAGER	F560W	FASTR1	30	2	1	Dither 1	2	4	338.555	
	3	LONG(C)	MRSLONG		FASTR1	30	2	1	Dither 1	2	4	338.555	
	3	LONG(C)	MRSSHORT		FASTR1	30	2	1	Dither 1	2	4	338.555	

Special Requirements

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