



# 2008 - A Blast From the Past: A Spectroscopic look at the Flash Heating of HD80606b

Cycle: 1, Proposal Category: GO

## INVESTIGATORS

<i>Name</i>	<i>Institution</i>	<i>E-Mail</i>
<b>Dr. Tiffany Kataria (PI)</b>	<b>Jet Propulsion Laboratory</b>	<b>tiffany.kataria@jpl.nasa.gov</b>
Dr. Nikole Lewis (CoI) (CoPI)	Cornell University	nikole.lewis@cornell.edu
Dr. Laura C Mayorga (CoI) (CoPI)	The Johns Hopkins University Applied Physics Laboratory	laura.mayorga@jhuapl.edu
Dr. Thomas P. Greene (CoI)	NASA Ames Research Center	tom.greene@nasa.gov
Dr. Charles A. Beichman (CoI)	Jet Propulsion Laboratory	charles.a.beichman@jpl.nasa.gov
Dr. Gregory Laughlin (CoI)	University of California - Santa Cruz	laughlin@ucolick.org
Dr. Mark S. Marley (CoI)	University of Arizona	marksmarley@email.arizona.edu

## OBSERVATIONS

<i>Folder</i>	<i>Observation</i>	<i>Label</i>	<i>Observing Template</i>	<i>Science Target</i>
Observation Folder				
	1	HD 80606b phase curve	MIRI Low Resolution Spectroscopy	(1) HD-80606

## ABSTRACT

High-eccentricity exoplanets are a unique class of planets that experience dramatic increases in insolation and tidal forcing near periastron passage. By observing the thermal response of the atmosphere before, during, and after periastron passage (the ‘flash-heating’ event), we can constrain physical processes within a single set of observations. We propose to utilize MIRI/LRS to observe a partial spectroscopic phase curve of the eccentric hot Jupiter HD 80606b. In conducting our spectroscopic observations from 12 hrs before through to 8 hrs after periastron, we will capture the time- and wavelength-dependent flux before, during and after HD 80606b’s flash-heating event. This information will allow us to constrain the internal

heating of HD 80606b, probe variations in clouds and chemistry throughout HD 80606b's orbit, and assess the role of dynamical mixing in HD 80606b's atmosphere. Our proposed MIRI/LRS observations of HD~80606b will build upon the rich legacy of Spitzer, Hubble, and now TESS and bring a new experimental look into fundamental processes that shape the phase-curve observations of not only eccentric hot Jupiters, but all hot Jupiters.

## **OBSERVING DESCRIPTION**

These observations aim to capture the temporal thermal variations in the highly eccentric exoplanet HD80606b's atmosphere during its periastron passage. The overall length of the observations and the achievable precision are key to meeting the science goals. We will use the MIRI LRS slitless mode to maximize both the signal from the planet and the achievable precision. Our observations will be taken in the standard MIRI LRS slitless configuration that uses the slitlessprism subarray, fast readmode, and no filter. We have selected 14 groups per integration so as to not exceed 80% full well and avoid saturation while maximizing the observing efficiency. The observations will consist of 32337 integrations with 2.3856 seconds per integration covering roughly 20 hours of HD80606b's orbit centered on the secondary eclipse event and inclusive of periastron passage. Given the target brightness and leveraging simulations from JWST PandExo, we expect to achieve a one-sigma precision of 100 ppm at 8 micron on a 1 hour timescale. We have allowed for a full 60 minute window/phase range at the start of the observation, which will still allow for capture of the relevant events/timescales during the course of the observations. This observation start window has been specified to ensure that sufficient time before and after the periastron passage of the planet are captured to allow for characterization of the planet's heating and subsequent response. We will consider discarding the first 30-45 minutes of the observation if necessary due to detector ramp effects with no impact on the overall science goals. Given the observability of the target and its orbital period, there are two opportunities to capture the periastron passage of HD80606b during cycle 1.

For target acquisition we have followed the procedure outlined in JDox for MIRI LRS slitless observations. We will acquire on the science target using the slitlessprism subarray with the fast readout pattern with 4 groups per integration and 1 integration resulting in a 0.64 second exposure time. We have selected the F1000W filter in our target acquisition strategy to avoid saturation while still achieving a SNR of 389 during the target acquisition.

Proposal 2008 - Targets - A Blast From the Past: A Spectroscopic look at the Flash Heating of HD80606b

Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous
	(1)	HD-80606	RA: 09 22 37.6680 (140.6569500d)	Proper Motion RA: 0.005875195586551243 sec of time/yr	
		Dec: +50 36 13.60 (50.60378d)	Proper Motion Dec: 0.01034 arcsec/yr		
		Equinox: J2000	Epoch of Position: 2015.5		
	<i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i>				
	Category=Star				
	Description=[Exoplanet Systems, G stars]				

Proposal 2008 - Observation 1 - A Blast From the Past: A Spectroscopic look at the Flash Heating of HD80606b

Thu Apr 28 23:00:45 GMT 2022

<b>Observation</b>	<p><b>Proposal 2008, Observation 1: HD 80606b phase curve</b></p> <p><b>Diagnostic Status: Warning</b></p> <p>Observing Template: MIRI Low Resolution Spectroscopy</p>								
<b>Diagnostics</b>	<p>(HD 80606b phase curve (Obs 1)) Warning (Form): Exposure Duration exceeds the limit of 10000.0 seconds. Above this limit it is possible that a High Gain Antenna move may occur during the exposure.</p> <p>(Visit 1:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.</p>								
<b>Fixed Targets</b>	<b>#</b>	<b>Name</b>	<b>Target Coordinates</b>	<b>Targ. Coord. Corrections</b>			<b>Miscellaneous</b>		
	(1)	HD-80606	RA: 09 22 37.6680 (140.6569500d) Dec: +50 36 13.60 (50.60378d) Equinox: J2000	Proper Motion RA: 0.005875195586551243 sec of time/yr Proper Motion Dec: 0.01034 arcsec/yr Epoch of Position: 2015.5					
	<p><i>Comments: This object was generated by the targetselector and retrieved from the SIMBAD database.</i></p> <p><i>Category=Star</i></p> <p><i>Description=[Exoplanet Systems, G stars]</i></p>								
<b>Acquisition</b>	<b>#</b>	<b>Target</b>	<b>Filter</b>	<b>Readout Pattern</b>	<b>Groups/Int</b>	<b>Integrations/Exp</b>	<b>Total Integrations</b>	<b>Total Exposure Time</b>	<b>ETC Wkbk.Calc ID</b>
	1	SAME	F1000W	FAST	4	1	1	0.636	60953
<b>Template</b>	<b>Subarray</b>				<b>Obtain Verification Image?</b>				
	SLITLESSPRISM				true				
<b>Dithers</b>	<b>#</b>	<b>Dither Type</b>	<b>No. Spectral Steps</b>	<b>Spectral Step Offset</b>	<b>No. Spatial Steps</b>	<b>Spatial Step Offset</b>			
	1	NONE							
<b>Pointing Verification</b>	<b>#</b>	<b>Readout Pattern</b>	<b>Groups/Int</b>	<b>Integrations/Exp</b>	<b>Total Integrations</b>	<b>Total Exposure Time</b>	<b>ETC Wkbk.Calc ID</b>		
	1	FAST	4	1	1	0.636	60953		

Proposal 2008 - Observation 1 - A Blast From the Past: A Spectroscopic look at the Flash Heating of HD80606b

Spectral Elements	#	Readout Pattern	Groups/Int	Integrations/Exp	Total Integrations	Exposures/Dith	Total Dithers	Total Exposure Time	ETC Wkbk.Calc ID
		1	FASTR1	14	10812	32436	3	1	77378.844
Special Requirements	Phase 0.9418 to 0.9422 with period 111.43639 Days and zero-phase 2455210.642 HJD Aperture PA Range 100 to 247 Degrees (V3 95.16574676 to 242.16574676) Aperture PA Range 282 to 65 Degrees (V3 277.16574676 to 60.16574676) Time Series Observation No Parallel								